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LAS PARTÍCULAS SUPERMASIVAS: NATURALEZA FENOMENOLÓGICA DE LA PARTÍCULA OSCURA

SUPERMASSIVE PARTICLES: PHENOMENOLOGICAL
NATURE OF THE DARK PARTICLE

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Las partículas supermasivas: naturaleza fenomenológica de la partícula oscura

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RESUMEN

Las partículas supermasivas, teorizadas por este autor a lo largo de los trabajos anteriores, comporta la existencia de partículas conexas al Modelo Estándar de Física de Partículas, las mismas que pueden ser de naturaleza fermiónica o bosónica, según sea el caso, que interactúan en un campo de gauge y en supersimetría, cuya masa, es extremadamente densa, capaz de deformar el espacio – tiempo cuántico en el que interactúa (a lo que llamamos gravedad endógena), afectando la configuración morfológica de las partículas circundantes (masa y energía) así como sus coordenadas en espacio y tiempo, y desdoblando multidimensiones o supermembranas, sin que sea necesario que la partícula supermasiva o masiva, deba interactuar con un campo gravitónico y en consecuencia, con un gravitón, como se ha propuesto en relación al modelo de gravedad exógena. Esta partícula, a lo largo de anteriores trabajos, ha sido denominada también como “partícula cosmológica”, “partícula supermasiva” o “superpartícula”, más, llámesela también “partícula oscura”, para propósitos de este artículo.

Palabras Clave: partícula supermasiva, partícula oscura, teoría cuántica de campos relativistas, agujeros negros cuánticos, supermembranas

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Supermassive particles: phenomenological nature of the dark particle

ABSTRACT

The supermassive particles, theorized by this author throughout the previous works, involve the existence of particles related to the Standard Model of Particle Physics, which can be of fermionic or bosonic nature, as the case may be, interacting in a gauge field and in supersymmetry, whose mass is extremely dense, capable of deforming the quantum space-time in which it interacts (what we call endogenous quantum gravity), affecting the morphological configuration of the surrounding particles (mass and energy) as well as their coordinates in space and time, and unfolding multidimensions or supermembranes, without it being necessary for the supermassive or massive particle to interact with a gravitonic field and consequently with a graviton, as has been proposed in relation to the exogenous gravity model. This particle, throughout previous works, has also been called “cosmological particle”, “supermassive particle” or “superparticle”, but, for the purposes of this article, let us also call it “dark particle”.

Keywords: supermassive particle, dark particle, relativistic quantum field theory, quantum black holes, supermembranes



INTRODUCCIÓN

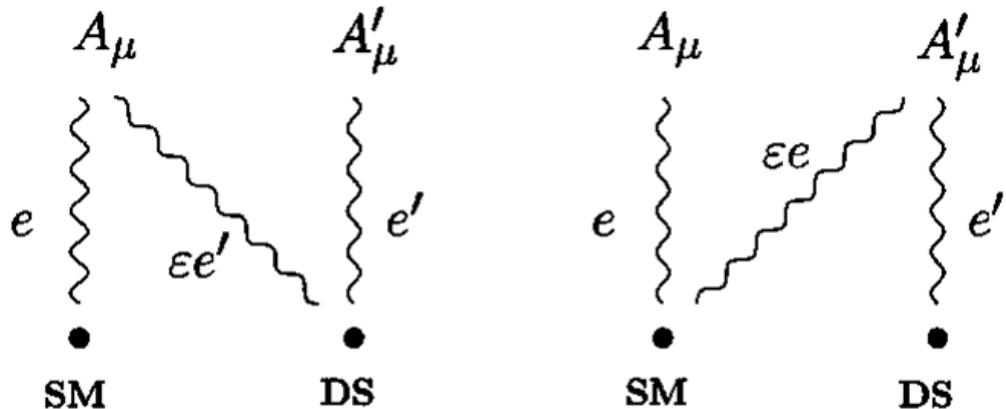
El propósito de mis trabajos anteriores, ha sido, establecer un marco teórico paralelo, cuyo objetivo es unificar la relatividad general y especial con la mecánica cuántica, de ahí nace lo que he denominado “Teoría Cuántica de Campos Relativistas” o “Teoría Cuántica de Campos Curvos”, cuyos postulados esenciales, son: 1) La existencia de supergravedad cuántica exógena, esto es, que una partícula deforma el espacio – tiempo en el que interactúa, por la permeabilidad de gravedad que le es inherente, al interactuar con un campo gravitónico y por ende, con un gravitón; 2) La existencia de gravedad cuántica endógena, esto es, cuando una partícula, a propósito de su masa extremadamente densa, curva el espacio – tiempo cuántico en el que interactúa, engendrando supermembranas y multidimensiones, así como deformaciones a los sistemas de referencia de las partículas interactuantes, entendiendo que todos los campos cuánticos, sean de naturaleza fermiónica o bosónica, según sea el caso, están interrelacionados, es decir, conectados por partículas comunes, verbigracia, el bosón de Higgs; y, 3) La existencia de agujeros negros cuánticos, a propósito de la creación de supermembranas y multidimensiones causados por la deformación masiva del espacio – tiempo cuántico, esto es, la singularidad cuántica de la gravitación por densidad.

Los postulados anteriormente mencionados, han sido desarrollados cuidadosamente en trabajos anteriores, sin embargo, las partículas supermasivas, como concepto esencial de la Teoría Cuántica de Campos Relativistas, requieren de un modelo matemático adicional que describa su comportamiento, fenomenología y características particulares, en relación a lo mencionado anteriormente, es decir, en relación a la gravedad cuántica que le es implícita, sus capacidades de deformación debido a sus interacciones de gauge y supersimetrías y el modelamiento de superespacios por la curvatura del tejido del espacio – tiempo cuántico.



RESULTADOS Y DISCUSIÓN

El campo de una partícula supermasiva, se presenta así:



$$\sin \theta = \varepsilon, \cos \theta = \sqrt{1 - \varepsilon^2}$$

$$\sin \theta = 0, \cos \theta = 1$$

Cuyos grupos de gauge en lagrangiano, son:

$$\mathcal{L}_0 = -\frac{1}{4}F_{a\mu\nu}F_a^{\mu\nu} - \frac{1}{4}F_{b\mu\nu}F_b^{\mu\nu} - \frac{\varepsilon}{2}F_{a\mu\nu}F_b^{\mu\nu}.$$

$$\mathcal{L} = eJ_\mu A_b^\mu + e'J'_\mu A_a^\mu,$$

$$\mathcal{L}' = \left[\frac{e'\cos \theta}{\sqrt{1 - \varepsilon^2}} J'_\mu + e \left(\sin \theta - \frac{\varepsilon \cos \theta}{\sqrt{1 - \varepsilon^2}} \right) J_\mu \right] A'^\mu + \left[-\frac{e'\sin \theta}{\sqrt{1 - \varepsilon^2}} J'_\mu + e \left(\cos \theta + \frac{\varepsilon \sin \theta}{\sqrt{1 - \varepsilon^2}} \right) J_\mu \right] A^\mu$$

$$\mathcal{L}' = \left[\frac{e'}{\sqrt{1 - \varepsilon^2}} J'_\mu - \frac{e\varepsilon}{\sqrt{1 - \varepsilon^2}} J_\mu \right] A'^\mu + eJ_\mu A^\mu$$

$$\mathcal{L}' = e'J'_\mu A'^\mu + \left[-\frac{e'\varepsilon}{\sqrt{1 - \varepsilon^2}} J'_\mu + \frac{e}{\sqrt{1 - \varepsilon^2}} J_\mu \right] A^\mu$$

Más en términos kinéticos, tenemos:

$$\begin{pmatrix} A_a^\mu \\ A_b^\mu \end{pmatrix} = \begin{pmatrix} \frac{1}{\sqrt{1 - \varepsilon^2}} & 0 \\ -\frac{\varepsilon^2}{\sqrt{1 - \varepsilon^2}} & 1 \end{pmatrix} \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} A'^\mu \\ A^\mu \end{pmatrix}$$

$$\sin \theta = \frac{\delta \sqrt{1 - \varepsilon^2}}{\sqrt{1 - 2\delta\varepsilon + \delta^2}} \quad \cos \theta = \frac{1 - \delta\varepsilon}{\sqrt{1 - 2\delta\varepsilon + \delta^2}}$$

Sin embargo, si se rompe la simetría de gauge, tenemos en lagrangiano:

$$\mathcal{L}_{stu} = -\frac{1}{2}M_a^2 A_{a\mu} A_a^\mu - \frac{1}{2}M_b^2 A_{b\mu} A_b^\mu - M_a M_b A_{a\mu} A_b^\mu$$



$$\mathcal{L}'' = \frac{1}{\sqrt{1 - 2\delta\varepsilon + \delta^2}} \left[\frac{e'(1 - \delta\varepsilon)}{\sqrt{1 - \varepsilon^2}} J'_\mu + \frac{e(\delta - \varepsilon)}{\sqrt{1 - \varepsilon^2}} J_\mu \right] A'^\mu + \frac{1}{\sqrt{1 - 2\delta\varepsilon + \delta^2}} [eJ_\mu - \delta e' J'_\mu] A^\mu$$

Cuyo acoplamiento viene dado por:

$$\mathcal{L} \supset -\frac{e\varepsilon}{\sqrt{1 - \varepsilon^2}} J_\mu A'^\mu \simeq -e\varepsilon J_\mu A'^\mu,$$

Por otro lado, la hiper – carga de la partícula supermasiva viene dada por:

$$\tilde{\mathcal{L}} = -\frac{\varepsilon}{2\cos\theta_W} \tilde{F}'_{\mu\nu} B^{\mu\nu}.$$

Cuya simetría de gauge viene dada por:

$$\begin{pmatrix} W_\mu^3 \\ B_\mu \\ \tilde{A}'_\mu \end{pmatrix} = \begin{pmatrix} c_W & s_W & -s_W\varepsilon \\ -s_W & c_W & -c_W\varepsilon \\ t_W\varepsilon & 0 & 1 \end{pmatrix} \begin{pmatrix} Z_\mu \\ A_\mu \\ A'_\mu \end{pmatrix},$$

Y en lagrangiano:

$$\mathcal{L} \supset -e\varepsilon J^\mu A'_\mu + e'\varepsilon t_W J'^\mu Z_\mu + e' J'^\mu A'_\mu,$$

Más sus operadores y osciladores vienen dados por:

$$\begin{aligned} \mathcal{L} &= \frac{e_D}{2\Lambda_5} \bar{\psi}_L^i \sigma_{\mu\nu} (\mathbb{D}_M^{ij} + i\gamma_5 \mathbb{D}_E^{ij}) \psi^j F'^{\mu\nu} \\ \mathcal{L} &= \frac{e_D}{2\Lambda^2} \bar{\psi}_L^i \sigma_{\mu\nu} (\mathbb{D}_M^{ij} + i\gamma_5 \mathbb{D}_E^{ij}) H \psi_R^j F'^{\mu\nu} + \text{H.c.} \\ \mathcal{L}' &= \frac{e_D}{2\Lambda^2} \bar{\psi}_L^i \gamma_\mu (\mathbb{R}_r^{ij} + i\gamma_5 \mathbb{R}_a^{ij}) D_\nu \psi^j F'^{\mu\nu} \end{aligned}$$

En los que la densidad de la partícula supermasiva, por isotropización, se mide así:

$$m_\chi \left(\frac{0.01}{\alpha_D} \right)^{2/3} \gtrsim 300 \text{GeV}$$

$$\frac{G_N m_\chi^4 N}{8\alpha_D^2} \gtrsim 50 \log \frac{G_N m_\chi^2 N}{2\alpha_D}$$

$$\chi\chi \rightarrow A'A'$$

$$\langle \sigma_{\chi\chi \rightarrow A'A'} v \rangle = \frac{2\pi\alpha_D^2}{m_\chi^2}$$

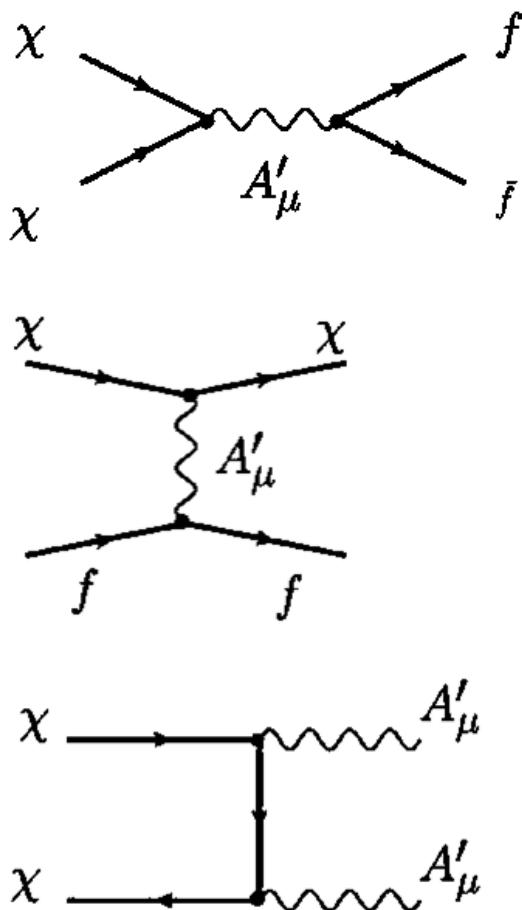
$$\Omega_\chi h^2 \approx \frac{2.5 \times 10^{-10} \text{GeV}^{-2}}{\langle \sigma_{\chi\chi \rightarrow A'A'} v \rangle}$$



$$\langle \sigma_{\chi\chi \rightarrow f\bar{f}v} \rangle \simeq \frac{2\pi\alpha_L^2}{m_S^2}$$

$$2\pi\alpha_L^2 \left(\frac{10\text{TeV}}{m_S}\right)^2 \simeq 0.1$$

En este punto y usando los diagramas de Feynman, pasa a explicarse las interacciones de las partículas supermasivas en distintos grupos de gauge, cuya aniquilación provoca un agujero negro cuántico supermasivo:



$$\chi\bar{\chi} \rightarrow A' \rightarrow \bar{f}f$$

$$\sigma_{\chi\chi \rightarrow f\bar{f}} = \frac{4\pi}{3} \varepsilon^2 \alpha \alpha_D m_\chi^2 \left(1 + \frac{2m_e^2}{s}\right) \left(1 + \frac{2m_\chi^2}{s}\right) \times \frac{s}{(s - m_{A'}^2)^2 + m_{A'}^2 \Gamma_{A'}^2} \sqrt{\frac{1 - \frac{4m_e^2}{m_{A'}^2}}{1 - \frac{4m_\chi^2}{m_{A'}^2}}}$$



$$\langle \sigma_{\chi\chi \rightarrow ff} v \rangle \simeq \varepsilon^2 \alpha \alpha_D \frac{16\pi m_\chi^2}{(4m_\chi^2 - m_{A'}^2)^2}$$

$$\Omega_\chi h^2 \approx \frac{2.5 \times 10^{-10} \text{GeV}^{-2}}{\langle \sigma_{\chi\chi \rightarrow ff} v \rangle}$$

$$y \equiv \varepsilon^2 \alpha_D \left(\frac{m_\chi}{m_{A'}} \right)^4$$

$$\langle \sigma_{\chi\chi \rightarrow ff} v \rangle \simeq \frac{16\pi \alpha y}{m_\chi^2}$$

Por otro lado, la producción iónica viene dada por:

$$\sigma_e = \frac{16\pi \mu_{\chi e}^2 \alpha \alpha_D \varepsilon^2}{(m_{A'}^2 + \alpha^2 m_e^2)^2} |F(q^2)|^2$$

$$F(q^2) = \frac{m_{A'}^2 + \alpha^2 m_e^2}{m_{A'}^2 + q^2}$$

$$\frac{dR}{d\ln E} = N_T \frac{\rho_\chi}{m_\chi} \frac{d\langle \sigma_e v \rangle}{d\ln E}$$

Desde una perspectiva relativista, la partícula supermasiva en cuanto a su densidad, queda expresada así:

$$\sigma_{p.e.} = 4\alpha^4 \sqrt{2} Z^5 \frac{8\pi r_e^2}{3} \left(\frac{m_e}{\omega} \right)^{7/2}$$

$$\sigma_{A'} = \varepsilon^2 \sigma_{p.e.}$$

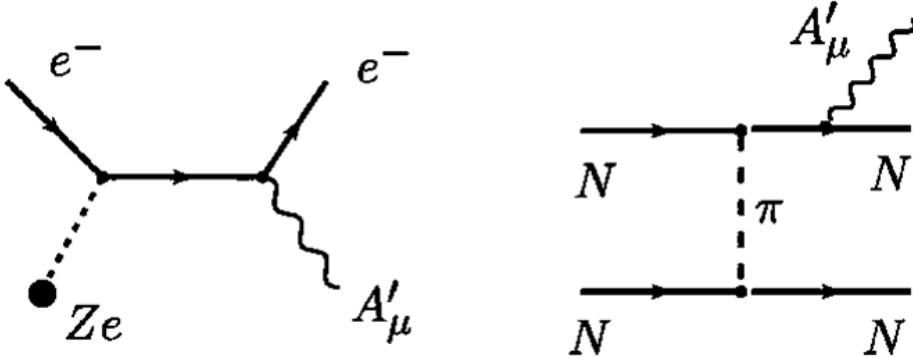
$$\Gamma_{A'} = \frac{\rho_{A'}}{m_{A'}} \sigma_{A'} v_{A'}$$

Ahora bien, la fenomenología relativista de la masa y energía de una partícula supermasiva (nucleosíntesis), se expresa así:

$$d_M^{ij} \equiv |\mathbb{D}_M^{ij}|$$

$$\mathcal{M}_{A'_\mu} \approx \bar{\psi}(\mathbf{k} \times \boldsymbol{\epsilon}) \cdot \boldsymbol{\sigma} \psi$$





$$\mathcal{M}_a \approx \bar{\psi} \mathbf{k} \cdot \boldsymbol{\sigma} \psi$$

$$\sum_{\text{spin}}|\mathcal{M}|^2 = \sum_j Z_j^2 n_j \frac{4\alpha^2\alpha'_{ae}}{\pi} \frac{|\mathbf{p}_1||\mathbf{p}_2|\omega^2}{(\mathbf{q}^2 + \kappa_F^2)^2} \left[2\omega^2 \frac{\mathbf{p}_1 \cdot \mathbf{p}_2 - m_e^2 + (\mathbf{p}_2 - \mathbf{p}_1) \cdot \mathbf{k}}{(\mathbf{p}_1 \cdot \mathbf{k})(\mathbf{p}_2 \cdot \mathbf{k})} + 2 - \frac{\mathbf{p}_1 \cdot \mathbf{k}}{\mathbf{p}_2 \cdot \mathbf{k}} - \frac{\mathbf{p}_2 \cdot \mathbf{k}}{\mathbf{p}_1 \cdot \mathbf{k}} \right]$$

$$\mathcal{Q}/\rho = \frac{\pi^2\alpha^2\alpha'_{ae}}{15} \frac{T^4}{m_e^2} \sum_j Z_j^2 n_j F(\kappa_F) \simeq \alpha'_{ae} 1.08 \times 10^{27} \left(\frac{T}{10^8 \text{ K}} \right)^4 \frac{Z^2}{A} F(\kappa_F)$$

$$F(\kappa_F) \simeq \frac{2+\kappa_F^2}{2} \ln \frac{2+\kappa_F^2}{\kappa_F^2} - 1$$

$$\alpha'_{ae} \leq 3.0 \times 10^{-27}$$

$$\alpha'_{ae} = 2 \frac{1}{4\pi} \left(2 e_D d_M^e \frac{v_h m_e}{\Lambda^2} \right)^2$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D}d_M^e} \gtrsim 4.5 \times 10^6 \text{TeV}^2$$

$$\sum_{\text{spin}}|\mathcal{M}|^2 = \frac{16(4\pi)^3\alpha_\pi^2\alpha'_{aN}}{3m_N^2} \left[\left(\frac{\mathbf{k}^2}{\mathbf{k}^2 + m_\pi^2} \right)^2 + \left(\frac{\mathbf{l}^2}{\mathbf{l}^2 + m_\pi^2} \right)^2 + \frac{\mathbf{k}^2\mathbf{l}^2 - 3(\mathbf{k} \cdot \mathbf{l})^2}{(\mathbf{k}^2 + m_\pi^2)(\mathbf{l}^2 + m_\pi^2)} \right]$$

$$\sum_{\text{spin}}|\mathcal{M}|^2 = \frac{32(4\pi)^3\alpha_\pi^2\alpha'_{aN}}{m_N^2}$$

$$\mathcal{Q}/\rho \simeq \alpha'_{aN} 1.74 \times 10^{33} \frac{\rho}{10^{15}} \left(\frac{T}{\text{MeV}} \right)^6$$

$$\alpha'_{aN} \leq 1.3 \times 10^{-18}$$

$$\alpha'_{aN} = 2 \frac{1}{4\pi} \left(2 e_D d_M^q \frac{v_h m_N}{\Lambda^2} \right)^2$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D}d_M^q} \gtrsim 4.3 \times 10^5 \text{TeV}^2$$

$$\alpha'_{aN} \geq 0.7\times 10^{-14}$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D}d_M^q}\lesssim 5.9\times 10^3 {\mathrm{TeV}}^2$$

$$N_{\mathrm{eff}}=2.878\pm0.278$$

$$H(T_d)=\frac{T_d^2}{M_{Pl}}\biggl(\frac{\pi^2}{90}g_*(T_d)\biggr)^{1/2}$$

$$\Gamma_{A'} = n_{A'} \langle \sigma v \rangle$$

$$n_{A'}=\frac{2\zeta(3)}{\pi^2}T^3$$

$$\langle \sigma v \rangle \simeq \frac{\alpha_D d_M^2 v_h^2}{\Lambda^4}$$

$$\frac{2\zeta(3)}{\pi^2}T_d^3\langle \sigma v \rangle < \frac{T_d^2}{M_{Pl}}\biggl(\frac{2\pi^2}{45}g_*(T_d)\biggr)^{1/2}$$

$$\left(\frac{T_{BBN}}{T_d}\right)^4=\left(\frac{g_*(T_{BBN})}{g_*(T_d)}\right)^{4/3}<\frac{7}{4}\Delta N_{\mathrm{eff}}$$

$$g_*(T_d)>(43/7)^{4/3}\Delta N_{\mathrm{eff}}^{-3/4}$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D}d_M^\ell}\geq 6.6\times 10^3 {\mathrm{TeV}}^2$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D}d_M^q}\geq 4.3\times 10^3 {\mathrm{TeV}}^2$$

$$V(r)=-\frac{\alpha_Dv^2d_M^ad_M^b}{4\Lambda^4r^3}[\boldsymbol{\sigma}_a\cdot\boldsymbol{\sigma}_b-3(\boldsymbol{\sigma}_a\cdot\hat{\boldsymbol{r}})(\boldsymbol{\sigma}_b\cdot\hat{\boldsymbol{r}})]$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D}d_M^e}\gtrsim 872 {\mathrm{GeV}}^2$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D}d_M^e}\gtrsim 1.61 {\mathrm{TeV}}^2$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D}\sqrt{d_M^ed_M^q}}\gtrsim 1.94 {\mathrm{TeV}}^2$$

$$a_\ell^{A'}=-\frac{3}{2}\frac{\alpha_D}{\pi}\biggl(\frac{m_\ell v_h d_M^\ell}{\Lambda^2}\biggr)^2\biggl[\frac{5}{4}+\log\frac{\mu^2}{m^2}\biggr]$$



$$\delta_{\Delta a_e} < 8.1 \times 10^{-13}$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D} d_M^e} \gtrsim 0.075 \text{TeV}^2$$

$$\Delta a_\mu < 2.74 \times 10^{-9}$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D} d_M^\mu} \gtrsim 0.5 \text{TeV}^2$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D} d_M^\mu} \simeq 0.27 \text{TeV}^2$$

$$\text{BR}(\mu \rightarrow e X^0) < 5.8 \times 10^{-5}$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D} d_M^{ue}} \gtrsim 5.1 \times 10^5 \text{TeV}^2$$

$$\text{BR}(K^+ \rightarrow \pi^+ \nu \bar{\nu}) < 1.85 \times 10^{-10}$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D} d_M^{sd}} \gtrsim 9.5 \times 10^6 \text{TeV}^2$$

$$6.5 \times 10^{-14} \leq \alpha'_{aN} \leq 8.0 \times 10^{-8}$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D} d_M^q} \gtrsim 1.9 \times 10^3 \text{TeV}^2$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D} d_M^e} \gtrsim 1.2 \text{TeV}^2$$

$$\frac{\Lambda^2}{\sqrt{\alpha_D} d_M^q} \gtrsim 4.3 \text{TeV}^2$$

En la que, la interacción y acoplamiento de Yukawa, en relación a la partícula supermasiva, en lagrangiano queda expresada así:

$$\mathcal{L} \supset -g_L (\phi_L^\dagger \bar{\chi}_R l_L + S_L^{U\dagger} \bar{Q}_R^U q_L + S_L^{D\dagger} \bar{Q}_R^D q_L) - g_R (\phi_R^\dagger \bar{\chi}_L e_R + S_R^{U\dagger} \bar{Q}_L^D u_R + S_R^{U\dagger} \bar{Q}_L^D d_R) + \text{H.c}$$

$$\mathcal{L} \supset -\lambda_s S_0 (H^\dagger \phi_R^\dagger \phi_L + \tilde{H}^\dagger S_R^{U\dagger} S_L^U + H^\dagger S_R^{D\dagger} S_L^D) + \text{h.c.}$$

Cuyos escalares, son:

$$m_\pm = m_{\phi,S} \sqrt{1 \pm \eta_s}$$

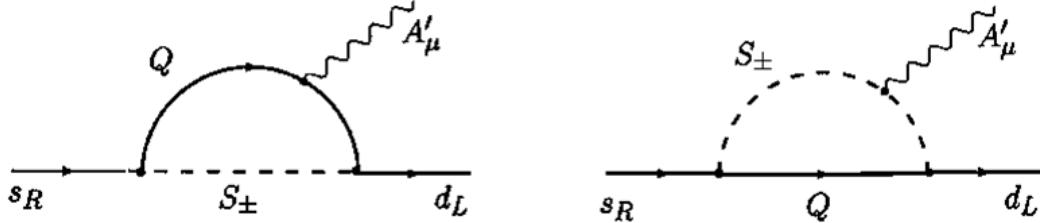
$$\eta_{\phi,S} \equiv \frac{\lambda_s \mu_{\phi,S} v_h}{m_S^2}.$$



$$\mathcal{L}^{(lep)} \supset -g_L \phi_{L\nu}^\dagger (\bar{\chi}_R \nu_L) - \frac{g_L}{\sqrt{2}} (\phi_+^\dagger + \phi_-^\dagger) (\bar{\chi}_R e_L) - \frac{g_R}{\sqrt{2}} (\phi_+^\dagger - \phi_-^\dagger) (\bar{\chi}_L e_R) + \text{h.c.}$$

$$\eta_{\phi,S} < 1 - \left(\frac{M}{m_{\phi,S}}\right)^2.$$

Y cuyos diagramas de vórtice, son:



$$S_L^{U_i^\dagger} \bar{Q}_R^{U_i} q_L^i \rightarrow S_L^{U_i\dagger} \bar{Q}_R^{U_i} (\rho_L^U)_{ij} q_L^j S_R^{U_i^\dagger} \bar{Q}_L^{U_i} q_R^i \rightarrow S_R^{U_i^\dagger\dagger} \bar{Q}_L^{U_i} (\rho_L^U)_{ij} q_R^j$$

$$\frac{v_h}{\Lambda^2} \simeq \frac{m_{Q^i}}{m_S^2}$$

$$\mathbb{D}_M^{ij} = \rho_{jj} \rho_{ij}^* \text{Re} \left[\frac{g_L g_R}{(4\pi)^2} \right] F_M(x, \eta_s)$$

$$F_M(x,y) = \frac{1}{2} [f(x,y) - f(x,-y)]$$

$$f(x,y) = \frac{1-x+y+(1+y)\log\left(\frac{x}{1+y}\right)}{(1-x+y)^2}$$

$$\rho_{nm} \rho_{mm}^* - \rho_{nm}^* \rho_{mm} = 2i \sin \delta_{\text{CP}}$$

$$m_S^i \gtrsim 940 \text{GeV}$$

$$m_\phi \gtrsim 290 \text{GeV}$$

$$\frac{m_\phi^2/m_{\chi^e}}{\sqrt{\alpha_D \alpha_L \alpha_R} |\rho_{ee}|^2 |F_M(x_e, \eta_\phi)|} \gtrsim 2.1 \times 10^6 \text{TeV},$$

$$\frac{m_S^2/m_{Q^u}}{\sqrt{\alpha_D \alpha_L \alpha_R} |\rho_{uu}|^2 F_M(x_u, \eta_S)} \gtrsim 2.0 \times 10^5 \text{TeV},$$

$$\frac{m_\phi^2/m_{\chi^e}}{\sqrt{\alpha_L \alpha_R} |\rho_{ee}|^2 G_M(x_e, \eta_\phi)} \gtrsim 9.8 \times 10^4 \text{TeV}$$

$$G_M(x,y) = \frac{1}{2} [g(x,y) - g(x,-y)]$$



$$g(x, y) = \frac{(1+y)^2 - x^2 + 2x(1+y)\log\left(\frac{x}{1+y}\right)}{2(x-1-y)^3}$$

$$\frac{m_\phi^2/m_{\chi^\mu}}{\sqrt{\alpha_L \alpha_R} |\rho_{\mu\mu}|^2 G_M(x_\mu, \eta_\phi)} \gtrsim 6.3 \times 10^3 \text{TeV}$$

$$\frac{m_\phi^2/m_{\chi^\mu}}{\sqrt{\alpha_L \alpha_R} |\rho_{\mu\mu} \rho_{\mu e}^*| G_M(x_\mu, \eta_\phi)} \gtrsim 4.9 \times 10^8 \text{TeV}.$$

$$\frac{m_S^2/m_{Q^b}}{\sqrt{\alpha_L \alpha_R} |\rho_{bb} \rho_{bs}^*| G_M(x_b, \eta_S)} \gtrsim 1.3 \times 10^4 \text{TeV},$$

$$\frac{m_S^2}{(\alpha_L^2 + \alpha_R^2) |\rho_{ss} \rho_{sd}^*|^2} \gtrsim 3 \times 10^5 \text{TeV}^2$$

La estructura morfológica de una partícula supermasiva, viene dada por:

1. Modelo Drell-Yan.

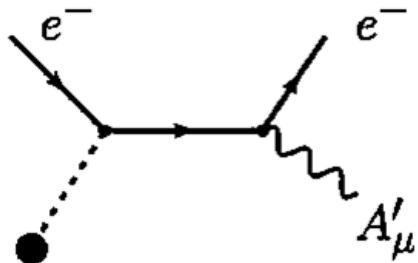
$$\mathcal{L} = -\varepsilon e J^\mu A'_\mu$$

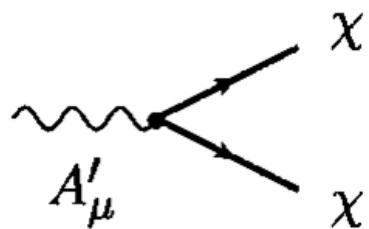
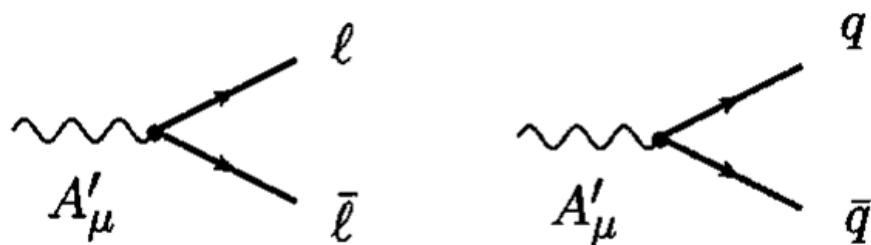
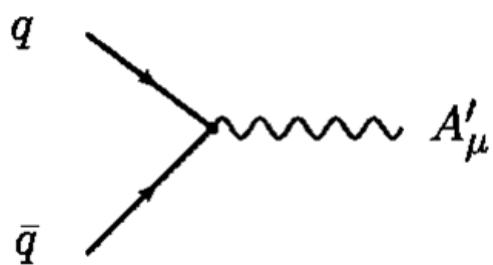
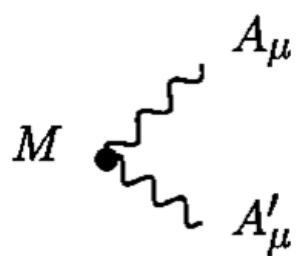
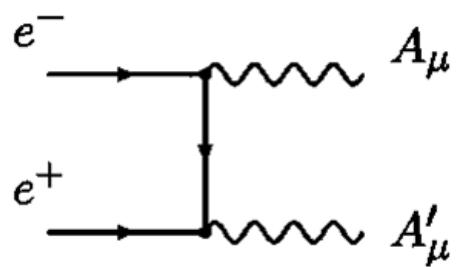
$$\Gamma(A' \rightarrow \ell^+ \ell^-) = \frac{1}{3} \alpha \varepsilon^2 m_{A'} \sqrt{1 - \frac{4m_\ell^2}{m_{A'}^2}} \left(1 + \frac{2m_\ell^2}{m_{A'}^2}\right),$$

$$\Gamma(A' \rightarrow \text{partícula-supermasiva}) = \frac{1}{3} \alpha \varepsilon^2 m_{A'} \sqrt{1 - \frac{4m_\mu^2}{m_{A'}^2}} \left(1 + \frac{2m_\mu^2}{m_{A'}^2}\right) R$$

$$\Gamma(A' \rightarrow \chi \bar{\chi}) = \frac{1}{3} \alpha_D m_{A'} \sqrt{1 - \frac{4m_\chi^2}{m_{A'}^2}} \left(1 + \frac{2m_\chi^2}{m_{A'}^2}\right)$$

2. Diagramas de caída de masa.





3. Singularidades cuánticas.



$$\frac{\varepsilon}{2}F_{\mu\nu}F^{\mu\nu},$$

$$\frac{a}{f_a}F_{\mu\nu}\tilde{F}^{\mu\nu}+\frac{1}{f_a}\partial_\mu a\bar{\psi}\gamma^\mu\gamma_5\psi,$$

$$(\mu S + \lambda S^2) H^\dagger H,$$

$$y_N \bar L H N,$$

4. Superdensidad de masa.

$$\Gamma = n \sigma v,$$

$$H(T)=\frac{\pi\sqrt{g_*(T)}}{\sqrt{90}}\frac{T^2}{m_{Pl}}$$

$$g_*(T)=\sum_{\text{bosons}} g_b\left(\frac{T_b}{T}\right)^4+\frac{7}{8}\sum_{\text{fermions}} g_f\left(\frac{T_f}{T}\right)^4,$$

$$n_{eq}(T)=g_*\int\,\frac{d^3p}{(2\pi)^3}\frac{1}{e^{E/T}\pm1}=\begin{cases}\displaystyle g_*\left(\frac{mT}{2\pi}\right)^{3/2}e^{-m/T}&\text{no relativista } (T\ll m)\\\displaystyle \frac{\zeta(3)}{\pi^2}g_*T^3&\text{bosones relativistas } (T\gg m)\\\displaystyle \frac{3\zeta(3)}{4\pi^2}g_*T^3&\text{fermiones relativistas } (T\gg m),\end{cases}$$

$$\dot{n}(t)+3H(t)n(t)=-\langle\sigma_{\chi\chi\rightarrow ff}v\rangle\big(n^2(t)-n_{eq}^2(t)\big),$$

$$\langle\sigma_{\chi\chi\rightarrow ff}v\rangle=\frac{\int_{\infty}^{4m_{\chi}^2}ds\sqrt{s}(s-4m_{\chi}^2)K_1\left(\frac{\sqrt{s}}{T}\right)\sigma_{\chi\chi\rightarrow ff}}{8m_{\chi}^4T\left[K_2\left(\frac{m_{\chi}}{T}\right)\right]^2}$$

$$\langle\sigma_{\chi\chi\rightarrow ff}v\rangle=\langle s_0+s_1v^2+O(v^4)\rangle$$

$$x=m_\chi/T=\sqrt{2tH(T=m_\chi)}$$

$$\frac{dY}{dx}=-\frac{\lambda(x)}{x^2}\big[Y^2(x)-Y_{eq}^2\big]$$

$$\lambda(x)=\frac{m_{\chi}^3\langle\sigma_{\chi\chi\rightarrow ff}v\rangle}{H(T=m_\chi)},$$

$$\langle\sigma_{\chi\chi\rightarrow ff}v\rangle=\sigma_{\chi\chi\rightarrow ff}v+O(v^2)$$

$$\lambda(x)=\frac{\sqrt{180}m_{Pl}m_\chi}{\pi\sqrt{g_*x}}\sigma_{\chi\chi\rightarrow ff}$$



$$H(T=m_\chi)=\frac{\pi\sqrt{g_*(T=m_\chi)}}{90}\frac{m_\chi^2}{m_{Pl}}.$$

$$Y(x')=\frac{x_d}{\lambda}$$

$$\rho_\chi = m_\chi n(x') = m_\chi^4 \frac{Y(x')}{28x_d}$$

$$\Omega_\chi h^2 \,\simeq 0.12 \frac{x_d}{23} \frac{\sqrt{g_*}}{10} \frac{1.7\times 10^{-9} {\rm GeV}^{-2}}{\langle \sigma_{\chi\chi \rightarrow ff} v\rangle} \simeq \frac{2.5\times 10^{-10} {\rm GeV}^{-2}}{\langle \sigma_{\chi\chi \rightarrow ff} v\rangle}$$

$$\begin{aligned} \mathcal{Q} = \prod_{i=1} \int \frac{d^3 \mathbf{p}_i}{2E_i(2\pi)^3} f_i(E_i) \prod_{f=1} \int \frac{d^3 \mathbf{p}_f}{2E_f(2\pi)^3} [1 \pm f_f(E_f)] \int \frac{d^3 \mathbf{p}_a}{2\omega_a(2\pi)^3} \omega_a \\ \times \frac{1}{\mathcal{S}} \sum_{\text{spin and pol.}} |\mathcal{M}|^2 (2\pi)^4 \delta^4 \left(\sum p_i - \sum p_f - p_a \right) \end{aligned}$$

$$n_j(E_j) = g_j \int \frac{d^2 \mathbf{p}_j}{(2\pi)^3} f(E_j)$$

$$L = \int \; dV \mathcal{Q} e^{-\tau}$$

$$\Pi^{\mu\nu}(k)\,=\,16\pi\alpha\int\;\frac{d^3\mathbf{p}_i}{(2\pi)^3}\frac{1}{2E}[n_e(E)+n_{\bar{e}}(E)]\times\frac{p\cdot k(p^\mu k^\nu+k^\mu p^\nu)-k^2p^\mu p^\nu-(p\cdot k)^2g^{\mu\nu}}{(p\cdot k)^2-(k^2)^2/4}$$

$$\Pi_T(\omega,{\bf k})=\frac{1}{2}\big(\delta^{ij}-k^ik^j\big)\Pi^{ij}(\omega,{\bf k})$$

$$\Pi_L(\omega,{\bf k})=\Pi^{00}(\omega,{\bf k})$$

$$D^{00}(\omega,k)=\frac{1}{k^2-\Pi_L(\omega,k)}$$

$$D^{ij}(\omega,k)=\frac{1}{k^2-\Pi_T(\omega,k)}\big(\delta^{ij}-k^ik^j\big)$$

$$\omega_T^2=k^2+\Pi_T(\omega_T,k)\;\;\text{and}\;\;\omega_L^2=\frac{\omega_L^2}{k^2}+\Pi_T(\omega_L,k)$$

5. Momentum de Fermi.

$$\Pi_T(\omega,{\bf k})=\omega_P^2\frac{3\omega^2}{2v_F^2k^2}\bigg(1-\frac{\omega^2-v_F^2k^2}{2v_F\omega k}\log\frac{\omega+v_Fk}{\omega-v_Fk}\bigg)$$

$$\Pi_L(\omega,{\bf k})=\omega_P^2\frac{3\omega}{2v_F^3k}\Big(\frac{\omega}{2v_Fk}\log\frac{\omega+v_Fk}{\omega-v_Fk}-1\Big)$$



$$\mathcal{Q} = - \int \frac{d^3\mathbf{k}}{(2\pi)^3} \frac{\text{Im}\Pi_L(\omega, \mathbf{k}) + \text{Im}\Pi_T(\omega, \mathbf{k})}{\omega(e^{\omega/T} - 1)}$$

6. Modelo inflacionario - relativista de una partícula supermasiva a propósito de la superdensidad de su masa y la superdensidad de su energía.

$$\frac{\Omega_A}{\Omega_{\text{DM}}} \simeq \sqrt{\frac{m}{6 \cdot 10^{-6} \text{eV}}} \left(\frac{H_I}{10^{14} \text{GeV}} \right)^2$$

$$S = \int dt d^3x \sqrt{-g} \left[-\frac{1}{4} g^{\mu\rho} g^{\nu\sigma} F_{\mu\nu} F_{\rho\sigma} + \frac{1}{2} m^2 g^{\mu\nu} A_\mu A_\nu \right]$$

$$\begin{aligned} S_T &= \int \frac{a^3 d^3 k dt}{(2\pi)^3} \frac{1}{2a^2} \left[|\partial_t \vec{A}_T|^2 - \left(\frac{k^2}{a^2} + m^2 \right) |\vec{A}_T|^2 \right] \\ S_L &= \int \frac{a^3 d^3 k dt}{(2\pi)^3} \frac{1}{2a^2} \left[\frac{a^2 m^2}{k^2 + a^2 m^2} |\partial_t A_L|^2 - m^2 |A_L|^2 \right] \end{aligned}$$

$$\langle X^*(t, \vec{k}) X(t, \vec{k}') \rangle \equiv (2\pi)^3 \delta^3(\vec{k} - \vec{k}') \frac{2\pi^2}{k^3} \mathcal{P}_X(t, k)$$

$$\langle \rho \rangle = \int d\log k \frac{1}{2a^2} \left[\frac{a^2 m^2}{k^2 + a^2 m^2} \mathcal{P}_{\partial_t A_L} + m^2 \mathcal{P}_{A_L} \right]$$

$$\left[\partial_t^2 + \frac{3k^2 + a^2 m^2}{k^2 + a^2 m^2} H \partial_t + \frac{k^2}{a^2} + m^2 \right] A_L = 0$$

$$\mathcal{P}_{A_L}(t, k) = \left(\frac{k H_I}{2\pi m} \right)^2 \left(\frac{A_L(t, k)}{A_{L,0}} \right)^2$$

$$\mathcal{P}_{A_L}(t, k) \simeq \left(\frac{k_* H_I}{2\pi m} \right)^2 \left(\frac{a_*}{a} \right) \frac{(k/k_*)^2}{1 + (k/k_*)^3}$$

$$a_0 \lambda_* = \frac{2\pi a_0}{m a_*} \simeq 10^{11} \text{ km} \left(\frac{10^{-5} \text{eV}}{m} \right)^{1/2}$$

$$\delta(x) \equiv \frac{\rho(x) - \bar{\rho}}{\bar{\rho}}$$

$$\mathcal{P}_\delta(t, k) \simeq \frac{\sqrt{3}(k/k_*)^3}{\pi((k/k_*)^{3/2} + 1)^{8/3}}$$



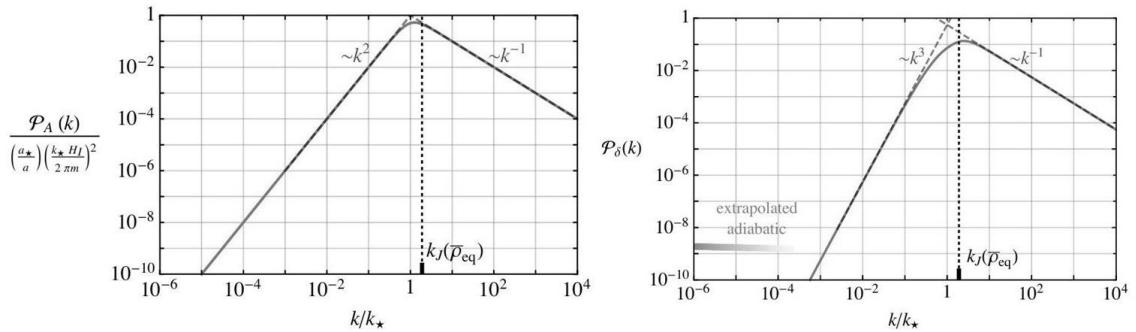


Figura 1. Fluctuaciones inflacionarias de la partícula supermasiva en un campo de gauge específico.

7. Modelo Schrödinger-Poisson para una partícula supermasiva.

$$A_i \equiv \frac{1}{\sqrt{2m^2a^3}}(\psi_i e^{-imt} + \text{c.c.}) \left(i\partial_t + \frac{\nabla^2}{2m} - m\Phi \right) \psi_i = \nabla^2 \Phi = \frac{4\pi G}{a} \sum_i (|\psi_i|^2 - \langle |\psi_i|^2 \rangle)$$

$$\begin{aligned} \partial_t \rho_i + 3H\rho_i + a^{-1}\nabla \cdot (\rho_i \vec{v}_i) &= 0 \\ \partial_t \vec{v}_i + H\vec{v}_i + a^{-1}(\vec{v}_i \cdot \nabla)\vec{v}_i &= -a^{-1}(\nabla\Phi + \nabla\Phi_{Qi}) \\ \nabla^2 \Phi &= 4\pi G a^2 (\rho - \bar{\rho}) \end{aligned}$$

$$\Phi_{Qi} \equiv -\frac{\hbar^2}{2a^2m^2} \frac{\nabla^2 \sqrt{\rho_i}}{\sqrt{\rho_i}}$$

$$k_J(\rho) \equiv a(16\pi G \rho m^2)^{1/4}$$

$$\left. \frac{k_J(\bar{\rho})}{k_*} \right|_{a=a_{\text{eq}}} = \frac{(16\pi G \bar{\rho}_{\text{eq}} m^2)^{1/4}}{m(a_*/a_{\text{eq}})} = g_R \left(12 \frac{\bar{\rho}_{\text{eq}}}{\rho_{\text{eq}}^{\text{tot}}} \right)^{1/4} = g_R \left(6 \frac{\Omega_A}{\Omega_M} \right)^{1/4}$$

$$\left. \frac{k_J(\bar{\rho})}{k_*} \right|_{a=a_{\text{eq}}} \simeq 1.9$$

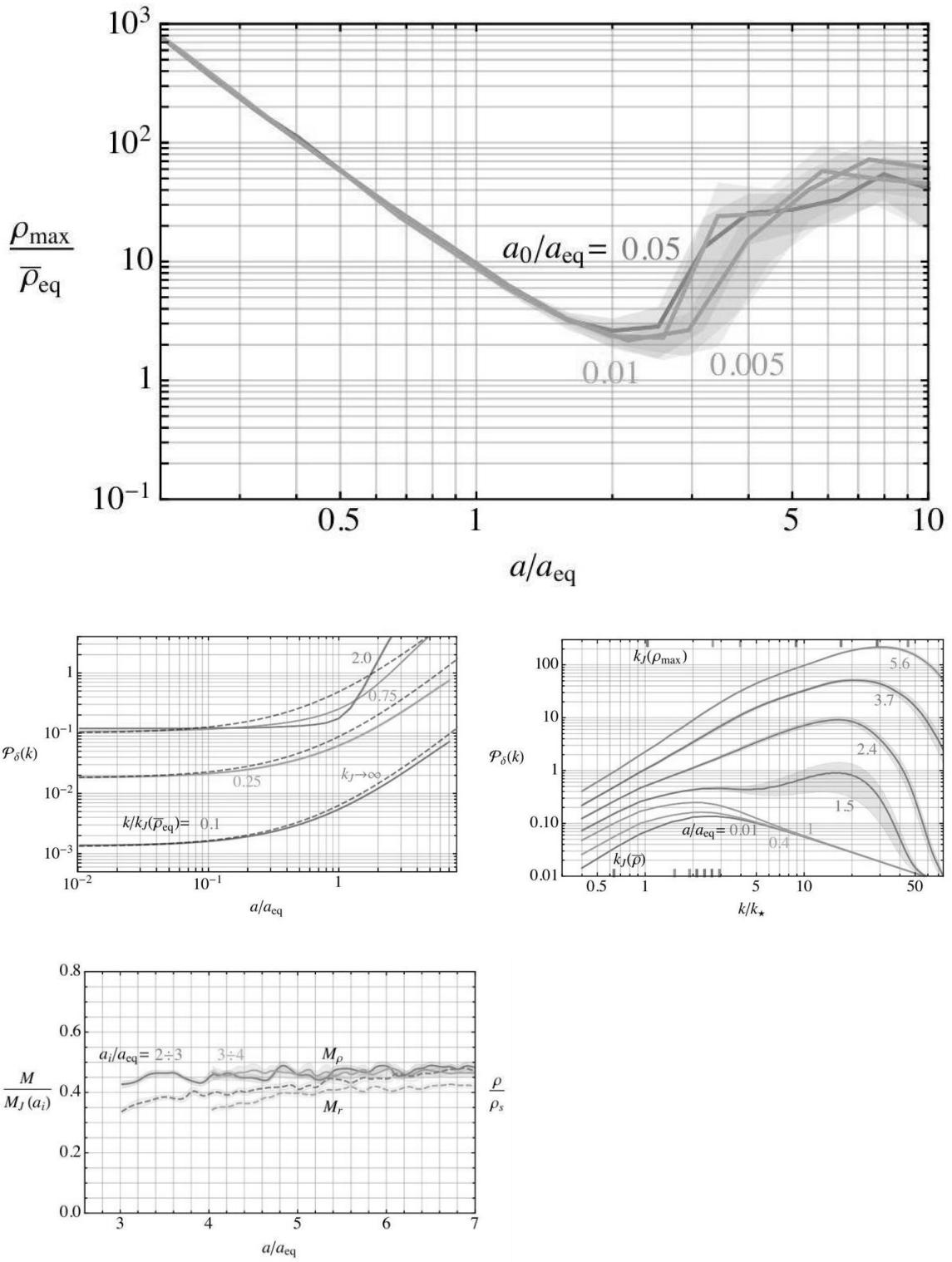
$$H^2 = \frac{8\pi G}{3} \rho_{\text{tot}} = \frac{8\pi G}{3} \frac{\bar{\rho}_{\text{eq}}}{2} \left[\left(\frac{a_{\text{eq}}}{a} \right)^3 + \left(\frac{a_{\text{eq}}}{a} \right)^4 \right] \equiv \frac{H_{\text{eq}}^2}{2} \left[\left(\frac{a_{\text{eq}}}{a} \right)^3 + \left(\frac{a_{\text{eq}}}{a} \right)^4 \right]$$

$$\frac{da}{dt} = a \frac{H_{\text{eq}}}{\sqrt{2}} \sqrt{\left(\frac{a_{\text{eq}}}{a} \right)^3 + \left(\frac{a_{\text{eq}}}{a} \right)^4}$$

$$\begin{aligned} \left(i\partial_{\tilde{t}} + \frac{\nabla'^2}{2} - \Phi' \right) \psi'_i &= 0 \\ \nabla'^2 \Phi' &= a \sum_i (|\psi'_i|^2 - \langle |\psi'_i|^2 \rangle) \end{aligned}$$

$$\frac{da}{d\tilde{t}} = \beta a^3 \sqrt{\left(\frac{a_{\text{eq}}}{a} \right)^3 + \left(\frac{a_{\text{eq}}}{a} \right)^4}$$

$$\langle |\psi'|^2 \rangle = \frac{\langle |\psi|^2 \rangle}{(\sqrt{4\pi G T})^{-2}} = \frac{f \bar{\rho}_{\text{eq}}/2}{(\sqrt{4\pi G} \sqrt{2} \beta H_{\text{eq}}^{-1})^{-2}} = \frac{3\Omega_A/\Omega_M}{2} \beta^2$$



Figuras 1, 2 y 3. Fluctuaciones del Modelo Schrödinger-Poisson para una partícula supermasiva.

8. Modelo Gravitacional de una partícula supermasiva a propósito de la superdensidad de su masa y energía respectivamente.



$$M(a) = c_M M_J(a), \text{ con } M_J(a) \equiv \frac{4\pi}{3} \bar{\rho} a^3 \lambda_J^3(a) \propto a^{-3/4}$$

$$\mathcal{N} \equiv \frac{1}{m} \int d^3x \sum_i |\psi_i|^2$$

$$E = M + \int d^3x \sum_i \left(\frac{1}{2m^2} |\nabla \psi_i|^2 + \frac{1}{2} \Phi |\psi_i|^2 \right)$$

$$\psi_i = \frac{m}{\sqrt{4\pi G}} e^{-i\gamma mt} \chi_1(mr) u_i, \Phi = \Phi_1(mr)$$

$$M \simeq \frac{2\alpha}{Gm}, R \simeq \frac{1.9}{\alpha m}$$

$$MR \simeq \frac{3.9}{Gm^2}$$

$$\rho_s = \frac{\alpha^4 m^2}{4\pi G} \simeq \frac{1}{Gm^2 R^4} \simeq \frac{G^3 m^6 M^4}{64\pi}$$

$$\begin{aligned} L_p &= \frac{i}{2m} \epsilon_{pqr} \int d^3x (\psi_m^* \partial_q \psi_m x^r - \text{c.c.}) \\ S_p &= \frac{i}{m} \epsilon_{pqr} \int d^3x \psi_q \psi_r^* \end{aligned}$$

$$\epsilon_{\hat{z}}^{(\pm 1)} = \frac{1}{\sqrt{2}} \begin{pmatrix} 1 \\ \pm i \\ 0 \end{pmatrix}, \epsilon_{\hat{z}}^{(0)} = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$$

$$\psi_i = \frac{m}{\sqrt{4\pi G}} e^{-i\gamma mt} \sum_{\lambda} \chi^{(\lambda)}(\vec{x}) \left(\epsilon_{\hat{z}}^{(\lambda)} \right)_i$$

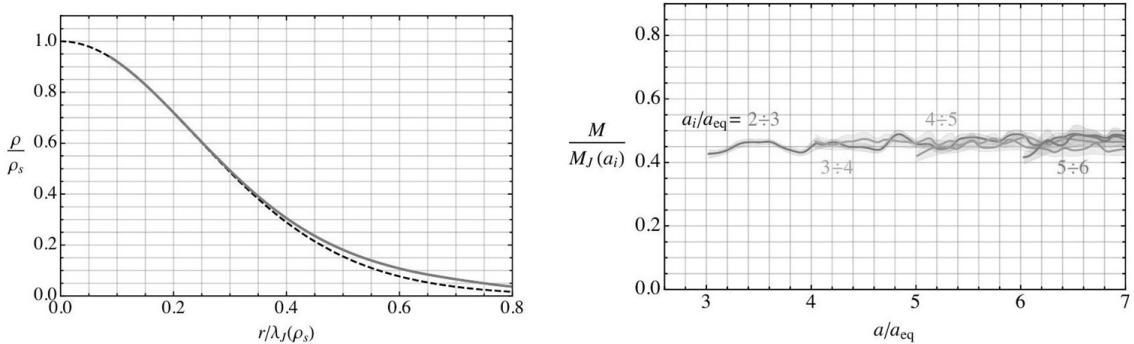


Figura 1. Fluctuaciones de la superdensidad de masa y energía de una partícula supermasiva en un campo de gauge específico.

9. Modelo de distribución de la masa en relación a una partícula supermasiva.



$$M_J^{\text{eq}} \equiv M_J(a_{\text{eq}}) = 5.2 \cdot 10^{-23} M_{\odot} \left(\frac{\text{eV}}{m} \right)^{3/2}$$

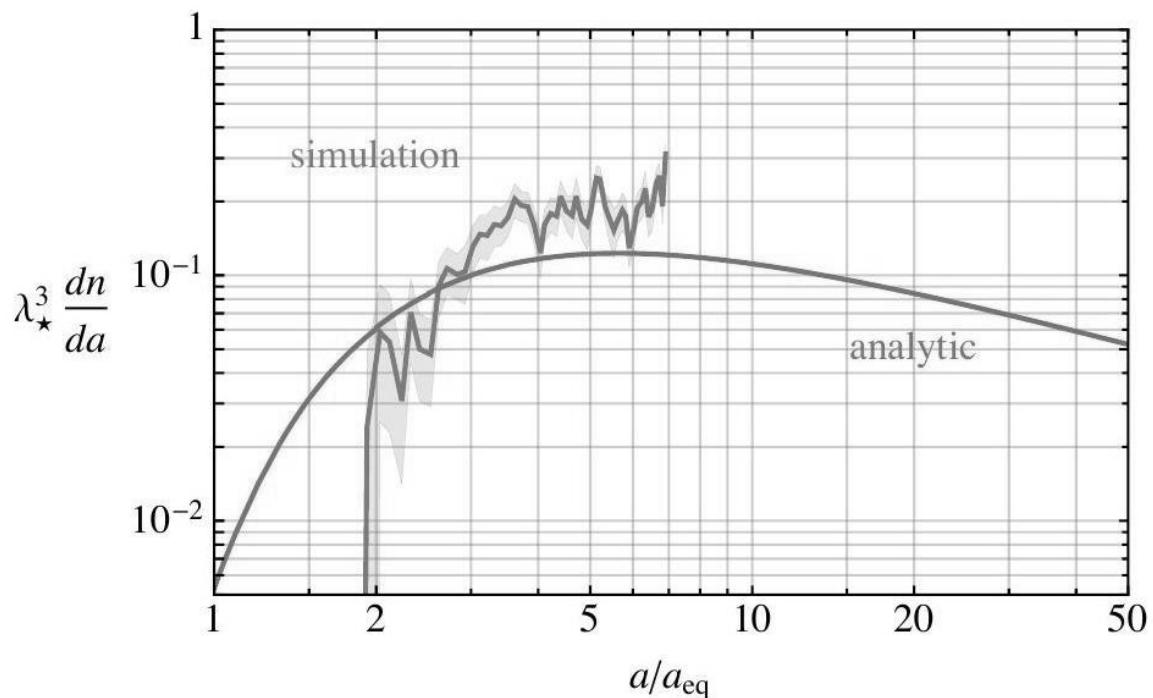
$$M(a) = 2.3 \cdot 10^{-23} M_{\odot} \left(\frac{c_M}{0.45} \right) \left(\frac{a_{\text{eq}}}{a} \right)^{3/4} \left(\frac{\text{eV}}{m} \right)^{3/2}$$

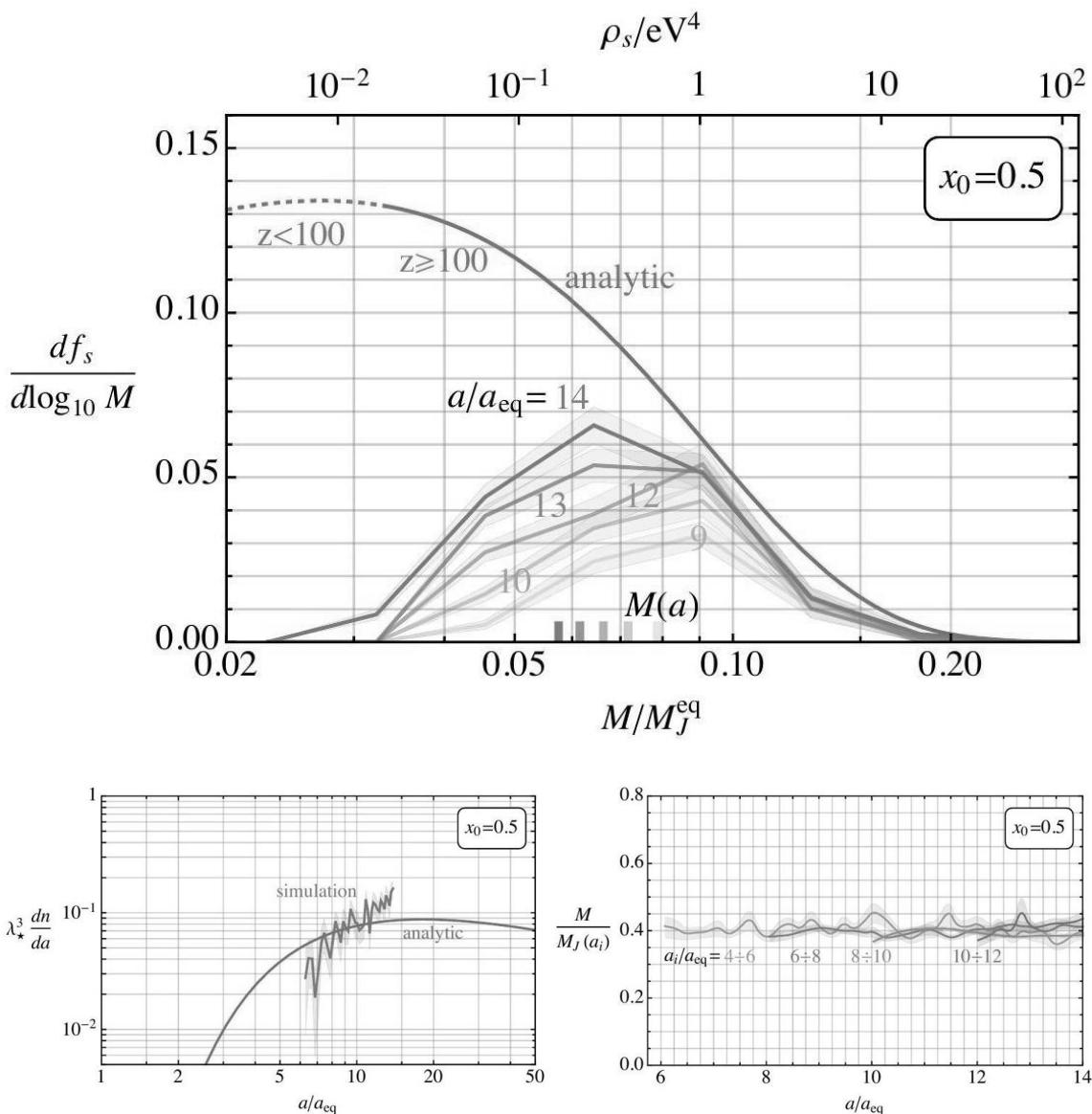
$$\frac{df_s(a,M)}{d\log M} = \frac{M}{\bar{\rho}} \frac{dn(a,M)}{d\log M}$$

$$\frac{\rho(r)}{\rho_s} = \frac{\rho_0}{\frac{r/\lambda_J(\rho_s)}{r_0} \left(1 + \frac{r/\lambda_J(\rho_s)}{r_0} \right)^2}$$

$$\frac{df_s(a_2,M)}{d\log M} = \frac{M^2}{\bar{\rho}(a_2)} \int_{a_1}^{a_2} da \frac{d\Pi_{\delta>\delta_c}(a_1,a)}{da} \frac{\bar{\rho}(a)}{M(a)} \left(\frac{a}{a_2} \right)^3 \delta[M - M(a)] = \frac{d\Pi_{\delta>\delta_c}(a_1,a_M)}{d\log M}$$

$$\frac{df_s}{d\log M} \simeq \frac{\delta_c}{\sqrt{2\pi}\sigma(M)} e^{-\frac{\delta_c^2}{2\sigma^2(M)}} \left| \frac{d\log \sigma(M)}{d\log M} \right|,$$





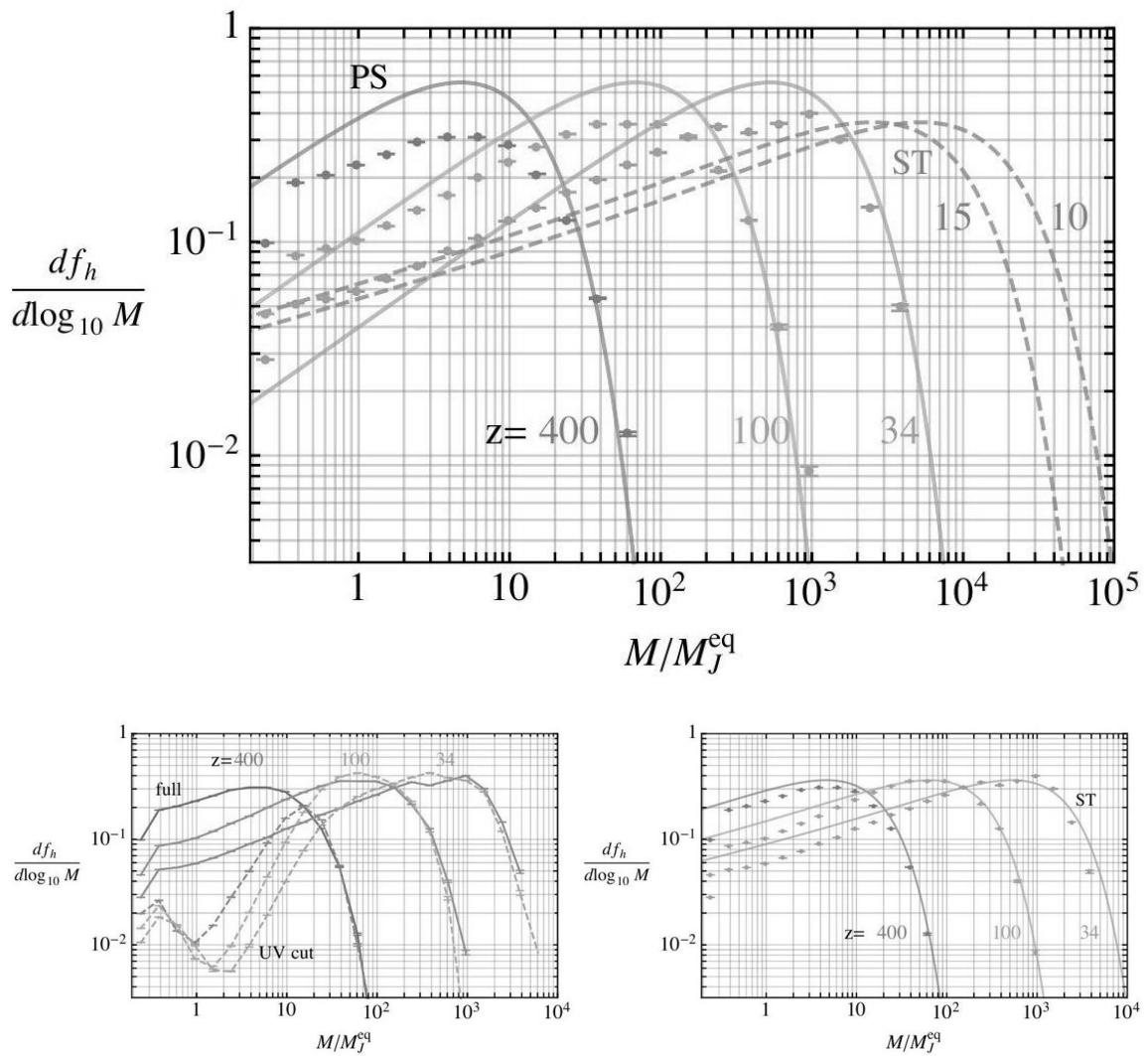
Figuras 1, 2 y 3. Fluctuaciones relativas a la distribución de masa y energía en relación a una partícula supermasiva.

10. Modelo de interacciones gravitacionales Press-Schechter de una partícula supermasiva.

$$\begin{aligned}\partial_t \rho + 3H\rho + a^{-1}\nabla \cdot (\rho \vec{v}) &= 0 \\ \partial_t \vec{v} + H\vec{v} + a^{-1}(\vec{v} \cdot \nabla)\vec{v} &= -a^{-1}\nabla\Phi \\ \nabla^2\Phi &= 4\pi G a^2(\rho - \bar{\rho})\end{aligned}$$

$$\frac{df_h(a, M)}{d \log M} = \sqrt{\frac{2}{\pi}} \nu e^{-\nu^2/2} \left| \frac{d \log \nu}{d \log M} \right| \simeq \sqrt{\frac{8M}{M_\star}} \frac{\pi \delta_c}{3^{3/4} D^2[a]} \exp \left[-\frac{\pi^3 \delta_c^2}{3\sqrt{3} D^2[a] M_\star} \frac{8M}{M_\star} \right]$$





Figuras 1 y 2. Fluctuaciones gravitacionales de una partícula supermasiva.

$$\frac{df_h}{d \log M} = A(p)(1 + (qv^2)^{-p})(qv^2)^{1/2} e^{-qv^2/2} \left| \frac{d \log v}{d \log M} \right|$$

$$\rho_0 = \frac{4.4 \cdot 10^{-3} c_\Delta^3 v_c^3 \Delta}{(c_\Delta + 1)^{-1} + \log(c_\Delta + 1) - 1} \left[\frac{M_J^{\text{eq}}}{M} \right]^{3/2} \bar{\rho}^{\text{eq}}$$

$$r_0 = \frac{4.2}{c_\Delta v_c \Delta^{1/3}} \left[\frac{M}{M_J^{\text{eq}}} \right]^{5/6} \lambda_J^{\text{eq}}$$

$$M \simeq 9.5 M_* \left[\frac{1.7}{\delta_c} \right]^2 \left[\frac{100}{z+1} \right]^2 \simeq 65 M_J^{\text{eq}} \left[\frac{1.7}{\delta_c} \right]^2 \left[\frac{100}{z+1} \right]^2$$

$$M_h = (3 \div 5) \cdot 10^3 M_J^{\text{eq}} \simeq (1.5 \div 3) \cdot 10^{-19} M_\odot (\text{eV/m})^{3/2}$$

$$\rho(r) = \frac{\rho_0}{r/r_0(1+r/r_0)^2}$$

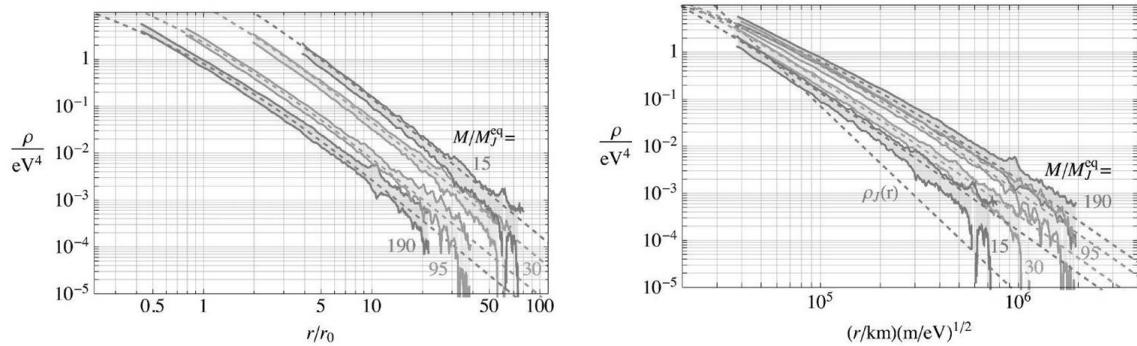


Figura 3: Fluctuaciones de masa a propósito de la gravedad cuántica endógena de la partícula supermasiva.

$$\begin{aligned}\rho_0 &\simeq 0.7 \left[\frac{10^3 M_J^{\text{eq}}}{M} \right]^{3/2} \bar{\rho}^{\text{eq}} \simeq 0.3 \text{eV}^4 \left[\frac{10^3 M_J^{\text{eq}}}{M} \right]^{3/2} \\ r_0 &\simeq 5.4 \left[\frac{M}{10^3 M_J^{\text{eq}}} \right]^{5/6} \lambda_J^{\text{eq}} \simeq 7 \cdot 10^5 \text{ km} \left[\frac{M}{10^3 M_J^{\text{eq}}} \right]^{5/6} \left[\frac{1 \text{eV}}{m} \right]^{1/2} \\ \langle \rho \rangle &\simeq \frac{3\rho_0}{c_{\Delta}^3} (\log c_{\Delta} - 1)\end{aligned}$$

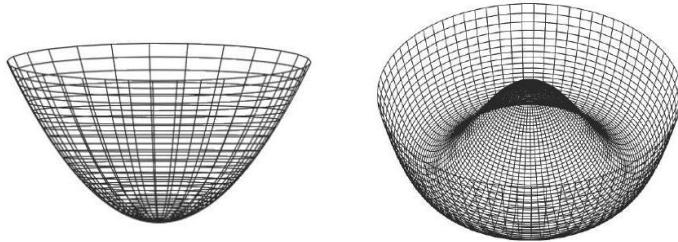


Figura. 4. Simulación de la curvatura del espacio – tiempo cuántico provocado por la interacción de una partícula supermasiva.

11. Modelo de aniquilación de una partícula supermasiva.

$$\begin{aligned}\rho_s &\simeq 1.51 \cdot 10^4 \text{eV}^4 \left(\frac{M}{M_J^{\text{eq}}} \right)^4 \\ \lambda_J(\rho_s) &= 4.6 \cdot 10^3 \text{ km} \left(\frac{\text{eV}}{m} \right)^{1/2} \left(\frac{M_J^{\text{eq}}}{M} \right) \\ n &= f_s \bar{\rho}(t_0)/M \simeq 10^{20} \text{pc}^{-3} \left(\frac{f_s}{0.05} \right) \left(\frac{\rho_{\text{local}}}{0.5 \text{GeV/cm}^3} \right) \left(\frac{0.1 M_J^{\text{eq}}}{M} \right) \left(\frac{m}{\text{eV}} \right)^{3/2}\end{aligned}$$

$$\Gamma \simeq n\pi R^2 v_{\text{rel}} \simeq \frac{0.1}{\text{yr}} \left(\frac{m}{\text{eV}}\right)^{1/2} \left(\frac{0.1M_J^{\text{eq}}}{M}\right)^3 \left(\frac{v_{\text{rel}}}{10^{-3}}\right) \left(\frac{f_s}{0.05}\right) \left(\frac{\rho_{\text{local}}}{0.5 \text{GeV/cm}^3}\right)$$

$$t_{\text{aniquilación}} \simeq 10^2 \text{ s} \left(\frac{0.1 M_J^{\text{eq}}}{M} \right) \left(\frac{\text{eV}}{m} \right)^{1/2}$$

$$\mathcal{N} \simeq \left(\frac{2\pi}{k_\star/a_{\text{eq}}} \right)^3 \frac{\rho_{\text{eq}}}{m} \simeq 90 \frac{\rho_{\text{eq}}^{1/4}}{G^{3/4} m^{5/2}} \simeq 30 \left(\frac{10^{17} \text{eV}}{m} \right)^{5/2} \gg 1$$

$$\mathcal{P}_{A_L}(t,k) = \left(\frac{k_\star H_I}{2\pi m}\right)^2 \left(\frac{a_\star}{a}\right) F_{A_L}^2[k/k_\star] \simeq \left(\frac{k_\star H_I}{2\pi m}\right)^2 \left(\frac{a_\star}{a}\right) \frac{(k/k_\star)^2}{1 + (k/k_\star)^3}$$

$$\mathcal{P}_\delta(t,k) = \frac{k^2}{8\langle A_L^2 \rangle^2} \int_0^\infty dq \int_{|q-k|}^{q+k} dp \frac{(k^2 - q^2 - p^2)^2}{q^4 p^4} \mathcal{P}_{A_L}(t,p) \mathcal{P}_{A_L}(t,q) \simeq \frac{\sqrt{3} (k/k_\star)^3}{\pi ((k/k_\star)^{3/2} + 1)^{8/3}}$$

$$\Delta E = \frac{4}{3} \frac{G^2 M_s^2 r^2}{v_{\text{rel}}^2 b^4} dm$$

$$b_{\text{crit}}^2 = \frac{M_s}{v_{\text{rel}}} \left(\frac{G}{\pi \bar{\rho}(r)} \right)^{1/2}$$

$$p_{\text{dest}} = \pi n \frac{S}{v_{\text{rel}}} \left(\frac{G}{\pi \bar{\rho}(r)} \right)^{1/2}$$

$$p_{\text{dest}} = 0.4 \left(\frac{n}{100} \right) \left(\frac{0.05 \text{eV}^4}{\bar{\rho}(r)} \right)^{1/2} \left(\frac{S}{140 M_\odot \text{pc}^{-2}} \right) \left(\frac{10^{-3}}{v_{\text{rel}}} \right)$$

12. Modelo de superdensidades de masa y energía de una partícula supermasiva.

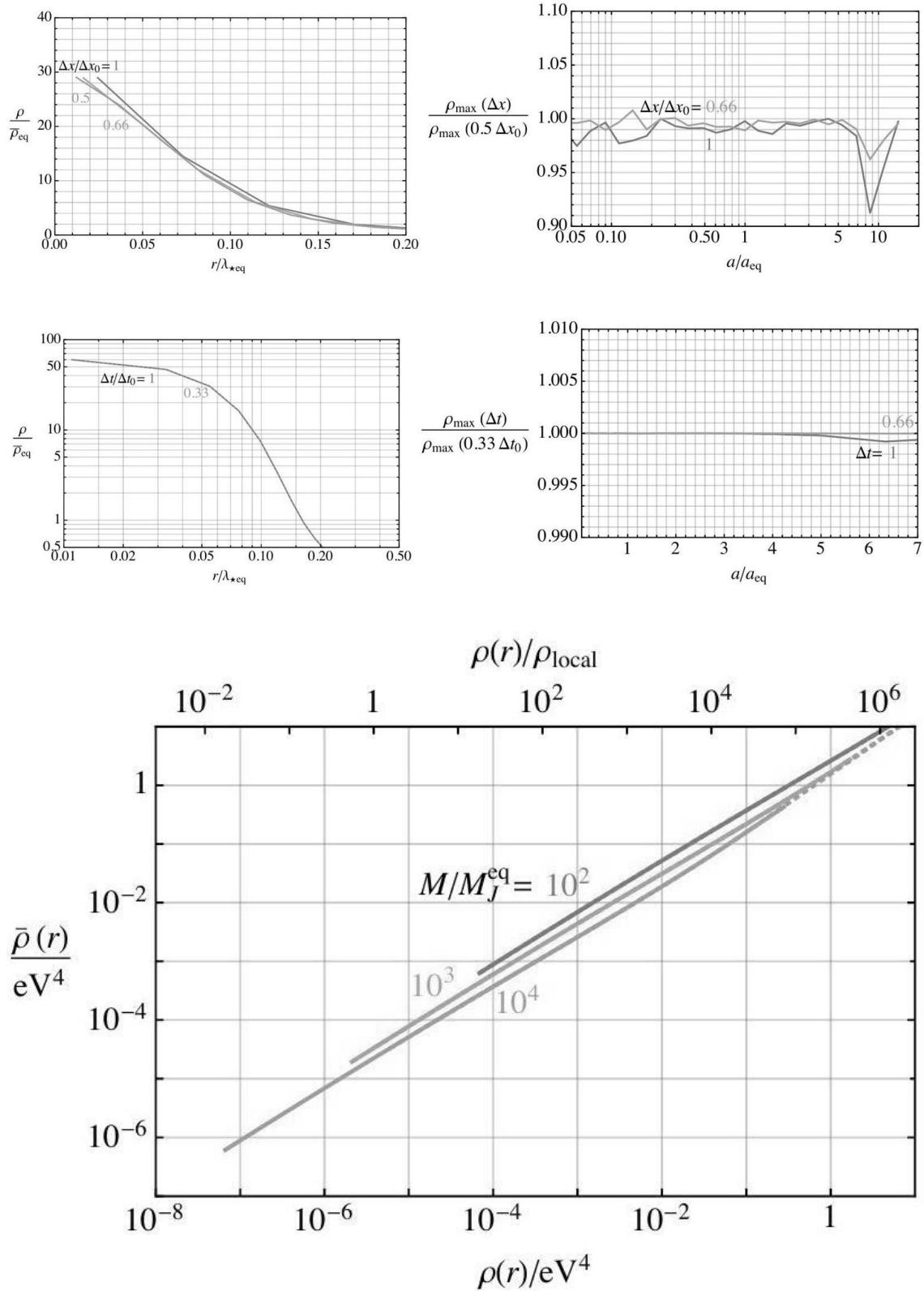
$$\ddot{\delta} + \frac{2}{a} \dot{a} \dot{\delta} - 4\pi G f \rho_{\text{nr}} \delta = 0$$

$$\frac{\partial^2 \delta}{\partial y^2} + \frac{2+3y}{2y(1+y)} \frac{\partial \delta}{\partial y} - \frac{3f}{2y(1+y)} \delta = 0$$

$$\frac{\partial^2 \delta}{\partial y^2} + \frac{2+3y}{2y(1+y)} \frac{\partial \delta}{\partial y} - \frac{3f}{2y(1+y)} \left(-1 + \left(\frac{k}{k_J} \right)^4 \right) \delta$$

$$\psi_i(t + \Delta t) = \left(\prod_{\alpha=1}^8 e^{-i d_\alpha \Delta \tilde{t} \Phi(\mathbf{x})} e^{-i c_\alpha \Delta \tilde{t} k^2 / 2} \right) \psi_i(t)$$





Figuras 1, 2 y 3: Fluctuaciones de superdensidad de masa y energía de una partícula supermasiva.

$$\Delta E = 2(2\pi G \sigma_s(r))^2 \frac{(\Delta z)^2}{v_z^2} dm (1 + a^2)^{-3/2}$$



$$\bar{\rho}(r) \lesssim 10^{-6} \text{eV}^4 \left(\frac{\sigma_s}{10^8 M_\odot / \text{kpc}^2} \right)^2 \left(\frac{10^{-3}}{v_z} \right)^2$$

13. Modelo de coordenadas de una partícula repercutida por una partícula supermasiva.

$$r_t = r_{\text{orbit}} \left(\frac{M_{\text{clump}}(r_t)}{M_{\text{host}}(r_{\text{orbit}})} \right)^{1/3} \left(3 - \frac{d \log M_{\text{host}}(r)}{d \log r} \Big|_{r=r_{\text{orbit}}} \right)^{-1/3}$$

$$t_{\text{decay}} = t_{\text{orbit}} \frac{M_{\text{host}}}{M_{\text{clump}}} \left(\log \left(\frac{M_{\text{host}}}{M_{\text{clump}}} \right) \right)^{-1} \simeq \frac{1}{\sqrt{G \bar{\rho}_{\text{host}}}} \frac{M_{\text{host}}}{M_{\text{clump}}} \left(\log \left(\frac{M_{\text{host}}}{M_{\text{clump}}} \right) \right)^{-1}$$

$$\Delta E \simeq \frac{4\pi}{3} \gamma_1 G \bar{\rho}_{\text{host}} r^2 dm$$

$$t_{\text{shock}} \simeq \frac{\bar{\rho}(r)}{\sqrt{G} \gamma_1 \gamma_2 \bar{\rho}_{\text{host}}^{3/2}} \simeq 10^9 \left(\frac{\bar{\rho}}{0.01 \text{eV}^4} \right) \left(\frac{1}{\gamma_1 \gamma_2} \right) \left(\frac{10^{-4} \text{eV}^4}{\bar{\rho}_{\text{host}}} \right)^{3/2}$$

$$t_{\text{dyn}} \simeq 3 \cdot 10^{11} \left(\frac{10^{-4} \text{eV}}{\bar{\rho}(r)_{\text{host}}} \right)^{1/2} \frac{M_{\text{host}}/M_s}{10^5} \frac{\log(10^5)}{\log(M_{\text{host}}/M_s)}$$

$$\rho_s/\bar{\rho}_{\text{host}}(r_{\text{orbit}}) \gtrsim 100$$

CONCLUSIONES.

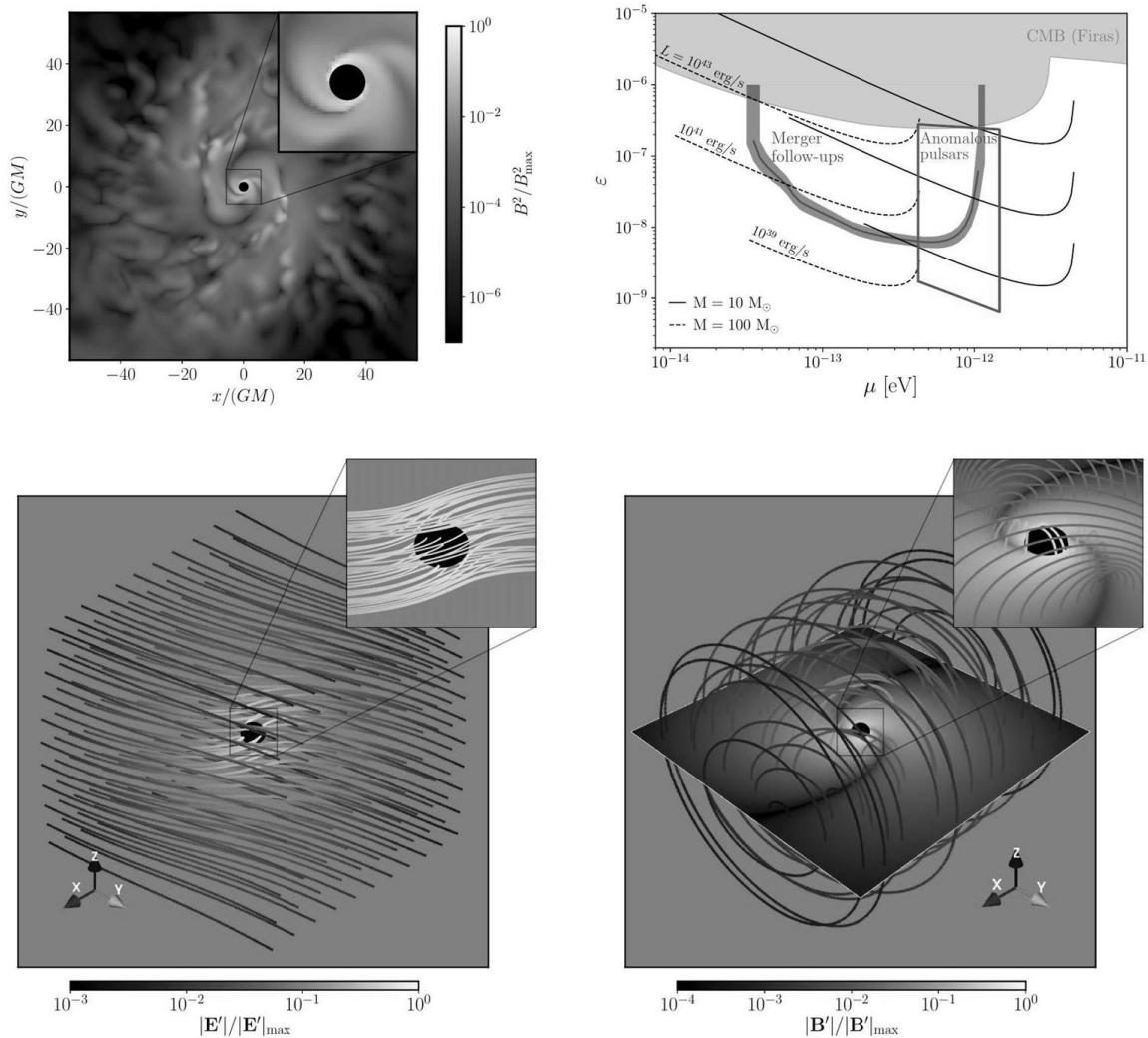
La partícula oscura o partícula supermasiva, es un componente esencial pero no principal de la Teoría Cuántica de Campos Relativistas o Teoría Cuántica de Campos Curvos, formulada por este investigador, sin embargo, se destacan ideas nucleares, (i) en primer término, que la partícula oscura o supermasiva, se caracteriza por tener una masa excesivamente densa, lo que la vuelve superlativamente pesada, (ii) por otro lado, que la partícula supermasiva u oscura, aunque se trate de una partícula subatómica extremadamente densa, produce energía cuya densidad es superior a cero, más sin embargo, este fenómeno no sucede en todos los casos, (iii) esta propiedad de la partícula supermasiva u oscura, le permite deformar el espacio – tiempo cuántico en el que interactúa, curvándolo, repercutiendo así en el sistema de coordenadas de las partículas repercutidas, lo que se identifica como gravedad cuántica endógena, o en su defecto, provocando agujeros negros cuánticos, a propósito de su aniquilación, que puede ser causada por colisiones o en su defecto, por implosiones, esto último cuando se está ante una partícula supermasiva cuya masa y energía son superdensas. Ahora bien, en este punto es importante reflexionar que si una partícula supermasiva, provoca un agujero negro cuántico, la singularidad

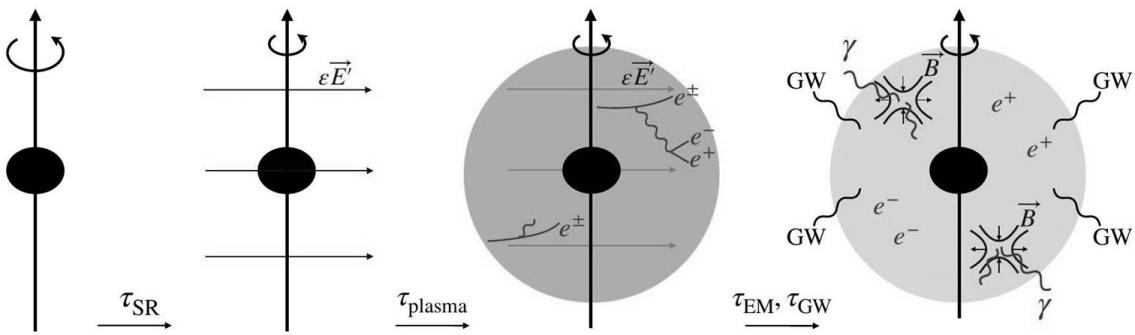


cuántica de éste, es la transformación de la materia y la energía, a propósito de la supergravedad cuántica y por ende, la existencia de multidimensiones o supermembranas de realidad microscópica. En esta singularidad, el tiempo abandona su condición de dimensión y pasa a convertirse en un objeto tridimensional, en tanto que el espacio, se funde con el tiempo, volviéndose multidimensional, infinitas veces, es decir, desplegando distintas capas de realidad, sí y sólo sí, se está ante un escenario de aniquilación y supergravedad cuántica.

Apéndice A.

Agujeros negros cuánticos provocados por una partícula supermasiva. Características funcionales y morfológicas y configuraciones de campo.





Figuras 1, 2 y 3. Simulación de un agujero negro cuántico provocado por la deformación del espacio – tiempo cuántico, a propósito de las interacciones de una partícula supermasiva.

$$\mathcal{L}' = -\frac{1}{4}F'_{\mu\nu}F'^{\mu\nu} - \frac{1}{2}\mu^2A'^{\mu}A'_{\mu}.$$

$$\omega \simeq \mu \left(1 - \frac{\alpha^2}{2} \right)$$

$$\Gamma_{\text{SR}} \equiv \tau_{\text{SR}}^{-1} \simeq 4\alpha^7(\Omega_{\text{BH}} - \omega) \simeq 4a_*\alpha^6\mu,$$

$$\Omega_{\text{BH}} = \frac{1}{2} \left(\frac{a_*}{1 + \sqrt{1 - a_*^2}} \right) r_g^{-1}.$$

$$\Omega_{\text{BH}} \leq \omega,$$

$$\alpha \lesssim 1/2$$

$$M_c \simeq 10^{-2} \left(\frac{\Delta a_*}{0.1} \right) \left(\frac{\alpha}{0.1} \right) M$$

$$A'_0 = \frac{\sqrt{M_c}}{\sqrt{\pi}\mu^2 r_c^{5/2}} e^{-r/r_c} \sin \theta \sin (\omega t - \phi)$$

$$\mathbf{A}' = -\frac{\sqrt{M_c}}{\sqrt{\pi}\mu r_c^{3/2}} e^{-r/r_c} \{ \cos \omega t, \sin \omega t, 0 \}$$

$$P_{\text{GW}} \simeq 17 \frac{\alpha^{10}}{G} \left(\frac{M_c(t)}{M} \right)^2$$

$$M_c(t) = \frac{M_c(t_0)}{1 + (t - t_0)/\tau_{\text{GW}}},$$

$$\tau_{\text{GW}} \simeq \frac{GM}{17\alpha^{11}\Delta a_*} \sim \zeta \left(\frac{0.1}{\Delta a_*} \right) \left(\frac{0.1}{\alpha} \right)^{11} \left(\frac{M}{10M_\odot} \right)$$

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{4}F'_{\mu\nu}F'^{\mu\nu} - \frac{\mu^2}{2}A'_\mu A'^\mu - \varepsilon\mu^2A'_\mu A^\mu + I_\mu A^\mu$$

$$\begin{array}{l} \nabla_{\alpha}F^{\alpha\beta}\, = -I^{\beta} + \varepsilon\mu^2A'^{\beta}\\ \nabla_{\alpha}F'^{\alpha\beta}\, = \mu^2A'^{\beta} + \varepsilon\mu^2A^{\beta} \end{array}$$

$$\begin{array}{l} \nabla_{\alpha}T^{\alpha\beta}\, = -F^{\beta\gamma}\big(I_{\gamma}-\varepsilon\mu^2A'_{\gamma}\big),\\ \nabla_{\alpha}T'^{\alpha\beta}\, = \varepsilon\mu^2F'^{\beta\gamma}A_{\gamma}. \end{array}$$

$$|\varepsilon {\bf E}'| \simeq \frac{\varepsilon \sqrt{\Delta a_*} \alpha^{5/2} \mu}{\sqrt{G}} \simeq 2 \cdot 10^{13} \; \mathrm{V/m} \sqrt{\Delta a_*} \left(\frac{\varepsilon}{10^{-7}} \right) \left(\frac{\alpha}{0.1} \right)^{5/2} \left(\frac{\mu}{10^{-12} \mathrm{eV}} \right)$$

$${\bf J}=\sigma({\bf E}+{\bf v}\times{\bf B})$$

$$\begin{array}{l} \rho\, \simeq \varepsilon \nabla \cdot {\bf E}' \simeq \pm \frac{\varepsilon \sqrt{\Delta a_*} \alpha^{7/2} \mu^2}{\sqrt{G}} \\ \simeq \pm 5 \cdot 10^7 \; \mathrm{cm}^{-3} \sqrt{\Delta a_*} \left(\frac{\varepsilon}{10^{-7}} \right) \left(\frac{\alpha}{0.1} \right)^{7/2} \left(\frac{\mu}{10^{-12} \mathrm{eV}} \right)^2 \end{array}$$

$$\partial_t {\bf B} = \frac{\varepsilon \mu^2 {\bf B}'}{\sigma} + \frac{1}{\sigma} \nabla^2 {\bf B} + \nabla \times ({\bf v} \times {\bf B}),$$

$$\gamma_e \simeq \frac{e\varepsilon |{\bf E}'|}{m_e \mu} \simeq e\varepsilon \alpha^{5/2} \sqrt{\Delta a_*} \frac{M_{\rm pl}}{m_e} \simeq 10^{12} \left(\frac{\varepsilon}{10^{-7}}\right) \left(\frac{\alpha}{0.1}\right)^{5/2} \left(\frac{\Delta a_*}{0.1}\right)^{1/2}$$

$$\Gamma_{\rm syn} \simeq \frac{2}{3} \frac{e^2 \gamma_e}{r_c} \simeq \frac{2}{3} \frac{e^3 \varepsilon |{\bf E}'|}{m_e} \simeq \frac{2}{3} e^3 \varepsilon \alpha^{5/2} \sqrt{\Delta a_*} \mu \frac{M_{\rm pl}}{m_e},$$

$$\varepsilon > \frac{1}{e^3 \alpha^{5/2} \sqrt{\Delta a_*} M_{\rm Pl}} \frac{m_e}{M_{\rm Pl}} \simeq 10^{-18} \left(\frac{0.1}{\alpha}\right)^{5/2} \left(\frac{0.1}{\Delta a_*}\right)^{1/2}$$

$$\frac{\Gamma_{e^\pm}}{V} = \frac{(e|{\bf E}|)^2}{4\pi^3} \sum_n^\infty \frac{1}{n^2} \exp\left(-\frac{\pi m_e^2}{e|{\bf E}|} n\right)$$

$$S_{\rm B}=-\gamma_\theta+\left(\frac{2m_e^2}{e\varepsilon|{\bf E}'|}+\frac{e\varepsilon|{\bf E}'|}{2m_e^2}\gamma_\theta^2\right)\arctan\frac{2m_e^2}{e\varepsilon|{\bf E}'|\gamma_\theta},$$

$$\Gamma_{e^\pm}^\gamma=\frac{\alpha_{\rm EM}}{2\pi}\frac{e\varepsilon|{\bf E}'|}{m_e}\exp\left(-\frac{2m_e^2}{e\varepsilon|{\bf E}'|}\frac{2m_e}{\omega_{\rm syn}}\right)=\frac{\alpha_{\rm EM}}{2\pi}\frac{e\varepsilon|{\bf E}'|}{m_e}\exp\left[-\frac{4m_e^6\mu^2}{(e\varepsilon|{\bf E}'|)^4}\right]$$

$$\varepsilon \gg \left(\ln \frac{\alpha_{\rm EM} \gamma_e}{2\pi \alpha}\right)^{-1/4} \frac{m_e^{3/2} \sqrt{2\mu}}{e|{\bf E}'|} \approx \left(\ln \frac{\alpha_{\rm EM} \gamma_e}{2\pi \alpha}\right)^{-1/4} \frac{1}{e\alpha^{5/2} \sqrt{\Delta a_*} M_{\rm pl}} \sqrt{\frac{2m_e}{\mu}} \simeq 10^{-10} \left(\frac{0.1}{\alpha}\right)^{5/2} \left(\frac{0.1}{\Delta a_*}\right)^{1/2} \left(\frac{10^{-12} \mathrm{eV}}{\mu}\right)^{1/2} \left(\frac{\log \frac{\alpha_{\rm EM} \gamma_e}{2\pi \alpha}}{20}\right)^{-1/4}$$

$$\tau_{\rm plasma} \simeq \frac{1}{2\Gamma_{e^\pm}^\gamma} \ln \left(\frac{n_e^f}{n_e^0} \right) \simeq \frac{1}{2\alpha\mu} \ln \frac{m_e^{3/2}}{e^2\alpha^2\mu^{3/2}}$$

$$\frac{\tau_{\rm plasma}}{\tau_{\rm SR}} \simeq 2a_*\alpha^5 \ln \frac{m_e^{3/2}}{e^2\alpha^2\mu^{3/2}}$$



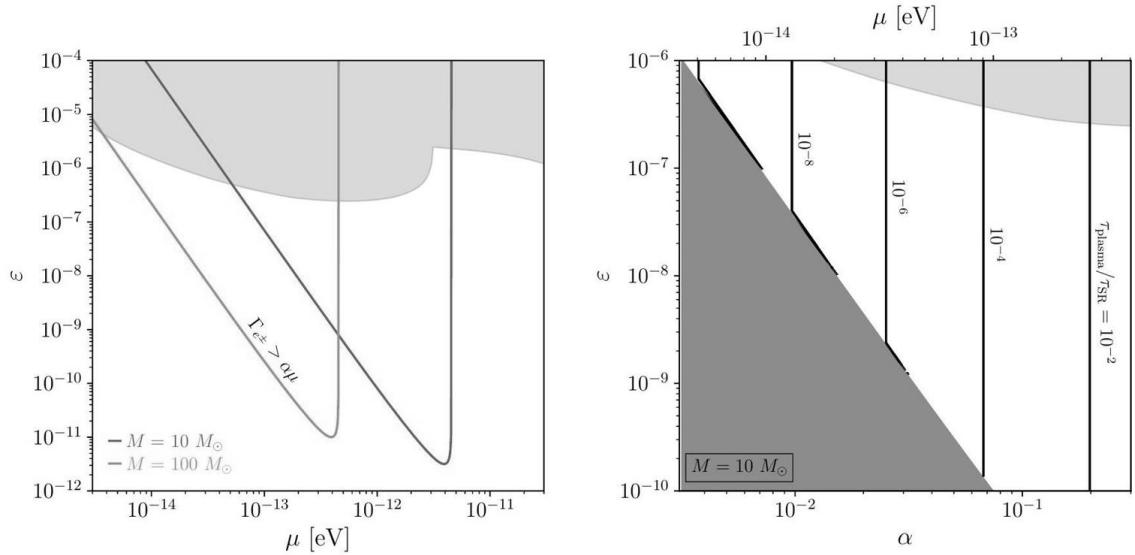


Figura 4. Espectro de radiación de un agujero negro cuántico.

$$M_{\text{plasma}} \approx \frac{m_e \rho}{e(\alpha \mu)^3} \simeq \varepsilon \alpha^{1/2} (\Delta a_*)^{1/2} \frac{m_e M_{\text{pl}}}{\mu} \simeq 10^{-29} M_{\odot} \left(\frac{\varepsilon}{10^{-7}} \right) \left(\frac{\alpha}{0.1} \right)^{1/2} \left(\frac{\Delta a_*}{0.1} \right)^{1/2} \left(\frac{10^{-12} \text{ eV}}{\mu} \right)$$

$$\tau_{\text{acc}} = \frac{M_{\text{plasma}}}{\dot{M}_{\text{Bondi}}} \approx \varepsilon \alpha^{-3/2} (\Delta a_*)^{1/2} \frac{c_s^3 M_{\text{pl}} \mu}{\pi n_M} \simeq 10 \text{ years} \left(\frac{\varepsilon}{10^{-7}} \right) \left(\frac{\alpha}{0.1} \right)^{-3/2} \left(\frac{\Delta a_*}{0.1} \right)^{1/2} \left(\frac{\mu}{10^{-12} \text{ eV}} \right) \left(\frac{1/\text{cm}^3}{n_M} \right) \left(\frac{c_s}{1} \right)^3$$

$$E^i \equiv n_\nu F^{i\nu}, B^i \equiv n_\nu (*F)^{i\nu} = \frac{1}{2} n_\nu \varepsilon^{i\nu\alpha\beta} F_{\alpha\beta}$$

$$J^i = I^i - \rho_q n^i, \quad \rho_q = -n_\mu I^\mu$$

$$\rho_q = D_i E^i - \varepsilon \mu^2 n_\mu A'^\mu$$

$$j^\mu = \sigma e^\mu$$

$$e^\mu \equiv u_\nu F^{\mu\nu}, j^\mu \equiv I^\mu + (u^\nu I_\nu) u^\mu$$

$$v_{d, \text{ideal}}^i = \frac{\varepsilon^{ijk} E_j B_k}{B^2},$$

$$v_d^i = \frac{\varepsilon^{ijk} E_j B_k}{B^2 + E_0^2}, E_0^2 = B_0^2 + E^2 - B^2$$

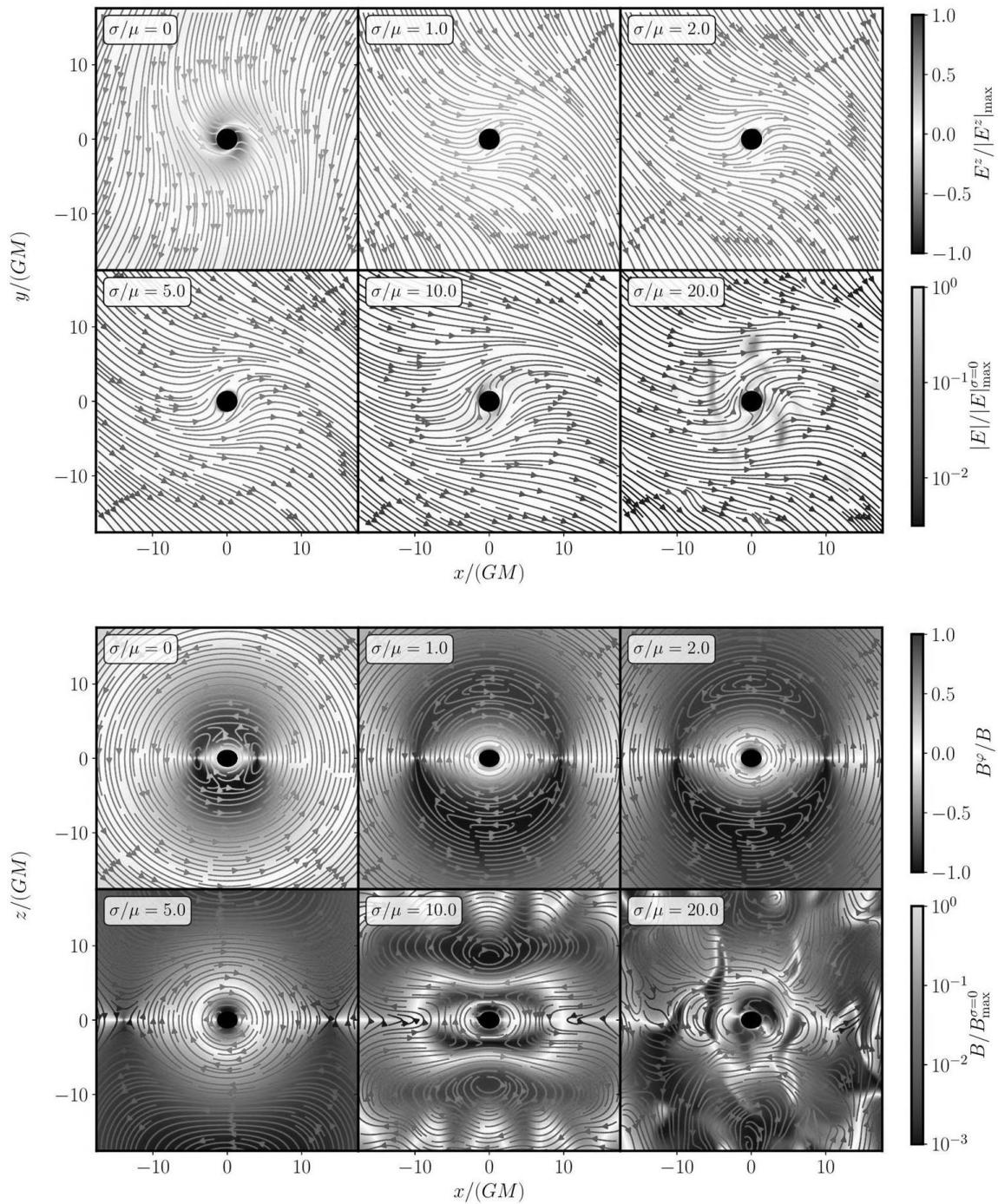
$$B_0^2 = \frac{1}{2} \left[B^2 - E^2 + \sqrt{(B^2 - E^2)^2 + 4(E_i B^i)^2} \right]$$

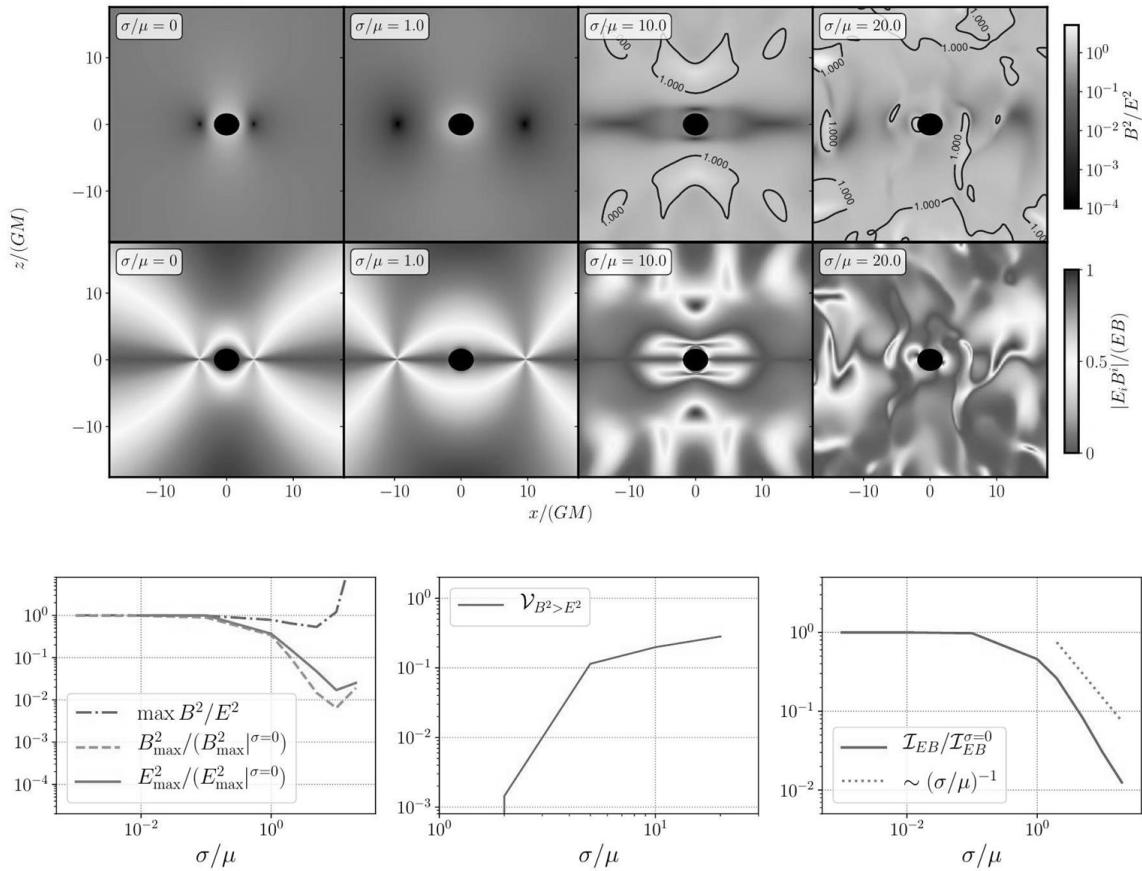
$$J^i = \rho_q v_d^i + \sigma E_0 \sqrt{\frac{B^2 + E_0^2}{B_0^2 + E_0^2}} \left(\frac{E_0 E^i + B_0 B^i}{B^2 + E_0^2} \right)$$

$$\partial_t E^i \approx -\sigma E^i + \varepsilon \mu^2 A'^i$$



$$\mathbf{E} \propto \begin{pmatrix} \sigma + i\mu \\ \mu - i\sigma \\ 0 \end{pmatrix} e^{-i\omega t} + c.c.$$

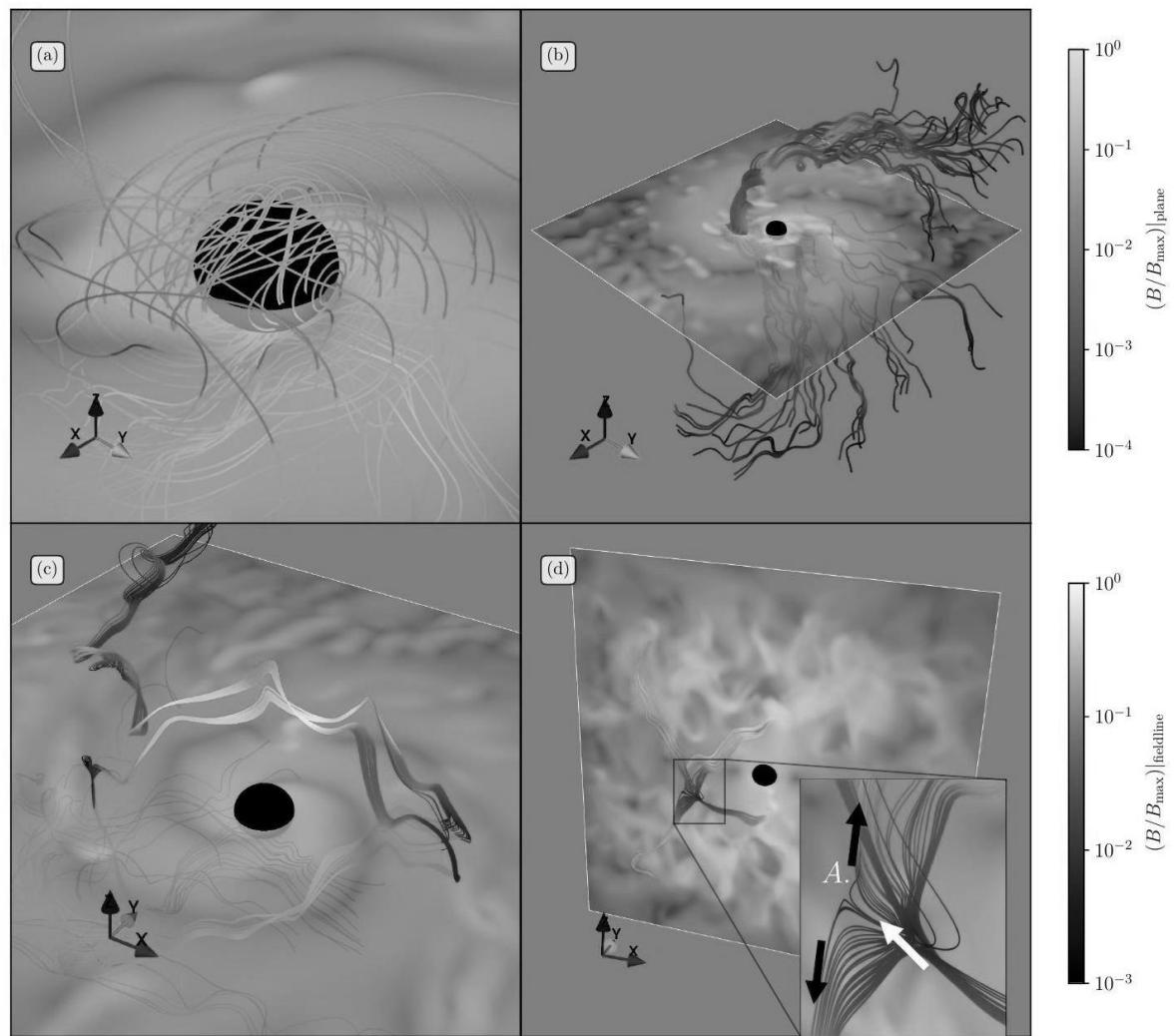
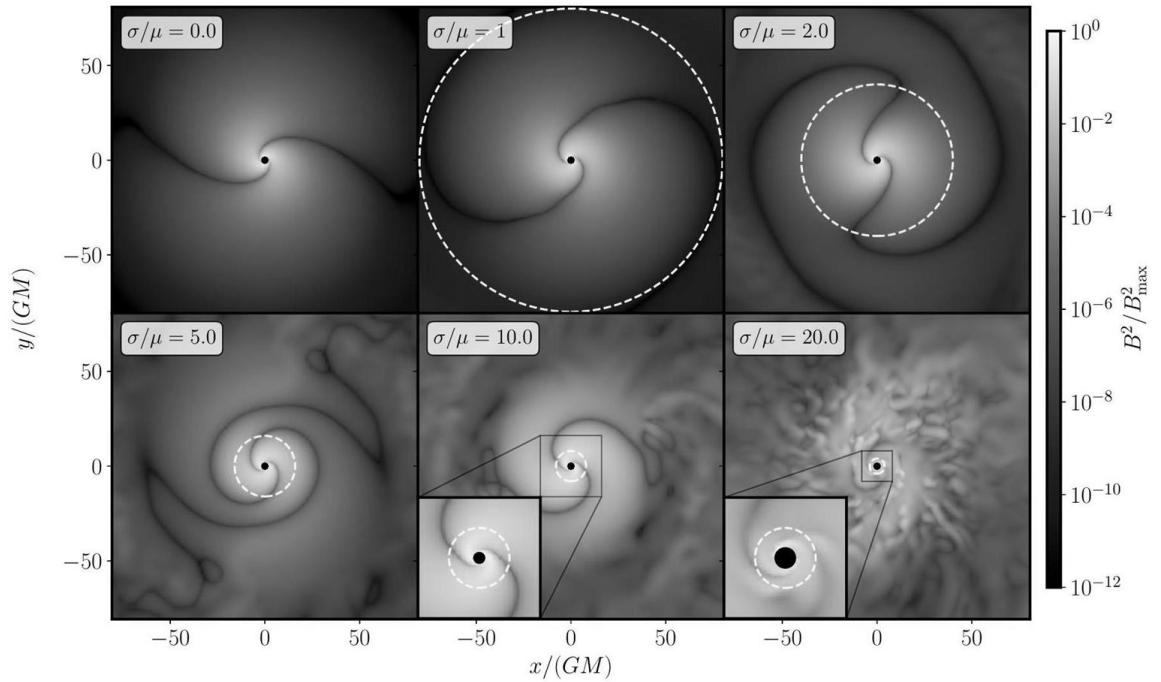




Figuras 5, 6, 7 y 8. Campos de energía desplegados por un agujero negro cuántico provocado por una partícula supermasiva.

$$\partial_t B^i = \frac{\varepsilon \mu^2 B'^i}{\sigma} + \frac{1}{\sigma} \partial_j \partial^j B^i + \varepsilon^{ijk} \varepsilon_{klm} \partial_j v^l B^m$$

$$r_* \approx 80 \mu GM / \sigma$$



Figuras 9 y 10. Configuraciones de campo del agujero negro cuántico – deformación del espacio – tiempo cuántico provocado por una partícula supermasiva.

$$\mathcal{E} = \int_D d^3x \sqrt{\gamma} T^\alpha{}_\mu n_\alpha \xi^\mu$$

$$T_{\mu\nu} = F_\mu{}^\lambda F_{\lambda\nu} - \frac{1}{4} g_{\mu\nu} F^{\alpha\beta} F_{\alpha\beta}$$

$$\partial_t \mathcal{E} = -P_{\text{EM}} - \dot{\mathcal{E}}_{\text{BH}} - L_{\text{diss}} + \dot{\mathcal{E}}_{A'}$$

$$P_{\text{EM}} = - \oint_{S_{\hat{\rho}}^2} d\Omega_\mu T^\mu_\nu \xi^\nu \stackrel{\text{curved}}{=} \oint_{S_{\hat{\rho}}^2} d\Omega \hat{\rho} \cdot (\mathbf{E} \times \mathbf{B})$$

$$\dot{\mathcal{E}}_{\text{BH}} = - \oint_{S_{\text{BH}}^2} d\Omega_\mu T^\mu{}_\nu \xi^\nu$$

$$L_{\text{diss}} = - \int_D d^3x \sqrt{-g} F^{\alpha\beta} \xi_\alpha I_\beta \stackrel{\text{curved}}{=} \int_D d^3x \mathbf{E} \cdot \mathbf{J}$$

$$\rho_{\text{diss}} = N F^{\alpha\beta} \xi_\alpha I_\beta$$

$$\dot{\mathcal{E}}_{A'} = -\varepsilon \mu^2 \int_D d^3x \sqrt{-g} F^{\alpha\beta} \xi_\alpha A'_\beta \stackrel{\text{flat}}{=} \varepsilon \mu^2 \int_D d^3x \mathbf{E} \cdot \mathbf{A}'$$

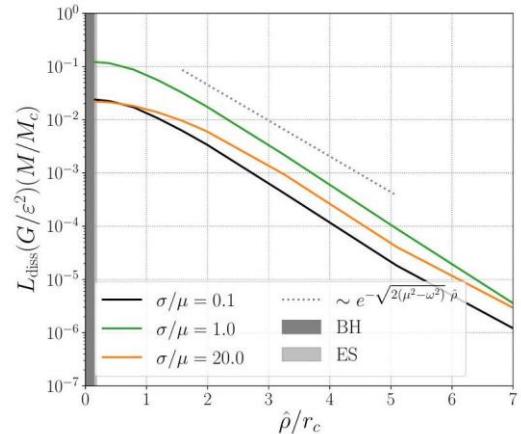
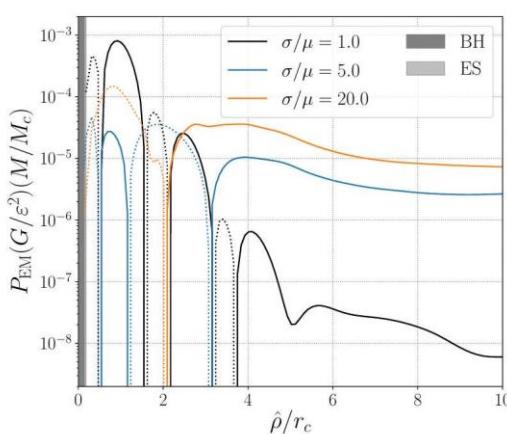


Figura 11. Coordenadas de un agujero negro cuántico.

$$L_{\text{diss}} = L_{\text{diss}}^{\text{bulk}} + L_{\text{diss}}^{\text{turb}}$$

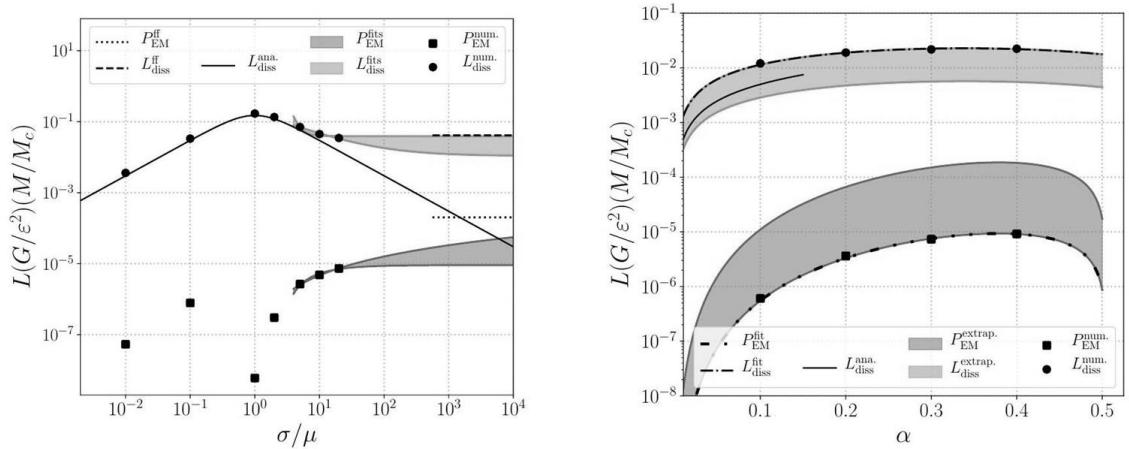


Figura 12. Disipación de un agujero negro cuántico.

$$\mathbf{J} = \sigma \mathbf{E} = -\frac{e^{-r/r_c} \sqrt{M_c \mu} \alpha^{3/2} \varepsilon \sigma \omega^2}{2\sqrt{\pi}(\sigma^2 + \omega^2)} \begin{pmatrix} \sigma + i\omega \\ -\omega + i\sigma \\ 0 \end{pmatrix} e^{-i\omega t} + c.c.$$

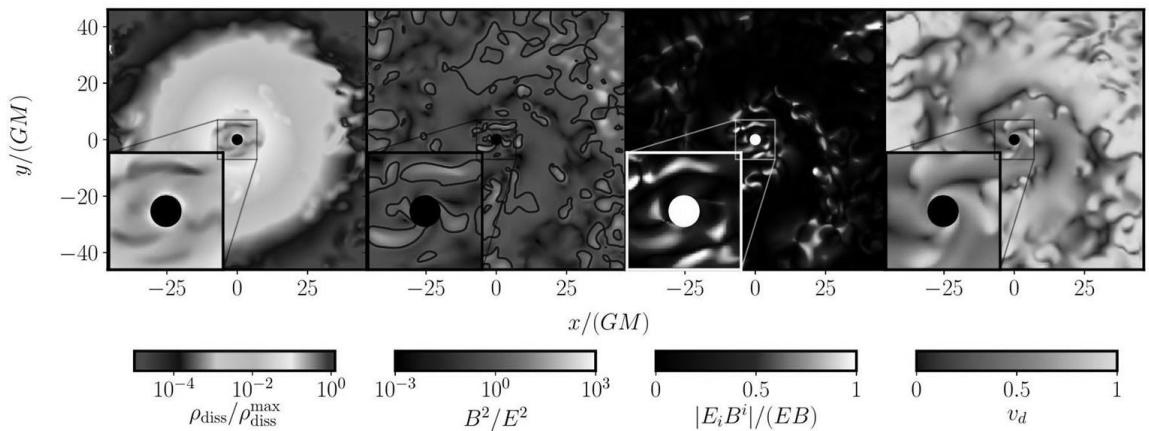
$$L_{\text{diss}}^{\text{bulk}} = \frac{\sigma \alpha \varepsilon^2}{\mu(1 + (\sigma/\mu)^2)} \frac{M_c}{GM}$$

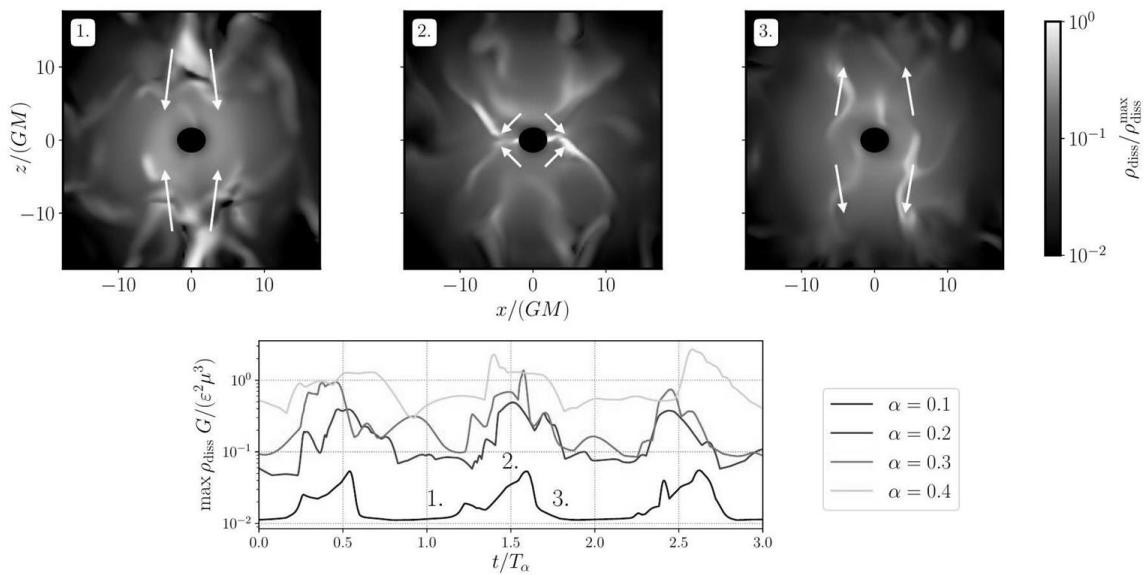
$$L_{\text{diss}}^{\text{fit}} = \varepsilon^2 F(\alpha) \frac{M_c}{GM},$$

$$P_{\text{EM}}^{\text{fit}} = \varepsilon^2 G(\alpha) \frac{M_c}{GM},$$

$$F(\alpha) = 1.31 \times 10^{-1} \alpha - 1.88 \times 10^{-1} \alpha^2$$

$$G(\alpha) = 6.86 \times 10^{-4} \alpha^3 - 1.36 \times 10^{-3} \alpha^4$$





Figuras 13, 14 y 15. Densidad de un agujero negro cuántico.

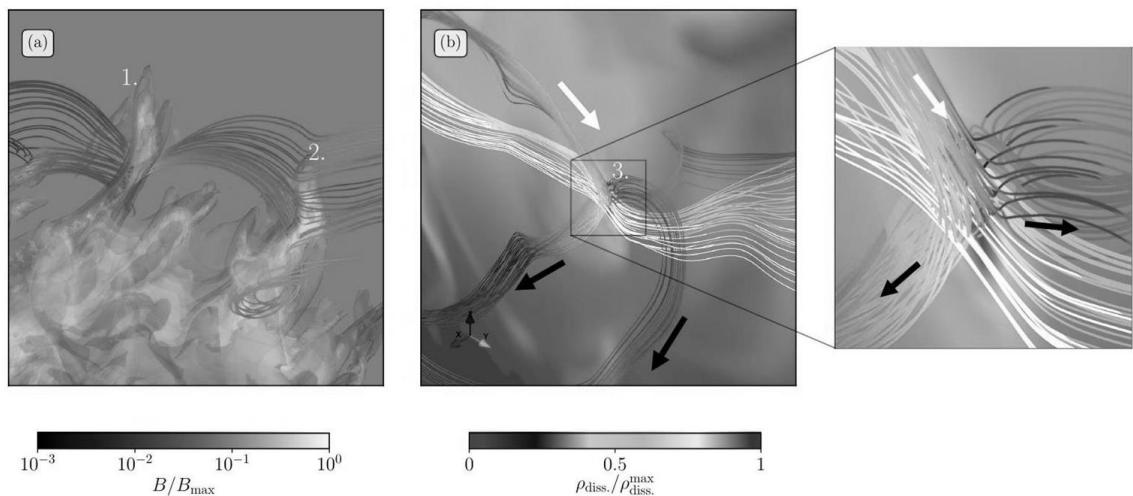


Figura 16. Comportamiento de la materia y la energía en un agujero negro cuántico.

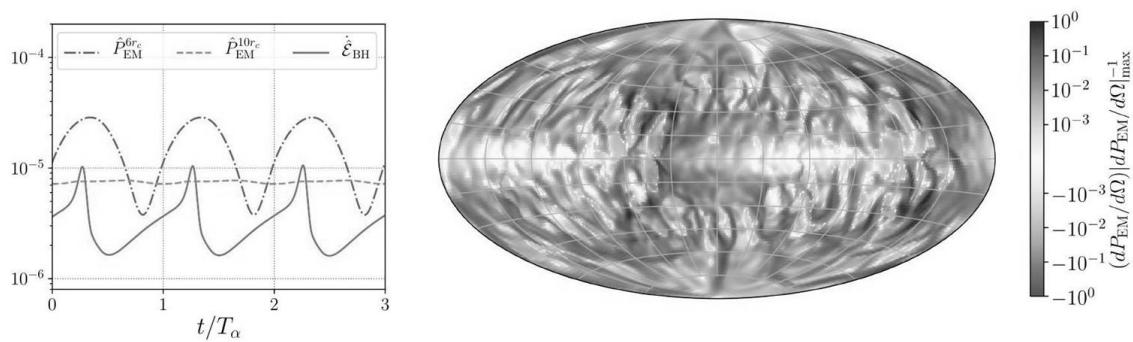


Figura 17. Radiación de un gujero negro cuántico

$$\langle B^2 \rangle^{1/2} = 2.5 \times 10^8 \text{ Gauss} \left(\frac{\varepsilon}{10^{-7}} \right) \left(\frac{M_\odot}{M} \right) \left(\frac{\alpha}{0.1} \right)^{5/2}$$

$$\gamma_c \approx 2.2 \times 10^7 \left(\frac{\varepsilon}{10^{-7}} \right) \left(\frac{\alpha}{0.1} \right)^{3/2}$$

$$\gamma_r = 3 \times 10^3 \left(\frac{10^{-7}}{\varepsilon} \right)^{1/2} \left(\frac{M}{M_\odot} \right)^{1/2} \left(\frac{0.1}{\alpha} \right)^{5/4}$$

$$\nu_{\text{peak}} = 12 \text{ keV} \left(\frac{\gamma}{10^2} \right)^2 \left(\frac{\varepsilon}{10^{-7}} \right) \left(\frac{M_\odot}{M} \right) \left(\frac{\alpha}{0.1} \right)^{5/2},$$

$$t_{\text{growth}} \sim \ln(M_c/\mu) \tau_{\text{SR}} \approx 10^4 \text{ s} \left(\frac{M}{10M_\odot} \right) \left(\frac{0.7}{a_*} \right) \left(\frac{0.1}{\alpha} \right)^7$$

$$L_{\text{EM}} = \varepsilon^2 F(\alpha) \frac{M_c}{GM} \simeq \varepsilon^2 \frac{\alpha^2 \Delta a_*}{G} \simeq 4 \times 10^{41} \text{ erg/s} \left(\frac{\varepsilon}{10^{-7}} \right)^2 \left(\frac{\alpha}{0.1} \right)^2 \left(\frac{\Delta a_*}{0.1} \right)$$

$$\tau_{\text{GW}} \approx \frac{GM}{17\alpha^{11}\Delta a_*} \approx 10^6 \text{ s} \left(\frac{M}{10M_\odot} \right) \left(\frac{0.1}{\alpha} \right)^{11} \left(\frac{0.1}{\Delta a_*} \right)$$

$$\tau_{\text{EM}} \approx \frac{G M \ln 2}{\varepsilon^2 F(\alpha)} \approx 10^{11} \text{ s} \left(\frac{M}{10M_\odot} \right) \left(\frac{10^{-7}}{\varepsilon} \right)^2 \left(\frac{10^{-2}}{F(0.1)} \right)$$

$$\frac{M_c(t)}{M_c(t_0)} = \begin{cases} [1 + (t - t_0)/\tau_{\text{GW}}]^{-1} & \tau_{\text{GW}} \ll \tau_{\text{EM}} \\ e^{-(t - t_0)/\tau_{\text{EM}} \ln 2} & \tau_{\text{GW}} \gg \tau_{\text{EM}} \end{cases}$$

$$\tau_{\text{SR}} \simeq 0.98 \tau_{\text{GW}}^{7/11} r_g^{4/11} \left[a_* \left(\frac{\Delta a_*}{0.5} \right)^{7/11} \right]$$

$$\dot{f}_{\text{int}} \simeq \frac{5}{8\pi} \alpha \mu^2 G P_{\text{GW}}$$

$$g^{\alpha\beta} \nabla_\alpha \nabla_\beta A'^\gamma = \mu^2 A'^\gamma$$

$$A'^\mu = B^{\mu\nu} \nabla_\nu Z, Z = R(r) S(\theta) e^{-i(\omega T - m\varphi)}$$

$$R_{\text{near}}(r) = \hat{r}^{i\kappa} (1 + \hat{a}_1 \hat{r} + \hat{a}_2 \hat{r}^2 + \dots),$$



$$\begin{aligned}\tau &= T + \frac{M^2 \log \frac{r - r_+}{r - r_-}}{\sqrt{M^2 - a^2}} + M \log \Delta \\ x &= \sin \theta (r \cos \bar{\phi} - a \sin \bar{\phi}) \\ y &= \sin \theta (a \cos \bar{\phi} + r \sin \bar{\phi}) \\ z &= r \cos \theta\end{aligned}$$

$$\begin{aligned}T &= t - \frac{M^2 \log \frac{r - r_+}{r - r_-}}{\sqrt{M^2 - a^2}} - M \log \Delta \\ r &= 2^{-1/2} [-a^2 + x^2 + y^2 + z^2 \\ &\quad + \sqrt{4a^2z^2 + (a^2 - x^2 - y^2 - z^2)^2}]^{1/2} \\ \phi &= \arctan \left[\frac{rx + ay}{-ax + ry} \right] - \frac{a \log \frac{r - r_+}{r - r_-}}{2\sqrt{M^2 - a^2}}, x > 0 \\ \theta &= \arccos \frac{z}{r}\end{aligned}$$

$$L_D(f) \equiv \int_D d^3x \sqrt{\gamma} |f|$$

$$\begin{aligned}ds^2 &= -dt^2 + dx^2 + dy^2 + dz^2 \\ &\quad + \frac{2Mr^3}{r^4 + a^2M^2z^2} \left[dt + \frac{z}{r} dz \right. \\ &\quad \left. + \frac{r(xdx + ydy)}{r^2 + a^2M^2} - \frac{aM(xdy - ydx)}{r^2 + a^2M^2} \right]^2\end{aligned}$$

$$\begin{aligned}D_t E^i &= NKE^i + \varepsilon^{ijk} D_j(NB_k) - NJ^i + N\varepsilon\mu^2\gamma^i{}_\mu A'^\mu \\ D_t B^i &= NKB^i - \varepsilon^{ijk} D_j(NE_k) \\ D_i E^i &= \rho_q + \varepsilon\mu^2 n_\mu A'^\mu \\ D_i B^i &= 0\end{aligned}$$

$$\begin{aligned}D_t \Psi &= -N(D_i E^i - \varepsilon\mu^2 n_\mu A'^\mu - \rho_q) - N\kappa\Psi \\ D_t \Phi &= -ND_i B^i - N\kappa\Phi\end{aligned}$$

$$\begin{aligned}E^i &\rightarrow E^j \left(\delta_j^i - \frac{B_j B^i}{B^2} \right) \\ E^i &\rightarrow E^i \left\{ 1 - \hat{\theta}(\lambda) + \hat{\theta}(\lambda) \frac{B}{E} \right\}\end{aligned}$$

$$J^a = \rho_q v_d^a + J_\perp^a$$

$$v_d^a = \frac{\varepsilon^{ajk} E_j B_k}{B^2}$$

$$J_{\perp,\text{FF}}^a = \frac{B^a}{B^2} [2KB_i E^i - 2K_{ij} E^i B^j + B_i \epsilon^{ijk} D_j B_k - E_i \epsilon^{ijk} D_j E_k + \varepsilon\mu^2 B_i \gamma_\mu^i A'^\mu]$$



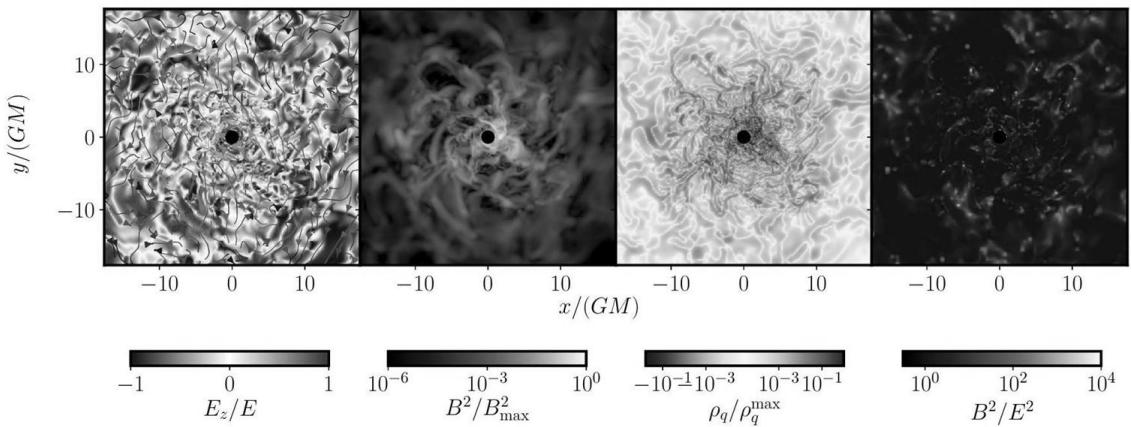


Figura 18. Radio y horizonte de eventos de un agujero negro cuántico.

$$v_d^i = \frac{\epsilon^{ijk} E_j B_k}{B^2 + E_0^2}, E_0^2 = B_0^2 + E^2 - B^2,$$

$$B_0^2 = \frac{1}{2} \left[B^2 - E^2 + \sqrt{(B^2 - E^2)^2 + 4(E_i B^i)^2} \right].$$

$$J_{\perp,(B)}^a = \sigma E_0 \sqrt{\frac{B^2 + E_0^2}{B_0^2 + E_0^2}} \frac{E_0 E^a + B_0 B^a}{B^2 + E_0^2}$$

$$J_{\perp,(C)}^a = \frac{\sigma}{(\sigma + \kappa)} \left(J_{\perp,\text{FF}}^a + \kappa E^i B_i \frac{B^a}{B^2} \right)$$

$$D_t(E_i B^i) = -\kappa N \left(E^i - \frac{1}{\sigma} J^i \right) B_i,$$

$$D_t \rho_q = N \rho_q K - D_i(N J^i) = N \rho_q K - D_i(N \rho_q v_d^i) = 0$$

$$\mathcal{I}_\rho = \int_D d^3x \sqrt{\gamma} |\rho_q|$$

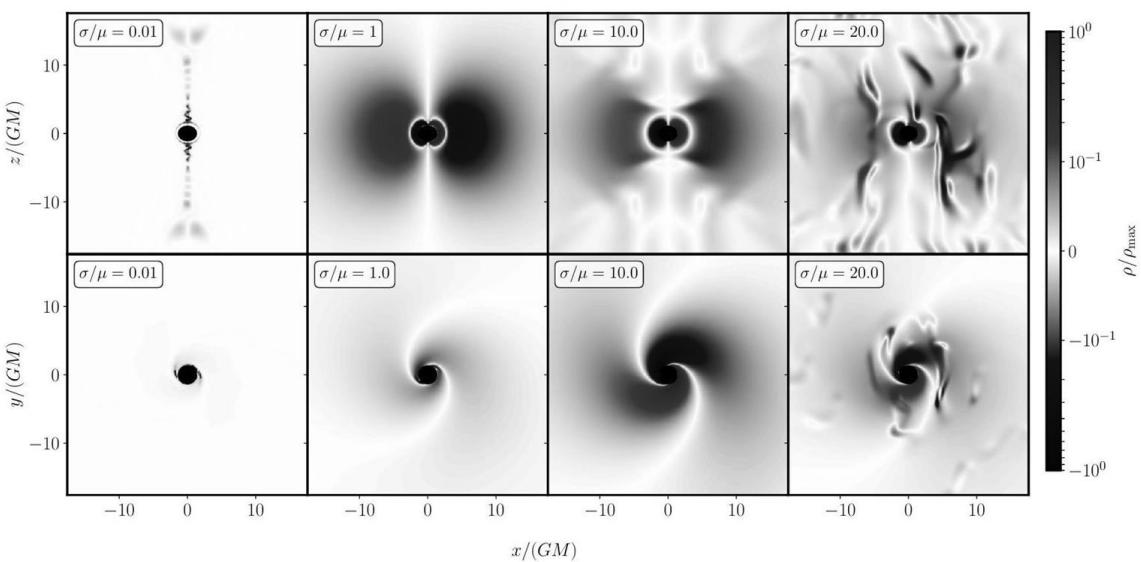


Figura 19. Fluctuaciones gravitacionales de un agujero negro cuántico.

$$\mathcal{I}_{E'} = \int_D d^3x \sqrt{\gamma} \varepsilon |D_i E'^i|$$

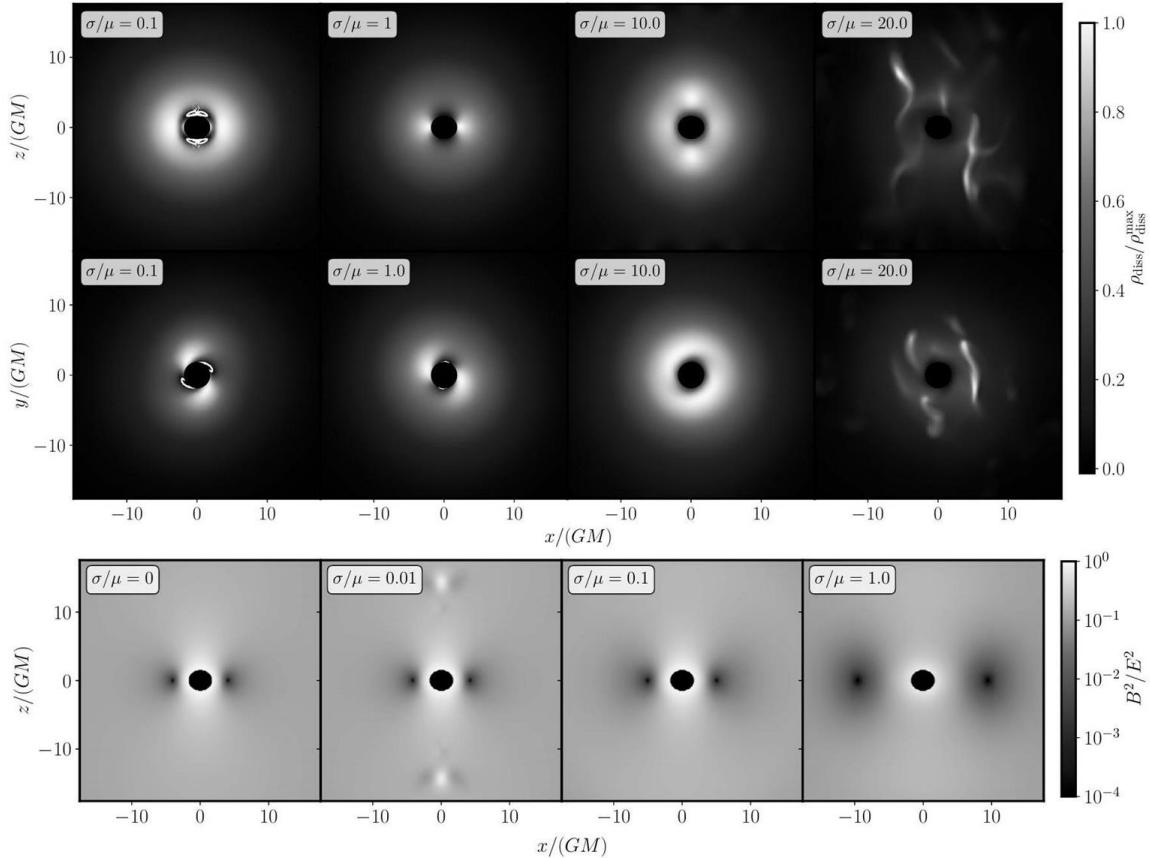


Figura 20. Simulación de un agujero negro cuántico.

$$\mathcal{L}_{\text{mass}} = -\frac{1}{4} \mathcal{F}_{\mu\nu} \mathcal{F}^{\mu\nu} - \frac{1}{4} \mathcal{F}'_{\mu\nu} \mathcal{F}'^{\mu\nu} - \frac{\mu^2}{2} \mathcal{A}'_\mu \mathcal{A}'^\mu + I_\mu (\mathcal{A}^\mu + \varepsilon \mathcal{A}'^\mu)$$

$$\begin{aligned} \mathcal{L}_{\text{inter}} = & -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \frac{1}{4} F'_{\mu\nu} F'^{\mu\nu} \\ & - \frac{\mu^2}{2} A'_\mu A'^\mu - \varepsilon \mu^2 A'_\mu A^\mu + I_\mu A^\mu \end{aligned}$$

$$\begin{aligned} \nabla_\alpha \mathcal{F}^{\alpha\beta} &= -I^\beta \\ \nabla_\alpha \mathcal{F}'^{\alpha\beta} &= \mu^2 \mathcal{A}'^\beta - \varepsilon I^\beta \end{aligned}$$

$$\begin{aligned} \nabla_\alpha F^{\alpha\beta} &= -I^\beta + \varepsilon \mu^2 A'^\beta \\ \nabla_\alpha F'^{\alpha\beta} &= \mu^2 A'^\beta + \varepsilon \mu^2 A^\beta \end{aligned}$$

$$\begin{aligned} \nabla_\alpha \mathcal{T}^{\alpha\beta} &= -\mathcal{F}^{\beta\gamma} I_\gamma \\ \nabla_\alpha \mathcal{T}'^{\alpha\beta} &= -\varepsilon \mathcal{F}'^{\beta\gamma} I_\gamma \end{aligned}$$



$$\begin{aligned}\nabla_\alpha T^{\alpha\beta} &= -F^{\beta\gamma}(I_\gamma - \varepsilon\mu^2 A'_\gamma) \\ \nabla_\alpha T'^{\alpha\beta} &= \varepsilon\mu^2 F'^{\beta\gamma} A_\gamma\end{aligned}$$

$$\nabla^\mu A'_\mu = 0, \quad \nabla_\mu I^\mu = 0$$

Apéndice B.

Electrodinámica cuántica de una partícula supermasiva. Modelo matemático.

$$\frac{dS_\alpha}{d\tau} = \Gamma_{\alpha\nu}^\lambda S_\lambda \frac{dx^\nu}{d\tau}$$

$$\frac{d\mathbf{S}}{dt} = -\mathbf{S} \frac{\partial \phi}{\partial t} - 2\mathbf{v} \cdot \mathbf{S} \nabla \phi - \mathbf{S}(\mathbf{v} \cdot \nabla \phi) + \mathbf{v}(\mathbf{S} \cdot \nabla \phi) + \frac{1}{2} \mathbf{S} \times (\nabla \times \mathbf{S})$$

$$\mathbf{S}_1 = (1 + \phi)\mathbf{S} - \frac{1}{2}\mathbf{v}(\mathbf{v} \cdot \mathbf{S}),$$

$$\frac{d\mathbf{S}_1}{dt} = \boldsymbol{\Omega} \times \mathbf{S}_1$$

$$\frac{\delta S}{S} \sim T \nu \frac{\phi}{R c^2}$$

$$\frac{\delta S}{S} \sim 4 \times 10^{-3} \left(\frac{\nu}{2 \times 10^{-3}} \right)^3 \frac{T}{13 \times 10^9 \text{yr}} \frac{8 \text{kpc}}{R}$$

$$\mathcal{L} \supset -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{4}X_{\mu\nu}X^{\mu\nu} + \frac{\sin \alpha}{2}F^{\mu\nu}X_{\mu\nu} + eJ_{\text{EM}}^\mu A_\mu + \frac{m_X^2 \cos^2 \alpha}{2}X^\mu X_\mu$$

$$\mathcal{L} \supset -\frac{1}{4}\tilde{F}_{\mu\nu}\tilde{F}^{\mu\nu} - \frac{1}{4}\tilde{X}_{\mu\nu}\tilde{X}^{\mu\nu} + \frac{e}{\cos \alpha}J_{\text{EM}}^\mu \tilde{A}_\mu + \frac{m_X^2 \cos^2 \alpha}{2}(\tilde{X}^\mu \tilde{X}_\mu + 2\chi \tilde{X}_\mu \tilde{A}^\mu + \chi^2 \tilde{A}^\mu \tilde{A}_\mu)$$

$$\mathcal{L} \supset -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{4}X_{\mu\nu}X^{\mu\nu} + eJ_{\text{EM}}^\mu A_\mu + \frac{m_X^2}{2}(X^\mu X_\mu + 2\chi X_\mu A^\mu)$$

$$-K^2 A^\nu = \chi m_X^2 X^\nu.$$

$$X_c^\mu(t, \mathbf{x}) = \sqrt{V} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} X^\mu(\mathbf{k}) e^{-i(\omega t - \mathbf{kx} + \delta(\mathbf{k}))}.$$

$$\rho = \frac{1}{V} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \frac{\omega(\mathbf{k})^2}{2} |\mathbf{X}(\mathbf{k})|^2$$

$$\langle X_c^\mu(t) \rangle = \frac{1}{V} \int d^3 \mathbf{x} X_c^\mu(t, \mathbf{x}) = \frac{X^\mu(\mathbf{k} = 0)}{\sqrt{V}} e^{-im_X t} \equiv X_0^\mu e^{-im_X t}$$



$$\rho = \rho \int ~\mathrm{d}^3\mathbf{v} f_{\mathrm{lab}}(\mathbf{v})$$

$$f_{\mathrm{lab}}(\mathbf{v})=\frac{m_X^3\omega^2}{2(2\pi)^3\rho}|\mathbf{X}(\mathbf{k})|^2$$

$$\rho = \frac{m_X^2}{2}\langle \left| {\mathbf{X}_c(t)} \right| \,\,\Bigr\rangle ^2$$

$$\rho\simeq\frac{m_X^2}{2}|\langle{\mathbf X}_c(t)|\rangle|^2=\frac{m_X^2}{2}|{\mathbf X}_0|^2$$

$$|{\mathbf E}_0|=\Bigl|\frac{\chi m_X}{\epsilon}{\mathbf x}_0\Bigr|.$$

$$|{\mathbf E}_0|=\Bigl|\frac{\chi m_X}{\epsilon}{\mathbf X}_0\cos\,\theta\Bigr|,$$

$$\begin{array}{l}\nabla\times\mathbf{B}_X=\mathbf{J}_X\\\nabla\times\mathbf{E}_X=-\frac{\partial\mathbf{B}_X}{\partial t}\end{array}$$

$$P(t)=P_X\mathrm{cos}^2~\theta(t)$$

$$\frac{S}{N}=\frac{P}{T_{\mathrm{sys}}}\sqrt{\frac{T}{\Delta\nu_{\mathrm{DP}}}}$$

$$\frac{S}{N}=2\big(\sqrt{n_s+n_d}-\sqrt{n_d}\big)$$

$$n_s=\eta\int~\mathrm{d}t\frac{P(t)}{\omega}$$

$$\frac{1}{T}\int~~\mathrm{d}tP(t)\equiv P_X\langle\mathrm{cos}^2~\theta\rangle_T$$

$$T\gg\tau=\frac{2\pi}{m_X v^2}\simeq 400\mu\,s\left(\frac{10\mu\mathrm{eV}}{m_X}\right)$$

$$\begin{aligned}\frac{S}{N}&\simeq\frac{S_1+S_2}{\sqrt{2N_1}}\\&\propto\frac{P_X}{T}\int_0^T~\mathrm{d}t\mathrm{cos}^2~\theta(t)+\frac{P_X}{T}\int_{T_{\mathrm{wait}}}^{T_{\mathrm{wait}}+T}~\mathrm{d}t\mathrm{cos}^2~\theta(t)\end{aligned}$$

$$\int_0^T~\mathrm{d}t\mathrm{cos}^2~\theta(t)\gg\int_{T_{\mathrm{wait}}}^{T_{\mathrm{wait}}+T}~\mathrm{d}t\mathrm{cos}^2~\theta(t)$$

$$P=P_a+N$$



$$\frac{P_a}{\sigma_N} > \Phi^{-1}[0.95] = 1.64$$

$$P=P_X+N\equiv P_X^0\langle\cos^2\theta\rangle_T+N$$

$$\Phi(Z) = \mathbb{P}(z \leq Z) = \int_{-\infty}^{\infty} \mathrm{d}x \int_{-\infty}^{Z-x} \mathrm{d}y f(x)f(y)$$

$$\Phi[0]\equiv\int_{-\infty}^{+\infty}\mathrm{d}P_X\int_{-\infty}^{0-P_X}\mathrm{d}Nf(P_X)f(N)=1-0.95$$

$$\int_0^1\mathrm{d}\langle\cos^2\theta\rangle_T\frac{f(\langle\cos^2\theta\rangle_T)}{2}\bigg[1+\mathrm{erf}\bigg(\frac{-P_X^0\langle\cos^2\theta\rangle_T}{\sqrt{2}\sigma_N}\bigg)\bigg]\\=1-0.95$$

$$\langle\cos^2\theta\rangle_T^\text{excl.}=\frac{1.64\sigma_N}{P_X^0}$$

$$\int_0^1\frac{f(\langle\cos^2\theta\rangle_T)}{2}\bigg[1+\mathrm{erf}\bigg(\frac{5\sigma_N-P_X^0\langle\cos^2\theta\rangle_T}{\sqrt{2}\sigma_N}\bigg)\bigg]\mathrm{d}\langle\cos^2\theta\rangle_T\\=1-0.95$$

$$\langle\cos^2\theta\rangle_T^\text{disc.}=\frac{(5+1.64)\sigma_N}{P_X^0}$$

$$\hat{Z}(t)=\begin{pmatrix} \cos\lambda_\text{lab}\cos\omega_\oplus t \\ \cos\lambda_\text{lab}\sin\omega_\oplus t \\ \sin\lambda_\text{lab} \end{pmatrix}$$

$$\hat{\mathbf{X}}=\begin{pmatrix} \sin\theta_X\cos\phi_X \\ \sin\theta_X\sin\phi_X \\ \cos\theta_X \end{pmatrix}$$

$$\cos^2\theta(t)=(\hat{\mathbf{X}}\cdot\hat{Z}(t))^2$$

$$\langle\cos^2\theta(t)\rangle_T\equiv\frac{1}{T}\int_0^T\cos^2\theta(t)\mathrm{d}t$$

$$\frac{1}{4\pi}\int~\langle\cos^2\theta(t)\rangle_T~\mathrm{d}\cos\theta_X~\mathrm{d}\phi=\frac{1}{3}$$

$$\langle\cos^2\theta(t)\rangle_{T=\eta}=\\\frac{1}{8}(3+\cos2\theta_X-(1+3\cos2\theta_X)\cos2\lambda_\text{lab})$$

$$\hat{\mathcal{W}}(t)=\begin{pmatrix} \sin\omega_\oplus t \\ -\cos\omega_\oplus t \\ 0 \end{pmatrix}$$



$$\hat{\mathcal{N}}(t)=\begin{pmatrix} \sin\lambda_{\text{lab}}\cos\omega_{\oplus}t \\ -\sin\lambda_{\text{lab}}\sin\omega_{\oplus}t \\ \cos\lambda_{\text{lab}} \end{pmatrix}$$

$$\langle \cos^2 \theta(t) \rangle_\eta = \begin{cases} \frac{1}{8}(3 + \cos 2\lambda_{\text{lab}} + (1 + 3\cos 2\lambda_{\text{lab}})\cos 2\theta_X) \\ \frac{\sin^2 \theta_X}{2} \\ \frac{1}{8}(3 + \cos 2\theta_X - (1 + 3\cos 2\theta_X)\cos 2\lambda_{\text{lab}}) \end{cases}$$

$$\lambda_{\text{lab}} = \frac{1}{2}\cos^{-1}\left(-\frac{1}{3}\right) \approx \pm 54.74^\circ \quad \quad \lambda_{\text{lab}} = \frac{1}{2}\cos^{-1}\left(\frac{1}{3}\right) \approx \pm 35.26^\circ$$

$$\cos \theta(t) = \sqrt{1 - (\hat{\mathbf{X}} \cdot \hat{\mathcal{N}}(t))^2},$$

$$\langle \cos^2 \theta(t) \rangle_\eta = \begin{cases} \frac{1}{8}(5 - \cos 2\lambda_{\text{lab}} - (1 + 3\cos 2\lambda_{\text{lab}})\cos 2\theta_X) \\ \frac{1}{4}(3 + \cos 2\theta_X) \\ \frac{1}{8}(5 + \cos 2\theta_X + (3\cos 2\lambda_{\text{lab}} - 1)\cos 2\theta_X) \end{cases}$$

$$\lambda_{\text{lab}} = \frac{1}{2}\cos^{-1}\left(-\frac{1}{3}\right) \approx \pm 54.74^\circ$$

$$\lambda_{\text{lab}} = \frac{1}{2}\cos^{-1}\left(\frac{1}{3}\right) \approx \pm 35.26^\circ$$

$$P_{\text{cav}} = \kappa \mathcal{G}_X V Q \rho_{\text{DM}} \chi^2 m_X \\ P_{\text{cav}} = \kappa \mathcal{G}_a V \frac{Q}{m_a} \rho_{\text{DM}} g_{a\gamma}^2 B^2$$

$$\mathcal{G}_X = \frac{\left(\int dV \mathbf{E}_\alpha \cdot \hat{\mathbf{X}}\right)^2}{V \frac{1}{2} \int dV \epsilon(\mathbf{x}) \mathbf{E}_\alpha^2 + \mathbf{B}_\alpha^2} \\ \mathcal{G}_a = \frac{\left(\int dV \mathbf{E}_\alpha \cdot \mathbf{B}\right)^2}{VB^2 \frac{1}{2} \int dV \epsilon(\mathbf{x}) \mathbf{E}_\alpha^2 + \mathbf{B}_\alpha^2}$$

$$\chi = g_{a\gamma} \frac{B}{m_X |\cos \theta|}$$

$$P_X \langle \cos^2 \theta \rangle_T^{\text{disc.}} > (3 + \Phi^{-1}[0.5])\sigma_N \approx 3\sigma_N,$$

$$\xi = \int_0^1 \frac{f(\langle \cos^2 \theta \rangle_T)}{2} \left[1 + \text{erf}\left(\frac{3\sigma_N - P_X \langle \cos^2 \theta \rangle_T}{\sqrt{2}\sigma_N}\right) \right] d\langle \cos^2 \theta \rangle_T$$



$$\Lambda(\theta_X, \phi_X) = 2 \left[\ln \mathcal{L}(d \mid \mathcal{M}_t, \{\hat{P}_X, \theta_X, \phi_X, \hat{P}_N\}) - \ln \mathcal{L}(d \mid \mathcal{M}_0, \{\hat{P}_X, \hat{P}_N\}) \right]$$

$$\begin{aligned} \ln \mathcal{L}(P^{\text{obs}} \mid \mathcal{M}_t, \{P_X^0, \theta_X, \phi_X, P_N\}) = \\ -\frac{1}{2\sigma_N^2} \sum_{j=1}^{N_t} \left[P_j^{\text{obs}} - P_X \langle c(\theta_X, \phi_X) \rangle_j - \frac{\Delta t}{T} P_N \right]^2 \end{aligned}$$

$$\langle c(\theta_X, \phi_X) \rangle_j \equiv \langle \cos^2 \theta \rangle_j = \frac{1}{\Delta t} \int_{t_j - \Delta t/2}^{t_j + \Delta t/2} \cos^2 \theta \, dt$$

$$\ln \mathcal{L}(P^{\text{obs}} \mid \mathcal{M}_t, \{P_X, P_N\}) = -\frac{1}{2\sigma_N^2} (P^{\text{obs}} - P_X - P_N)^2$$

$$P_j^{\text{obs}} = P_j^{\text{Asi}} \equiv P_X \langle c(\theta_X, \phi_X) \rangle_j + \frac{\Delta t}{T} P_N$$

$$\text{TS}_{\text{mod}}(\theta_X, \phi_X) = \frac{1}{\sigma_N^2} \left(\sum_j P_X \langle c(\theta_X, \phi_X) \rangle_j - \hat{P}_X \right)^2$$

$$\text{TS}_{\text{mod}}(\theta_X, \phi_X) = \left(\frac{P_X}{\sigma_N} \right)^2 \left(\sum_j \langle c \rangle_j - \langle c \rangle_T \right)^2 = \left(\frac{P_X}{\sigma_N} \right)^2 \left(\frac{1}{\Delta t} \int_0^T c \, dt - \langle c \rangle_T \right)^2$$

$$= \left(\frac{T \langle c \rangle_T P_X}{\sigma_N} \right)^2 \left(\frac{1}{\Delta t} - \frac{1}{T} \right)^2 \approx \left(\frac{T \langle c \rangle_T P_X}{\Delta t \sigma_N} \right)^2$$

$$\text{TS}_{\text{mod}} \approx \left(\frac{3 T \langle c \rangle_T}{\Delta t \langle \cos^2 \theta \rangle_{\Delta t}^{\text{disc}}} \right)^2$$

$$\Lambda(\theta_X, \phi_X) = \left(\frac{3}{\langle \cos^2 \theta \rangle_{\Delta t}^{\text{disc}}} \right)^2 \left[\left(T \langle c_{\text{true}} \rangle_T \left(\frac{1}{\Delta t} - \frac{1}{T} \right) \right)^2 - \frac{1}{\Delta t} \int_0^T (c_{\text{true}} - c)^2 \, dt \right]$$

$$\begin{pmatrix} \sin \theta_X \cos \phi_X \\ \sin \theta_X \sin \phi_X \\ \cos \theta_X \end{pmatrix} = R_{\text{gal}} \begin{pmatrix} \cos l \cos b \\ \sin l \cos b \\ \sin b \end{pmatrix}$$

$$R_{\text{gal}} = \begin{pmatrix} -0.05487556 & +0.49410943 & -0.86766615 \\ -0.87343709 & -0.44482963 & -0.19807637 \\ -0.48383502 & +0.74698225 & +0.45598378 \end{pmatrix}$$



Apéndice C.

Modelo de Multidimensiones y Supermembranas a propósito de la deformación dfel espacio – tiempo cuántico por interacción de una partícula supermasiva. Modelo Sigma.

$$S_{\text{sigma-model}} = -T \int d^2\sigma \frac{1}{2} \sqrt{-\gamma} \gamma^{ab} \partial_a X^\mu \partial_b X^\nu g_{\mu\nu}$$

$$\begin{aligned} T_{ab} &= \left(\partial_a X^\mu \partial_b X^\nu - \frac{1}{2} \gamma_{ab} \gamma^{cd} \partial_c X^\mu \partial_d X^\nu \right) g_{\mu\nu} \\ &\quad \frac{1}{\sqrt{-\gamma}} \partial_a (\sqrt{-\gamma} \gamma^{ab} \partial_b X^\mu) + \gamma^{ab} \partial_a X^\nu \partial_b X^\lambda \Gamma_{\nu\lambda}^\mu \end{aligned}$$

$$\Phi(\varphi) = \frac{1}{2} \epsilon_{ij} \epsilon^{ab} \partial_a \varphi^i \partial_b \varphi^j$$

$$S = - \int d^2\sigma \Phi(\varphi) \left(\frac{1}{2} \gamma^{ab} \partial_a X^\mu \partial_b X^\nu g_{\mu\nu} - \frac{\epsilon^{ab}}{2\sqrt{-\gamma}} F_{ab}(A) \right)$$

$$\begin{aligned} \gamma_{ab} &\rightarrow J \gamma_{ab} \\ \varphi^i &\rightarrow \varphi'^i = \varphi'^i(\varphi^i) \end{aligned}$$

$$\Phi \rightarrow \Phi' = J\Phi$$

$$\epsilon^{ab} \partial_b \varphi^i \partial_a \left(\gamma^{cd} \partial_c X^\mu \partial_d X^\nu g_{\mu\nu} - \frac{\epsilon^{cd}}{\sqrt{-\gamma}} F_{cd} \right) = 0$$

$$\gamma^{cd} \partial_c X^\mu \partial_d X^\nu g_{\mu\nu} - \frac{\epsilon^{cd}}{\sqrt{-\gamma}} F_{cd} = M = \text{const.}$$

$$T_{ab} = \partial_a X^\mu \partial_b X^\nu g_{\mu\nu} - \frac{1}{2} \gamma_{ab} \frac{\epsilon^{cd}}{\sqrt{-\gamma}} F_{cd} = 0$$

$$\epsilon^{ab} \partial_b \left(\frac{\Phi(\varphi)}{\sqrt{-\gamma}} \right) = 0$$

$$\frac{\Phi(\varphi)}{\sqrt{-\gamma}} = T$$

$$S_{\text{sigma-model}} = -T \int d^d\sigma \frac{1}{2} \sqrt{-\gamma} (\gamma^{ab} \partial_a X^\mu \partial_b X^\nu g_{\mu\nu} + 2\Lambda)$$

$$\begin{aligned} T_{ab} &= \left(\partial_a X^\mu \partial_b X^\nu - \frac{1}{2} \gamma_{ab} \gamma^{cd} \partial_c X^\mu \partial_d X^\nu \right) g_{\mu\nu} - \gamma_{ab} \Lambda \\ &\quad \frac{1}{\sqrt{-\gamma}} \partial_a (\sqrt{-\gamma} \gamma^{ab} \partial_b X^\mu) + \gamma^{ab} \partial_a X^\nu \partial_b X^\lambda \Gamma_{\nu\lambda}^\mu \end{aligned}$$

$$\Phi(\varphi) = \epsilon_{ijk\dots m} \epsilon^{abc\dots d} \partial_a \varphi^i \partial_b \varphi^j \dots \partial_d \varphi^m$$



$$S=-\int \,\,d^d\sigma \Phi(\varphi)\left(\frac{1}{2}\gamma^{ab}\partial_aX^\mu\partial_bX^\nu g_{\mu\nu}-\frac{\epsilon^{abcd...}}{2\sqrt{-\gamma}}F_{abcd...}(A)\right)$$

$$\begin{array}{c} \gamma_{ab} \rightarrow J \gamma_{ab} \\ \varphi^i \rightarrow \varphi'^i = \lambda^{ij} (\varphi^j) \end{array}$$

$$\Phi \rightarrow \Phi' = J \Phi$$

$$A_{bcd.....}\rightarrow J^{\frac{d-2}{2}}A_{bcd.....}$$

$$K^a_b\partial_a\bigg(\gamma^{cd}\partial_cX^\mu\partial_dX^\nu g_{\mu\nu}-\frac{\epsilon^{cdef}}{\sqrt{-\gamma}}\cdots F_{cdef}\cdots\bigg)$$

$$\gamma^{cd}\partial_cX^\mu\partial_dX^\nu g_{\mu\nu}-\frac{\epsilon^{cdef}}{\sqrt{-\gamma}}\cdots F_{cdef}\cdots=M=\textrm{ const.}$$

$$T_{ab}=\partial_a X^\mu\partial_b X^\nu g_{\mu\nu}-\frac{1}{2}\gamma_{ab}\frac{\epsilon^{cdef}}{\sqrt{-\gamma}}\cdots F_{cdef}\cdots=0$$

$$\gamma_{ab}=\frac{1-p}{M}\partial_a X^\mu\partial_b X^\nu g_{\mu\nu}$$

$$\epsilon^{abc..d}\partial_d\left(\frac{\Phi(\varphi)}{\sqrt{-\gamma}}\right)=0$$

$$\frac{\Phi(\varphi)}{\sqrt{-\gamma}}=T$$

$$S_{\text{current}}\,=\int \,\,d^{p+1}\sigma A_{a_2...a_{p+1}} j^{a_2...a_{p+1}}$$

$$\epsilon^{a_1...a_{p+1}}\partial_{a_1}\left(\frac{\Phi}{\sqrt{-\gamma}}\right)=j^{a_2...a_{p+1}}$$

$$j^{a_1...a_{p+1}}=e\partial_\mu\phi\frac{\partial X^\mu}{\partial\sigma^a}\epsilon^{aa_2...a_{p+1}}\equiv e\partial_a\phi\epsilon^{aa_2...a_{p+1}}$$

$$T=\frac{\Phi}{\sqrt{-\gamma}}=e\phi+T_i$$

$$g_{\mu\nu}\rightarrow \omega g_{\mu\nu}$$

$$\begin{array}{c} A_a \rightarrow \omega A_a \\ \Phi(\varphi) \rightarrow \omega^{-1} \Phi(\varphi) \end{array}$$

$$\phi \rightarrow \omega^{-1} \phi$$

$$\Phi=\sqrt{-\gamma}(e\phi+T_i)$$

$$S_i=-\int \,\,d^2\sigma(e\phi+T_i)\frac{1}{2}\sqrt{-\gamma}\gamma^{ab}\partial_aX^\mu\partial_bX^\nu g_{\mu\nu}+\int \,\,d^2\sigma A_{\mu\nu}\epsilon^{ab}\partial_aX^\mu\partial_bX^\nu+\int \,\,d^2\sigma\sqrt{-\gamma}\varphi R$$



$$g^i_{\mu\nu}=(e\phi+T_i)g_{\mu\nu}$$

$$R_{\mu\nu}\big(g^1_{\alpha\beta}\big)=0$$

$$R_{\mu\nu}\big(g^2_{\alpha\beta}\big)=0$$

$$ds_2^2=-(1-2GM/r)dt^2+\frac{dr^2}{1-2GM/r}+r^2d\Omega^2+dy_4^2+dy_5^2+\cdots...+dy_{25}^2$$

$$ds_2^2=-dt^2+t^{2p_1}dx_1^2+t^{2p_2}dx_2^2+\cdots...+t^{2p_{25}}dx_{25}^2$$

$$e\phi+T_1=\Omega^2(e\phi+T_2)$$

$$e\phi=\frac{\Omega^2 T_2-T_1}{1-\Omega^2}$$

$$e\phi+T_1=\frac{\Omega^2(T_2-T_1)}{1-\Omega^2}$$

$$e\phi+T_2=\frac{(T_2-T_1)}{1-\Omega^2}$$

$$ds^2=-dt^2+t^2(d\chi^2+\sinh^2~\chi d\Omega_2^2)$$

$$ds_2^2=-dt^2+t^2(d\chi^2+\sinh^2~\chi d\Omega_{D-2}^2)$$

$$ds_1^2=\frac{\sigma}{t^4}\Bigl(-dt^2+t^2(d\chi^2+\sinh^2~\chi d\Omega_{D-2}^2)\Bigr)$$

$$ds^2=g_{\mu\nu}dx^\mu dx^\nu=\left(\frac{1-\Omega^2}{T_2-T_1}\right)\Bigl(-dt^2+t^2(d\chi^2+\sinh^2~\chi d\Omega_{D-2}^2)\Bigr)$$

$$dT=\sqrt{\frac{1+\frac{K}{t^4}}{T_1-T_2}}dt$$

$$ds^2=l^2dt^2-l^2\cosh^2~t\left(\frac{dr^2}{1-r^2}+r^2d\Omega_2^2\right)\Bigr)-dl^2$$

$$ds^2=\frac{\Lambda l^2}{3}\Biggl(dt^2\Biggl(1-\frac{2M}{r}-\frac{\Lambda r^2}{3}\Biggr)-\frac{dr^2}{1-\frac{2M}{r}-\frac{\Lambda r^2}{3}}-r^2d\Omega_2^2\Biggr)-dl^2$$

$$ds_2^2=l^2\bar g_{\mu\nu}(x)dx^\mu dx^\nu-dl^2$$

$$ds_1^2=\sigma l^{-2}\bar g_{\mu\nu}(x)dx^\mu dx^\nu-\sigma \frac{dl^2}{l^4}=\sigma l^{-4}ds_2^2$$

$$ds^2=\left(\frac{1-\Omega^2}{T_2-T_1}\right)\bigl(l^2\bar g_{\mu\nu}(x)dx^\mu dx^\nu-dl^2\bigr)$$



$$dL = \sqrt{\frac{1+\frac{K}{l^4}}{T_1-T_2}}dl$$

$$ds_1^2=\eta_{\alpha\beta}dx^\alpha dx^\beta$$

$$ds_2^2=\Omega(x)^2\eta_{\alpha\beta}dx^\alpha dx^\beta$$

$$x''^\mu = \frac{(x^\mu + a^\mu x^2)}{(1+2a_\nu x^\nu + a^2 x^2)}$$

$$g_{\alpha\beta}^2=\Omega^2\eta_{\alpha\beta}=\frac{1}{\left(1+2a_\mu x^\mu+a^2x^2\right)^2}\eta_{\alpha\beta}$$

$$\begin{aligned} e\phi+T_1 &= \frac{(T_2-T_1)\left(1+2a_\mu x^\mu+a^2x^2\right)^2}{\left(1+2a_\mu x^\mu+a^2x^2\right)^2-1}=\frac{(T_2-T_1)\left(1+2a_\mu x^\mu+a^2x^2\right)^2}{(2a_\mu x^\mu+a^2x^2)(2+2a_\mu x^\mu+a^2x^2)} \\ e\phi+T_2 &= \frac{(T_2-T_1)}{\left(1+2a_\mu x^\mu+a^2x^2\right)^2-1}=\frac{(T_2-T_1)}{(2a_\mu x^\mu+a^2x^2)(2+2a_\mu x^\mu+a^2x^2)} \end{aligned}$$

$$e\phi+T_1=e\phi+T_2=\frac{K}{4t}$$

$$g_{\mu\nu}=\frac{1}{(e\phi+T_1)}\,g_{\mu\nu}^1=\frac{4t}{K}\,\eta_{\mu\nu}$$

$$(T_2-T_1)\Delta t \approx \text{ constant}$$

$$2a_\mu x^\mu+a^2x^2=0$$

$$2+2a_\mu x^\mu+a^2x^2=0$$

$$(2a_\mu x^\mu+a^2x^2)(2+2a_\mu x^\mu+a^2x^2)=(2At+A^2(t^2-x^2))(2+2At+A^2(t^2-x^2))$$

$$x_1^2+x_2^2+x_3^2\dots\dots+x_{D-1}^2-\left(t+\frac{1}{A}\right)^2=-\frac{1}{A^2}$$

$$A(x_1^2+x_2^2+x_3^2\dots\dots+x_{D-1}^2)-At^2-2t=0$$

$$x_1^2+x_2^2+x_3^2\dots\dots+x_{D-1}^2-\left(t+\frac{1}{A}\right)^2=\frac{1}{A^2}$$

$$\sqrt{\frac{1}{A^2}+\left(t+\frac{1}{A}\right)^2}-\sqrt{-\frac{1}{A^2}+\left(t+\frac{1}{A}\right)^2}\rightarrow\frac{1}{tA^2}\rightarrow 0$$

$$\pi T=(\kappa^2)^{-\frac{1}{d-2}}$$

$$\kappa^2=8\pi G=\frac{1}{M_P^2}$$



$$S = \int d^4x \left(\sqrt{-g} \left(-\frac{\epsilon}{2} \phi^2 R + X - V(\phi) \right) \right.$$

$$V(\phi)=\frac{1}{8}\lambda(\phi^2-v^2)^2$$

$$\Phi(\varphi)\rightarrow\omega^{-1}\Phi(\varphi)$$

$$\Delta\phi/M_P < {\cal O}(1)$$

$$M_P \frac{dV/d\phi}{V} > {\cal O}(1)$$

$$-M_P^2 \frac{d^2V/d\phi^2}{V} > {\cal O}(1)$$

$$e\phi+T_1=\Omega_{12}^2(e\phi+T_2)$$

$$e\phi=\frac{\Omega_{12}^2 T_2-T_1}{1-\Omega_{12}^2}$$

$$e\phi+T_1=\frac{\Omega_{12}^2(T_2-T_1)}{1-\Omega_{12}^2}$$

$$e\phi+T_2=\frac{(T_2-T_1)}{1-\Omega_{12}^2}$$

$$e\phi+T_2=\Omega_{23}^2(e\phi+T_3)$$

$$e\phi=\frac{\Omega_{23}^2 T_3-T_2}{1-\Omega_{23}^2}$$

$$\Omega_{23}=\frac{T_1+\Omega_{12}^2(T_3-T_2)}{\Omega_{12}^2(T_3-T_2)-T_3}$$

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Anexo A.

Modelo matemático para calcular agujeros negros cuánticos en campos cuánticos relativistas o curvos con aproximaciones en supergravedad cuántica.

$$S_{\text{EH}}[g_{\mu\nu}] = \frac{1}{16\pi G_N} \int_x \sqrt{g}(R - 2\Lambda),$$

$$g_{\mu\nu} = \bar{g}_{\mu\nu} + \sqrt{Z_h G} h_{\mu\nu}$$

$$Z_h = Z_h(\Delta), G = G(\Delta)$$

$$\Delta = -g^{\mu\nu}\nabla_\mu\nabla_\nu$$

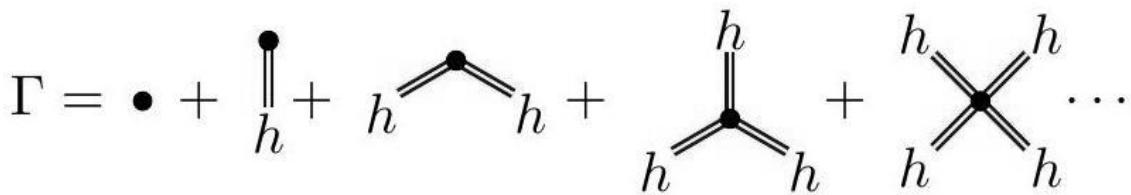
$$g_{\mu\nu} = \eta_{\mu\nu} = \delta_{\mu\nu}$$

$$\Gamma[g] = \Gamma[g, h = 0],$$

$$\Gamma[g_{\mu\nu}] = \frac{1}{16\pi} \int_x \sqrt{g} [\mathcal{R}(\Delta, R) + R f_{R^2}(\Delta) R + R_{\mu\nu} f_{R_{\mu\nu}^2}(\Delta) R^{\mu\nu}]$$

$$\int_x \sqrt{g} \frac{1}{G(\Delta)} R \simeq \int_x \sqrt{g} \frac{1}{G(0)} R$$

$$\begin{aligned} \text{IR: } & \mathcal{R}(\Delta, R) \xrightarrow{G_N \Delta \rightarrow 0} \frac{1}{G_N} R, \\ \text{UV: } & \mathcal{R}(\Delta, R) \xrightarrow{G_N \Delta \rightarrow \infty} 0, \end{aligned}$$



$$\Gamma^{(n)}[g] = \Gamma^{(n,0)}[g, 0] \approx (Z_h G)^{-n/2} \Gamma^{(0,n)}[g, 0]$$



$$\Gamma^{(n,m)}[\bar{g},h]=\frac{\delta^{n+m}\Gamma[\bar{g},h]}{\delta \bar{g}^n\delta h^m}$$

$$\begin{aligned}\Gamma_{\rm tt}^{(2,0)}(p) &= \frac{1}{16\pi G_N} Z_g(p)p^2\Pi_{\rm tt}\\ \Gamma_{\rm tt}^{(0,2)}(p) &= \frac{1}{16\pi} Z_h(p)p^2\Pi_{\rm tt}\end{aligned}$$

$$G(p)=\frac{G_N}{Z_g(p)}$$

$$\Gamma_{\rm tt}^{(n)}(\pmb p)=\gamma_g^{(n)}(p)\mathcal T_{R,{\rm tt}}^{(n)}(\pmb p)$$

$$\Gamma_{\rm tt}^{(0,n)}(\pmb p)=\prod_{i=1}^n\sqrt{Z_h(p_i)G(p_i)}\gamma_g^{(n)}(p)\mathcal T_{R,{\rm tt}}^{(n)}(\pmb p)$$

$$\gamma_h^{(n)}(\pmb p)=\prod_{i=1}^n\sqrt{G(p_i)}\gamma_g^{(n)}(\pmb p)$$

$$\bar{\Gamma}^{(0,n)}(\pmb p)=\frac{\Gamma^{(0,n)}(\pmb p)}{\prod_{i=1}^n\sqrt{Z_h(p_i)}}$$

$$G(p)=\left[\gamma_h^{(3)}(p)\right]^2=\left[\gamma_g^{(3)}(p)\right]^{-1}$$

$$\bar g_{\mu\nu} = \delta_{\mu\nu},$$

$$\partial_t\Gamma_k\left[\bar{g},h\right]=\frac{1}{2}\;\;\textcolor{black}{\circlearrowleft\!\!\!\circlearrowright}\;\;-\;\;\textcolor{black}{\circlearrowleft\!\!\!\circlearrowright\!\!\!\circlearrowleft\!\!\!\circlearrowright}$$

$$\partial_t\Gamma_k[\bar{g},h]=\frac{1}{2}\mathrm{Tr}\mathcal{G}_k[\bar{g},h]\partial_t\mathcal{R}_k$$



$$\mathcal{G}_k[\bar{g}, h] = \frac{1}{\Gamma_k^{(0,2)}[\bar{g}, h] + \mathcal{R}_k},$$

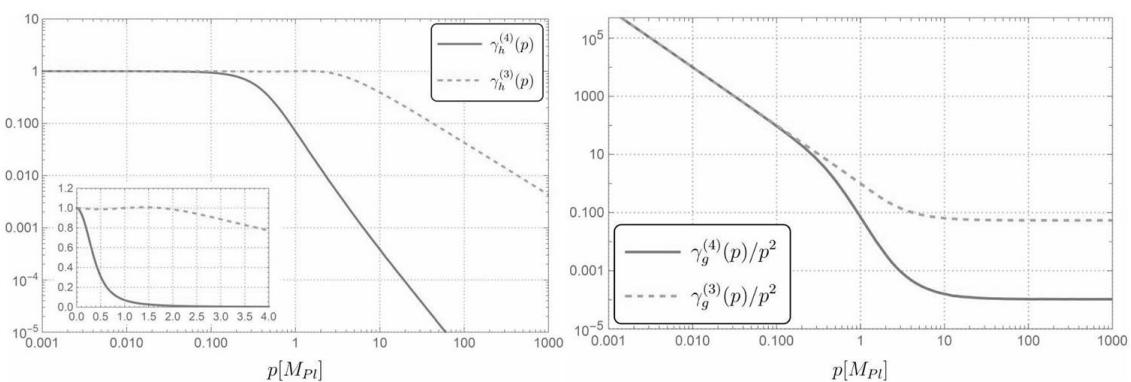
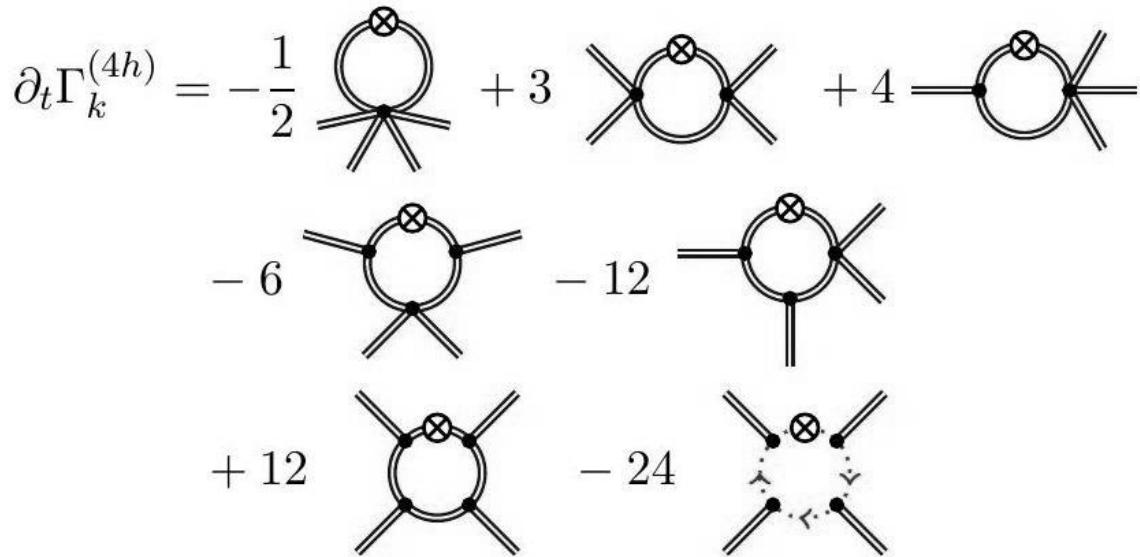


Figura 1. Fluctuaciones de materia y energía de un agujero negro cuántico causado por una partícula supermasiva.

$$\Gamma_{tt}^{(n)} \circ \mathcal{T}_{R,tt}^{(n)} / p^2 = C_R^{(n)} \gamma_g^{(n)} p^2$$

$$C_R^{(n)} \gamma_g^{(n)} p^2 = C_{\mathcal{R}}^{(n)} g_{\mathcal{R}} p^2 + C_{R^2}^{(n)} f_{R^2} p^4 + C_{R_{\mu\nu}^2}^{(n)} f_{R_{\mu\nu}^2} p^4$$

$$f_{R_{\mu\nu}^2}(p^2) = \frac{1}{C_{R_{\mu\nu}^2}^{(3)}} \frac{C_R^{(3)} \gamma_g^{(3)} - C_{\mathcal{R}}^{(3)} g_{\mathcal{R}}}{p^2}$$

$$f_{R^2}(p^2) = \frac{1}{C_{R^2}^{(4)}} \frac{C_R^{(4)} \gamma_g^{(4)} - C_{\mathcal{R}}^{(4)} g_{\mathcal{R}} - C_{R_{\mu\nu}^2}^{(4)} f_{R_{\mu\nu}^2} p^2}{p^2}$$

$$\mathcal{R}(\Delta, R) = R \frac{\gamma_g^{(3)}(\Delta) - \bar{\gamma}_3 \Delta}{\Delta + R} R$$

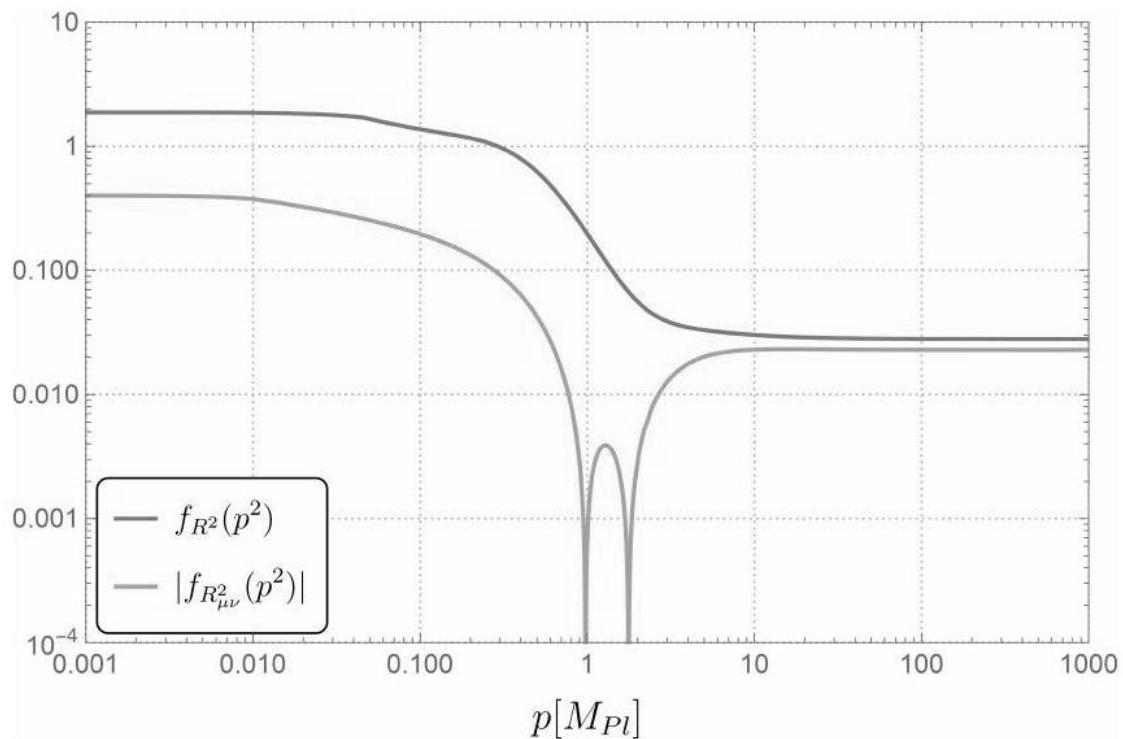


Figura 2. Rangos de momentum del agujero negro cuántico.

$$f_{R_{\mu\nu}^2}(\Delta) \simeq \tilde{g}_{R_{\mu\nu}^2} + \frac{g_{R_{\mu\nu}^2} - \tilde{g}_{R_{\mu\nu}^2}}{1 + p_0^{-2}\Delta}$$

$$f_{R^2}(\Delta) \simeq \tilde{g}_{R^2} + \frac{g_{R^2} - \tilde{g}_{R^2}}{1 + p_1^{-2}\Delta}$$

$$f_{R_{\mu\nu}^2}(\Delta) \simeq g_{R_{\mu\nu}^2} + c_1\Delta$$

$$f_{R^2}(\Delta) \simeq g_{R^2} + c_2\Delta$$

$$g_{R_{\mu\nu}^2} = c_3 \left[\gamma_g^{(3)} \right]'(0)$$

$$g_{R^2} = c_4 \left[\gamma_g^{(4)} \right]'(0) - c_5 \left[\gamma_g^{(3)} \right]'(0)$$

$$2 \left(R_{\mu\nu} - \frac{1}{4} R g_{\mu\nu} + g_{\mu\nu} \square - \nabla_\mu \nabla_\nu \right) f_{R^2}(\square) R = 0$$

$$\left(2R_{\mu\sigma}g_{\nu\rho} - \frac{1}{2}g_{\mu\nu}R_{\rho\sigma} + g_{\rho\mu}g_{\sigma\nu} \square \right) f_{R_{\mu\nu}^2}(\square) R^{\rho\sigma}$$

$$+ (g_{\mu\nu}\nabla_\rho\nabla_\sigma - 2g_{\mu\sigma}\nabla_\rho\nabla_\nu) f_{R_{\mu\nu}^2}(\square) R^{\rho\sigma} = 0$$

$$(\square + R)\mathcal{G}(x, x') = \frac{1}{\sqrt{-g}}\delta(x - x')$$

$$\Gamma_{\text{IR}}[g_{\mu\nu}] = \frac{1}{16\pi} \int_x \sqrt{-g} \left(G_N^{-1} R + g_{R_{\mu\nu}^2} R_{\mu\nu} R^{\mu\nu} + g_{R^2} R^2 + c_1 R_{\mu\nu} \square R^{\mu\nu} + c_2 R \square R \right)$$

$$ds^2 = -f(r)dt^2 + \frac{1}{g(r)}dr^2 + r^2 d\Omega^2$$

$$d\Omega^2 = d\theta^2 + \sin^2(\theta)d\varphi^2$$



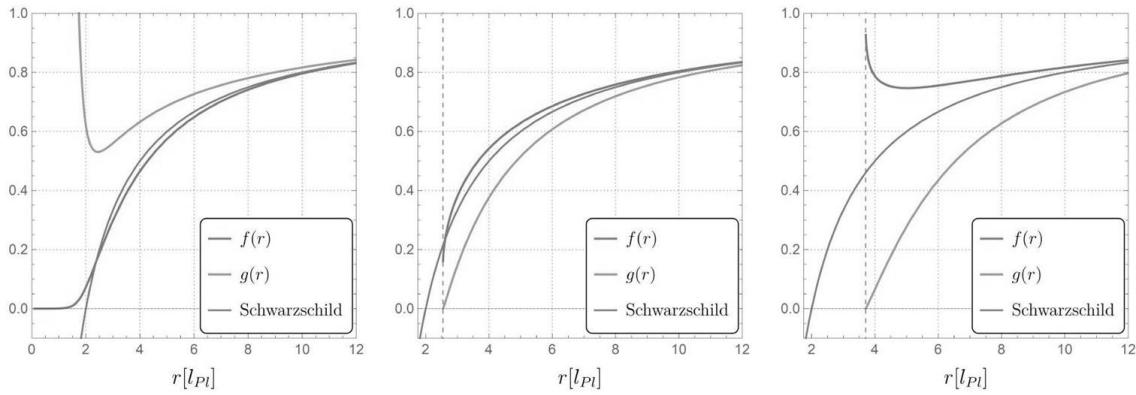


Figura 3. Fluctuaciones de simetría para agujeros negros cuánticos.

$$f(r \gg 1) = 1 - \frac{2M}{r} + S_0 \frac{e^{-m_0 r}}{r} + S_2 \frac{e^{-m_2 r}}{r}$$

$$g(r \gg 1) = 1 - \frac{2M}{r} - S_0 \frac{e^{-m_0 r}}{r} (1 + m_0 r) + \frac{1}{2} S_2 \frac{e^{-m_2 r}}{r} (1 + m_2 r)$$

$$T = \frac{1}{4\pi} \sqrt{|f'(r_h)g'(r_h)|}$$

$$S_{gf}[\bar{g}, h] = \frac{1}{2\alpha} \int_x \sqrt{\bar{g}} \bar{g}^{\mu\nu} F_\mu F_\nu$$

$$F_\mu[\bar{g}, h] = \bar{\nabla}^\nu h_{\mu\nu} - \frac{1+\beta}{4} \bar{\nabla}_\mu h^\nu$$

$$S_{gh}[\bar{g}, h, c, \bar{c}] = \int_x \sqrt{\bar{g}} \bar{c}^\mu \mathcal{M}_{\mu\nu} c^\nu$$

$$\mathcal{M}_{\mu\nu} = \bar{\nabla}^\rho (g_{\mu\nu} \nabla_\rho + g_{\rho\nu} \nabla_\mu) - \frac{1+\beta}{2} \bar{g}^{\sigma\rho} \bar{\nabla}_\mu g_{\nu\sigma} \nabla_\rho$$

$$\Gamma_k[\bar{g}, \phi] = \sum_{n=0}^{\infty} \frac{1}{n!} \Gamma_k^{(0, \phi_{i_1} \dots \phi_{i_n})}[\bar{g}, 0] \prod_{l=1}^n \phi_{i_l}$$

$$\phi = (h, \bar{c}, c)$$

$$\Gamma_k^{(0, \phi_1 \dots \phi_n)}(\mathbf{p}) = \left(\prod_{i=1}^n Z_{\phi_i}^{\frac{1}{2}}(p_i^2) \right) \gamma_h^{(n)}(\mathbf{p}) \mathcal{T}^{(\phi_1 \dots \phi_n)}(\mathbf{p}; \Lambda_n)$$



$$\mathcal{T}_{\text{EH}}^{(\phi_1 \dots \phi_n)}(\boldsymbol{p}; \Lambda_n) = G_N S_{\text{EH}}^{(\phi_1 \dots \phi_n)}(\boldsymbol{p}; \Lambda \rightarrow \Lambda_n).$$

$$\mathfrak{F}^{(n)}(\hat p)\!:=\!\frac{\partial_t\Gamma^{(n)}(\hat p)}{Z_h^{n/2}k^{4-n}},$$

$$g_n\!:=\!\left(\gamma_h^{(n)}\right)^{\frac{2}{n-2}}k^2,\lambda_n\!:=\!\Lambda_n k^{-2},\mu_h\!:=\!-2\lambda_2$$

$$\Pi^{\mu\nu\alpha\beta}_{\tt tt}\!:=\!\frac{1}{2}\Big(\Pi^{\mu\alpha}_\perp\Pi^{\nu\beta}_\perp+\Pi^{\mu\beta}_\perp\Pi^{\nu\alpha}_\perp\Big)-\frac{1}{3}\Pi^{\mu\nu}_\perp\Pi^{\alpha\beta}_\perp$$

$$\Pi^{\mu\nu}_\perp\!:=\delta^{\mu\nu}-\frac{p^\mu p^\nu}{p^2}$$

$$\mathcal{T}_{R,\text{tt}}^{(n)}(\boldsymbol{p})=\prod_{i=1}^n~\Pi_{\text{tt}}(p_i)\mathcal{T}_{\text{EH}}^{(n)}(\boldsymbol{p},0)$$

$$\begin{aligned}\Gamma^{(0,n)}(\boldsymbol{p})&=\prod_{i=1}^n~\sqrt{Z_h(p_i)G(p_i)}\gamma_g^{(n)}(\boldsymbol{p})\mathcal{T}_{R,\text{tt}}^{(n)}(\boldsymbol{p})+\cdots\\ \Gamma^{(n)}(\boldsymbol{p})&=\gamma_g^{(n)}(\boldsymbol{p})\mathcal{T}_{R,\text{tt}}^{(n)}(\boldsymbol{p})+\cdots\end{aligned}$$

$$\gamma_g^{(n)}(\boldsymbol{p})=\prod_{i=1}^n~\frac{1}{\sqrt{Z_h(p_i)G(p_i)}}\times\frac{\Gamma_{\mu_1\nu_1\cdots\mu_n\nu_n}^{(0,n)}\mathcal{T}_{R,\text{tt}}^{(n)\mu_1\nu_1\cdots\mu_n\nu_n}}{\mathcal{T}_{R,\text{tt}}^{(n)}{}_{\mu_1\nu_1\cdots\mu_n\nu_n}\mathcal{T}_{R,\text{tt}}^{(n)\mu_1\nu_1\cdots\mu_n\nu_n}}$$

$$\gamma_g^{(n)}(\boldsymbol{p})=\frac{\Gamma_{\mu_1\nu_1\cdots\mu_n\nu_n}^{(n)}\mathcal{T}_{R,\text{tt}}^{(n)\mu_1\nu_1\cdots\mu_n\nu_n}}{\mathcal{T}_{R,\text{tt}}^{(n)}{}_{\mu_1\nu_1\cdots\mu_n\nu_n}\mathcal{T}_{R,\text{tt}}^{(n)\mu_1\nu_1\cdots\mu_n\nu_n}}$$

$$\Gamma_G^{(0,n)}(p^2):=\mathcal{T}_{R,\text{tt}}^{(n)}/p^2\circ\Gamma_k^{(0,n)}=Z_h^{\frac{n}{2}}g_n^{\frac{n}{2}-1}k^{2-n}\left(C_{\Lambda_n}^{(n)}\lambda_nk^2+C_R^{(n)}p^2\right)$$

$$\mathcal{T}_{R,\text{tt}}^{(n)}/p^2\circ S_{R^2+R^2_{\mu\nu}}^{(n)}=C_{R^2}^{(n)}g_{R^2}p^4+C_{R^2_{\mu\nu}}^{(n)}g_{R^2_{\mu\nu}}p^4$$



$$\begin{array}{ll} C_{\Lambda_3}^{(3)}=-\frac{9}{4096\pi^2}, & C_{\Lambda_4}^{(4)}=\frac{222485}{60466176\pi^2}\\ C_R^{(3)}=\frac{171}{32768\pi^2}, & C_R^{(4)}=\frac{6815761}{544195584\pi^2}\\ C_{R_{\mu\nu}^2}^{(3)}=-\frac{405}{32768\pi^2}, & C_{R_{\mu\nu}^2}^{(4)}=-\frac{3676621}{51018336\pi^2}\\ C_{R^2}^{(3)}=0 & , \,\,\, C_{R^2}^{(4)}=-\frac{96203921}{1632586752\pi^2}\end{array}$$

$$\mathcal{R}_k(p^2)=\Gamma_k^{(0, hh)}(p^2)\Big|_{\mu_h=0}r(\hat{p}^2)=Z_h(p^2)\mathcal{T}^{(2)}(p^2;0)r(\hat{p}^2)$$

$$r(\hat{p}^2) = \left(\frac{1}{\hat{p}^2}-1\right)\Theta(1-\hat{p}^2)$$

$$\frac{\dot{g}_4(p^2)}{g_4(p^2)}=2\big(1+\eta_h(p^2)\big)+2C_4\frac{k^2}{p^2}(\eta_h(p^2)-\eta_h(0))\lambda_4+\frac{1}{C_R^{(4)}}\frac{k^2}{p^2}\bigg(\frac{\mathfrak{F}_G^{(4)}(p^2)}{g_4(p^2)}-\frac{\mathfrak{F}_G^{(4)}(0)}{g_4(0)}\bigg)$$

$$\eta_{\phi_i}(p^2)\!:= -\partial_t\!\ln\left(Z_{\phi_i}(p^2)\right)$$

$$g_6=g_5=g_4 \text{ and } \lambda_6=\lambda_5=\lambda_3$$

$$\mu_h=-0.26, \lambda_3=0.145 \text{ and } \lambda_4=0.025$$

$$\mathfrak{F}_G^{(4)}(\hat{p})=\frac{g_3^2}{g_4}(c_1+c_2\eta_h)+\sqrt{g_3g_4}(c_3+c_4\eta_h)+g_3(c_5+c_6\eta_h)+g_4(c_7+c_8\eta_h)$$

$$(g_3^*, g_4^*)=(2.15, 1.48)$$



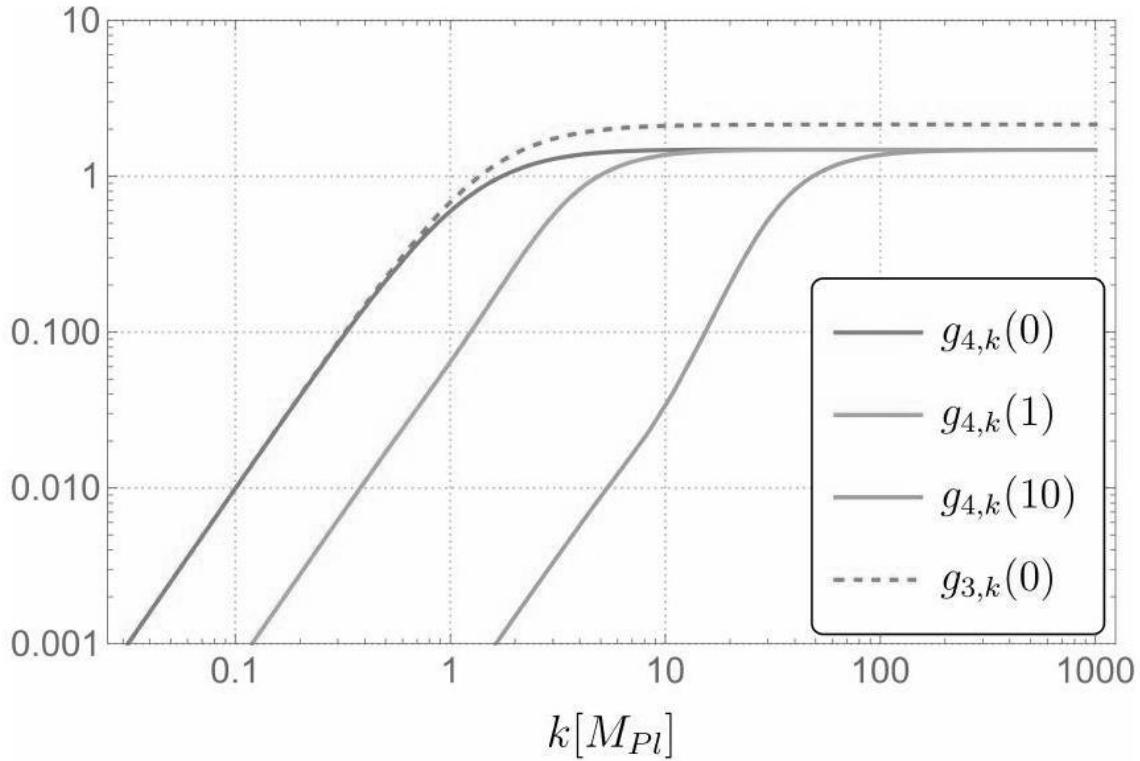


Figura 4. Coordenadas de un agujero negro cuántico en curvatura.

$$\bar{\gamma}_3 = \lim_{p^2 \rightarrow \infty} \left(\frac{\gamma_g^{(3)}(p^2)}{p^2} \right)$$

$$\bar{\gamma}_3 \approx 0.054$$

$$R^{(n)} \left(\frac{\gamma_g^{(3)} - \bar{\gamma}_3 \Delta}{\Delta + R} R \right) \Bigg|_{\Delta \rightarrow 0, R \rightarrow 0}$$

$$g_{R_{\mu\nu}^2} \approx -0.40, g_{R^2} \approx 1.9$$

$$\tilde{g}_{R_{\mu\nu}^2} \approx -0.023, \tilde{g}_{R^2} \approx 0.028$$

$$c_1 = 344.09, c_2 = -136.75$$

$$c_3 = \frac{C_R^{(3)}}{C_{R_{\mu\nu}^2}^{(3)}}, c_4 = \frac{C_R^{(4)}}{C_{R^2}^{(4)}}, c_5 = \frac{C_{R_{\mu\nu}^2}^{(4)}}{C_{R^2}^{(4)}} c_3$$



$$m_2^2=\frac{-G_N^{-1}}{g_{R_{\mu\nu}^2}}\approx 2.5M_{\rm pl}^2$$

$$m_0^2=\frac{G_N^{-1}}{6g_{R^2}+2g_{R_{\mu\nu}^2}}\approx 0.095M_{\rm pl}^2$$

$$\gamma^{(4)}_{h,\,\text{simple}}\left(p^2\right)=\frac{G_4^*}{G_4^*+p^2}$$

$$\gamma^{(4)}_{h,\text{fit}}(p^2) = \frac{a_1 + a_2 p^2 + a_3 p^4 + a_4 p^6 + a_5 p^8}{b_1 + b_2 p^2 + b_3 p^4 + b_4 p^6 + b_5 p^8 + b_6 p^{10}}$$

$$\begin{array}{lll} a_1=0.0075542, & a_2=7.3967, & a_3=109.42, \\ a_4=122.40, & a_5=8.4140, & \\ b_1=0.0075403, & b_2=7.4773, & b_3=153.00, \\ b_4=876.09, & b_5=2261.9, & b_6=228.87. \end{array}$$

$$p_0\approx 0.091916, p_1\approx 0.27551$$

$$f_{R_{\mu\nu}^2,\,\text{fit}}\left(p^2\right)=a_0+\sum_{i=1}^4\frac{a_i}{(p/p_i)^2+1}$$

$$\begin{array}{lll} a_0=-0.023601, & a_1=-0.13727, & a_2=0.13138 \\ a_3=-0.22100, & a_4=-0.15080, & p_1=0.12436 \\ p_2=1.2476, & p_3=0.56405, & p_4=0.021230 \end{array}$$

$$f_{R^2,\text{fit}}(p^2)=a_0+\sum_{i=1}^3\frac{a_i}{((p/p_i)^2+1)^2}$$

$$\begin{array}{lll} a_0=0.028373, & a_1=0.012637, & a_2=1.2661 \\ a_3=0.57040, & p_1=5.7131, & p_2=0.73200 \\ p_3=0.092956 & & \end{array}$$

$$\Gamma_{\mathrm{QG}}=\Gamma_{\mathrm{L}}+\Gamma_{\mathrm{NL}}+\Gamma_{\mathrm{matter}}$$

$$\begin{aligned}\Gamma_{\mathrm{L}}=&\int d^4x\sqrt{|g|}\Big[\frac{M_P^2}{2}(R-2\Lambda)+c_1(\mu)R^2+c_2(\mu)R_{\mu\nu}R^{\mu\nu}\\&+c_3(\mu)R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma}+c_4(\mu)\Box R+O(M_P^{-2})\Big]\end{aligned}$$



$$\Gamma_{\text{NL}} = - \int d^4x \sqrt{|g|} [\alpha R \ln \left(\frac{\square}{\mu^2} \right) R + \beta R_{\mu\nu} \ln \left(\frac{\square}{\mu^2} \right) R^{\mu\nu} + \gamma R_{\mu\nu\rho\sigma} \ln \left(\frac{\square}{\mu^2} \right) R^{\mu\nu\rho\sigma} + O(M_P^{-2})]$$

$$M_P = \sqrt{\hbar c / (8\pi G_N)} = 2.4 \times 10^{18} \text{GeV}$$

	α	β	γ
Scalar	$5(6\xi - 1)^2$	-2	2
Fermion	-5	8	7
Vector	-50	176	-26
Boson	250	-244	424

$$\begin{aligned} c_1(\mu) &= c_1(\mu_*) - \alpha \ln \left(\frac{\mu^2}{\mu_*^2} \right), \\ c_2(\mu) &= c_2(\mu_*) - \beta \ln \left(\frac{\mu^2}{\mu_*^2} \right), \\ c_3(\mu) &= c_3(\mu_*) - \gamma \ln \left(\frac{\mu^2}{\mu_*^2} \right). \end{aligned}$$

$$\begin{aligned} \Gamma_L &= \int d^4x \sqrt{|g|} \left[\frac{R}{16\pi G_N} + \bar{c}_1(\mu)R^2 + \bar{c}_2(\mu)R_{\mu\nu}R^{\mu\nu} \right] \\ \Gamma_{\text{NL}} &= - \int d^4x \sqrt{|g|} \left[\bar{\alpha}R \ln \left(\frac{\square}{\mu^2} \right) R + \bar{\beta}R_{\mu\nu} \ln \left(\frac{\square}{\mu^2} \right) R^{\mu\nu} \right] \end{aligned}$$

$$E_{\mu\nu} \equiv R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} + 16\pi G_N(H_{\mu\nu}^L + H_{\mu\nu}^{\text{NL}}) = 8\pi G_N T_{\mu\nu},$$

$$\begin{aligned} H_{\mu\nu}^L &= \bar{c}_1(\mu) \left(2R_{\mu\nu}R - \frac{1}{2}g_{\mu\nu}R^2 + 2g_{\mu\nu}\nabla^2R - \nabla_\mu\nabla_\nu R - \nabla_\nu\nabla_\mu R \right) \\ &\quad + \bar{c}_2(\mu) \left(-\frac{1}{2}g_{\mu\nu}R_{\rho\sigma}R^{\rho\sigma} + 2R_\mu^\rho R_{\nu\rho} + \nabla^2R_{\mu\nu} - \nabla_\rho\nabla_\mu R_\nu^\rho - \nabla_\rho\nabla_\nu R_\mu^\rho + g_{\mu\nu}\nabla_\sigma\nabla_\rho R^{\rho\sigma} \right) \end{aligned}$$

$$\begin{aligned} H_{\mu\nu}^{\text{NL}} &= -\bar{\alpha} \left(2R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R + 2g_{\mu\nu}\nabla^2 - \nabla_\mu\nabla_\nu - \nabla_\nu\nabla_\mu \right) \ln \left(\frac{\square}{\mu^2} \right) R \\ &\quad - \bar{\beta} \left(-\frac{1}{2}g_{\mu\nu}R^{\rho\sigma} + 2\delta_\nu^\sigma R_\mu^\rho + \delta_\mu^\rho\delta_\nu^\sigma\nabla^2 - \delta_\nu^\sigma\nabla^\rho\nabla_\mu - \delta_\mu^\sigma\nabla^\rho\nabla_\nu + g_{\mu\nu}\nabla^\sigma\nabla^\rho \right) \ln \left(\frac{\square}{\mu^2} \right) R_{\rho\sigma} \end{aligned}$$



$$ds^2 = h(r)dt^2 - \frac{1}{f(r)}dr^2 - r^2d\theta^2 - r^2\sin^2\theta d\phi^2$$

$$\begin{aligned} S = & \int d^4x \sqrt{|g|} \left(\frac{M_P^2}{2}R + \tilde{c}_1(\mu)R^2 + \tilde{c}_2(\mu)C_{\mu\nu\alpha\beta}C^{\mu\nu\alpha\beta} \right. \\ & \left. + \tilde{\alpha}R\ln\left(\frac{\square}{\mu^2}\right)R + \tilde{\beta}C_{\mu\nu\alpha\beta}\ln\left(\frac{\square}{\mu^2}\right)C^{\mu\nu\alpha\beta} \right) \end{aligned}$$

$$\begin{aligned} h(r) &= k((r - r_0) + h_2(r - r_0)^2 + h_3(r - r_0)^3) + \mathcal{O}((r - r_0)^4) \\ f(r) &= f_1(r - r_0) + f_2(r - r_0)^2 + f_3(r - r_0)^3 + \mathcal{O}((r - r_0)^4) \end{aligned}$$

$$\begin{aligned} \bar{c}_1(\mu) &\rightarrow \bar{c}_1(\mu) + 2\bar{\alpha}\ln\mu, \\ \bar{c}_2(\mu) &\rightarrow \bar{c}_2(\mu) + 2\bar{\beta}\ln\mu, \\ \bar{c}_3(\mu) &\rightarrow \bar{c}_3(\mu) + 2\bar{\gamma}\ln\mu, \end{aligned}$$

$$\begin{aligned} h_2 &= \frac{1 - 2f_1r_0}{f_1r_0^2} + \frac{1 - f_1r_0}{8(\tilde{c}_2(\mu) + 2\tilde{\beta}\ln\mu)f_1^2r_0} \\ f_2 &= \frac{1 - 2f_1r_0}{r_0^2} - \frac{3(1 - f_1r_0)}{8(\tilde{c}_2(\mu) + 2\tilde{\beta}\ln\mu)f_1r_0} \end{aligned}$$

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ANEXO B.

Materia y energía oscuras inherentes a un espacio – tiempo cuántico relativista por interacciones de una partícula oscura. Métrica de Einstein - Friedmann-Lemaître-Robertson-Walker. Modelo Inflacionario.

$$ds^2 = -dt^2 + a^2(t) \left[\frac{dr^2}{1-Kr^2} + r^2(d\theta^2 + \sin^2 \theta d\phi^2) \right]$$

$$G^{\mu\nu} = R^{\mu\nu} - \frac{1}{2}g^{\mu\nu}R = \frac{8\pi G}{c^4}T^{\mu\nu}$$

$$\begin{aligned} R_0^0 &= \frac{3\ddot{a}}{a} \\ R_j^i &= \left(\frac{\ddot{a}}{a} + \frac{2\dot{a}^2}{a^2} + \frac{2K}{a^2} \right) \delta_j^i \\ R &= 6 \left(\frac{\ddot{a}}{a} + \frac{\dot{a}^2}{a^2} + \frac{K}{a^2} \right) \end{aligned}$$

$$T_\xi^\alpha = \text{diag}(\rho, -p, -p, -p)$$

$$\begin{aligned} H^2 &\equiv \left(\frac{\dot{a}}{a} \right)^2 = \frac{8\pi G\rho}{3} - \frac{K}{a^2} \\ \dot{H} &= -4\pi G(p + \rho) + \frac{K}{a^2} \end{aligned}$$

$$\dot{\rho} + 3H(\rho + p) = 0$$

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3}(\rho + 3p).$$

$$\omega = p/\rho$$

$$\begin{aligned} a &\propto (t - t_0)^{\frac{2}{3(1+w)}} \\ \rho &\propto a^{-3(1+w)} \end{aligned}$$

$$\begin{aligned} a &\propto e^{H_0 t} \\ \rho &= \text{Const} \end{aligned}$$

$$\begin{aligned} \text{Radiation: } a &\propto (t - t_0)^{1/2}, \rho \propto a^{-4} \\ \text{Dust: } a &\propto (t - t_0)^{2/3}, \rho \propto a^{-3} \end{aligned}$$

$$S = \int d^4x \sqrt{-g} \left[\frac{1}{16\pi G} \mathcal{R} - \frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - V(\phi) \right]$$

$$\square \phi - V'(\phi) = \frac{1}{\sqrt{-g}} \partial_\mu (g^{\mu\nu} \sqrt{-g} \partial_\nu \phi) - V'(\phi) = 0$$



$$\ddot{\phi}+3H\dot{\phi}+\frac{dV}{d\phi}=0.$$

$$T_{\mu\nu}=-\frac{2}{\sqrt{-g}}\frac{\delta S}{\delta g^{\mu\nu}}=\partial_\mu \phi \partial_\nu \phi -g_{\mu\nu}\Big[\frac{1}{2}g^{\alpha\beta}\partial_\alpha \phi \partial_\beta \phi +V(\phi)\Big]$$

$$\rho=T_0^0=\frac{1}{2}\dot{\phi}^2+V(\phi), p=-T_i^i=\frac{1}{2}\dot{\phi}^2-V(\phi)$$

$$\begin{gathered} H^2=\frac{8\pi G}{3}\Big[\frac{1}{2}\dot{\phi}^2+V(\phi)\Big]\\ \ddot{a}=-\frac{8\pi G}{3}\big[\dot{\phi}^2-V(\phi)\big].\end{gathered}$$

$$\frac{1}{\kappa^2}G_{\mu\nu}=\Lambda(\phi)g_{\mu\nu}-\frac{1}{2}\nabla_\mu\phi\nabla_\nu\phi.$$

$$\begin{gathered} \rho_{\rm sca}=\frac{1}{2}\dot{\phi}^2+\Lambda(\phi)\\ p_{\rm sca}=-\Lambda(\phi)\end{gathered}$$

$$\ddot{\phi}+\frac{3}{2}H\dot{\phi}+\Lambda'(\phi)=0.$$

$$\begin{gathered} \rho_{\rm sca}=\frac{C}{a^3}+\Lambda,\\ p_{\rm sea}=-\Lambda.\end{gathered}$$

$$\frac{1}{2}mv^2+V(x)=E=\text{ Const}\,,$$

$$\mathcal{S}=\int~dt\mathcal{L}=\int~dt\frac{dB}{dt}\Bigl(\frac{1}{2}mv^2+V(x)\Bigr),$$

$$m\dot{v}=-V'(x),$$

$$m\frac{d}{dt}\Bigl(\frac{dB}{dt}\frac{dx}{dt}\Bigr)=\frac{dV(x)}{dx}\frac{dB}{dt}.$$

$$\mathcal{H}=\frac{dB}{dt}p_a+\frac{dx}{dt}p_x-\mathcal{L}=\frac{dx}{dt}p_x=mv^2\frac{dB}{dt}=\text{ Const}\,,$$

$$\frac{dB}{dt}=C/v^2.$$

$$m\frac{d}{dt}\Bigl(\frac{C}{v}\Bigr)=V'(x)\frac{C}{v^2},$$

$$\mathcal{S}_{(\chi)}=\int~d^4x\sqrt{-g}\chi_{\mu;\nu}T_{(\chi)}^{\mu\nu}$$

$$\Gamma^\rho_{\mu\nu}=\left\{{}^\rho_{\mu\nu}\right\}=\frac{1}{2}g^{\rho\lambda}(g_{\lambda\mu,\nu}+g_{\lambda\nu,\mu}-g_{\mu\nu,\lambda}).$$



$$\pi_{\chi_0}=\frac{\partial \mathcal{L}}{\partial \dot{\chi}^0}=T_0^0(\chi)$$

$$\nabla_\mu T_{(\chi)}^{\mu\nu}=0$$

$$G^{\mu\nu} = \frac{2}{\sqrt{-g}} \frac{\delta \sqrt{-g}}{\delta g^{\mu\nu}} \big[\mathcal{L}_\chi + \mathcal{L}_m \big], \nabla_\mu G^{\mu\nu} = 0$$

$$\nabla_\mu T^{\mu\nu}=3\sigma j^\nu$$

$$S_{(\chi,A)}=\int~d^4x\sqrt{-g}\chi_{\mu;\nu}T_{(\chi)}^{\mu\nu}+\frac{\sigma}{2}\int~d^4x\sqrt{-g}\big(\chi_\mu+\partial_\mu A\big)^2$$

$$\nabla_\nu T_{(\chi)}^{\mu\nu}=\sigma(\chi^\mu+\partial^\mu A)=f^\mu$$

$$\nabla_\mu f^\mu=\nabla_\mu (\chi^\mu+\partial^\mu A)=0$$

$$f_\mu=\sigma\big(\chi_\mu+\partial_\mu A\big)=0\,\Rightarrow\,\chi_\mu=-\partial_\mu A$$

$$\mathcal{S}=-\int~d^4x\sqrt{-g}T_{(\chi)}^{\mu\nu}\nabla_\mu\nabla_\nu A$$

$$T^{\mu\nu}=-\frac{1}{2}\phi^\mu\phi^\nu+U(\phi)g^{\mu\nu},$$

$$-\frac{1}{2}\nabla_\mu(\phi^\mu\phi^\nu)=0\Rightarrow\dot{\phi}^2\sim\frac{1}{a^3}.$$

$$\mathcal{S}_{(\chi)}=\int~d^4x\sqrt{-g}\chi_{\mu;\nu}T_{(\chi)}^{\mu\nu}$$

$$\pi_{\chi_0}=\frac{\partial \mathcal{L}}{\partial \dot{\chi}^0}=T_0^0(\chi).$$

$$\nabla_\mu T_{(\chi)}^{\mu\nu}=0.$$

$$G^{\mu\nu} = \frac{2}{\sqrt{-g}} \frac{\delta \sqrt{-g}}{\delta g^{\mu\nu}} \big[\mathcal{L}_\chi + \mathcal{L}_m \big], \nabla_\mu G^{\mu\nu} = 0$$

$$\chi_\mu\rightarrow\chi_\mu+k_\mu,T_{(\chi)}^{\mu\nu}\rightarrow T_{(\chi)}^{\mu\nu}+\Lambda g^{\mu\nu},$$

$$\sqrt{-g}\chi^\mu_{;\mu}\mathcal{L}_1=\partial_\mu\big(\sqrt{-g}\chi^\mu\big)\mathcal{L}_1=\Phi\mathcal{L}_1,$$

$$\partial_\alpha \mathcal{L}_1=0\Rightarrow \mathcal{L}_1=M=\text{ const.}$$

$$\Phi=\frac{1}{4!}\varepsilon^{\alpha\beta\gamma\delta}\varepsilon_{abcd}\partial_\alpha\varphi^{(a)}\partial_\beta\varphi^{(b)}\partial_\gamma\varphi^{(c)}\partial_\delta\varphi^{(d)},$$

$$S=\int~d^4x\Phi\mathcal{L}_1+\int~d^4x\sqrt{-g}\mathcal{L}_2$$



$$\mathcal{L} = -\frac{1}{2}R + \chi_{\mu;\nu}T^{\mu\nu}_{(\chi)}-\frac{1}{2}g^{\alpha\beta}\phi_{,\alpha}\phi_{,\beta}-V(\phi).$$

$$T^{\mu\nu}_{(\chi)}=-\frac{1}{2}\phi^\mu\phi^\nu+U(\phi)g^{\mu\nu}.$$

$$ds^2=-dt^2+a(t)^2\left(\frac{dr^2}{1-Kr^2}+r^2d\Omega^2\right)$$

$$\ddot{\phi}+\frac{3}{2}\mathcal{H}\dot{\phi}+U'(\phi)=0.$$

$$\begin{gathered}\chi^\lambda_{;\lambda}U'(\phi)-V'(\phi)=\nabla_\mu j^\mu,\\ j^\mu=\frac{1}{2}\phi_{,\nu}(\chi^{\mu;\nu}+\chi^{\nu;\mu})+\phi^{,\mu},\end{gathered}$$

$$\begin{aligned}\ddot{\phi}(\dot{\chi}_0-1)+\dot{\phi}[\ddot{\chi}_0+3\mathcal{H}(\dot{\chi}_0-1)]\\=U'(\phi)(\dot{\chi}_0+3\mathcal{H}\chi_0)-V'(\phi).\end{aligned}$$

$$\begin{aligned}[1-2\dot{\chi}_0-3\mathcal{H}\chi_0]\ddot{\phi}-\Big[\ddot{\chi}_0-3\mathcal{H}+\frac{9}{2}\mathcal{H}(\dot{\chi}_0+\chi_0\mathcal{H})\Big]\dot{\phi}\\+V'(\phi)=0.\end{aligned}$$

$$\begin{aligned}G^{\mu\nu}= & g^{\mu\nu}\left(\frac{1}{2}\phi_{,\alpha}\phi^{,\alpha}+V(\phi)+\frac{1}{2}\chi^{\alpha;\beta}\phi_{,\alpha}\phi_{,\beta}+\chi^\lambda\phi_{,\lambda}U'(\phi)\right)\\&-\frac{1}{2}\phi^\mu\left((\chi^\lambda_{;\lambda}+2)\phi^\nu+\chi^{\lambda;\nu}\phi_{,\lambda}+\chi^\lambda\phi^\nu_{;\lambda}\right)\\&-\frac{1}{2}\left(\chi^\lambda\phi^\mu_{;\lambda}\phi^\nu+\chi^{\lambda;\mu}\phi_{,\lambda}\phi^\nu\right)\end{aligned}$$

$$\rho=\dot{\phi}^2\left(\dot{\chi}_0\left(1-\frac{3}{2}\mathcal{H}\right)-\frac{1}{2}\right)+V(\phi)-\dot{\phi}\dot{\chi}_0\big(U'(\phi)+\ddot{\phi}\big)$$

$$p=\frac{1}{2}\dot{\phi}^2(\dot{\chi}_0-1)-V(\phi)-\chi_0\dot{\phi}U'(\phi)$$

$$\rho=\left(\dot{\chi}_0-\frac{1}{2}\right)\dot{\phi}^2+V(\phi),$$

$$U(\phi)=C,V(\phi)=\Omega_\Lambda.$$

$$\dot{\phi}^2=\frac{2\Omega_m}{a^3},$$

$$\dot{\chi}_0=1-\kappa a^{-1.5},$$

$$\begin{aligned}\rho &= \Omega_\Lambda-\frac{\Omega_\kappa}{a^{4.5}}+\frac{\Omega_m}{a^3},\\ p &= -\Omega_\Lambda-\frac{1}{2}\frac{\Omega_\kappa}{a^{4.5}},\end{aligned}$$

$$U(\phi)=C,V(\phi)=\Omega_\Lambda e^{-\beta\phi},$$

$$\ddot{\phi} = -\frac{3}{2}\mathcal{H}\dot{\phi}$$

$$\delta=\dot{\chi}_0-1,$$

$$\dot{\phi}\left(\dot{\delta}+\frac{3}{2}\mathcal{H}\delta\right)=\beta V(\phi)$$

$$x = \frac{\dot{\phi}}{\sqrt{6}H}, y = \frac{\sqrt{V(\phi)}}{\sqrt{3}H}$$

$$(1+2\delta)x^2+y^2=1$$

$$\begin{aligned}\frac{dx}{d\tau} &= -\frac{3x}{4}(x^2 - 1 + 3y^2) \\ \frac{dy}{d\tau} &= -\frac{y}{4}(-9 + 3x^2 + 9y^2 + 2\sqrt{6}x\beta)\end{aligned}$$

$$\omega_\chi = \frac{1}{2}(1 - x^2 - 3y^2)$$

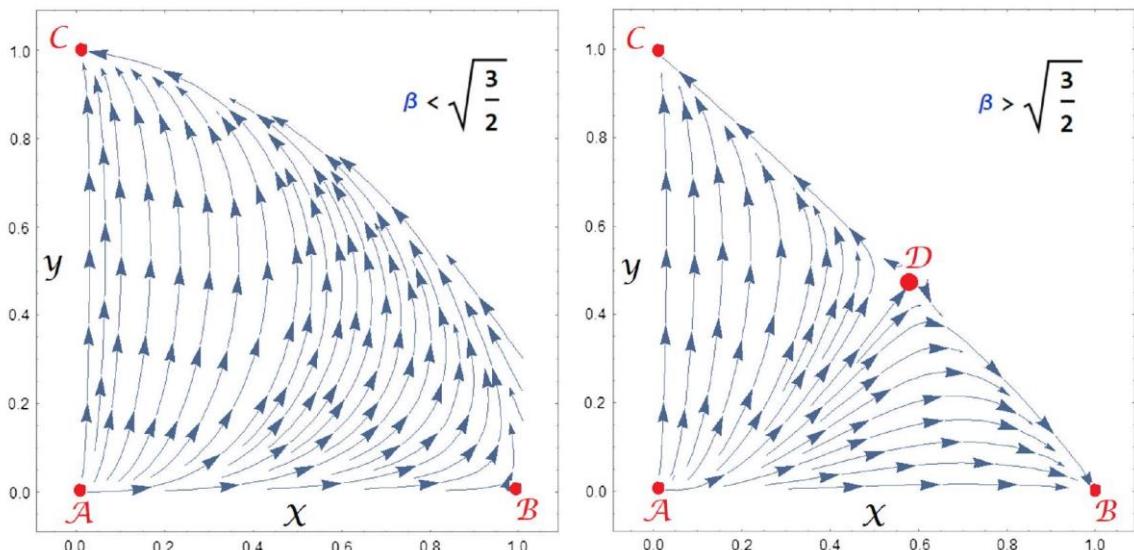


Figura 1. Deformaciones del espacio – tiempo cuántico por interacción de la partícula oscura.

$$\frac{d}{dt} = -\mathcal{H}(z)(z+1) \frac{d}{dz}$$

$$\left(x = \sqrt{\frac{3}{2}}\frac{1}{\beta}, y = \frac{\sqrt{2\beta^2 - 3}}{\sqrt{6}\beta}\right)$$

$$\mathcal{S} = \int \sqrt{-g} \left[\frac{1}{2\kappa^2} R + \chi_{\mu;\nu} T_{(\phi)}^{\mu\nu} - \frac{1}{2} g^{\alpha\beta} \phi_{,\alpha} \phi_{,\beta} - V(\phi) \right] d^4x$$

$$T_{(\phi)}^{\mu\nu} = -\frac{1}{2} \phi^{\mu}_{\cdot\alpha} \phi^{\nu}_{\cdot\beta} + U(\phi) g^{\mu\nu}.$$



$$\nabla_\mu T_{(\phi)}^{\mu\nu}=0$$

$$\begin{aligned}\chi_{;\lambda}^\lambda U'(\phi) - V'(\phi) &= \nabla_\mu j^\mu \\ j^\mu &= \frac{1}{2} \phi_{,\nu} (\chi^{\mu;\nu} + \chi^{\nu;\mu}) + \phi^{,\mu}\end{aligned}$$

$$\begin{aligned}\frac{1}{\kappa^2} G^{\mu\nu} &= g^{\mu\nu} \left(\frac{1}{2} \phi_{,\alpha} \phi^{,\alpha} + V(\phi) + \frac{1}{2} \chi^{\alpha;\beta} \phi_{,\alpha} \phi_{,\beta} + \chi^\lambda \phi_{,\lambda} U'(\phi) \right) \\ - \frac{1}{2} \phi^{,\mu} [(\chi_{;\lambda}^\lambda + 2) \phi^{,\nu} + \chi^{\lambda;\nu} \phi_{,\lambda} + \chi^\lambda \phi_{;\lambda}^\nu] &- \frac{1}{2} (\chi^\lambda \phi_{;\lambda}^\mu \phi^{,\nu} + \chi^{\lambda;\mu} \phi_{,\lambda} \phi^{,\nu}).\end{aligned}$$

$$\mathrm{d}s^2 = \mathrm{d}t^2 - a^2(t) \left[\frac{\mathrm{d}r^2}{1 - Kr^2} + r^2 (\mathrm{d}\theta^2 + \sin^2 \theta \mathrm{d}\phi^2) \right]$$

$$\ddot{\phi} + \frac{3}{2} H \dot{\phi} + U'(\phi) = 0,$$

$$j^\mu = (\dot{\phi}(1-\dot{\chi}_0), 0, 0, 0)$$

$$j_{;\mu}^\mu = \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} j^\mu) = \frac{1}{a^3} \partial_\mu (a^3 j^0) = -\ddot{\phi} (\dot{\chi}_0 - 1) + \dot{\phi} [3H(\dot{\chi}_0 - 1) - \ddot{\chi}_0]$$

$$\begin{aligned}\chi_{;\lambda}^\lambda U'(\phi) - V'(\phi) &= U'(\phi) \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} \chi^\mu) - V'(\phi) \\ &= -U'(\phi) \frac{1}{a^3} \partial_\mu (a^3 \chi^0) - V'(\phi) = U'(\phi) [\dot{\chi}_0 + 3H\chi_0] + V'(\phi).\end{aligned}$$

$$\ddot{\phi} (\dot{\chi}_0 - 1) + \dot{\phi} [3H(\dot{\chi}_0 - 1) + \ddot{\chi}_0] = U'(\phi) [\dot{\chi}_0 + 3H\chi_0] - V'(\phi).$$

$$\begin{aligned}\rho &= \dot{\phi}^2 \left(\dot{\chi}_0 \left(1 - \frac{3}{2} \mathcal{H} \right) - \frac{1}{2} \right) + V(\phi) - \dot{\phi}_0 (U'(\phi) + \ddot{\phi}) \\ p &= \frac{1}{2} \dot{\phi}^2 (\dot{\chi}_0 - 1) - V(\phi) - \chi_0 \dot{\phi} U'(\phi)\end{aligned}$$

$$\begin{aligned}\rho &= \left(\dot{\chi}_0 - \frac{1}{2} \right) \dot{\phi}^2 + V(\phi) \\ p &= \frac{1}{2} \dot{\phi}^2 (\dot{\chi}_0 - 1) - V(\phi) - \chi_0 \dot{\phi} U'(\phi)\end{aligned}$$

$$U(\phi) = \text{Const.}$$

$$\dot{\phi}^2 = \frac{2C_{m0}}{a^3}$$

$$\delta = \dot{\chi}_0 - 1$$

$$\dot{\phi} \left(\dot{\delta} + \frac{3}{2} H \delta \right) = -V'(\phi)$$



$$\begin{aligned}\rho &= \left(\delta + \frac{1}{2}\right)\dot{\phi}^2 + V(\phi) \\ p &= \frac{\delta}{2}\dot{\phi}^2 - V(\phi)\end{aligned}$$

$$\delta = \frac{1}{2}\xi a^{-3/2}$$

$$\begin{aligned}\rho &= \Lambda + \frac{\xi C_{m0}}{a^{9/2}} + \frac{C_{m0}}{a^3} \\ p &= -\Lambda + \frac{\xi C_{m0}}{2a^{9/2}}\end{aligned}$$

$$H(z)=H_0\big[\Omega_\Lambda+\Omega_{m0}(1+z)^3+\Omega_{\xi0}(1+z)^{9/2}\big]^{1/2}$$

$$\delta=\xi=0$$

$$\begin{aligned}\rho &= \Lambda + \frac{C_{m0}}{a^3} \\ p &= -\Lambda\end{aligned}$$

$$\begin{aligned}\frac{d\delta}{dz} &= \frac{V'(\phi)}{(z+1)^{5/2}\sqrt{C_{m0}}H(\phi,\delta)} + \frac{3\delta}{2(z+1)} \\ \frac{d\phi}{dz} &= -\frac{\sqrt{2C_{m0}(z+1)}}{H(\phi,\delta)}\end{aligned}$$

$$H(\phi,\delta)=H_0[(2\delta+1)\Omega_{m0}(z+1)^3+\Omega_{DE}(\phi)]^{1/2}$$

$$x=\frac{\phi}{\sqrt{6}H}, y=\frac{\sqrt{V(\phi)}}{\sqrt{3}H}, z=-\frac{V'(\phi)}{V(\phi)}.$$

$$\begin{aligned}x' &= -\frac{3}{4}x(x^2+3y^2-1) \\ y' &= -\frac{1}{4}y(3x^2+9y^2-9+2\sqrt{6}xz) \\ z' &= -\sqrt{6}z^2x(\Gamma-1)\end{aligned}$$

$$\Gamma=\frac{V(\phi)V''(\phi)}{V'(\phi)^2}$$

$$(1+2\delta)x^2+y^2=1$$

$$\beta>\sqrt{\frac{3}{2}}\left(\beta>\sqrt{\frac{3}{2}}\right)\left(\sqrt{\frac{3}{2}}\frac{1}{\beta},\sqrt{\frac{2\beta^2-3}{\sqrt{6}\beta}}\right)$$

$$z'=\sqrt{6}x(1+z^2)$$

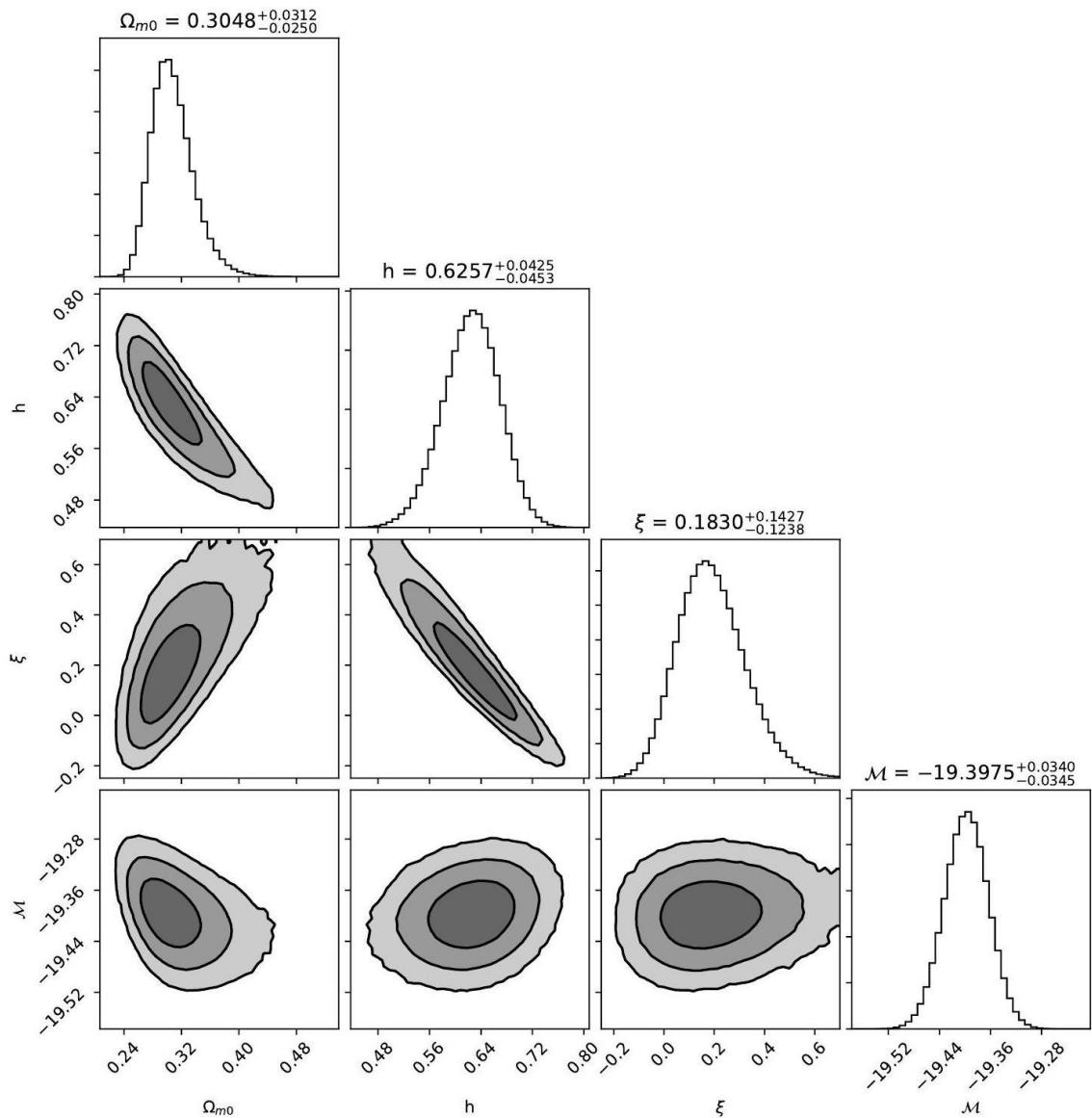
$$z'=\sqrt{6}x(z-\alpha)^2$$

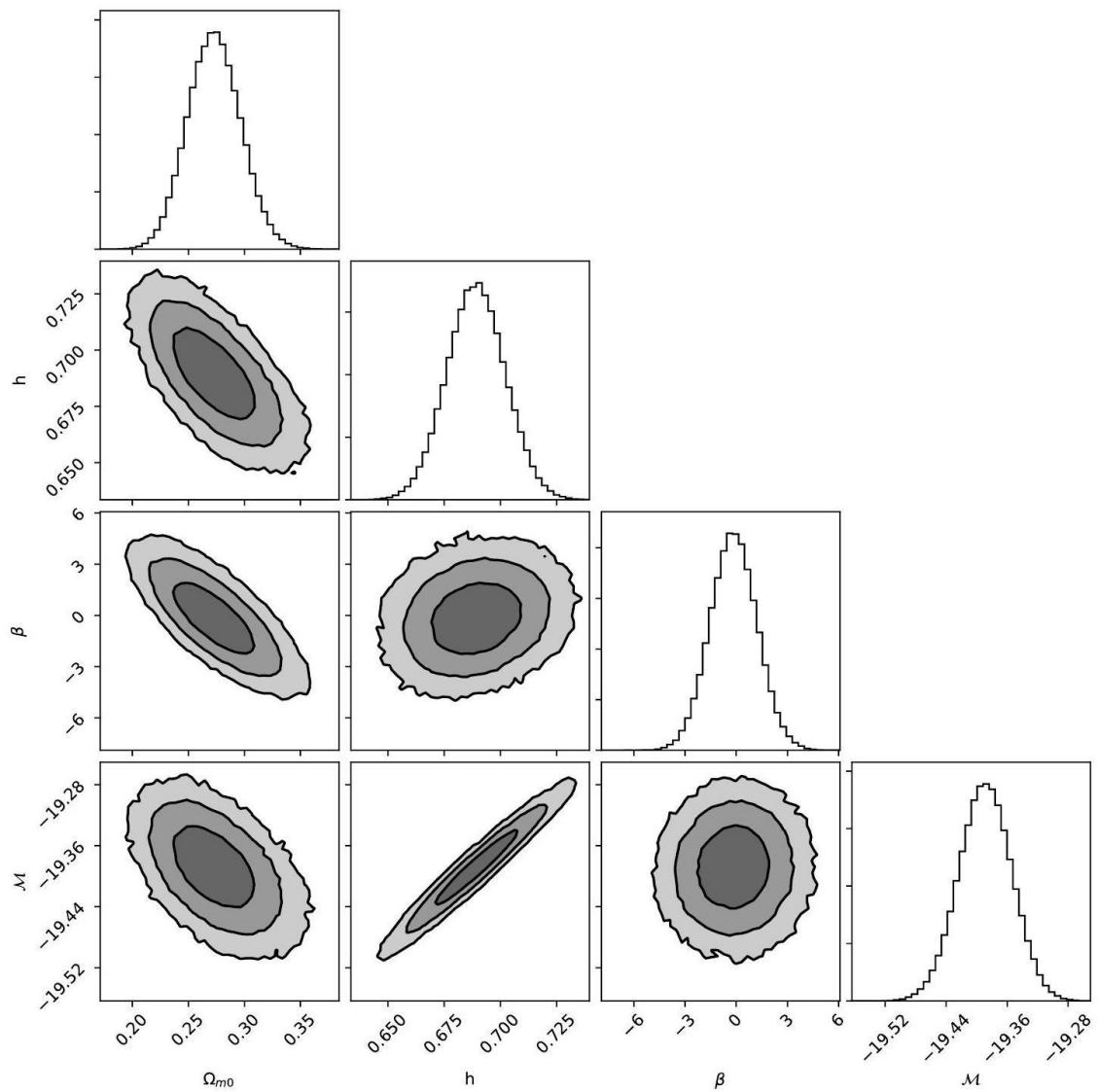
$$\left(\sqrt{\frac{3}{2\alpha}},\sqrt{\frac{2\alpha-3}{6\alpha}}\right)\alpha<\sqrt{\frac{5}{6}}\alpha>\sqrt{\frac{5}{6}}$$

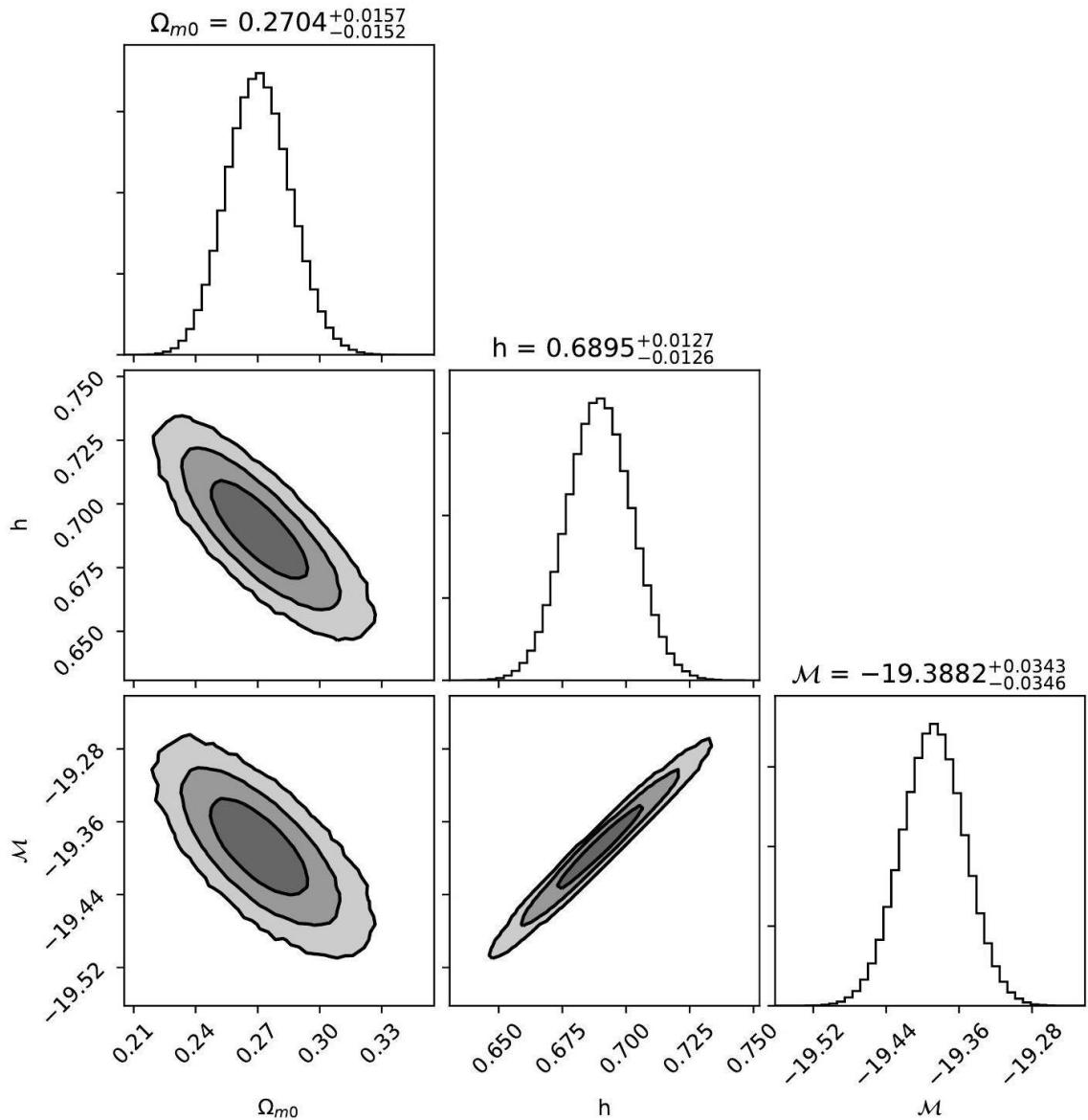


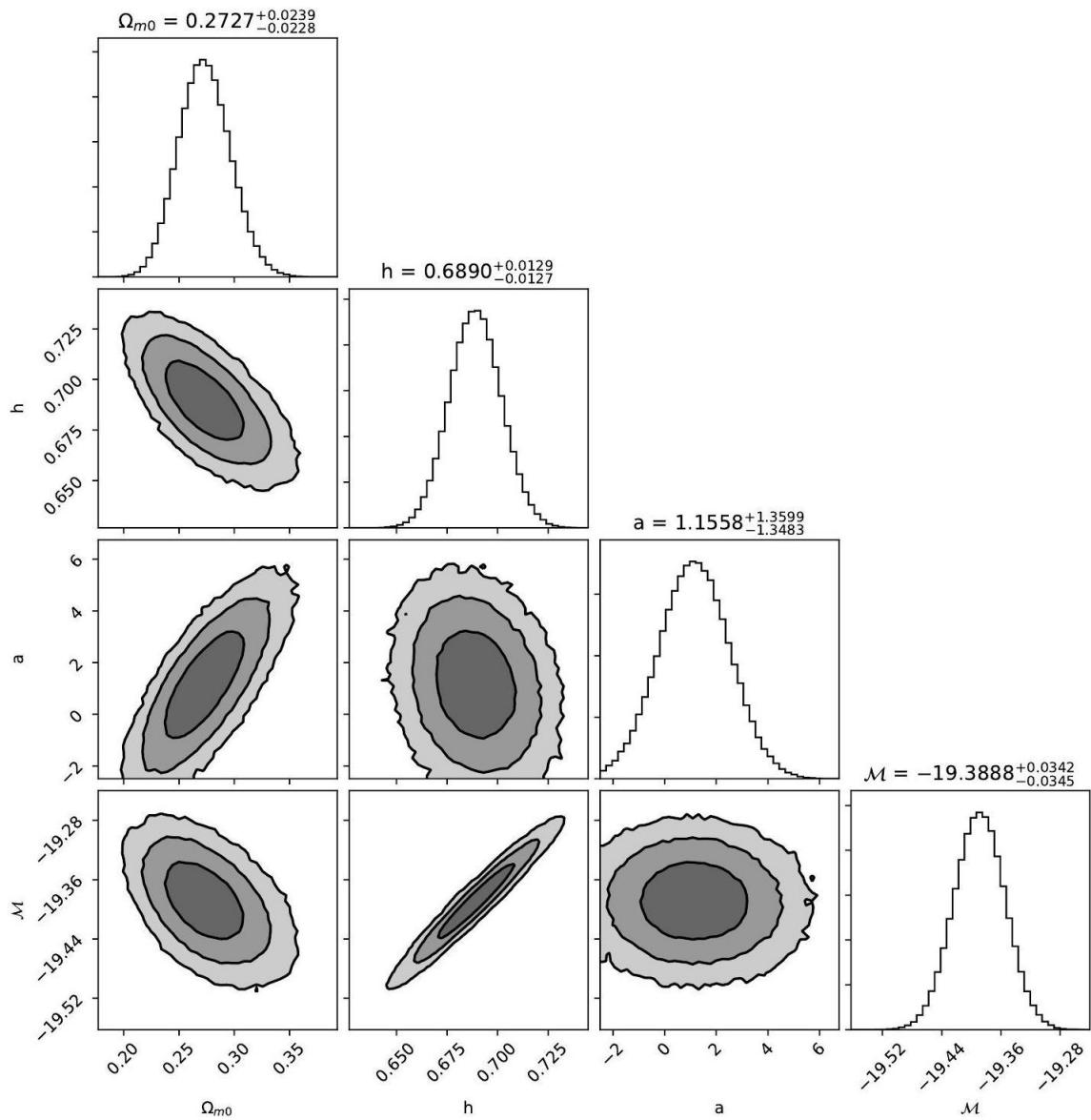
$$\chi^2_H(\phi^v) = \mathcal{H} \mathbf{C}_{H,\text{cov}}^{-1} \mathcal{H}^T$$

$$\chi^2_s(\phi_s^v) = \mu_s \mathbf{C}_{s,\text{cov}}^{-1} \mu_s^T,$$









Figuras 2, 3, 4 y 5. Fluctuaciones de materia y energía de la partícula oscura.

$$\mu_{\text{th}} = 5 \log \left(\frac{D_L(z)}{\text{Mpc}} \right) + 25$$

$$D_L(z) = c(1+z) \int_0^z \frac{dx}{H(x, \theta^\nu)}$$

$$\mathcal{L}_{\text{tot}}(\phi^\psi) = \prod_{p=1}^P \exp(-\chi_p^2)$$

$$\chi_{\text{tot}}^2 = \sum_{p=1}^P \chi_p^2,$$



Model	Ω_{m0}	h	α or ξ	β	\mathcal{M}	χ^2_{\min}
$V, U \text{const.}$	$0.305^{+0.031}_{-0.025}$	$0.6257^{+0.0428}_{-0.0455}$	$0.183^{+0.143}_{-0.125}$	—	$-19.397^{+0.034}_{-0.035}$	84.114
V_1	$0.277^{+0.024}_{-0.023}$	$0.6885^{+0.0130}_{-0.0128}$	—	$-0.593^{+1.367}_{-1.355}$	$-19.390^{+0.034}_{-0.035}$	88.100
$V_2(\text{cosine})$	0.270 ± 0.015	$0.6895^{+0.0128}_{-0.0127}$	— —	1	-19.388 ± 0.035	87.954
V_3	$0.273^{+0.024}_{-0.023}$	$0.6890^{+0.0130}_{-0.0127}$	$1.152^{+1.370}_{-1.352}$	1	-19.389 ± 0.034	87.942
ΛCDM	$0.281^{+0.016}_{-0.015}$	0.686 ± 0.013	—	—	-19.403 ± 0.035	85.700

$$\text{AIC} = -2\ln (\mathcal{L}_{\max}) + 2\psi + \frac{2\psi(\psi+1)}{N_{\text{tot}} - \psi - 1}$$

$$\text{BIC} = -2\ln (\mathcal{L}_{\max}) + \psi \log (N_{\text{tot}}).$$

$$\text{DIC} = D\left(\overline{\phi^\psi}\right) + 2C_B$$

$$100\% \times \frac{\left(H_{\Lambda CDM}(z_{\text{BBN}}, \Omega_{m0}, h) - H_i(z_{\text{BBN}}, \theta_i^v)\right)^2}{H_{\Lambda CDM}(z_{\text{BBN}}, \Omega_{m0}, h)^2} < 10\%$$

$$ds^2=-dt^2+R(t)^2\frac{\Sigma dx^jdx^j}{f_D(x)^2}+R(t)^2\frac{\Sigma dx^pdx^p}{f_d(x)^2}$$

$$f_D=1+\frac{k_D}{4}\Sigma(x^i)^2, f_d=1+\frac{k_d}{4}\Sigma(y^p)^2$$

$$\mathcal{H}_R=\frac{\dot{R}}{R}\,\mathcal{H}_r=\frac{\dot{r}}{r}.$$

$$V=R^Dr^d,$$

$$\mathcal{H}=\frac{\dot{V}}{V}.$$

$$\mathcal{H}=D\mathcal{H}_R+d\mathcal{H}_d.$$

$$T_v^\mu=\text{diag}(\rho,-p,-p,\dots,-p',-p',\dots).$$

$$\begin{aligned} &\frac{1}{2}D(D-1)\left[\frac{\dot{R}^2}{R^2}+\frac{k_D}{R^2}\right]+\frac{1}{2}d(d-1)\left[\frac{\dot{R}^2}{R^2}+\frac{k_d}{R^2}\right] \\ &+Dd\frac{\dot{R}}{R}\frac{\dot{r}}{r}=8\pi\rho \\ &(D-1)\frac{\ddot{R}}{R}+d\frac{\ddot{r}}{r}-d\frac{\dot{R}}{R}\frac{\dot{r}}{r}-(D-1)\left[\frac{\dot{R}^2}{R^2}+\frac{k_D}{R^2}\right] \\ &=-8\pi(\rho+p) \\ &D\frac{\ddot{R}}{R}+(d-1)\frac{\ddot{r}}{r}-D\frac{\dot{R}}{R}\frac{\dot{r}}{r}-(d-1)\left[\frac{\dot{r}^2}{r^2}+\frac{k_d}{r^2}\right] \\ &=-8\pi(\rho+p'). \end{aligned}$$

$$D\frac{\ddot{R}}{R}+d\frac{\ddot{r}}{r}=\frac{8\pi\rho}{D+d-1}[1-(D+d)\gamma].$$



$$\frac{\ddot{V}}{V}=\frac{D+d}{D+d-1}8\pi(\rho-p).$$

$$E = \frac{1}{2}\dot{V}^2 - \frac{D+d}{D+d-1}\Omega V^2,$$

$$\begin{aligned}\frac{\dot{R}}{R}=&\frac{1}{(D+d)V}\Bigg[\dot{V}+\sqrt{\frac{2Ed}{D}(D+d-1)}\Bigg]\\\frac{\dot{r}}{r}=&\frac{1}{(D+d)V}\Bigg[\dot{V}-\sqrt{\frac{2ED}{d}(D+d-1)}\Bigg].\end{aligned}$$

$$R(t)=V^{\frac{1}{D+d}}\text{exp}\left[+\frac{1}{D+d}\sqrt{\frac{2Ed(D+d-1)}{D}}\int\,\frac{dt}{V}\right]$$

$$r(t)=V^{\frac{1}{D+d}}\text{exp}\left[-\frac{1}{D+d}\sqrt{\frac{2ED(D+d-1)}{d}}\int\,\frac{dt}{V}\right].$$

$$\begin{aligned}p=&\frac{1}{D+d}\Bigg(1+\sqrt{\frac{d(D+d+1)}{D}}\Bigg)\\q=&\frac{1}{D+d}\Bigg(1-\sqrt{\frac{D(D+d+1)}{d}}\Bigg),\end{aligned}$$

$$Dp + dq = Dp^2 + dq^2 = 1.$$

$$\mathcal{L}=-\frac{1}{2}R+\chi_{\mu;\nu}T_{(\phi)}^{\mu\nu}-\frac{1}{2}g^{\alpha\beta}\phi_{,\alpha}\phi_{,\beta}-V(\phi),$$

$$\nabla_\mu T_{(\phi)}^{\mu\nu}=0.$$

$$T_{(\phi)}^{\mu\nu}=-\frac{1}{2}\phi^\mu\phi^\nu+U(\phi)g^{\mu\nu},$$

$$\ddot{\phi}+\frac{1}{2}\mathcal{H}\dot{\phi}+U'(\phi)=0.$$

$$\begin{aligned}\chi_{;\lambda}^\lambda U'(\phi)-V'(\phi)&=\nabla_\mu j^\mu\\j^\mu&=\frac{1}{2}\phi_{,\nu}(\chi^{\mu;\nu}+\chi^{\nu;\mu})+\phi^\mu,\end{aligned}$$

$$\begin{aligned}\ddot{\phi}(\dot{\chi}_0-1)+\dot{\phi}(\mathcal{H}(\dot{\chi}_0-1)+\ddot{\chi}_0)\\=U'(\phi)(\dot{\chi}_0+\mathcal{H}\chi_0)-V'(\phi).\end{aligned}$$



$$\begin{aligned} G^{\mu\nu}=&g^{\mu\nu}\left(\frac{1}{2}\phi_{,\alpha}\phi^{,\alpha}+V(\phi)+\frac{1}{2}\chi^{\alpha;\beta}\phi_{,\alpha}\phi_{,\beta}+\chi^\lambda\phi_{,\lambda}U'(\phi)\right)\\ &-\frac{1}{2}\phi^\mu\left((\chi^\lambda_{;\lambda}+2)\phi^\nu+\chi^{\lambda;\nu}\phi_{,\lambda}+\chi^\lambda\phi^\nu_{;\lambda}\right)\\ &-\frac{1}{2}(\chi^\lambda\phi^\mu_{;\lambda}\phi^\nu+\chi^{\lambda;\mu}\phi_{,\lambda}\phi^\nu) \end{aligned}$$

$$\rho=\Bigl(\dot\chi_0-\frac{1}{2}\Bigr)\dot\phi^2+V(\phi)$$

$$U(\phi)=C, V(\phi)=\Omega_\Lambda.$$

$$\dot{\phi}^2 = \frac{2\Omega_m}{V},$$

$$\dot{\chi}_0=1+\frac{\kappa}{V^2},$$

$$\Omega=\Omega_\Lambda+\frac{\Omega_m}{V}+\frac{\Omega_\kappa}{V^{3/2}},$$

$$E=\frac{1}{2}\dot{V}^2+U_{\rm eff}(V)$$

$$U_{\rm eff}(V)=-\frac{D+d}{D+d-1}\big(\Omega_\Lambda V^2+\Omega_m V+\Omega_\kappa \sqrt{V}\big)$$

$$R(t)=V_c^{\frac{1}{D+d}}\text{exp}\left[+\frac{1}{D+d}\sqrt{\frac{2Ed(D+d-1)}{D}}\frac{t}{V_c}\right]$$

$$r(t)=V_c^{\frac{1}{D+d}}\text{exp}\left[-\frac{1}{D+d}\sqrt{\frac{2ED(D+d-1)}{d}}\frac{t}{V_c}\right]$$

$$V(\phi)=\frac{\Lambda_{+\infty}-\Lambda_{-\infty}}{2}\tanh{(\beta\phi)}+\frac{\Lambda_{+\infty}+\Lambda_{-\infty}}{2},$$

$$V(\phi)=\frac{\Lambda_{+\infty}-\Lambda_{-\infty}}{2}\text{Sign}(\phi)+\frac{\Lambda_{+\infty}+\Lambda_{-\infty}}{2}.$$

$$\xi=\dot{\chi}_0-1,$$

$$\phi\left(\dot\xi+\frac{1}{2}\mathcal H\xi\right)=-V'(\phi).$$

$$\frac{2\Omega_m}{V^{\frac{3}{2}}}d\left(\xi V^{\frac{1}{2}}\right)=-dV(\phi)$$



$$\frac{2\Omega_m}{V}\Big(\frac{d}{d\phi}\xi+\frac{1}{2V}\frac{d}{d\phi}V\Big)=-(\Lambda_{+\infty}-\Lambda_{-\infty})\delta(\phi)$$

$$\xi(\phi<0)=\frac{\kappa_-}{V_{(\phi=0)}^{\frac{3}{2}}}, \xi(\phi>0)=\frac{\kappa_+}{V_{(\phi=0)}^{\frac{3}{2}}}.$$

$$\Delta\xi=-\frac{V_{(\phi=0)}}{2\Omega_m}(\Lambda_{+\infty}-\Lambda_{-\infty}).$$

$$\kappa_+-\kappa_-=-\frac{V_{(\phi=0)}^{\frac{3}{2}}}{2\Omega_m}(\Lambda_{+\infty}-\Lambda_{-\infty}).$$

$$E=\frac{1}{2}\dot{V}^2-\frac{D+d}{D+d-1}V^2\Big[\frac{2\Omega_m}{V}\Big(\xi+\frac{1}{2}\Big)\\+\frac{\Lambda_{+\infty}-\Lambda_{-\infty}}{2}\text{Sign}(\phi)+\frac{\Lambda_{+\infty}+\Lambda_{-\infty}}{2}\Big].$$

$$\frac{1}{2}\Delta\dot{V}^2=\frac{D+d}{D+d-1}\bigg(\frac{2\Omega_m}{V_{(\phi=0)}}\Delta\xi+\Lambda_{+\infty}-\Lambda_{-\infty}\bigg)V_{(\phi=0)}^2=0,$$

$$V(t)=V_0\mathrm{exp}\left(\chi t\right) ,$$

$$\chi^2 = 2\frac{D+d}{D+d-1}\Omega$$

$$R(t)\sim e^{\frac{\chi t}{D+d}}\mathrm{exp}\left[-\sqrt{\frac{2Ed(D+d-1)}{D}}\frac{V_0^{\frac{1}{D+d}}}{\chi}e^{-\chi t}\right]\\r(t)\sim e^{\frac{\chi t}{D+d}}\mathrm{exp}\left[+\sqrt{\frac{2ED(D+d-1)}{d}}\frac{V_0^{\frac{1}{D+d}}}{\chi}e^{-\chi t}\right].$$

$$R(t)\rightarrow R_0\mathrm{exp}\left(\frac{\chi}{D+d}t\right)\\r(t)\rightarrow r_0\mathrm{exp}\left(\frac{\chi}{D+d}t\right)$$

$$\dot{\rho}+(p+\rho)D\frac{\dot{R}}{R}+(p'+\rho)\frac{\dot{r}}{r}=0.$$



$$\begin{gathered} R_{00}=-\left(D \frac{\ddot{R}}{R}+d \frac{\dot{r}}{r}\right) \\ R_{DD}=\dot{H}_D+(D H_D+d H_d) H_D+(D-1) \frac{k_D}{R^2} \\ R_{dd}=\dot{H}_d+(D H_D+d H_d) H_d+(d-1) \frac{k_d}{r^2} . \end{gathered}$$

$$\begin{aligned} R= & 2 D \frac{\ddot{R}}{R}+2 d \frac{\dot{r}}{r}+2 D d H_D H_d \\ & +D(D-1)\left(H_D^2+\frac{k_D}{R^2}\right)+d(d-1)\left(H_d^2+\frac{k_d}{R^2}\right) . \end{aligned}$$

$$\nabla_\mu T^{\mu\nu}=3\sigma j^\nu$$

$$-\nabla_\mu T^{\mu\nu}_{(\Lambda)} = \nabla_\mu T^{\mu\nu}_{(\mathrm{Dust})} = J^\nu, \nabla_\mu J^\mu = 0$$

$$S=\int\,\,\mathrm{d}^4x\Phi\mathcal{L}_1+\int\,\,\mathrm{d}^4x\sqrt{-g}\mathcal{L}_2$$

$$A_a^\alpha \partial_\alpha {\mathcal L}_1 = 0$$

$$\begin{aligned} S=\mathcal{S}_{(\chi)}+\mathcal{S}_{(R)}=& \int\,\,\mathrm{d}^4x\sqrt{-g}\chi_{\mu;\nu}T^{\mu\nu}_{(\chi)} \\ & +\frac{1}{16\pi G} \int\,\,\mathrm{d}^4x\sqrt{-g}R \end{aligned}$$

$$S_{(\chi)}=\int\,\,\mathrm{d}^4x\sqrt{-g}\chi,\mu;\nu T^{\mu\nu}_{(\chi)}$$

$$\nabla_\mu T^{\mu\nu}_{(\chi)}=f^\nu; \nabla_\nu f^\nu=0$$

$$T^{\mu\nu}_{(G)}=\frac{-2}{\sqrt{-g}}\frac{\delta(\sqrt{-g}\mathcal{L}_M)}{\delta g^{\mu\nu}};\,\nabla_\mu T^{\mu\nu}_{(G)}=0$$

$$\chi\rightarrow\chi+C_\chi; T^{\mu\nu}_{(\chi)}\rightarrow T^{\mu\nu}_{(\chi)}+g^{\mu\nu}C_T$$

$$\begin{aligned} S_{(\chi,A)}=& \int\,\,\mathrm{d}^4x\sqrt{-g}\chi_{\mu;\nu}T^{\mu\nu}_{(\chi)}\mathrm{d}^4x \\ & +\frac{\sigma}{2} \int\,\,\mathrm{d}^4x\sqrt{-g}\big(\chi_\mu+\partial_\mu A\big)^2\,\,\mathrm{d}^4x \end{aligned}$$

$$\nabla_\mu T^{\mu\nu}_{(\chi)}=\sigma(\chi^\mu+\partial^\mu A)$$

$$f^\mu=\sigma(\chi^\mu+\partial^\mu A)$$



$$S=\int~\ddot{B}\left[\frac{1}{2}m\dot{x}^2+V(x)\right]\mathrm{d}t$$

$$S=\int~\ddot{B}\left[\frac{1}{2}m\dot{x}^2+V(x)\right]\mathrm{d}t$$

$$\frac{1}{2}m\dot{x}^2+V(x)=E(t)=Pt+E_0$$

$$m\ddot{x}\frac{\mathrm{d}^2B}{\mathrm{d}t^2}+m\dot{x}\frac{\mathrm{d}^3B}{\mathrm{d}t^3}=V'(x)\frac{\mathrm{d}^2B}{\mathrm{d}t^2}$$

$$\frac{\dot{\tilde{B}}}{\ddot{B}}=\frac{2V'(x)}{\sqrt{2m(E(t)-V(x))}}-\frac{P}{2(E(t)-V(x))}$$

$$\pi_x = \frac{\partial \mathcal{L}}{\partial \dot{x}} = m\dot{x}\ddot{B}$$

$$\pi_B = \frac{\partial \mathcal{L}}{\partial \dot{B}} - \frac{\mathrm{d}}{\mathrm{d} t} \frac{\partial \mathcal{L}}{\partial \ddot{B}} = -\frac{\mathrm{d}}{\mathrm{d} t} E(t)$$

$$\Pi_B = \frac{\partial \mathcal{L}}{\partial \ddot{B}} = E(t)$$

$$\begin{aligned}\mathcal{H} &= \dot{x}\pi_x + \dot{B}\pi_B + \ddot{B}\Pi_B - \mathcal{L} \\ &= \pi_x\sqrt{\frac{2}{m}(\Pi_B-V(x))}+\dot{B}\pi_B=\text{ Const.}\end{aligned}$$

$$\begin{aligned}\mathcal{S} &= \frac{1}{16\pi G}\int~\mathrm{d}^4x\sqrt{-g}R+\int~\mathrm{d}^4x\sqrt{-g}\mathcal{L}(\phi,X) \\ &+ \int~\mathrm{d}^4x\sqrt{-g}\chi,\mu;\nu T_{(\chi)}^{\mu\nu}\end{aligned}$$

$$\mathcal{L}(\phi,X)=\sum_{N=1}^\infty A_n(\phi)X^n-V(\phi)$$

$$T_{(\chi)}^{\mu\nu}=g^{\mu\nu}\Lambda(\phi,X)\Rightarrow\mathcal{S}_{(\chi)}=\int~\mathrm{d}^4x\Lambda\Box\chi$$

$$j_\alpha=2(\Box\chi+1)\phi_{,\alpha}$$

$$T_{(G)}^{\mu\nu}=g^{\mu\nu}\bigl(-\Lambda+\chi^{\sigma}\Lambda_{,\sigma}\bigr)+j^{\mu}\phi^{\cdot\nu}-\chi^{\cdot\mu}\Lambda^{\cdot\nu}-\chi^{\cdot\nu}\Lambda^{\cdot\mu}$$



$$\mathrm{d} s^2 = -\mathrm{d} t^2 + a^2(t) \left[\frac{\mathrm{d} r^2}{1-k r^2} + r^2 \; \mathrm{d} \Omega^2 \right]$$

$$2\dot{\phi}\ddot{\phi}=\frac{C_2}{a^3}$$

$$\dot{\phi}^2=C_1+C_2\int\;\frac{\mathrm{d} t}{a^3}$$

$$2\dot{\phi}(\Box\,\chi+1)=\frac{C_3}{a^3}$$

$$\dot{\chi}=\frac{1}{a^3}\int\;a^3\;\mathrm{d} t+\frac{C_4}{a^3}-\frac{C_3}{2a^3}\int\;\frac{\mathrm{d} t}{\dot{\phi}}$$

$$\rho_{de}=\dot{\phi}^2+2\dot{\chi}\dot{\phi}\ddot{\phi}$$

$$\rho_{dm}=\frac{C_3}{a^3}\dot{\phi}-4\dot{\chi}\dot{\phi}\ddot{\phi}$$

$$\lim_{t\rightarrow\infty}\dot{x}=\frac{1}{3H_0}$$

$$\rho_{de}=C_1+O\left(\frac{1}{a^6}\right);$$

$$\rho_{dm}=\Big(C_3\sqrt{C_1}-\frac{2C_2}{3H_0}\Big)\frac{1}{a^3}+O\left(\frac{1}{a^6}\right)$$

$$\mathcal{S}=\frac{1}{16\pi G}\int\;\;\mathrm{d}^4x\sqrt{-g}R+\int\;\;\mathrm{d}^4x(\Phi+\sqrt{-g})\mathcal{L}(X,\phi)$$

$$\rho_{DE}=\dot{\phi}^2=C_1$$

$$\rho_{\mathrm{Dust}}=\frac{\sqrt{C_1}C_3}{a^3}$$

$$a_0(t)=\left(\frac{C_3}{\sqrt{C_1}}\right)^{\frac{1}{3}}\sinh^{2/3}\left(\alpha t\right)$$

$$\Omega_\Lambda=\frac{C_1}{H};\;\Omega_m=\frac{C_1\sqrt{C_3}}{H}$$



$$\lambda_1(t,t_0)=\frac{C_2}{C_1}\int_{t_0}^t \frac{\mathrm{d} t}{a^3}$$

$$\lambda_2(t,t_0)=\frac{C_2}{\sqrt{C_1}C_3}\dot{\chi}(t,t_0)$$

$$\rho_{de}=C_1\left(1+\lambda_1+\frac{C_3}{\sqrt{C_1}}\lambda_2\right)+O_2(\lambda_1,\lambda_2)$$

$$\rho_{dm}=\frac{\sqrt{C_1}C_3}{a^3}\bigg(1-\frac{1}{2}(\lambda_1+\lambda_2)\bigg)+O_2(\lambda_1,\lambda_2)$$

$$\begin{aligned}\mathcal{S}=&\frac{1}{16\pi G}\int\sqrt{-g}R+\int\sqrt{-g}\Lambda+\int\sqrt{-g}\chi_{\mu;\nu}T_{(\chi)}^{\mu\nu}\\&+\frac{\sigma}{2}\int\sqrt{-g}\big(\chi_\mu+\partial_\mu A\big)^2\end{aligned}$$

$$\nabla_\mu \Lambda = f_\mu = \sigma \big(\chi_\mu + \partial_\mu A \big)$$

$$f^\mu_{;\mu}=0$$

$$j_\alpha=2\bigl(\chi^\lambda_{;\lambda}+1\bigr)\phi_{,\alpha}; j^\alpha_\alpha=0$$

$$\begin{aligned}T_{(G)}^{\mu\nu}=&g^{\mu\nu}\left(-\Lambda+\chi^{\lambda}\Lambda_{,\lambda}-\frac{1}{2\sigma}\Lambda^{\lambda}\Lambda_{,\lambda}\right)\\&+j^\mu\phi^{,\nu}-\chi^{\mu}\Lambda^{,\nu}-\chi^{,\nu}\Lambda^{\mu}+\frac{1}{\sigma}\Lambda^{\mu}\Lambda^{\mu}\end{aligned}$$

$$\begin{aligned}\mathcal{S}=&\frac{1}{16\pi G}\int\sqrt{-g}R+\int\sqrt{-g}g^{\alpha\beta}\phi_{,\alpha}\phi_{,\beta}\\&-\int\sqrt{-g}\chi_\mu\nabla_\nu T_{(\chi)}^{\mu\nu}+\frac{\sigma}{2}\int\sqrt{-g}\big(\chi_\mu+\partial_\mu A\big)^2.\end{aligned}$$

$$0=-\nabla_\nu T_{(\chi)}^{\mu\nu}+\sigma \big(\chi_\mu + \partial_\mu A \big)$$

$$\begin{aligned}\mathcal{S}=&\frac{1}{16\pi G}\int\sqrt{-g}R+\int\sqrt{-g}g^{\alpha\beta}\phi_{,\alpha}\phi_{,\beta}\\&-\frac{1}{2\sigma}\int\sqrt{-g}\left(\nabla_\nu T_{(\chi)}^{\mu\nu}\right)^2+\int\sqrt{-g}\partial_\nu A\nabla_\alpha T_{(\chi)}^{\nu\alpha}\end{aligned}$$

$$\begin{aligned}\mathcal{Z}=&\int\mathcal{D}\phi\delta(\nabla_\nu f^\nu)\\&\exp\left[\frac{1}{2\sigma}\int\mathrm{d}^4x\sqrt{g}f_\mu f^\mu-\int\mathrm{d}^4x\sqrt{g}g^{\mu\nu}\phi_{,\mu}\phi_{,\nu}\right]\end{aligned}$$

$$\rho_{de}=C_1+C_2\int\frac{\mathrm{d} t}{a^3}$$



$$\rho_{dm}=\frac{C_3}{a^3}-\frac{C_2t}{a^3}$$

$$\rho_{de}=C_1+C_2\int~\frac{\mathrm{d}t}{a^3}$$

$$\rho_{dm}=\frac{C_3}{a^3}\dot{\phi}$$

$$\frac{\partial g^{\alpha \beta}}{\partial g_{\mu \nu}} = -\frac{1}{2} \big(g^{\alpha \mu} g^{\beta \nu} + g^{\alpha \nu} g^{\beta \mu} \big)$$

$$\frac{\partial \Gamma^\tau_{\lambda\sigma}}{\partial g_{\mu\nu}}=-\frac{1}{2}\big(g^{\mu\tau}\Gamma^\nu_{\lambda\sigma}+g^{\nu\tau}\Gamma^\mu_{\lambda\sigma}\big)$$

$$\frac{\partial \Gamma^\tau_{\lambda\alpha}}{\partial g_{\mu\nu,\sigma}}=\frac{1}{4}\big[g^{\mu\tau}(\delta^\nu_\alpha\delta^\sigma_\lambda+\delta^\nu_\lambda\delta^\sigma_\alpha)+g^{\tau\nu}\big(\delta^\mu_\alpha\delta^\sigma_\lambda+\delta^\mu_\lambda\delta^\sigma_\alpha\big)-g^{\tau\sigma}\big(\delta^\mu_\alpha\delta^\nu_\lambda+\delta^\mu_\lambda\delta^\nu_\alpha\big)\big]$$

$$T_{(G)}^{\alpha\beta}=\frac{-2}{\sqrt{-g}}\frac{\partial\left(\sqrt{-g}\chi_{\mu;\nu}T_{(\chi)}^{\mu\nu}\right)}{\partial g_{\alpha\beta}}$$

$$+\frac{2}{\sqrt{-g}}\frac{\partial}{\partial x^\sigma}\frac{\partial\left(\sqrt{-g}\chi_{\mu;\nu}T_{(\chi)}^{\mu\nu}\right)}{\partial g_{\alpha\beta,\sigma}}$$

$$\dot\rho+3\frac{\dot a}{a}(\rho+p)=\frac{C_2}{a^3}$$

$$\nabla_\mu T_{\rm (Dust)}^{\mu\nu}=\gamma^2 j^\nu$$

$$\nabla_\mu T_{\rm (Dust)}^{\mu\nu}=-\nabla_\mu T_{(\Lambda)}^{\mu\nu}=\gamma^2 j^\nu$$

$$\nabla_\mu \left(T_{\rm (Dust)}^{\mu\nu}+T_{(\Lambda)}^{\mu\nu}\right)=0$$

$$T^\mu_\nu=\textbf{Diag}(\rho,-p,-p,-p)$$

$$\dot{\rho}_{\rm dust}+3H(1+\tilde{\omega})\rho_{\rm dust}=\frac{\gamma^2}{a^3}$$

$$\dot{\rho}_\Lambda+3H(1+\omega)\rho_\Lambda=-\frac{\gamma^2}{a^3}.$$

$$x=\frac{\rho_{\rm dust}}{3H^2}, y=\frac{\rho_\Lambda}{3H^2}, \delta=\frac{\gamma^2}{a^3 H \rho_{\rm dust}}$$

$$x'=6x^2(\tilde{\omega}-\omega)+x(\gamma\delta+3+6\omega+3\tilde{\omega})$$



$$\delta' = \delta(\gamma\delta + 3(x-1)(\omega - \tilde{\omega})).$$

Name	The point	Eigenvalues	Densities fraction
A	$(0, \frac{3}{\gamma}(\omega - \tilde{\omega}))$	$3(3\omega + 1), 3(\omega - \tilde{\omega})$	0
B	$\left(\frac{\omega+1/3}{\omega-\tilde{\omega}}, -\frac{3\tilde{\omega}+1}{\gamma}\right)$	$\frac{1}{2} \left(\pm \sqrt{36\omega^2 - 72\omega\tilde{\omega} + 9(\tilde{\omega}-2)\tilde{\omega}} - 3\omega - 3\tilde{\omega} - 3 \right)$	$-\frac{3\omega+1}{3\tilde{\omega}+1}$
C	$(0,0)$	$3(\tilde{\omega} - \omega), 3(2\omega + \tilde{\omega} + 1)$	0
D	$\left(\frac{2\omega+\tilde{\omega}+1}{2(\omega-\tilde{\omega})}, 0\right)$	$-3(2\omega + \tilde{\omega} + 1), \frac{3}{2}(1 + 3\tilde{\omega})$	$-\frac{2\omega+\tilde{\omega}+1}{3\tilde{\omega}+1}$

$$\Phi = \varepsilon^{\alpha\beta\gamma\delta} \varepsilon_{abcd} \partial_\alpha \varphi_a \partial_\beta \varphi_b \partial_\gamma \varphi_c \partial_\delta \varphi_d.$$

$$S = \int d^4x \Phi \mathcal{L}_1 + \int d^4x \sqrt{-g} \mathcal{L}_2$$

$$A_a^\alpha \partial_\alpha \mathcal{L}_1 = 0,$$

$$\mathcal{S} = \int d^4x \sqrt{-g} \chi_{\mu;\nu} T_{(\chi)}^{\mu\nu}$$

$$\delta \chi_\mu : \nabla_\mu T_{(\chi)}^{\mu\nu} = 0.$$

$$\begin{aligned} \int d^4x \sqrt{-g} \chi_{\mu;\nu} T_{(\chi)}^{\mu\nu} &= \int d^4x \sqrt{-g} \chi_{;\lambda}^\lambda \mathcal{L}_m \\ &= \int d^4x \partial_\mu (\sqrt{-g} \chi^\mu) \mathcal{L}_m = \int d^4x \Phi \mathcal{L}_m \end{aligned}$$

$$\pi_{\chi^0} = \frac{\partial \mathcal{L}}{\partial \dot{\chi}^0} = T_0^0(\chi) := \rho_{(\chi)}$$

$$\begin{aligned} S_{(\chi,A)} &= \int d^4x \sqrt{-g} \chi_{\mu;\nu} T_{(\chi)}^{\mu\nu} \\ &\quad + \frac{\kappa}{2} \int d^4x \sqrt{-g} (\chi_\mu + \partial_\mu A)^2 \end{aligned}$$

$$\nabla_\nu T_{(\chi)}^{\mu\nu} = \kappa(\chi^\mu + \partial^\mu A) = f^\mu,$$

$$\nabla_\mu f^\mu = \kappa \nabla_\mu (\chi^\mu + \partial^\mu A) = 0.$$

$$T_{(G)}^{\mu\nu} = \frac{-2}{\sqrt{-g}} \frac{\delta(\sqrt{-g} \mathcal{L}_M)}{\delta g^{\mu\nu}}, \nabla_\mu T_{(G)}^{\mu\nu} = 0.$$

$$f_\mu = \kappa(\chi_\mu + \partial_\mu A) = 0 \Rightarrow \chi_\mu = -\partial_\mu A$$

$$\mathcal{S} = - \int d^4x \sqrt{-g} \nabla_\mu \nabla_\nu \chi T_{(\chi)}^{\mu\nu}$$

$$\mathcal{L} = \frac{1}{2} \mathcal{R} + \chi_{,\mu;L} T_{(\chi)}^{\mu\nu} - \frac{1}{2} \phi^{,\mu} \phi_{,\mu} - V(\phi)$$



$$T_{(\chi)}^{\mu\nu}=-\frac{\lambda_1}{2}\phi^\mu\phi^\nu-\frac{\lambda_2}{2}g^{\mu\nu}\big(\phi_{,\alpha}\phi^{,\alpha}\big)+g^{\mu\nu}U(\phi),$$

$$\begin{aligned}\rho_{(\chi)}&=(\lambda_1+\lambda_2)\frac{\dot{\phi}^2}{2}+U(\phi),\\ p_{(\chi)}&=-\lambda_2\frac{\dot{\phi}^2}{2}-U(\phi).\end{aligned}$$

$$\nabla_\mu \nabla_\nu T_{(\chi)}^{\mu\nu}=0$$

$$j^\mu=\frac{\lambda_1}{2}(\chi^{\mu;\nu}+\chi^{\nu;\mu})\phi_{,\nu}+(1+\lambda_2\Box\chi)\phi^{,\mu},$$

$$\nabla_\mu j^\mu=V'(\phi)-\Box\,\chi U'(\phi)$$

$$\begin{aligned}G^{\mu\nu}=&g^{\mu\nu}\left(-\chi_{,\alpha;\beta}T_{(\chi)}^{\alpha\beta}+\frac{1}{2}\phi^{,\alpha}\phi_{,\alpha}+V(\phi)\right)\\&-\phi^{,\mu}\phi^\nu+\chi_{,\alpha;\beta}\frac{\partial T_{(\chi)}^{\alpha\beta}}{g_{\mu\nu}}\\&+\nabla_\lambda\left(\chi^\mu T_{(\chi)}^{\nu\lambda}+\chi^\nu T_{(\chi)}^{\mu\lambda}-\chi^\lambda T_{(\chi)}^{\mu\nu}\right)\end{aligned}$$

$$\begin{aligned}\chi_{,\alpha;\beta}\frac{\partial T^{\alpha\beta}}{\partial g_{\mu\nu}}=&-\frac{\lambda_1}{2}\chi^{(\mu}\phi^{\nu)}\Box\phi+\left(\frac{\lambda_1}{2}+\lambda_2\right)\phi^{,\mu}\phi^{,\nu}\Box\chi\\&+\frac{\lambda_1}{2}\chi^{\gamma;\mu}\phi^{\nu}\phi_{,\gamma}-\lambda_2\phi^{\mu;\lambda}\chi^{\nu}\phi_{,\lambda}-\lambda_2\chi^{\mu}\phi^{\gamma;\nu}\phi_{,\nu}\\&+\frac{\lambda_1}{2}\phi^{,\mu}\chi^{\gamma;\nu}\phi_{,\gamma}-\frac{\lambda_1}{2}\chi^{(\nu}\phi^{,\mu);\gamma}\phi_{,\gamma}\\&+\frac{\lambda_1}{2}\pi^{(\nu}\phi^{,\mu);\gamma}\chi_{,\gamma}+\chi^{(\mu}\phi^{\nu)}U'(\phi).\end{aligned}$$

$$ds^2 = -dt^2 + a(t)^2 \left(\frac{dr^2}{1-Kr^2} + r^2 d\Omega^2 \right)$$

$$(\lambda_1+\lambda_2)\dot{\phi}\ddot{\phi}+U'(\phi)\dot{\phi}+3H\lambda_1\dot{\phi}^2=\frac{\sigma}{a^3}$$

$$\dot{\phi}^2=\dot{\phi}_{(0)}^2a^{-\frac{3\lambda_1}{\lambda_1+\lambda_2}}+\frac{\sigma}{\lambda_1+\lambda_2}a^{-\frac{3\lambda_1}{\lambda_1+\lambda_2}}\int_0^tdsa^{-\frac{3\lambda_2}{\lambda_1+\lambda_2}}$$

$$\left(\frac{\lambda_1}{2}-\lambda_2\right)\ddot{\chi}+(1-3H\dot{\chi})\lambda_2=\frac{\tilde{\sigma}}{\dot{\phi}a^3}$$

$$\begin{aligned}\rho=&\frac{3}{2}H(\lambda_1-2\lambda_2)\dot{\chi}\dot{\phi}^2+\frac{1}{2}\dot{\phi}^2(1-2(\lambda_1+\lambda_2)\ddot{\chi})\\&+\dot{\chi}\dot{\phi}\left((\lambda_1+\lambda_2)\ddot{\phi}\right)+V,\end{aligned}$$

$$p=\frac{1}{2}\dot{\phi}^2-\frac{1}{2}\lambda_1\ddot{\chi}\dot{\phi}^2+\lambda_2\dot{\chi}\dot{\phi}\ddot{\phi}-V$$

$$a\sim t^\alpha.$$



$$\dot{\phi} = \sqrt{\frac{2\sigma}{3\alpha(\lambda_1 - \lambda_2) + \lambda_1 + \lambda_2}} t^{\frac{1}{2} - \frac{3\alpha}{2}},$$

$$\dot{\chi} = \mathcal{C}t$$

$$C = \frac{2\lambda_2}{-6\alpha\lambda_2 + \lambda_1 - 2\lambda_2}.$$

$$\rho = \frac{\alpha_1}{a^3} + \frac{\alpha_2 t}{a^3} + V$$

$$\begin{aligned}\alpha_1 &= \frac{18\alpha^2\lambda_2(2\lambda_2 - \lambda_1)}{2(\lambda_1 - 2\lambda_2(3\alpha + 1))} \\ \alpha_2 &= \frac{(6\alpha + 2)\lambda_1\lambda_2 + 2(3\alpha + 1)(\lambda_2 - 1)\lambda_2 + \lambda_1}{2(\lambda_1 - 2\lambda_2(3\alpha + 1))}\end{aligned}$$

$$a \sim t.$$

$$\dot{\phi}^2 = \dot{\phi}_0^2 a^{-\frac{3\lambda_1}{\lambda_1 + \lambda_2}} - \sigma_0 H_0 \frac{\lambda_1 + \lambda_2}{3\lambda_2} \frac{1}{a^3}$$

$$\dot{x} = \frac{1}{3H_0} + \mathcal{O}\left(\frac{1}{a^3}\right).$$

$$\begin{aligned}\rho &= H_0(3\lambda_2 - 1)\sigma \frac{\lambda_1 + \lambda_2}{6\lambda_2} \frac{1}{a^3} + V \\ &\quad + \frac{1}{2}\dot{\phi}_0^2(1 - 2\lambda_2)a^{-\frac{3\lambda_1}{\lambda_1 + \lambda_2}}.\end{aligned}$$

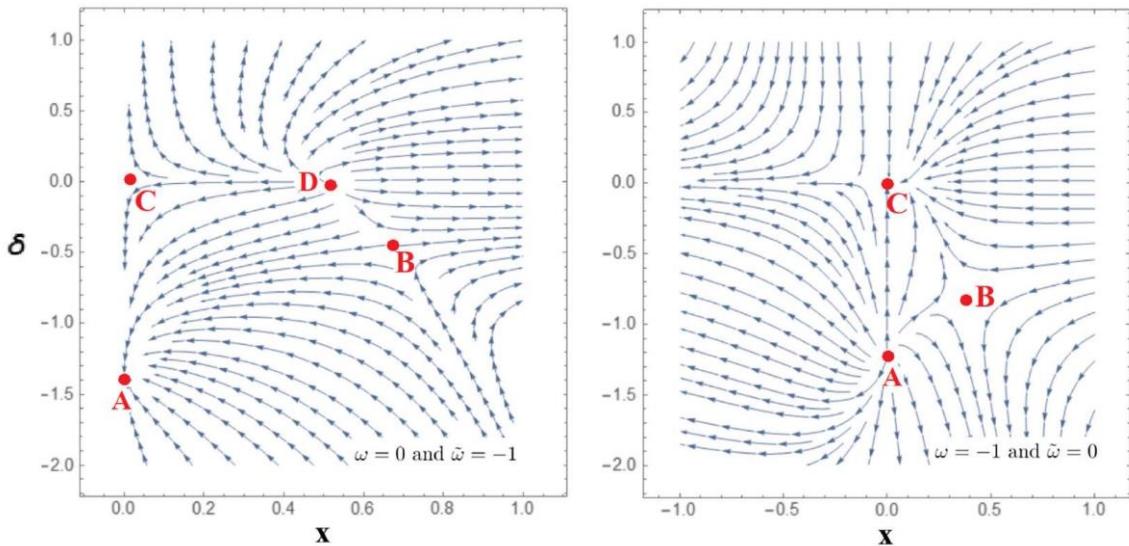


Figura 6. Comportamiento vectorial de la partícula oscura en un espacio – tiempo cuántico relativista.

$$T_{(\chi)}^{\mu\nu} = -\frac{\lambda_1}{2}\phi^{\mu}\phi^{\nu} + g^{\mu\nu}U(\phi)$$



$$\dot{\phi}^2=\frac{\dot{\phi}_{(0)}^2}{a^3}+\frac{\sigma}{\lambda_1}\frac{t}{a^3}$$

$$\dot{\chi}(t)=\dot{\chi}(0)-\frac{2}{\lambda_1}t+\int~dt\frac{2\tilde{\sigma}}{\lambda_1\dot{\phi}a^3}$$

$$\dot{\chi}(t\rightarrow \infty)\rightarrow -\frac{2t}{\lambda_1}$$

$$\rho=V+\frac{\alpha_1}{a^3}+\frac{\alpha_2}{a^{4.5}}$$

$$\begin{aligned}\alpha_1&=\frac{5\dot{\phi}_0^2\lambda_1+\lambda_1\sigma\chi_0+3\sigma t}{2\lambda_1}\\ \alpha_2&=-\frac{2\tilde{\sigma}}{3\dot{\phi}_0H_0\lambda_1}\big(3\dot{\phi}_0^2H_0\lambda_1+3H_0\sigma t+\sigma\big)\end{aligned}$$

$$\chi\rightarrow\chi+ct$$

$$\chi^0\rightarrow\chi^0+c$$

$$\partial_t\rho_{\mathrm{dust}}+3H\rho_{\mathrm{dust}}=\frac{\gamma^2}{a^3},$$

$$\rho_{\mathrm{dust}}=\frac{C_1}{a^3}+\frac{\gamma^2t}{a^3},$$

$$\partial_t\rho_\Lambda=-\frac{\gamma^2}{a^3},$$

$$\rho_\Lambda=C_2-\gamma^2\int~\frac{dt}{a^3}$$

$$\rho=\frac{C_1+\gamma^2t}{a^3}+C_2+\mathcal{O}\left(\frac{1}{a^6}\right),$$

$$\begin{aligned}C_1&=\frac{5\dot{\phi}_0^2+\sigma\chi_0}{2}\\ \gamma^2&=\frac{3\sigma}{2\lambda_1}, C_2=V.\end{aligned}$$

$$\frac{d}{dt}=-H(z)(z+1)\frac{d}{dz}$$

$$\left(\frac{\lambda_1}{2}-\lambda_2\right)\ddot{\chi}+(1-3H\dot{\chi})\lambda_2=\frac{\tilde{\sigma}}{\dot{\phi}a^3}$$

$$\lambda_1\ddot{\chi}-\lambda_2(\ddot{\chi}-3H\dot{\chi})+1=\tilde{\sigma}/(\dot{\phi}a^3).$$

$$\dot{\chi}=\frac{1}{3H_0}+\mathcal{O}\left(\frac{1}{a^3}\right),$$



$$\dot{\chi} = \frac{1}{3\lambda_2 H_0} + \mathcal{O}\left(\frac{1}{a^3}\right).$$

$$\rho = H_0(3\lambda_2 - 1)\sigma \frac{\lambda_1 + \lambda_2}{6\lambda_2} \frac{1}{a^3} + V + \frac{1}{2}\dot{\phi}_0^2(1 - 2\lambda_2)a^{-\frac{3\lambda_1}{\lambda_1 + \lambda_2}}$$

$$\rho = \frac{H_0\sigma_0(\lambda_1 + \lambda_2)}{3\lambda_2} \frac{1}{a^3} + V - \frac{1}{2}\dot{\phi}_0^2a^{-\frac{3\lambda_1}{\lambda_1 + \lambda_2}}.$$

$$\dot{\chi}(t) = \dot{\chi}(0) - \frac{1}{\lambda_1}t + \int dt \frac{\tilde{\sigma}}{\lambda_1 \dot{\phi} a^3}$$

$$\dot{\chi}(t) = \dot{\chi}(0) - \frac{2}{\lambda_1}t + \int dt \frac{2\tilde{\sigma}}{\lambda_1 \dot{\phi} a^3}.$$

$$\dot{\chi}(t \rightarrow \infty) \rightarrow -\frac{t}{\lambda_1}$$

$$\dot{\chi}(t \rightarrow \infty) \rightarrow -\frac{2t}{\lambda_1}.$$

$$\begin{aligned}\alpha_1 &= \sigma\dot{\chi}_0 + \frac{\sigma t}{\lambda_1} + 3\dot{\phi}_0^2 \\ \alpha_2 &= \frac{\tilde{\sigma}}{\sqrt{\lambda_1(\lambda_1\dot{\phi}_0^2 + \sigma t)}}(3H_0(\lambda_1\dot{\phi}_0^2 + \sigma t) + 2\sigma)\end{aligned}$$

$$\alpha_1 = \frac{5\dot{\phi}_0^2\lambda_1 + \lambda_1\sigma\chi_0 + 3\sigma t}{2\lambda_1}, \alpha_2 = -\frac{2\tilde{\sigma}}{3\dot{\phi}_0 H_0 \lambda_1}(3\dot{\phi}_0^2 H_0 \lambda_1 + 3H_0\sigma t + \sigma)$$

$$\begin{aligned}C_1 &= \frac{3\dot{\phi}_0^2 + \sigma\dot{\chi}_0}{2}, \\ \gamma^2 &= \frac{\sigma}{\lambda_1}, C_2 = V\end{aligned}$$

$$C_1 = \frac{5\dot{\phi}_0^2 + \sigma\chi_0}{2}, \gamma^2 = \frac{3\sigma}{2\lambda_1}, C_2 = V$$

SECCIÓN II.

$$p_{dm} = 0, \rho_{dm} = nm$$

$$m = \lambda\varphi$$

$$\rho_{dm} = \lambda n\varphi.$$

$$\rho_\varphi = \frac{1}{2}\dot{\varphi}^2 + V(\varphi), \rho_\varphi = \frac{1}{2}\dot{\varphi}^2 - V(\varphi).$$

$$\begin{aligned}\dot{\varphi} + 3H\dot{\varphi} + \frac{dV_{eff}}{d\varphi} &= 0 \\ V_{eff} &= V(\varphi) + \lambda n\varphi\end{aligned}$$



$$\ddot{\varphi}+3H\dot{\varphi}+\frac{dV_{eff}}{d\varphi}=-\lambda n_0a^{-3}.$$

$$m_\psi=\lambda \langle \varphi \rangle.$$

$$V(\varphi)=u_0\varphi^{-p}\;(p>0).$$

$$V_{eff}(\varphi) = V(\varphi) + \lambda n_\psi \varphi$$

$$\langle \varphi \rangle = \left(\frac{p u_0}{\lambda n_\psi} \right)^{1/1+p}$$

$$\nabla_\mu \left(T^{\mu}_{(dm)_\nu^\mu}+T^{\mu}_{(\varphi)_\nu^\mu}\right)=\nabla_\mu T^{\mu}_{(dm)_\nu^\mu}+\nabla_\mu T^{\mu}_{(\varphi)_\nu^\mu}=0,$$

$$\nabla_\mu T^{\mu}_{(dm)_\nu^\mu}=-\nabla_\mu T^{\mu}_{(\varphi)_\nu^\mu}$$

$$T^{\mu}_{(\varphi)_\nu^\mu}=\partial^\mu\varphi\partial_\nu\varphi-\delta^\mu_\nu\left[\frac{1}{2}\partial^\alpha\varphi\partial_\alpha\varphi-V(\varphi)\right]$$

$$\nabla_\mu T^{\mu}_{(\varphi)_\nu^\mu}=\partial_\mu T^{\mu}_{(\varphi)_\nu^\mu}+\Gamma^\mu_{\mu\beta}T^\beta_{(\varphi)_\nu^\beta}-\Gamma^\beta_{\mu\nu}T^\mu_{(\varphi)_\beta^\mu}=-\left(\ddot{\varphi}+3H\dot{\varphi}+\frac{dV}{d\varphi}\right)\partial_\nu\varphi.$$

$$\nabla_\mu T^{\mu}_{(\varphi)_\nu^\mu}=\lambda n\partial_\nu\varphi$$

$$\nabla_\mu T^{\mu}_{(dm)_0^\mu}=\dot{\rho}_{dm}+3H\rho_{dm}=\lambda n\dot{\varphi}$$

$$T^\mu_\nu=Vg^\mu_\nu$$

$$T^\mu_\nu=(\rho+p)u^\mu u_\nu-p g^\mu_\nu$$

$$\nabla_\mu \hat{T}^\mu_\nu=F_\nu, F_\mu\equiv \nabla_\mu V$$

$$\nabla_\mu \left(T^{\mu}_{(de)_\nu^\mu}+T^{\mu}_{(dm)_\nu^\mu}\right)=0.$$

$$\nabla_\mu T^{\mu}_{(de)_\nu^\mu}=-T^{\mu}_{(dm)_\nu^\mu}=F_\nu,$$

$$R_{\mu\nu}-\frac{1}{2}Rg_{\mu\nu}=8\pi G\left(T_{(de)_{\mu\nu}}+T_{(dm)_{\mu\nu}}\right)$$

$$\begin{aligned}u^\mu\nabla^\nu T_{(dm) \mu\nu}&=-u^\mu F_\mu\\u^\mu\nabla^\nu T_{(de) \mu\nu}&=u^\mu F_\mu\end{aligned}$$

$$\begin{aligned}h^{\mu\beta}\nabla^\nu T_{(dm) \mu\nu}&=-h^{\mu\beta}F_\mu\\h^{\mu\beta}\nabla^\nu\nabla^\nu T_{(de) \mu\nu}&=h^{\mu\beta}F_\mu\end{aligned}$$

$$\begin{aligned}h^{\mu\beta}\nabla_\mu p_{dm}+(\rho_{dm}+p_{dm})u^\mu\nabla_\mu u^\beta&=-h^{\mu\beta}F_\mu\\h^{\mu\beta}\nabla_\mu p_{de}+(\rho_{de}+p_{de})u^\mu\nabla_\mu u^\beta&=h^{\mu\beta}F_\mu\end{aligned}$$

$$\begin{array}{l}\nabla_\mu u^\mu=3H\\ u^\mu\nabla_\mu u^\nu=0\end{array}$$

$$\begin{array}{l}\dot{\rho}_{dm}+3H\rho_{dm}=Q\\\dot{\rho}_{de}+3H(\rho_{de}+p_{de})=-Q\end{array}$$

$$L=i\bar{\psi}\gamma_{\mu}\nabla^{\mu}\psi-m_{\psi}\bar{\psi}\psi-\frac{1}{2}\nabla_{\mu}\varphi\nabla^{\mu}\varphi-\frac{1}{2}m_{\varphi}\varphi^2+g\varphi\bar{\psi}\psi$$

$$V(r) = - \frac{G m_{\psi}^2}{r} \Big[1 + \alpha {\rm exp} \left(- \frac{r}{r_s} \right) \Big]$$

$$\mathcal{L}=\frac{1}{2}\Big[\frac{R}{\kappa}-\nabla^a\varphi\nabla_a\varphi\Big]+V(\varphi)-C(\varphi)\mathcal{L}_{\text{DM}}+\mathcal{L}_{\text{S}}$$

$$T_{ab}=\nabla_a\varphi\nabla_b\varphi-g_{ab}\left[\frac{1}{2}\nabla^c\nabla_c\varphi-V(\varphi)\right]+C(\varphi)T_{ab}^{\text{DM}}+T_{ab}^S$$

$$\frac{d^2\mathbf{r}}{dt^2}=-\vec{\nabla}\Phi-\frac{C_{\varphi}(\varphi)}{C(\varphi)}\vec{\nabla}\varphi$$

$$\mathcal{S}=\int\,\,d^4x\sqrt{-g}\left[\frac{1}{2}\partial^{\mu}\phi\partial_{\mu}\phi+U(\phi)+m(\phi)\bar{\psi}\psi-\mathcal{L}_{kin}[\psi]\right],$$

$$\nabla_\nu T^\nu_{(\alpha)\mu}=Q_{(\alpha)\mu},$$

$$\sum_\alpha\, Q_{(\alpha)\mu}=0$$

$$\begin{aligned}\frac{d\rho_\phi}{d\eta}&=-3\mathcal{H}(p_\phi+\rho_\phi)+\beta(\phi)\frac{d\phi}{d\eta}(1-3w_\alpha)\rho_\alpha\\ \frac{d\rho_\alpha}{d\eta}&=-3\mathcal{H}(p_\phi+\rho_\phi)-\beta(\phi)\frac{d\phi}{d\eta}(1-3w_\alpha)\rho_\alpha\end{aligned}$$

$$Q_{(\phi)\mu}=\frac{\partial \mathrm{ln}~m(\phi)}{\partial \phi}T_\alpha\partial_\mu\phi~,m_\alpha=\bar{m}_\alpha e^{-\beta(\phi)\phi}.$$

$$\frac{d\mathbf{v}_\alpha}{d\eta}+\Big(\mathcal{H}-\beta(\phi)\frac{d\phi}{d\eta}\Big)\mathbf{v}_\alpha-\nabla[\Phi_\alpha+\beta\phi]=0.$$

$$\dot{\mathbf{v}}_\alpha=-\tilde{H}\mathbf{v}_\alpha-\nabla\frac{\tilde{G}_\alpha m_\alpha}{r}.$$

$$\nabla[\Phi_\alpha+\beta\phi]\tilde{G}_\alpha=G_N[1+2\beta^2(\phi)]$$

$$\tilde{H}\mathbf{v}_\alpha\equiv H\bigg(1-\beta(\phi)\frac{\dot{\phi}}{H}\bigg)\mathbf{v}_\alpha$$

$$\begin{array}{l}\dot{\rho}_{dm}+3H\rho_{dm}=Q\\\dot{\rho}_{de}+3H(1+w_{de})\rho_{de}=-Q\end{array}$$



$Q \begin{cases} > 0 \rightarrow \text{energy transfer is} \\ < 0 \end{cases} \begin{cases} \text{dark energy} \rightarrow \text{dark matter} \\ \text{dark matter} \rightarrow \text{dark energy} \end{cases}$

$$w_{de} = w_\varphi = \frac{p_\varphi}{\rho_\varphi} = \frac{\frac{1}{2}\dot{\varphi}^2 - V(\varphi)}{\frac{1}{2}\dot{\varphi}^2 + V(\varphi)}$$

$$\ddot{\varphi}^2 + 3H\dot{\varphi} + \frac{dV}{d\varphi} = -\frac{Q}{\dot{\varphi}}.$$

$$\dot{\rho}_i + 3H(1 + w_{eff,i})\rho_i = 0, i = de, dm$$

$$w_{eff,dm} = -\frac{Q}{3H\rho_{dm}}, w_{eff,de} = w_{de} + \frac{Q}{3H\rho_{de}},$$

$$\begin{aligned} Q > 0 \rightarrow & \begin{cases} w_{eff,dm} < 0 & \text{dark matter redshifts slower than } a^{-3} \\ w_{eff,de} > w_{de} & \text{dark energy has less accelerating power} \end{cases} \\ Q < 0 \rightarrow & \begin{cases} w_{eff,dm} > 0 & \text{dark matter redshifts faster than } a^{-3} \\ w_{eff,de} < w_{de} & \text{dark energy has more accelerating power} \end{cases} \end{aligned}$$

$$w_{eff,dm} = w_{dm} = 0, w_{eff,de} = w_{de}$$

$$Q \equiv -3H\Pi_{dm} = +3H\Pi_{de}$$

$$\begin{aligned} \dot{\rho}_{dm} + 3H(\rho_{dm} + \Pi_{dm}) &= 0 \\ \dot{\rho}_{de} + 3H(\rho_{de} + p_{de} + \Pi_{de}) &= 0 \end{aligned}$$

$$\dot{r} = \frac{\rho_{dm}}{\rho_{de}} \left(\frac{\dot{\rho}_m}{\rho_m} - \frac{\dot{\rho}_{de}}{\rho_{de}} \right) = 3Hr \left(w_{de} + \frac{1+r}{\rho_{dm}} \frac{Q}{3H} \right)$$

$$\frac{Q}{3H\rho_{dm}} = -\frac{w_{de} + \frac{\xi}{3}}{1+r}$$

$$\begin{aligned} \frac{\rho'_{dm}}{\rho_{dm}} &= -1 - \frac{\Pi}{\rho_{dm}} \\ \frac{\rho'_{de}}{\rho_{de}} &= -(1+w_{de}) + \frac{\Pi}{\rho_{de}} \end{aligned}$$

$$\begin{aligned} \rho' &= -\left(1 + \frac{w_{de}}{1+r}\right)\rho \\ r' &= r \left[w_{de} - \frac{(1+r)^2}{r\rho} \Pi \right] \end{aligned}$$

$$\begin{aligned} \Omega'_{dm} &= w_{de}\Omega_{dm}\Omega_{de} \\ \Omega'_{de} &= -w_{de}\Omega_{dm}\Omega_{de} \end{aligned}$$

$$r_c = -(1+w_{de})$$

$$\rho_c = -\frac{w_{de}}{1+w_{de}}\Pi_c$$



$$\frac{d}{dt} = \frac{d}{dz} \frac{dz}{da} \frac{da}{dt} = -(1+z)H(z) \frac{d}{dz}$$

$$\begin{aligned}\frac{d\rho_{dm}}{dz}-\frac{3}{1+z}\rho_{dm}&=-\frac{Q(z)}{(1+z)H(z)}\\\frac{d\rho_{de}}{dz}-\frac{3}{1+z}(1+w_{de})\rho_{de}&=\frac{Q(z)}{(1+z)H(z)}\end{aligned}$$

$$I_Q(z)\equiv\frac{1}{\rho_{crit}^0(1+z)^3H(z)}Q(z)$$

$$\begin{aligned}\frac{d\Omega_{dm}}{dz}-\frac{3}{1+z}\Omega_{dm}&=-(1+z)^2I_Q(z)\\\frac{d\Omega_{de}}{dz}-\frac{3}{1+z}(1+w_{de})\Omega_{de}&=(1+z)^2I_Q(z)\end{aligned}$$

$$q=q_0^*+q_{dm}^*(\rho_{dm}-\rho_{dm,0})+q_{de}^*(\rho_{de}-\rho_{de,0})$$

$$q=q_0+q_{dm}\rho_{dm}+q_{de}\rho_{de}$$

$$\begin{aligned}q &\propto \rho_{dm}, q_0 = q_{de} = 0 \\q &\propto \rho_{de}, q_0 = q_{dm} = 0 \\q &\propto \rho_{total}, q_0 = 0, q_{dm} = q_{de}\end{aligned}$$

$$\delta(a)=\frac{d\ln m_\psi(a)}{d\ln a}$$

$$\dot{\rho}_{dm} + 3H\rho_{dm} - \delta(a)H\rho_{dm} = 0$$

$$\rho_{dm}(a)=\rho_{dm}^{(0)}a^{-3}\exp\left(-\int_a^1\delta(a')d\ln a'\right)$$

$$\rho_{dm}(a)=\rho_{dm,0}a^{-3+\delta}$$

$$\dot{\rho}_{de} + 3H(\rho_{de} + p_{de}) + \delta(a)H\rho_{dm} = 0$$

$$\rho_{de}(a)=\rho_{de,0}a^{-3(1+w_{de})}+\frac{\delta\rho_{dm,0}}{\delta+3w_{de}}\left(a^{-3(1+w_{de})}-a^{-3+\delta}\right)$$

$$\rho_{\Lambda}(a)=\rho_{\Lambda,0}-\frac{\delta\rho_{dm,0}}{3-\delta}a^{-3+\delta}$$

$$\begin{aligned}\rho_{de}(a)&=\rho_{de,0}a^{-[3(1+w_{de})+\delta]}\\\rho_{dm}(a)&=\frac{-\delta\rho_{de,0}}{3w_{de}+\delta}a^{-[3(1+w_{de})+\delta]}+\left(\rho_{dm,0}+\frac{\delta\rho_{de,0}}{3w_{de}+\delta}\right)a^{-3}\end{aligned}$$

$$\begin{aligned}p_1&=(\gamma_1-1)\rho_1\\p_2&=(\gamma_2-1)\rho_2\end{aligned}$$

$$\begin{aligned}\dot{\rho}_1+3H\gamma_1\rho_1&=-\beta H\rho_1+\alpha H\rho_2\\\dot{\rho}_2+3H\gamma_2\rho_2&=\beta H\rho_1-\alpha H\rho_2\end{aligned}$$



$$\begin{aligned}\ddot{H}+H\dot{H}(\alpha+\beta+3\gamma_1+3\gamma_2)+\frac{3}{2}H^3(\alpha\gamma_1+\beta\gamma_2+3\gamma_1\gamma_2)\\=\ddot{H}+AH\dot{H}+BH^3=0\\A\equiv\alpha+\beta+3\gamma_1+3\gamma_2,B\equiv\frac{3}{2}(\alpha\gamma_1+\beta\gamma_2+3\gamma_1\gamma_2)\end{aligned}$$

$$H=\frac{h}{t}, h\neq 0$$

$$Bh^2-Ah+2=0$$

$$h_{\pm}=\frac{A\pm\sqrt{A^2-8B}}{2B}$$

$$H^2=a^{-A/2}\left(c_1a^{\sqrt{A^2-8B}/2}+c_2a^{-\sqrt{A^2-8B}/2}\right)$$

$$H^2 \rightarrow a^{-(A-\sqrt{A^2-8B})/2}$$

$$H^2 \rightarrow a^{-(A+\sqrt{A^2-8B})/2}$$

$$a_{\pm}\propto t^{(A\pm\sqrt{A^2-8B})/2B}$$

$$\frac{\rho_2''}{\rho_2}+A\frac{\rho_2'}{\rho_2}+2B=0$$

$$\rho_2=\rho_{20}a^M$$

$$\rho_1=\rho_{10}a^M$$

$$\frac{\rho_2}{\rho_1}=\frac{\beta}{N+3\gamma_1+\alpha}$$

$$\begin{aligned}A&=\beta+4\\B&=2\beta\\\delta&=\frac{B}{A^2}=\frac{2\beta}{(\beta+4)^2}\\h_+&=\frac{1}{2},h_- =\frac{2}{\beta}\end{aligned}$$

$$\rho_{dm}=\frac{r}{1+r}\rho,\rho_{de}=\frac{1}{1+r}\rho,r\equiv\frac{\rho_{dm}}{\rho_{de}},\rho=\rho_{dm}+\rho_{de}$$

$$\Pi=-\gamma\rho^mr^n(1+r)^s$$

$$Q=3\gamma H\rho_{de}^{m-n}\rho_{dm}^n=3\gamma H\rho_{de}^mr^n$$

$$\begin{aligned}\rho'&=-\Big(1+\frac{w_{de}}{1+r}\Big)\rho,\\r'&=r(w_{de}+\gamma).\end{aligned}$$



$$r = r_0 a^{3(w_{de} + \gamma)},$$

$$\rho = \rho_0 a^{-3(1+w_{de})} \left[\frac{1 + r_0 a^{3(w_{de} + \gamma)}}{1 + r_0} \right]^{\frac{w_{de}}{w_{de} + \gamma}},$$

$$\rho_{dm} = \rho_{dm0} a^{-3(1-\gamma)} \left[\frac{1 + r_0 a^{3(w_{de} + \gamma)}}{1 + r_0} \right]^{\frac{\gamma}{w_{de} + \gamma}},$$

$$\rho_{de} = \rho_{de0} a^{-3(1+w)} \left[\frac{1 + r_0 a^{3(w_{de} + \gamma)}}{1 + r_0} \right]^{\frac{\gamma}{w_{de} + \gamma}}.$$

$$r = r_0 \frac{w_{de}}{(w_{de} + \gamma r_0)^{-3w_{de} - \gamma r_0}},$$

$$\rho = \rho_0 a^{-3\left(1 - \frac{\gamma w_{de}}{w_{de} - \gamma}\right)} \left[\frac{(w_{de} + \gamma r_0) a^{-3w_{de}} + r_0 (w_{de} - \gamma)}{w_{de} (1 + r_0)} \right]$$

$$r = \left(r_0 - \frac{\gamma}{|w_{de}|} \right) a^{-3|w_{de}|} + \frac{\gamma}{|w_{de}|}$$

$$\rho = \rho_0 a^{-3\left(1 - \frac{w_{de}^2}{|w_{de}| + \gamma}\right)} \left[\frac{|w_{de}| + \gamma + (|w_{de}| r_0 - \gamma) a^{-3|w_{de}|}}{|w_{de}| (1 + r_0)} \right]^{|w_{de}|}.$$

$$r = \frac{\rho_{dm}}{\rho_{de}} = f(a).$$

$$\dot{\rho}_{dm} = \dot{\rho}_{def} + \rho_{def'} \dot{a}$$

$$\dot{\rho}_{de} = \frac{\dot{\rho}_{dm}}{f} - \frac{\rho_{dm} f' \dot{a}}{f^2}, f' = \frac{df}{da}$$

$$Q = \frac{f}{1+f} \left(\frac{f'}{f} a - 3w_{de} \right) H \rho_{de}$$

$$Q = \left(\frac{f'}{f} a - 3w_{de} \right) H \rho_{de} \Omega_{dm}$$

$$Q = (\xi - 3w_{de}) H \rho_{de} \Omega_{dm}$$

$$Q = q(\alpha \dot{\rho} + 3\beta H \rho)$$

$$Q = q(\alpha \dot{\rho}_m + 3\beta H \rho_m),$$

$$Q = q(\alpha \dot{\rho}_{tot} + 3\beta H \rho_{tot}),$$

$$Q = q(\alpha \dot{\rho}_{DE} + 3\beta H \rho_{DE}).$$

$$\dot{\rho}_{\Lambda} = -Q$$

$$H^2 = \frac{\kappa^2}{3} \rho_{tot} = \frac{\kappa^2}{3} (\rho_{\Lambda} + \rho_m)$$

$$\dot{H} = -\frac{\kappa^2}{2} (\rho_{tot} + p_{tot}) = -\frac{\kappa^2}{2} \rho_m$$

$$\dot{\rho}_m = \frac{\beta q - 1}{1 - \alpha q} \cdot 3H\rho_m$$



$$Q=\frac{\beta-\alpha}{1-\alpha q}\cdot 3qH\rho_m$$

$$\rho_m=-\frac{2}{\kappa^2}\dot{H}$$

$$\ddot{H}=\frac{\beta q-1}{1-\alpha q}\cdot 3H\dot{H}$$

$$aH''+\frac{a}{H}H'^2+H'=\frac{\beta q-1}{1-\alpha q}\cdot 3H'$$

$$q=-1-\frac{\dot{H}}{H^2}=-1-\frac{a}{H}H'$$

$$Q=3\beta qH\rho_m$$

$$H(a)=C_{12}\big[3C_{11}(1+\beta)-(2+3\beta)a^{-3(1+\beta)}\big]^{1/(2+3\beta)}$$

$$\Omega_m \equiv \frac{\kappa^2 \rho_m}{3H^2} = -\frac{2\dot{H}}{3H^2} = -\frac{2aH'}{3H}.$$

$$\Omega_m=\frac{2(1+\beta)}{2+3\beta-3C_{11}(1+\beta)a^{3(1+\beta)}}$$

$$\begin{gathered} C_{11}=\frac{\Omega_{m0}(2+3\beta)-2(1+\beta)}{3\Omega_{m0}(1+\beta)} \\ C_{12}=H_0[3C_{11}(1+\beta)-(2+3\beta)]^{-1/(2+3\beta)} \end{gathered}$$

$$E\equiv\frac{H}{H_0}=\Bigl\{1-\frac{2+3\beta}{2(1+\beta)}\Omega_{m0}\bigl[1-(1+z)^{3(1+\beta)}\bigr]\Bigr\}^{1/(2+3\beta)}$$

$$q(z)=-\frac{(1+z)}{E(z)}\frac{d}{dz}\Big(\frac{1}{E(z)}\Big)-1$$

$$q(z)=-1+\frac{3}{2}\Omega_{m0}\frac{(1+z)^{3(1+\beta)}}{E^{(2+3\beta)}}$$

$$w_{\rm eff}\equiv\frac{p_{tot}}{\rho_{tot}}=\frac{(2q-1)}{3}$$

$$Q=\frac{3qH^3}{\kappa^2}\bigg(2\alpha\frac{\dot{H}}{H^2}+3\beta\bigg)$$

$$aH''+\frac{a}{H}H'^2+(4+3\alpha q)H'+\frac{9\beta qH}{2a}=0$$

$$H(a)=C_{22}\cdot a^{-3(2-3\beta+r_1)/8}\cdot\left(a^{3r_1/2}+C_{21}\right)^{1/2}$$

$$\Omega_m=\frac{1}{4}\bigg[2-3\beta+\Big(\frac{2C_{21}}{a^{3r_1/2}+C_{21}}-1\Big)r_1\bigg]$$



$$\mathcal{C}_{21}=-1+\frac{2r_1}{2-3\beta-4\Omega_{m0}+r_1}, \mathcal{C}_{22}=H_0(1+\mathcal{C}_{21})^{-1/2}$$

$$E\equiv\frac{H}{H_0}=(1+z)^{3(2-3\beta+r_1)/8}\cdot\left[\frac{(1+z)^{-3r_1/2}+\mathcal{C}_{21}}{1+\mathcal{C}_{21}}\right]^{1/2}$$

$$Q=\frac{3\beta q H\rho_\Lambda}{1+\alpha q}$$

$$\rho_\Lambda = \frac{3}{\kappa^2}H^2 - \rho_m = \frac{1}{\kappa^2}(3H^2 + 2\dot{H})$$

$$aH''+\frac{a}{H}H'^2+\Big(4+\frac{3\beta q}{1+\alpha q}\Big)H'+\frac{9\beta qH}{2a(1+\alpha q)}=0$$

$$H(a)=\mathcal{C}_{32}\cdot a^{-3(2-5\beta+r_2)/[4(2-3\beta)]}\cdot\left(a^{3r_2/2}+\mathcal{C}_{31}\right)^{1/(2-3\beta)},$$

$$\Omega_m=\frac{1}{2(2-3\beta)}\Big[2-5\beta+\Big(\frac{2\mathcal{C}_{31}}{a^{3r_2/2}+\mathcal{C}_{31}}-1\Big)r_2\Big].$$

$$\mathcal{C}_{31}=-1+\frac{2r_2}{2-5\beta+r_2+2\Omega_{m0}(3\beta-2)}, \mathcal{C}_{32}=H_0(1+\mathcal{C}_{31})^{1/(3\beta-2)}$$

$$E\equiv\frac{H}{H_0}=(1+z)^{3(2-5\beta+r_2)/[4(2-3\beta)]}\cdot\left[\frac{(1+z)^{-3r_2/2}+\mathcal{C}_{31}}{1+\mathcal{C}_{31}}\right]^{1/(2-3\beta)}$$

$$w(z)=\frac{H(z)^2-\frac{2}{3}H(z)H'(z)(1+z)}{H_0^2\Omega_m(1+z)^3-H^2(z)}, H'(z)=\frac{dH}{dz}$$

$$\begin{aligned}\dot{\rho}_{dm}+3H\rho_{dm}&=Q(t)\\\dot{\rho}_{de}+3H(1+w)\rho_{de}&=-Q(t),\\Q(t)&=\gamma H\rho_{dm},\gamma=\text{const}\end{aligned}$$

$$\begin{aligned}\rho_{dm}&=\rho_{dm0}(1+z)^{3-\gamma}\\\rho_{de}&=\left(\rho_{deo}+\rho_{dm0}\frac{\gamma}{\gamma+3w}\right)(1+z)^{3(1+w)}\\H^2&=H_0^2\left[\Omega_{dm0}\left(1-\frac{\gamma}{\gamma+3w}\right)(1+z)^{3-\gamma}+\left(1-\frac{3\Omega_{dm0}w}{\gamma+3w}\right)(1+z)^{3(1+w)}\right]\end{aligned}$$

$$H\rightarrow \bar{H}=-H,\rho\rightarrow \bar{\rho}=\rho,p\rightarrow \bar{p}=-2\rho-p.$$

$$\gamma=\frac{\rho+p}{\rho}\rightarrow\bar{\gamma}=\frac{\bar{\rho}+\bar{p}}{\bar{\rho}}=-\bar{\gamma}.$$

$$\begin{gathered}3H^2=\rho_1+\rho_2\\\dot{\rho}_1+\dot{\rho}_2+3H(\rho_1+\rho_2+p_1+p_2)=0\end{gathered}$$

$$\begin{aligned}\bar{\rho}_1 &= \alpha\rho_1 + (1-\beta)\rho_2 \\ \bar{\rho}_2 &= (1-\alpha)\rho_1 + \beta\rho_2 \\ \bar{H} &= -H\end{aligned}$$

$$\alpha=\frac{\bar{\gamma}_2+\gamma_1}{\bar{\gamma}_1+\bar{\gamma}_2}, \beta=-\frac{\gamma_2+\bar{\gamma}_1}{\bar{\gamma}_1+\bar{\gamma}_2}.$$

$$\begin{aligned}3H^2 &= \rho_1 + \rho_2 \\ \dot{\rho}_1 + 3H\gamma_1\rho_1 &= -3H\Pi, \\ \dot{\rho}_2 + 3H\gamma_2\rho_2 &= 3H\Pi\end{aligned}$$

$$w = \frac{p}{\rho} = \frac{\frac{1}{2}\dot{\phi}^2 - V(\phi)}{\frac{1}{2}\dot{\phi}^2 + V(\phi)}$$

$$\rho = \rho_s + \rho_{dm}, p = p_s + p_{dm}$$

$$\begin{aligned}\dot{\rho}_{dm} + 3H(\rho_{dm} + p_{dm}) &= Q \\ \dot{\rho}_s + 3H(\rho_s + p_s) &= -Q\end{aligned}$$

$$Q \equiv -3H\Pi_{dm} = 3H\Pi_s.$$

$$\begin{aligned}\dot{\rho}_{dm} + 3H(\rho_{dm} + p_{dm} + \Pi_{dm}) &= 0 \\ \dot{\rho}_s + 3H(\rho_s + p_s + \Pi_s) &= 0\end{aligned}$$

$$\dot{r} = r \left(\frac{\dot{\rho}_{dm}}{\rho_{dm}} - \frac{\dot{\rho}_s}{\rho_s} \right)$$

$$\dot{r} = -3Hr \left[\gamma_{dm} - \gamma_s + \frac{1+r}{r} \Pi_{dm} \right]$$

$$\Pi_{dm} = (\gamma_s - \gamma_{dm}) \frac{r}{1+r}$$

$$Q = -3H(\gamma_s - 1) \frac{r}{1+r} \rho_s$$

$$Q = -\sqrt{3}(\gamma_s - 1) \frac{\rho_s \rho_{dm}}{\sqrt{\rho_s + \rho_{dm}}}$$

$$w_\varphi = \frac{p_\varphi}{\rho_\varphi} = \frac{\dot{\varphi}^2 + 2V(\varphi)}{\dot{\varphi}^2 - 2V(\varphi)}$$

$$a(t) = a(t_m) \left[-w_\varphi + (1+w_\varphi) \left(\frac{t}{t_m} \right) \right]^{\frac{2}{3(1+w_\varphi)}}$$

$$\begin{aligned}\dot{\rho}_{dm} + 3H\rho_{dm} &= \delta H\rho_{dm} \\ \dot{\rho}_{de} + 3H(1+w_{de})\rho_{de} &= -\delta H\rho_{dm}\end{aligned}$$

$$\rho_{dm} = \rho_{dm0} a^{-3} e^{\int \delta(a) d\log a}$$



$$r\equiv \frac{\rho_{dm}}{\rho_{de}}=\frac{\rho_{dm0}}{\rho_{deo}}a^{-\xi}=A^{-1}a^{-\xi}, A\equiv \frac{\rho_{deo}}{\rho_{dm0}}=\frac{\Omega_{deo}}{\Omega_{dm0}}$$

$$\rho_{de}=\frac{Aa^{\xi}}{1+Aa^{\xi}}\rho_{tot}, \rho_{dm}=\frac{1}{1+Aa^{\xi}}\rho_{tot}$$

$$\frac{d\rho_{tot}}{da}+\frac{3}{a}\frac{1+(1+w_{de})Aa^{\xi}}{1+Aa^{\xi}}\rho_{tot}=0$$

$$\rho_{tot}=\rho_{tot0}a^{-3}\big[1-\Omega_{deo}(1-a^\xi)\big]^{-3w_{de}/\xi}, \rho_{tot0}=\rho_{deo}+\rho_{dm0}$$

$$H^2=H_0^2a^{-3}\big[1-\Omega_{deo}(1-a^\xi)\big]^{-3w_{de}/\xi}$$

$$\delta=3+\frac{1}{H}\frac{\dot{\rho}_{dm}}{\rho_{dm}}=-\frac{(\xi+3w_{de})Aa^{\xi}}{1+Aa^{\xi}}=-(\xi+3w_{de})\frac{\rho_{de}}{\rho_{tot}}$$

$$\delta=\frac{\delta_0}{\Omega_{deo}+(1-\Omega_{deo})a^{-\xi}}, \delta_0\equiv -\Omega_{deo}(\xi+3w_{de})$$

$$q=-\frac{\ddot{a}}{aH^2}=-1+\frac{\dot{H}}{H^2}=-1+\frac{3}{2}\frac{1-\Omega_{deo}+(1+w_{de})\Omega_{deo}a^\xi}{1-\Omega_{deo}(1-a^\xi)}.$$

$$L=-V(T)\sqrt{1-g_{00}\dot{T}^2}$$

$$\rho_T=\frac{V(T)}{\sqrt{1-\dot{T}^2}}$$

$$p_T=-V(T)\sqrt{1-\dot{T}^2}$$

$$\frac{\ddot{T}}{1-\dot{T}^2}+3H\dot{T}+\frac{1}{V(T)}\frac{dV}{dT}=0.$$

$$\dot{\rho}_T=-3H\dot{T}^2\rho_T$$

$$Q=3H\frac{r}{(r+1)^2}(1-\dot{T}^2)(\rho_T+\rho_{dm})$$

$$p=-\frac{A}{\rho},$$

$$\rho = \sqrt{A + \frac{B}{a^6}},$$

$$L=\frac{1}{2}\dot{\varphi}^2-V(\varphi),$$



$$\begin{aligned}\rho_\varphi &= \frac{1}{2}\dot{\varphi}^2 + V(\varphi) = \rho = \sqrt{A + \frac{B}{a^6}} \\ p_\varphi &= \frac{1}{2}\dot{\varphi}^2 - V(\varphi) = -\frac{A}{\rho} = -\frac{A}{\sqrt{A + \frac{B}{a^6}}}\end{aligned}$$

$$\dot{\varphi}^2 = \frac{B}{a^6\sqrt{A + \frac{B}{a^6}}}, V(\varphi) = \frac{1}{2}\sqrt{A}\left(\cosh 3\varphi + \frac{1}{\cosh 3\varphi}\right),$$

$$p=-\frac{A}{\rho^\alpha}$$

$$\rho=\left[A+\frac{B}{a^{3(1+\alpha)}}\right]^{\frac{1}{1+\alpha}}$$

$$p=\frac{A(a)}{\rho^\alpha}$$

$$\rho=\left[Aa^{-3(1+w_{de})(1+\alpha)}+Ba^{-3(1+\alpha)}\right]^{\frac{1}{1+\alpha}}$$

$$A(a)=-w_{de}Aa^{-3(1+w_{de})(1+\alpha)}$$

$$\rho=\rho_{de}+\rho_{dm}$$

$$\rho_{de}=\frac{p}{w_{de}}=\frac{Aa^{-3(1+w_{de})(1+\alpha)}}{[Aa^{-3(1+w_{de})(1+\alpha)}+Ba^{-3(1+\alpha)}]^{\frac{\alpha}{1+\alpha}}}$$

$$\rho_{dm}=\frac{Ba^{-3(1+\alpha)}}{[Aa^{-3(1+w_{de})(1+\alpha)}+Ba^{-3(1+\alpha)}]^{\frac{\alpha}{1+\alpha}}}$$

$$\frac{\rho_{dm}}{\rho_{de}}=\frac{B}{A}a^{3w_{de}(1+\alpha)}$$

$$\begin{aligned}w_{eff,de} &= w_{de}-\frac{\alpha w_{de}(1-\Omega_{0de})a^{3w}d_{de}(1+\alpha)}{\Omega_{0de}(1-\Omega_{0de})a^{3w}d_{e^{(1+\alpha)}}} \\ w_{eff,dm} &= \frac{\alpha w_{de}\Omega_{0de}}{\Omega_{0de}+(1-\Omega_{0de})a^{3w}d_{e^{(1+\alpha)}}}\end{aligned}$$

$$\rho_{de}(z)=\rho_{de,0}(1+z)^{3(1+w_{de})}+\frac{\delta\rho_{dm,0}}{\delta+3w_{de}}\left[(1+z)^{3(1+w_{de})}-(1+z)^{3-\delta}\right]$$

$$w_{eff,de}=\frac{p_{eff,de}}{\rho_{de}}=\frac{p_{de}+Q/3H}{\rho_{de}}=-1+\Delta,\Delta\equiv\frac{1}{3}\frac{d\ln \rho_{de}}{d\ln (1+z)}.$$



$$\begin{aligned}\frac{d\rho_{de}(z)}{dz} &= 3(1+w_{de})\rho_{de,0}(1+z)^{2+3w_{de}} \\ &+ \frac{\delta\rho_{dm,0}}{\delta+3w_{de}}\left[3(1+w_{de})(1+z)^{2+3w_{de}}-(3-\delta)(1+z)^{2-\delta}\right]\end{aligned}$$

$$\begin{aligned}\dot{x} &= f(x,y,t) \\ \dot{y} &= g(x,y,t)\end{aligned}$$

$$f(x_c,y_c)=g(x_c,y_c)=0$$

$$(x(t),y(t))\rightarrow (x_c,y_c) \,\,\,{\rm for}\,\, t\rightarrow\infty.$$

$$x=x_c+\delta x, y=y_c+\delta y.$$

$$\frac{d}{dN}\binom{\delta x}{\delta y}=\hat{M}\binom{\delta x}{\delta y}$$

$$\hat{M} = \begin{pmatrix} \frac{\partial f}{\partial x} & \frac{\partial f}{\partial y} \\ \frac{\partial g}{\partial x} & \frac{\partial g}{\partial y} \end{pmatrix}$$

$$\begin{aligned}\delta x &= C_1e^{\lambda_1 N} + C_2e^{\lambda_2 N} \\ \delta y &= C_3e^{\lambda_1 N} + C_4e^{\lambda_2 N}\end{aligned}$$

$$\begin{aligned}\rho'_{de}+3(1+w_{de})\rho_{de}&=-Q \\ \rho'_{dm}+3(1+w_{dm})\rho_{dm}&=Q\end{aligned}$$

$$Q_1=3\gamma_m\rho_{dm}, Q_2=3\gamma_d\rho_{de}, Q_3=3\gamma_{tot}\rho_{tot}$$

$$\begin{aligned}Q &= Q_1, w_{eff,de}=w_{de}(\Omega_{de})+\gamma_m\frac{1-\Omega_{de}}{\Omega_{de}}, w_{eff,dm}=w_{dm}-\gamma_m \\ Q &= Q_2, w_{eff,de}=w_{de}(\Omega_{de})+\gamma_d, w_{eff,dm}=w_{dm}-\gamma_d\frac{\Omega_{de}}{1-\Omega_{de}}\end{aligned}$$

$$Q=Q_3, w_{eff,de}=w_{de}(\Omega_{de})+\gamma_m\frac{1}{\Omega_{de}}, w_{eff,dm}=w_{dm}-\frac{\gamma_{tot}}{1-\Omega_{de}}.$$

$$\begin{aligned}\Omega'_{dm} &= 3f_j\Omega_{dm}\Omega_{de} \\ \Omega'_{de} &= -3f_j\Omega_{dm}\Omega_{de}\end{aligned}$$

$$\begin{aligned}f_j &= w_{eff,dej}-w_{eff,dmj} \\ f_1 &= f_0+\frac{\gamma_m}{\Omega_{de}} \\ f_2 &= f_0+\frac{\gamma_d}{1-\Omega_{\Omega_d}} \\ f_3 &= f_0+\frac{\gamma_{tot}}{\Omega_{de}(1-\Omega_{de})}\end{aligned}$$



$$\hat{M} = \begin{pmatrix} 3f(\bar{\Omega}_{de})\bar{\Omega}_{de} & 3(f(\bar{\Omega}_{de})\bar{\Omega}_{dm} + f'\bar{\Omega}_{dm}\bar{\Omega}_{de}) \\ -3f(\bar{\Omega}_{de})\bar{\Omega}_{de} & -3(f(\bar{\Omega}_{de})\bar{\Omega}_{dm} + f'\bar{\Omega}_{dm}\bar{\Omega}_{de}) \end{pmatrix}$$

$$\begin{array}{l} \lambda_1=0 \\ \lambda_2=3f(2\Omega_{de}-1)-3f'\Omega_{de}(1-\Omega_{de}) \end{array}$$

$$\begin{aligned} \lambda^2 + \left[2 + w_{de} - w_{de}(1+w_{de}) \frac{\partial_r \Pi}{\Pi} \right] \lambda + \left(1 + w_{de} + w_{de} \partial_\rho \Pi \right) &= 0 \\ \partial_r \Pi \equiv \frac{\partial \Pi}{\partial r}, \partial_\rho \Pi \equiv \frac{\partial \Pi}{\partial \rho} \end{aligned}$$

$$\lambda_{\pm}=\frac{1}{2}\left\{\begin{aligned}&\left[w_{de}(1+w_{de}) \frac{\partial_r \Pi}{\Pi}-(2+w_{de})\right] \\&\pm \sqrt{\left(2+w_{de}-w_{de}(1+w_{de}) \frac{\partial_r \Pi}{\Pi}\right)^2-4(1+w_{de}+w_{de} \partial_\rho \Pi)}\end{aligned}\right\}$$

$$G^{\mu\nu}=8\pi G T^{\mu\nu}\rightarrow G^{\mu\nu}=8\pi G T^{\mu\nu}+\Lambda g^{\mu\nu}$$

$$u_\mu T^{\mu\nu}_{dm;\nu}=-u_\mu\left(\frac{\Lambda}{8\pi G}g^{\mu\nu}\right)_{;\nu}$$

$$\dot{\rho}_{dm}+3H\rho_{dm}=-\dot{\rho}_\Lambda$$

$$T\frac{dS}{dt}=-\frac{\dot{\Lambda}a^3}{8\pi G}$$

$$\begin{aligned}\dot{\rho}_{dm}+3H\rho_{dm} &= Q \\ \dot{\rho}_\Lambda+3H(\rho_\Lambda+p_\Lambda) &= -Q\end{aligned}$$

$$Q=3H\rho_{dm}\frac{\Lambda'+2\Lambda''\rho_{dm}}{1+3\Lambda'+2\Lambda''\rho_{dm}}$$

$$\begin{aligned}H^2 &= \frac{1}{3}\rho+f(t) \\ \ddot{a} &= -\frac{1}{6}(\rho+3p)+g(t)\end{aligned}$$

$$\dot{\rho}+3H(\rho+p)=6H\left(-f(t)+\frac{\dot{f}(t)}{2H}+g(t)\right)$$

$$\dot{\rho}+3H(\rho+p)=6H\left(-f(t)+\frac{\dot{f}(t)}{2H}+g(t)\right)\rightarrow \dot{\rho}+3H(\rho+p)=0$$

$$\dot{\rho}+3H(\rho+p)=6H\left(-f(t)+\frac{\dot{f}(t)}{2H}+g(t)\right)\rightarrow \dot{\rho}+3H(\rho+p)=\dot{\Lambda}(t)$$

$$S_{\text{EH}}=\int~d^4x\sqrt{-g}\frac{1}{16\pi G}R=\int~d^4x\sqrt{-g}\frac{M_{\text{pl}}^2}{2}R$$



$$S_\phi = -\int~d^4x \sqrt{-g} \left\{ \frac{1}{2} (\partial \phi)^2 + V(\phi) \right\}$$

$$S_m=-\int~d^4x L_m\left(\psi^{(i)}_m,g^{(i)}_{\mu\nu}\right)$$

$$g^{(i)}_{\mu\nu}=e^{\frac{2\beta_i\varphi}{M_{Pl}}}g_{\mu\nu}$$

$$S=\int~d^4x \sqrt{-g} \left\{ \frac{M_{Pl}^2}{2} R + \frac{1}{2} \partial_\mu \varphi \partial^\mu \varphi - V(\varphi) - \frac{1}{\sqrt{-g}} L_m\left(\psi^{(i)}_m,g^{(i)}_{\mu\nu}\right) \right\}$$

$$\nabla^2\varphi=\frac{dV(\varphi)}{d\varphi}+\sum_i\frac{1}{\sqrt{-g}}\frac{\partial L_m\left(\psi^{(i)}_m,g^{(i)}_{\mu\nu}\right)}{\partial g^{(i)}_{\mu\nu}}\frac{2\beta_i}{M_{Pl}}g^{(i)}_{\mu\nu}.$$

$$\frac{1}{\sqrt{-g}}\frac{\partial L_m\left(\psi^{(i)}_m,g^{(i)}_{\mu\nu}\right)}{\partial g^{(i)}_{\mu\nu}}g^{(i)}_{\mu\nu}=\frac{1}{2}\rho_i(1-3w_i)e^{(1-3w_i)\beta_i\varphi/M_{Pl}}.$$

$$\nabla^2\varphi=\frac{dV(\varphi)}{d\varphi}+\sum_i~(1-3w_i)\frac{\beta_i}{M_{Pl}}\rho_ie^{(1-3w_i)\beta_i\varphi/M_{Pl}}.$$

$$\nabla^2\varphi=\frac{dV_{eff}(\varphi)}{d\varphi};~V_{eff}=V(\varphi)+\sum_i~\rho_ie^{(1-3w_i)\beta_i\varphi/M_{Pl}}$$

$$V_{eff}=V(\varphi)+\sum_i~\rho_ie^{\beta_i\varphi/M_{Pl}}$$

$$V_{eff}=V(\varphi)+U(\beta\varphi/M_{Pl})\rho.$$

$$V(\phi)=\frac{M^{4+n}}{\phi^n}$$

$$V(\phi)=M^4\text{exp}\left(\frac{M^n}{\phi^n}\right)$$

$$V(\phi)=M^4\text{exp}\left(\frac{M^n}{\phi^n}\right)$$

$$g_{\mu\nu}=\text{diag}(-1,a^2,a^2,a^2)$$

$$\begin{aligned}\nabla^2\phi &= g^{\mu\nu}\nabla_\mu\nabla_\nu\phi \\&= g^{\mu\nu}\partial_\mu\partial_\nu\phi-g^{\mu\nu}\Gamma^\rho_{\nu\mu}\phi_{,\rho}\\&= g^{00}\partial_0\partial_0\phi-(a^{-2}\Gamma^0_{11}+a^{-2}\Gamma^0_{22}+a^{-2}\Gamma^0_{33})\phi_{,0}\\&= -\ddot{\phi}-a^{-2}(3a\dot{a})\dot{\phi}\\&= -(\ddot{\phi}+3H\dot{\phi})\end{aligned}$$

$$\ddot{\phi}+3H\dot{\phi}=-V_{\text{eff},\phi}(\phi),$$



$$3H^2M_{\rm pl}^2=\frac{1}{2}\dot{\phi}^2+V(\phi)+\rho_{\rm m}e^{\beta\phi/M_{\rm pl}}+\rho_{\rm r}$$

$$\rho_{\rm critical}\equiv \frac{1}{2}\dot{\phi}^2+V(\phi)+\rho_{\rm m}e^{\beta\phi/M_{\rm pl}}+\rho_{\rm r}$$

$$\Omega_{\rm m} \equiv \frac{\rho_{\rm m} e^{\beta \phi_{\rm min}/M_{\rm pl}}}{\rho_{\rm critical}}$$

$$\ddot{x}^\rho + \tilde{\Gamma}_{\mu\nu}^\rho \dot{x}^\mu \dot{x}^\nu = 0$$

$$\tilde{g}_{\mu\nu,\sigma}=\left(\frac{2\beta_i}{M_{\rm pl}}\phi_{,\sigma}g_{\mu\nu}+g_{\mu\nu,\sigma}\right)e^{2\beta_i\phi/M_{\rm pl}}$$

$$\begin{aligned}\tilde{\Gamma}_{\mu\nu}^\rho &= \frac{1}{2}\tilde{g}^{\sigma\rho}(\tilde{g}_{\sigma\nu,\mu}+\tilde{g}_{\sigma\mu,\nu}-\tilde{g}_{\mu\nu,\sigma}) \\&= \frac{1}{2}e^{-2\beta_i\phi/M_{\rm pl}}g^{\sigma\rho}\left(\frac{2\beta_i}{M_{\rm pl}}\phi_{,\mu}g_{\sigma\nu}+g_{\sigma\nu,\mu}+\frac{2\beta_i}{M_{\rm pl}}\phi_{,\nu}g_{\sigma\mu}\right. \\&\quad \left.+g_{\sigma\mu,\nu}-\frac{2\beta_i}{M_{\rm pl}}\phi_{,\sigma}g_{\mu\nu}-g_{\mu\nu,\sigma}\right)e^{2\beta_i\phi/M_{\rm pl}} \\&= \frac{1}{2}g^{\sigma\rho}(g_{\sigma\nu,\mu}+g_{\sigma\mu,\nu}-g_{\mu\nu,\sigma})+\frac{\beta_i}{M_{\rm pl}}g^{\sigma\rho}(\phi_{,\mu}g_{\sigma\nu}+\phi_{,\nu}g_{\sigma\mu}-\phi_{,\sigma}g_{\mu\nu}) \\&= \Gamma_{\mu\nu}^\rho+\frac{\beta_i}{M_{\rm pl}}(\phi_{,\mu}\delta_\nu^\rho+\phi_{,\nu}\delta_\mu^\rho-g^{\sigma\rho}\phi_{,\sigma}g_{\mu\nu})\end{aligned}$$

$$\begin{aligned}0 &= \ddot{x}^\rho + \Gamma_{\mu\nu}^\rho \dot{x}^\mu \dot{x}^\nu + \frac{\beta_i}{M_{\rm pl}}(\phi_{,\mu}\delta_\nu^\rho+\phi_{,\nu}\delta_\mu^\rho-g^{\sigma\rho}\phi_{,\sigma}g_{\mu\nu})\dot{x}^\mu \dot{x}^\nu \\&= \ddot{x}^\rho + \Gamma_{\mu\nu}^\rho \dot{x}^\mu \dot{x}^\nu + \frac{\beta_i}{M_{\rm pl}}(\phi_{,\mu}\dot{x}^\mu \dot{x}^\rho+\phi_{,\nu}\dot{x}^\rho \dot{x}^\nu-g^{\sigma\rho}\phi_{,\sigma}g_{\mu\nu}\dot{x}^\mu \dot{x}^\nu) \\&= \ddot{x}^\rho + \Gamma_{\mu\nu}^\rho \dot{x}^\mu \dot{x}^\nu + \frac{\beta_i}{M_{\rm pl}}(2\phi_{,\mu}\dot{x}^\mu \dot{x}^\rho+g^{\sigma\rho}\phi_{,\sigma})\end{aligned}$$

$$\frac{\vec{F}_\phi}{m}=-\frac{\beta_i}{M_{\rm pl}}\vec{\nabla}\phi$$

$$L_{\text{scalar + matter}} = w \rho f(\phi),$$

$$p=w\rho$$

$$H^2=\varepsilon$$

$$\varepsilon=\frac{\dot{\phi}^2}{2}+V(\phi)+\rho f(\phi)$$

$$\ddot{\phi}+3H\dot{\phi}+V'(\phi)+w\rho f'(\phi)=0,$$

$$\dot{\varepsilon}+3H(\varepsilon+P)=0$$



$$P=w\rho+\frac{\dot{\phi}^2}{2}-V(\phi)$$

$$\begin{aligned}H^2 &= \frac{\dot{\phi}^2}{2} + V + \frac{\rho_0}{f^{1-w}a^{3(1+w)}} \\ \ddot{\phi} + 3H\dot{\phi} + V' + \frac{w\rho_0 f'}{f^{1-w}a^{3(1+w)}} &= 0\end{aligned}$$

$$\begin{aligned}V(\phi) &= \frac{8\cosh^4 \frac{\phi}{2\phi_0}}{3(1+w)} \left(6\alpha^2(1+w) + 3\phi_0^2(1-w) + 4\alpha\tanh \frac{\phi}{2\phi_0} \right), \\ f(\phi) &= \left(-\frac{16\cosh^4 \frac{\phi}{2\phi_0} \exp \left(3\alpha(1+w) \frac{\phi}{\phi_0} \right)}{3Mt_R^2(1+w)} \left(3\phi_0^2 + 2\alpha\tanh \frac{\phi}{2\phi_0} \right) \right)^{\frac{1}{w}},\end{aligned}$$

$$\begin{array}{l}\phi_0 \geq \sqrt{\frac{2\alpha}{3}} \\ w < -1\end{array}$$

$$\phi_0 > \sqrt{\frac{4\alpha - 6\alpha^2(1+w)}{3(1-w)}}$$

$$\sqrt{\frac{2\alpha}{3}} < \phi_0 < \sqrt{\frac{4\alpha - 6\alpha^2(1+w)}{3(1-w)}}$$

$$\phi = 2\phi_0 \operatorname{arctanh} \frac{5\alpha^2(1+w) + 3\phi_0^2(1-w)}{4\alpha}.$$

$$w=\frac{2m+1}{n},$$

$$\left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi}{3\phi}\epsilon_m + \frac{\omega}{6}\left(\frac{\dot{\phi}}{\phi}\right)^2 - \frac{\dot{a}}{a}\frac{\dot{\phi}}{\phi}$$

$$\begin{aligned}\frac{\ddot{a}}{a} &= -\frac{8\pi}{3\phi} \left[\rho_m \left(\frac{3+\omega}{3+2\omega} \right) + 3p_m \left(\frac{\omega}{3+2\omega} \right) \right] - \frac{\omega}{3} \left(\frac{\dot{\phi}}{\phi} \right)^2 + \frac{\dot{a}}{a} \frac{\dot{\phi}}{\phi} + \frac{4\pi}{3+2\omega} S \\ \frac{\ddot{\phi}}{\phi} + 3 \frac{\dot{a}}{a} \frac{\dot{\phi}}{\phi} &= \frac{8\pi}{(3+2\omega)\phi} (\rho_m - 3p_m) - \frac{8\pi}{3+2\omega} S\end{aligned}$$

$$\begin{aligned}\dot{\rho}_m + 3\gamma \frac{\dot{a}}{a} \rho_m &= \frac{1}{2} S \dot{\phi} \\ \dot{\rho}_\phi + 6 \frac{\dot{a}}{a} \rho_\phi &= -\frac{1}{16\pi} R \dot{\phi} - \frac{1}{2} S \dot{\phi}\end{aligned}$$

$$\rho_\phi \equiv \frac{\omega \dot{\phi}^2}{16\pi \phi}$$



$$X = \frac{\dot{a}}{a}, Y = \frac{\dot{\phi}}{\phi} \text{ and } Z = \frac{\rho}{\phi}$$

$$\begin{aligned}\ddot{X} &= \left(Q_3 - \frac{2Q_1Q_4}{Q_2}\right)^{-1} \left[\left(\frac{\omega A}{6} - \frac{\omega}{3} + \frac{Q_4}{Q_2} - \frac{\omega B}{6\alpha^2} - \frac{\omega B Q_4}{6\alpha^2 Q_2} - \frac{\gamma \omega X}{2\alpha Y} \right) Y^2 \right. \\ &\quad + \left(1 - A + \frac{B}{\alpha^2} + \frac{3\gamma X}{2\alpha Y} \right) X^2 \\ &\quad \left. + \left(1 - A - \frac{3Q_4}{Q_2} + \frac{B}{\alpha^2} + \frac{B Q_4}{\alpha^2 Q_1} + \frac{3\gamma X}{\alpha Y} \right) XY \right] \\ \dot{Y} &= \left(1 - 2 \frac{Q_1}{Q_2 Q_3}\right)^{-1} \left\{ \left[\frac{1}{Q_2} \left(1 - \frac{BC}{3} - \frac{\omega B}{6\alpha^2} - \frac{3\omega B X}{6\alpha^2 Y} \right) + \frac{\omega Q_1}{Q_2 Q_3} \left(\frac{A}{3} - \frac{2}{3} - \frac{B}{3\alpha^2} - \frac{\gamma B X}{2\alpha Y} \right) \right] Y^2 \right. \\ &\quad + \left[\frac{1}{Q_2} \left(\frac{2BC}{\alpha} + \frac{B}{\alpha^2} + \frac{3B X}{2\alpha^2 Y} \right) + \frac{Q_1}{Q_2 Q_3} \left(2 - 2A + \frac{2B}{\alpha^2} + \frac{3(\gamma+1)BX}{\alpha^2 Y} \right) \right] X^2 \\ &\quad \left. + \left[\frac{1}{Q_2} \left(-3 + \frac{2BC}{\alpha} + \frac{B}{\alpha^2} + \frac{3B X}{2\alpha^2 Y} \right) + \frac{Q_1}{Q_2 Q_3} \left(-2A + \frac{2B}{\alpha^2} + \frac{3\gamma BX}{\alpha^2 Y} \right) \right] XY \right\}\end{aligned}$$

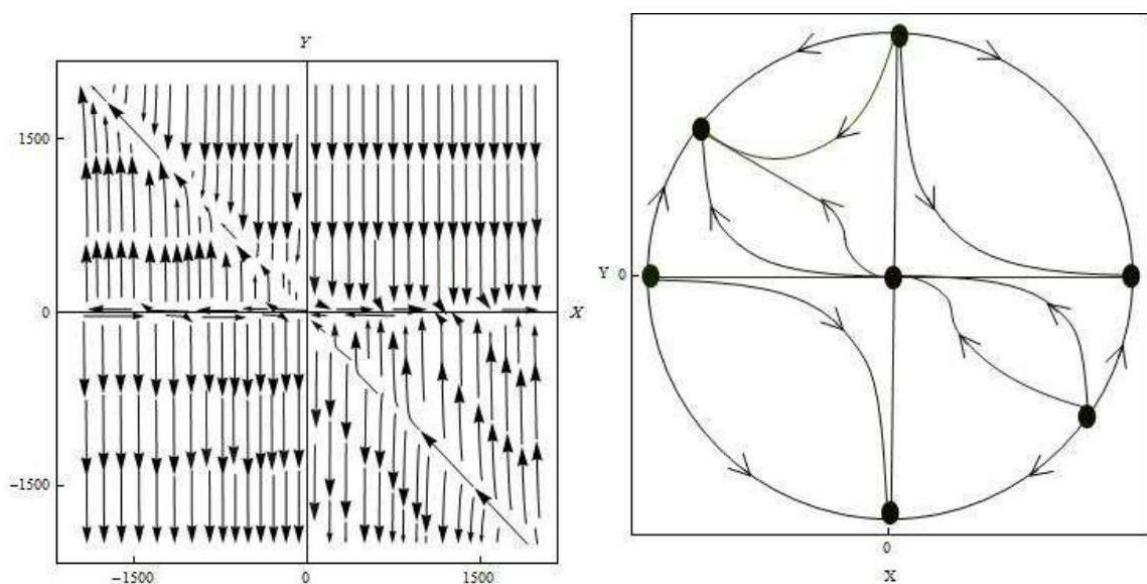
$$\alpha = \frac{8\pi}{3}, A = \frac{3 - 2\omega + 3\gamma}{3 + 2\omega}, B = \frac{4\pi}{3 + 2\omega} \text{ and } C = 2 - 3\gamma$$

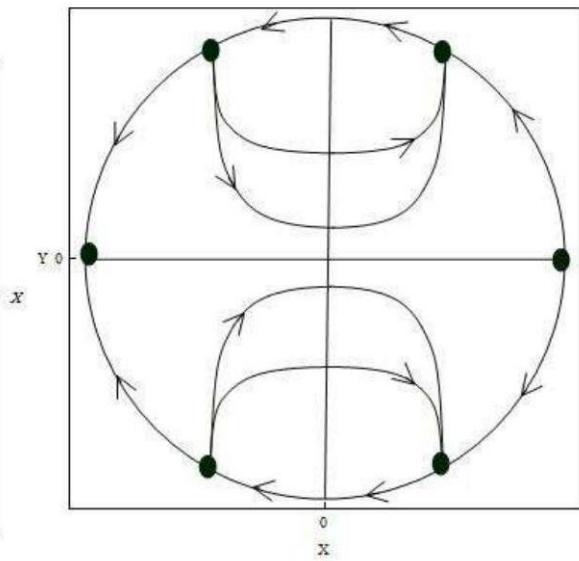
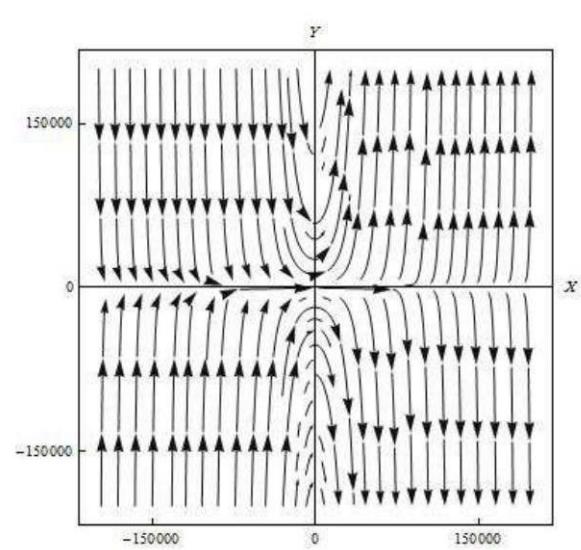
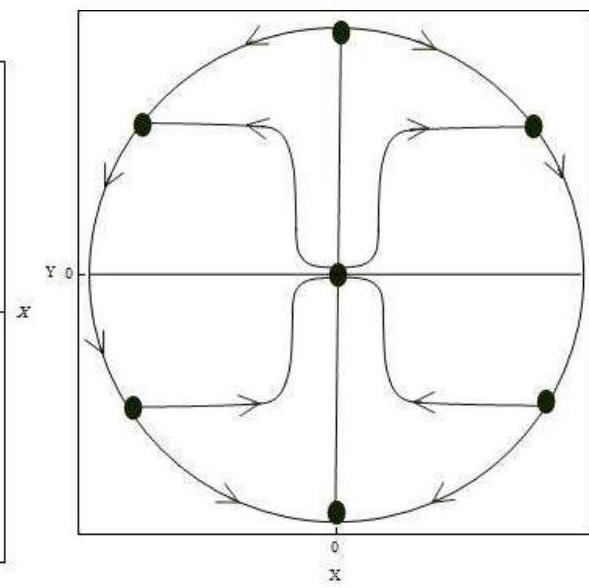
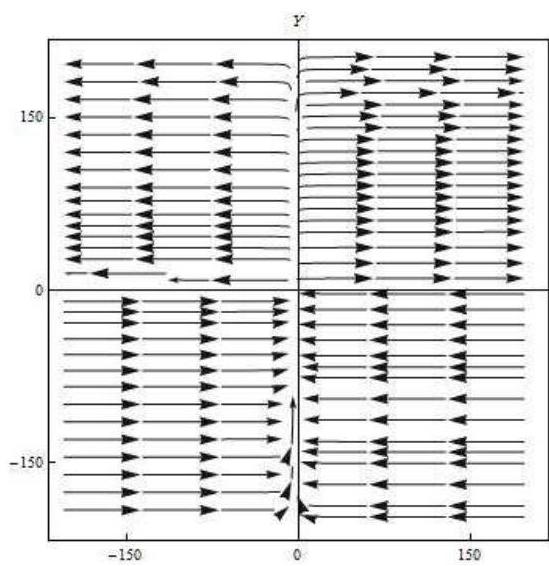
$$Q_1 = \frac{B}{\alpha} \left(1 - 2 \frac{X}{Y} \right)$$

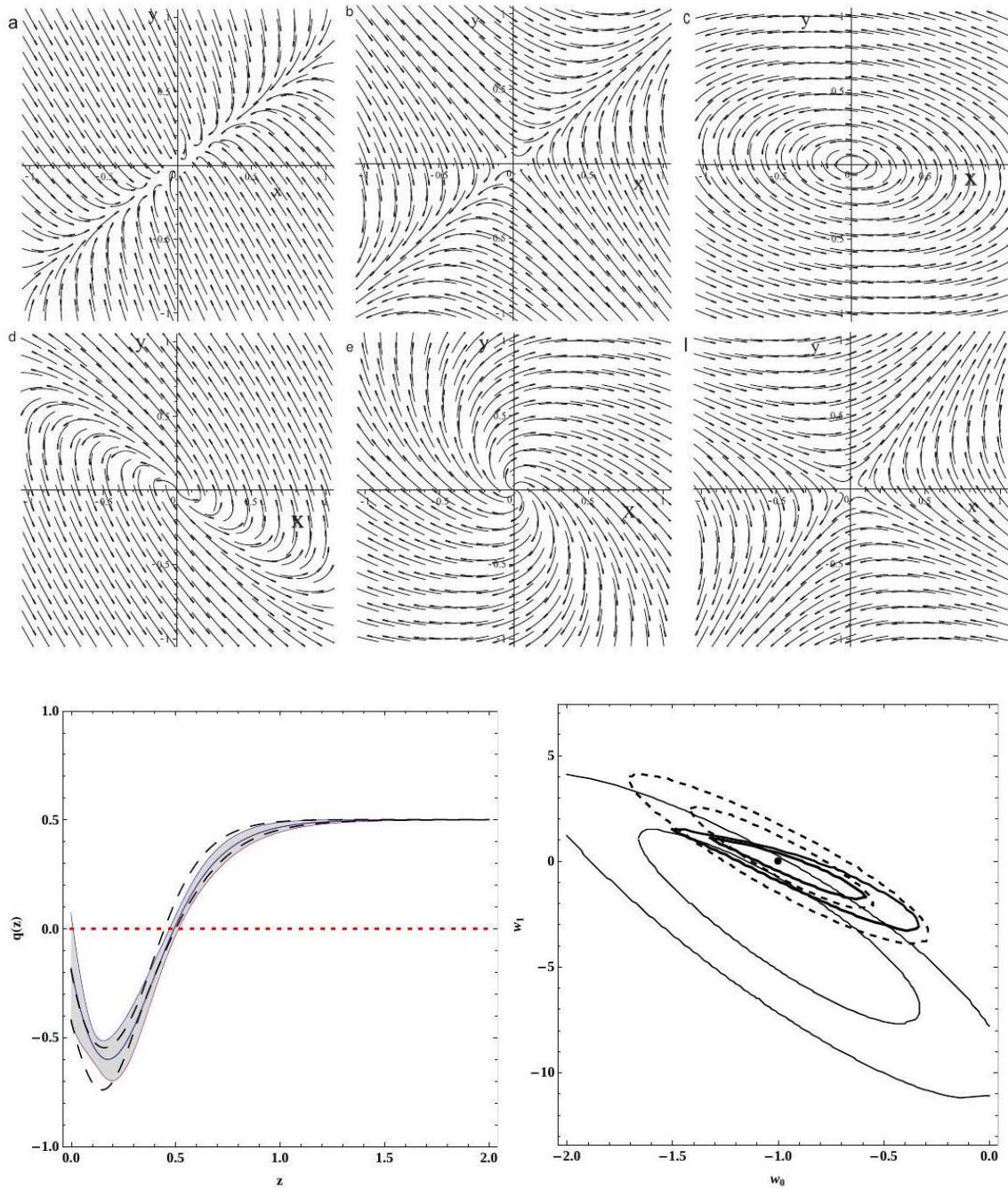
$$Q_2 = 1 + \frac{2B\omega}{3\alpha} - \frac{2BX}{\alpha Y}$$

$$Q_3 = 1 - \frac{B}{\alpha} - \frac{2BX}{\alpha Y}$$

$$Q_4 = \frac{B}{\alpha} \left(\frac{X}{Y} - \frac{\omega}{3} \right).$$







Figuras 6, 7, 8, 9 y 10. Deformación del espacio – tiempo cuántico, por interacción de la partícula oscura.

$$S = \frac{1}{2\kappa} \int d^4x \sqrt{-g} f(R) + S_m(g^{\mu\nu}, \psi), S_m = \int d^4x \sqrt{-g} L_m(g^{\mu\nu}, \psi)$$

$$f'R_{\mu\nu} - \frac{1}{2} f g_{\mu\nu} - \nabla_\mu \nabla_\nu f' + g_{\mu\nu} \square f'' = k T_{\mu\nu}^{(m)}, f' \equiv \frac{df}{dR}$$

$$T_{\mu\nu}^{(m)} = -\frac{2}{\sqrt{-g}} \frac{\delta(\sqrt{-g}L_m)}{\delta(g^{\mu\nu})}$$

$$\begin{gathered}\phi\equiv f'(R)\\ U(\phi)\equiv R(\phi)f'-f[R(\phi)]\end{gathered}$$

$$S=\frac{1}{2\kappa}\int\,\,d^4x\sqrt{-g}(\phi R-U(\phi)+L_m)$$

$$g_{\mu\nu}\rightarrow \tilde g_{\mu\nu}=f'g_{\mu\nu}$$

$$d\tilde{\phi} = \sqrt{\frac{3}{2k}}\frac{d\phi}{\phi}.$$

$$S=\int\,\,d^4x\sqrt{-g}\Biggl[\frac{\tilde{R}}{2\kappa}-\frac{1}{2}\partial^\mu\tilde{\phi}\partial_\mu\tilde{\phi}-V(\tilde{\phi})\Biggr]+S_m\bigl(e^{-2\beta\tilde{\phi}}\tilde{g}_{\mu\nu},\psi\bigr)$$

$$V(\tilde{\phi})=\frac{Rf'-f}{2\kappa f'^2}$$

$$\begin{gathered}\tilde{G}_{\mu\nu}=\kappa\left(\tilde{T}_{\mu\nu}^{\tilde{\phi}}+\tilde{T}_{\mu\nu}^m\right).\\\Box\,\tilde{\phi}-\frac{dV(\tilde{\phi})}{d\tilde{\phi}}=-\beta\sqrt{\kappa}\tilde{T}^m,\end{gathered}$$

$$\ddot{\phi}+3H\dot{\phi}+\frac{dV}{d\phi}=-\beta\sqrt{\kappa}\rho_m$$

$$\dot{\rho}_\phi + 3H(1+w_\phi)\rho_\phi = -Q, Q = \beta\sqrt{\kappa}\dot{\phi}\rho_m.$$

$$d\tilde{t}=\sqrt{F}dt, \tilde{a}=\sqrt{F}a, \tilde{H}=\frac{1}{\tilde{a}}\frac{d\tilde{a}}{d\tilde{t}}=\frac{1}{\sqrt{F}}\bigg(H+\frac{\dot{F}}{2F}\bigg), F=e^{-2\beta\sqrt{\kappa}\phi}$$

$$\dot{\rho}_m+3H\rho_m=Q$$

$$(2q-1)H=-\beta\sqrt{k}\dot{\phi}$$

$$\frac{f''}{f'}\dot{R}=-2(2q-1)H$$

$$I=\frac{1}{16\pi G}\int\,\,d^4x\sqrt{-g}[f(T)+L_m]$$

$$H^2+\frac{k}{a^2}=\frac{1}{3}(\rho_m+\rho_T)$$

$$\dot{H}-\frac{k}{a^2}=-\frac{1}{2}(\rho_m+\rho_T+p_T),$$

$$\begin{gathered}\rho_T=\frac{1}{2}(2Tf'-f-T),\\ p_T=-\frac{1}{2}[-8\dot{H}Tf''+(2T-4\dot{H})f'-f+4\dot{H}-T],\end{gathered}$$



$$T=-6\left(H^2+\frac{k}{a^2}\right)$$

$$\rho_m=\frac{1}{2}[f-2Tf'].$$

$$\dot{\rho_m}+3H\rho_m=Q,$$

$$\dot{\rho_T}+3H(\rho_T+p_T)=-Q,$$

$$\dot{\rho_T}+3H\rho_T(1+w_{eff})=0.$$

$$w_{eff}=w_T+\frac{Q}{3H\rho_T}.$$

$$w_{eff}=-1+\left(\frac{4k}{a^2}-\frac{\dot{T}}{3H}\right)\left(\frac{2Tf''+f'-1}{2Tf'-f-T}\right)+\frac{Q}{3H\rho_T}.$$

$$\dot{\rho_T}=\frac{\dot{T}}{2}[f'+2Tf''-1],$$

$$w_T=-\left[1+\frac{Q}{3H\rho_T}+\frac{\dot{T}}{3H}\frac{(2Tf''+f'-1)}{(2Tf'-f-T)}\right].$$

$$\dot{T}=\frac{12H}{(f'+2Tf'')}\Bigg[\frac{(f-2Tf')}{4}+\frac{k}{a^2}(f'+2Tf''-1)\Bigg].$$

$$w_T=-\left[1+\frac{Q}{3H\rho_T}+\frac{4}{(f'+2Tf'')} \frac{(2Tf''+f'-1)}{(2Tf'-f-T)}\left(\frac{(f-2Tf')}{4}+\frac{k}{a^2}(f'+2Tf''-1)\right)\right]$$

$$q=\frac{1}{2}-\frac{k}{2a^2}\biggl[\frac{T}{6}+\frac{k}{a^2}\biggr]^{-1}+\biggl[\frac{T}{6}+\frac{k}{a^2}\biggr]^{-1}\biggl[\frac{(2Tf'-f-T)}{4}+\frac{Q}{6H}+\frac{(2Tf''+f'-1)}{(f'+2Tf'')} \times \\ \biggl(\frac{(f-2Tf')}{4}+\frac{k}{a^2}(f'+2Tf''-1)\biggr)\biggr]$$

$$S=S_g+S_\mathrm{m}$$

$$S_g = \frac{M_p^2}{2} \int \;\; {\rm d}\varrho(x) \sqrt{-g} \big(R - 2\lambda - \omega \partial_\mu v \partial^\mu v \big)$$

$$S_\mathrm{m} = \int \;\; {\rm d}\varrho \sqrt{-g} \mathcal{L}_\mathrm{m}$$

$$\left(\frac{D}{2}-1\right)H^2+H\frac{\dot{v}}{v}-\frac{1}{2}\frac{\omega}{D-1}\dot{v}^2=\frac{1}{M_p^2(D-1)}\rho+\frac{\lambda}{D-1}-\frac{k}{a^2}\\ \frac{\Box\, v}{v}-(D-2)\left(H^2+\dot{H}-H\frac{\dot{v}}{v}+\frac{\omega}{D-1}\dot{v}^2\right)+\frac{2\lambda}{D-1}=\frac{1}{M_p^2(D-1)}\left[(D-3)\rho+(D-1)p\right]$$

$$\dot{H}+(D-1)H^2+\frac{2k}{a^2}+\frac{\Box\, v}{v}+H\frac{\dot{v}}{v}+\omega(v\Box\, v-\dot{v}^2)=0$$



$$\dot{\rho}+\left[(D-1)H+\frac{\dot{v}}{v}\right](\rho+p)=0$$

$$H^2=\frac{1}{3M_p^2}\rho+\frac{\lambda}{3}-\frac{k}{a^2}\\ H^2+\dot{H}=-\frac{1}{6M_p^2}(3p+\rho)+\frac{\lambda}{3}$$

$$\nu=t^{-\beta}$$

$$\begin{array}{l}\dot{\rho}_m+(3H-\beta t^{-1})\rho_m=Q\\\dot{\rho}_x+(1+w)(3H-\beta t^{-1})\rho_x=-Q\end{array}$$

$$\Omega_m=\frac{\rho_m}{3M_p^2H^2}, \Omega_x=\frac{\rho_x}{3M_p^2H^2}$$

$$\begin{array}{l}\dot{\Omega}_m+(3H-\beta t^{-1})\Omega_m+2\Omega_m\frac{\dot{H}}{H}=\frac{Q}{3M_p^2H^2},\\\dot{\Omega}_x+(1+w_x)(3H-\beta t^{-1})\Omega_x+2\Omega_x\frac{\dot{H}}{H}=-\frac{Q}{3M_p^2H^2},\end{array}$$

$$\dot{H}+H^2-\frac{\beta H}{2t}+\frac{\beta(\beta+1)}{2t^2}+\frac{\omega\beta^2}{3t^{2(\beta+1)}}=-\frac{1}{2}((1+3w)\Omega_x+\Omega_m)H^2.$$

$$\Omega_f=\frac{\omega\dot{v}^2}{6H^2}-\frac{\dot{v}}{H\nu}.$$

$$\Omega_f=\frac{\omega\beta^2}{6H^2t^{2(\beta+1)}}+\frac{\beta}{Ht}$$

$$\sum_{\alpha=k,f,x,m}\Omega_\alpha\equiv 1.$$

$$Q\equiv H(\delta\rho_x+\gamma\rho_m)$$

$$\begin{array}{l}\dot{\Omega}_m+(3H-\beta t^{-1})\Omega_m+2\Omega_m\frac{\dot{H}}{H}=H(\delta\Omega_x+\gamma\Omega_m),\\\dot{\Omega}_x+(1+w_x)(3H-\beta t^{-1})\Omega_x+2\Omega_x\frac{\dot{H}}{H}=-H(\delta\Omega_x+\gamma\Omega_m),\\\dot{\Omega}_f+\left(\frac{\dot{H}}{H}+2(1+\beta)t^{-1}\right)\Omega_f-\frac{(1+2\beta)\beta}{Ht}=0.\end{array}$$

$$\begin{array}{l}\Omega'_m=\theta\Omega_m+\sigma\delta\Omega_x\\\Omega'_x=-\delta\gamma\Omega_m+\nu\Omega_x\end{array}$$

$$\tau^2-(\theta+\nu)\tau+\delta^2\sigma\gamma+\theta\nu=0$$

$$\tau_{\pm}=\frac{\theta+\nu}{2}\Bigg[1\pm\sqrt{1-4\frac{(\delta^2\sigma\gamma+\theta\nu)}{(\theta+\nu)^2}}\Bigg]$$



$$\tau_{\pm}\in \Re, \tau_{\pm}<0, \tau_+>\tau_->0, \theta+\nu<0, 4(\delta^2\sigma\gamma+\theta\nu)<(\theta+\nu)^2, \delta^2\sigma\gamma+\theta\nu>0.$$

$$\tau_{\pm}\in \Re, \tau_{\pm}>0, \tau_+>\tau_->0, \theta+\nu>0, 4(\delta^2\sigma\gamma+\theta\nu)<(\theta+\nu)^2, \delta^2\sigma\gamma+\theta\nu>0.$$

$$\tau_{\pm}\in \Re, \tau_+\tau_-<0, \delta^2\sigma\gamma+\theta\nu<0.$$

$$\tau_{\pm}\in \mathbb{C}, \tau_{\pm}=\tau_1\pm i\tau_2, \tau_1,\tau_2\in \Re \tau_1, \tau_2>0, \theta+\nu<0, (\theta+\nu)^2<4(\delta^2\sigma\gamma+\theta\nu).$$

$$\tau_{\pm}\in \mathbb{C}, \tau_{\pm}=\tau_1\pm i\tau_2, \tau_1,\tau_2\in \Re \tau_1, \tau_2\langle 0, \theta+\nu\rangle 0, (\theta+\nu)^2<4(\delta^2\sigma\gamma+\theta\nu).$$

$$\tau_{\pm}\in \Im, \tau_{\pm}=\pm i\tau, \tau\in \Re, \theta=\nu, \delta^2\sigma\gamma+\theta\nu>0.$$

$$\begin{array}{l}\gamma=-2,\sigma=3,\beta=-1,\delta=3,\gamma=-3,\sigma=1,\beta=2,\delta=3,\gamma=3,\sigma=3,\beta=-1,\delta=1,\gamma=\\3,\sigma=1,\beta=2,\delta=3,\gamma=3,\sigma=-3,\beta=2,\delta=3,\gamma=3,\sigma=-3,\beta=1,\delta=-2.\end{array}$$

$$L^3\Lambda^3\leq S_{BH}\equiv\frac{1}{4}\frac{A_{BH}}{l_p^2}=\pi L^2M_p^2$$

$$L\sim \Lambda^{-3} M_p^2$$

$$\rho_{vac}\approx \frac{\Lambda^4}{16\pi^2}$$

$$L^3\rho_\Lambda\leq LM_{Pl}^2\equiv 2M_{BH}$$

$$L^3\rho_\Lambda\leq M_{BH}\sim LM_{Pl}^2$$

$$\rho_\Lambda\sim L^{-2}M_{Pl}^2\sim H^2M_{Pl}^2$$

$$M_{Pl}\simeq 1.2\times 10^{19}{\rm GeV};\;H_0\simeq 1.6\times 10^{-42}{\rm GeV}$$

$$\rho_L = 3c^2 M_p^2 L^{-2}$$

$$\begin{array}{l}\dot{\rho}_{dm}+3H\rho_{dm}=Q\\\dot{\rho}_L+3H(\rho_L+p_L)=-Q\end{array}$$

$$\Omega_L=\frac{8\pi\rho_L}{3M_{Pl}^2H^2}, \Omega_m=\frac{8\pi\rho_m}{3M_{Pl}^2H^2}, \Omega_k=\frac{k}{H^2a^2}$$

$$H^2=\frac{8\pi G}{3}(\rho_L+\rho_m)-\frac{k}{a^2},$$

$$\Omega_L+\Omega_m=1+\Omega_k$$

$$r=\frac{1-\Omega_L+\Omega_k}{\Omega_L}$$

$$\dot{r}=3Hr\left[w_L-w_m+\frac{1+r}{r}\frac{\Gamma}{3H}\right]=3Hr\big[w_L^\mathrm{eff}-w_m^\mathrm{eff}\big].$$

$$\frac{1}{H}\frac{dH}{dx}=-\frac{3}{2}-\frac{1}{2}\Omega_k-\frac{3}{2}w_L\Omega_L$$



$$\frac{d\Omega_L}{dx} = 3\Omega_L \left[\frac{1}{3}\Omega_k + w_L(\Omega_L - 1) - \frac{\Gamma}{3H} \right]$$

$$\frac{d\Omega_k}{dx} = \Omega_k(1 + \Omega_k + 3w_L\Omega_L)$$

$$\Gamma=3H(-1-w_L)+2\frac{\dot{L}}{L}$$

$$\frac{d\Omega_L}{dx}=-3\Omega_L(1-\Omega_L)\big(w_L^{\rm eff}-w_m^{\rm eff}\big)+\Omega_k\Omega_L\big(1+3w_m^{\rm eff}\big)$$

$$\frac{d\Omega_k}{dx}=3\Omega_k\Omega_L\big(w_L^{\rm eff}-w_m^{\rm eff}\big)+\Omega_k(1+\Omega_k)\big(1+3w_m^{\rm eff}\big)$$

$$Q=3\alpha H\rho$$

$$Q=3\alpha H\rho_L$$

$$\dot{\rho}_L=-2\rho_L\frac{\dot{L}}{L},$$

$$w_L\equiv\frac{p_L}{\rho_L}=\frac{2}{3}\frac{\dot{L}}{LH}-\alpha-1$$

$$\rho_H=3c^2M_p^2H^2$$

$$w_H=-\frac{2}{3}\frac{\dot{H}}{H^2}-\alpha-1$$

$$H=\frac{2(1-\alpha c^2)}{3-2\alpha c^2}\frac{A}{t}$$

$$a(t)=a_0 t^{\frac{2(1-\alpha c^2)}{3-2\alpha c^2}}$$

$$q(t)=\frac{1}{2(1-2\alpha c^2)}$$

$$\frac{Q}{3H\rho_m}=\mu\left(\frac{H}{H_0}\right)^{-n}$$

$$\frac{H}{H_0}=\left(\frac{1}{3}\right)^{1/n}\left[1-2q_0+2(1+q_0)a^{-3n/2}\right]^{1/n},$$

$$\mu=\frac{1}{3}(1-2q_0)$$

$$L_f=a(t)\int_t^\infty\frac{1}{a(t')}dt'$$

$$\dot{L_f}=HL_f-1$$



$$w=-\frac{1}{3}-\frac{2}{3c}\sqrt{\Omega_f}-\alpha$$

$$q = - \frac{\ddot{a}}{aH^2} = \frac{1}{2} + \frac{3}{2} w \Omega_f = \frac{1}{2} - (1+3\alpha) \Omega_f - \frac{1}{c} \Omega_f^{\frac{3}{2}}.$$

$$\rho_{\mathcal{R}}=3\alpha M_p^2\left(\dot{H}+2H^2+\frac{k}{a^2}\right)$$

$$\mathcal{R}=-6\left(\dot{H}+2H^2+\frac{k}{a^2}\right)$$

$$H^2=\frac{1}{3M_p^2}\big(\rho_{\mathcal{R}}+\rho_m+\rho_\gamma+\rho_k\big)$$

$$Q=\gamma H\rho_{\mathcal{R}}$$

$$\begin{aligned}\frac{\alpha}{2}\frac{d^2H^2}{dx^2}-\Big(1-\frac{7\alpha}{2}-\frac{\alpha\gamma}{2}\Big)\frac{dH^2}{dx}-(3-6\alpha-2\alpha\gamma)H^2\\-\frac{\rho_{\gamma0}}{3M_p^2}e^{-4x}-\{1-\alpha(1+\gamma)\}ke^{-2x}=0\end{aligned}$$

$$\frac{H^2}{H_0^2}=A_+ e^{\sigma_+ x}+A_- e^{\sigma_- x}+A_\gamma e^{-4x}+A_k e^{-2x}$$

$$\sigma_{\pm}=\frac{2-7\alpha-\alpha\gamma\pm\sqrt{(2-\alpha)^2-2\alpha(\alpha+2)\gamma+\alpha^2\gamma^2}}{2\alpha},$$

$$\begin{gathered} A_\gamma=\Omega_{\gamma0}, \\ A_k=\Omega_{k0}, \\ A_\pm=\pm\frac{\alpha(\sigma_{\mp}+3)\Omega_{k0}+2\Omega_{\Lambda0}-\alpha(1-\Omega_{\gamma0})(\sigma_{\mp}+4)}{\alpha(\sigma_+-\sigma_-)}. \end{gathered}$$

$$\rho_{\mathcal{R}}=\rho_{c0}\sum_{i=+,-}\alpha\left(\frac{\sigma_i}{2}+2\right)A_ie^{\sigma_ix}.$$

$$\rho_m=\rho_{c0}\sum_{i=+,-}\left\{1-\alpha\left(\frac{\sigma_i}{2}+2\right)\right\}A_ie^{\sigma_ix}$$

$$w_{\mathcal{R}}=-1-\frac{1}{3}\Big(\gamma+\frac{1}{\rho_{\mathcal{R}}}\frac{d\rho_{\mathcal{R}}}{dx}\Big)$$

$$3H^2=\rho=\rho_c+\rho_x$$

$$\rho'=d\rho/d\eta=-\rho_c-(1+\omega_x)\rho_x$$

$$\rho'=-\alpha\rho_c-\beta\rho_x$$

$$\omega_x=(\alpha-1)r+\beta-1$$



$$\rho_c=-(\beta\rho+\rho')/\Delta,\rho_x=(\alpha\rho+\rho')/\Delta$$

$$\rho''+(\alpha+\beta)\rho'+\alpha\beta\rho=Q\Delta.$$

$$Q=c_1\frac{(\gamma_s-\alpha)(\gamma_s-\beta)}{\Delta}\rho+c_2(\gamma_s-\alpha)\rho_c\\-c_3(\gamma_s-\beta)\rho_x-c_4\frac{(\gamma_s-\alpha)(\gamma_s-\beta)}{\gamma_s\Delta}\rho'$$

$$Q=\frac{u\rho+\gamma_s^{-1}[u-(\gamma_s-\alpha)(\gamma_s-\beta)]\rho'}{\Delta}$$

$$\rho''+(\gamma_s+\gamma^+)\rho'+\gamma_s\gamma^+\rho=0$$

$$\rho_c=\frac{(\gamma_s-\beta)b_1a^{-3\gamma_s}+(1-\beta)b_2a^{-3}}{\Delta}\\ \rho_x=\frac{(\alpha-\gamma_s)b_1a^{-3\gamma_s}+(\alpha-1)b_2a^{-3}}{\Delta}$$

$$Q=\frac{(\alpha\beta-1)}{\Delta\gamma}\rho+\frac{(\alpha+\beta-\nu-2)}{\Delta\gamma}\rho'-\frac{\nu\rho'^2}{\rho\Delta\gamma},$$

$$\rho = \left[\rho_{10}a^{-3} + \rho_{20}a^{-3(1+\nu)}\right]^{1/(1+\nu)}$$

$$\rho_c=\frac{-\rho}{\alpha-\beta}\Big[\beta-1+\frac{\nu}{(1+\nu)(1+\rho_{20}a^{-3\nu}/\rho_{10})}\Big]\\ \rho_x=\frac{\rho}{\alpha-\beta}\Big[\alpha-1+\frac{\nu}{(1+\nu)(1+\rho_{20}a^{-3\nu}/\rho_{10})}\Big]\\ p=-\frac{\nu\rho_{10}}{1+\nu}\frac{a^{-3}}{\rho^\nu}$$

$$\omega=-\frac{\nu\rho_{10}}{(1+\nu)(\rho_{10}+\rho_{20}a^{-3\nu})}$$

$$\delta t=\beta t_p^{2/3}t^{1/3}$$

$$E_{\delta t^3}\sim t^{-1}$$

$$\rho_q\sim\frac{E_{\delta t^3}}{\delta t^3}\sim\frac{1}{t_p^2t^2}$$

$$\rho_q=\frac{3n^2M_p^2}{T^2}$$

$$T=\int_0^a\frac{da'}{Ha'}$$

$$\Omega_q=\frac{n^2}{H^2T^2}$$



$$\Omega'_q=\Omega_q\left(-2\frac{\dot{H}}{H^2}-\frac{2}{n}\sqrt{\Omega_q}\right).$$

$$-\frac{\dot{H}}{H^2}=\frac{3}{2}\big(1-\Omega_q\big)+\frac{\Omega_q^{3/2}}{n}-\frac{Q}{6M_p^2H^3}.$$

$$\Omega'_q=\Omega_q\left[\big(1-\Omega_q\big)\Big(3-\frac{2}{n}\sqrt{\Omega_q}\Big)-\frac{Q}{3M_p^2H^3}\right],$$

$$\frac{Q}{3M_p^2H^3}=\begin{cases} 3\alpha \Omega_q & \text{for } Q=3\alpha H\rho_q \\ 3\beta\big(1-\Omega_q\big) & \text{for } Q=3\beta H\rho_m \\ 3\gamma & \text{for } Q=3\gamma H\rho_{tot} \end{cases}$$

$$w_q=-1+\frac{2}{3n}\sqrt{\Omega_q}-\frac{Q}{3H\rho_q},$$

$$\frac{Q}{3H\rho_q}=\begin{cases} \alpha & \text{for } Q=3\alpha H\rho_q \\ \beta\big(\Omega_q^{-1}-1\big) & \text{for } Q=3\beta H\rho_m \\ \gamma\Omega_q^{-1} & \text{for } Q=3\gamma H\rho_{tot} \end{cases}$$

$$q\equiv -\frac{\ddot{a}a}{\dot{a}^2}=-1-\frac{\dot{H}}{H^2}=\frac{1}{2}-\frac{3}{2}\Omega_q+\frac{\Omega_q^{3/2}}{n}-\frac{Q}{6M_p^2H^3}.$$

$$\begin{aligned}\dot{\rho}_{dm}+3H\rho_{dm}&=\frac{\dot{f}}{f}\rho_{dm}\\\dot{\rho}_{de}+3H(1+w)\rho_{de}&=-\frac{\dot{f}}{f}\rho_{dm}\end{aligned}$$

$$\begin{gathered}3H^2=8\pi G\rho,\\\frac{\dot{H}}{H^2}=-\frac{3}{2}\Big(1+\frac{p}{\rho}\Big),\end{gathered}$$

$$\rho_{dm}=\rho_{dm,0}\left(\frac{a_0}{a}\right)^3\frac{f}{f_0}$$

$$\dot{\rho}_{de}+3H\big(1+w_{eff}\big)\rho_{de}=0, w_{eff}\equiv w+\frac{\dot{f}}{3Hf}r$$

$$w_{eff}=-\frac{\dot{f}}{3Hf}, w=(1+r)w_{eff}$$

$$\frac{p}{\rho}=\frac{p_{de}}{\rho_{de}+\rho_{dm}}=\frac{w}{1+r}=w_{eff}$$

$$q=\frac{1}{2}\Big(\frac{3p}{\rho}+1\Big)$$

$$q=\frac{1}{2}\Bigg(1-\frac{\dot{f}}{Hf}\Bigg)$$



$$\dot{r}=3Hr\left[w_{de}+\frac{Q}{9H^3}\frac{(r+1)^2}{r}\right].$$

$$\dot{r}=\dot{H}\frac{dr}{dH}, \dot{H}=-\frac{1}{2}(\rho_{dm}+\rho_{de}+p_{de})=-\frac{3}{2}\frac{1+w_{de}+r}{1+r}H^2$$

$$\frac{dr}{dH}=\frac{I}{H}, I\equiv -2r\frac{1+r}{1+w_{de}+r}\Bigg[w_{de}+\frac{Q}{9H^3}\frac{(r+1)^2}{r}\Bigg]$$

$$Q=3\alpha H(\rho_{dm}+\rho_{de}), Q=3\beta H\rho_{dm}, Q=3\gamma H\rho_{de},$$

$$r_s^\pm = -1 + 2b \pm 2\sqrt{b(b-1)}, b = -\frac{w}{4\alpha} > 1$$

$$r(x)=\frac{r_s^-+xr_s^+}{1+x}$$

$$w(z) = w_0 + \frac{w_a z}{1+z}$$

$$w(z)=-\frac{1+\tanh\left[(z-z_t)\Delta\right]}{2}$$

$$\dot{\rho}_m+3\frac{\dot{a}}{a}\rho_m=-\dot{\rho}_{\Lambda}$$

$$\rho_m=\rho_{m,0}a^{-3+\epsilon(a)}$$

$$\rho_\Lambda=\rho_{m0}\int_a^1\frac{\epsilon(\tilde{a})+\tilde{a}\epsilon'\mathrm{ln}\,(\tilde{a})}{\tilde{a}^{4-\epsilon(a)}}d\tilde{a}+\mathrm{X}$$

$$H=H_0\big[\Omega_{b,0}a^{-3}+\Omega_{m0}\varphi(a)+\Omega_{\mathrm{X},0}\big]^{1/2}$$

$$\varphi(a)=a^{-3+\epsilon(a)}+\int_a^1\frac{\epsilon(\tilde{a})+\tilde{a}\epsilon'\mathrm{ln}\,(\tilde{a})}{\tilde{a}^{4-\epsilon(a)}}d\tilde{a}$$

$$\epsilon(a)=\epsilon_0a^\xi=\epsilon_0(1+z)^{-\xi}$$

$$\rho_\Lambda=\rho_{m0}\epsilon_0\int_a^1\frac{\left[1+\mathrm{ln}\,(\tilde{a}^\xi)\right]}{\tilde{a}^{4-\xi-\epsilon_0}\tilde{a}^\xi}d\tilde{a}+\mathrm{X}$$

$$\begin{aligned}\Omega_b(a)&=\frac{a^{-3}}{A+a^{-3}+\mathrm{B}^{-1}\varphi(a)}\\ \Omega_m(a)&=\frac{a^{-3+\epsilon(a)}}{\mathrm{D}+\mathrm{B}a^{-3}+\varphi(a)}\\ \Omega_\Lambda(a)&=\frac{\mathrm{D}+\varphi(a)-a^{-3+\epsilon(a)}}{\mathrm{D}+\mathrm{B}a^{-3}+\varphi(a)}\end{aligned}$$

$$q(a)=\frac{3}{2}\frac{\Omega_{b,0}a^{-3}+\Omega_{m0}a^{\epsilon(a)-3}}{\Omega_{b,0}a^{-3}+\Omega_{m0}\varphi(a)+\Omega_{\mathrm{X},0}}-1$$



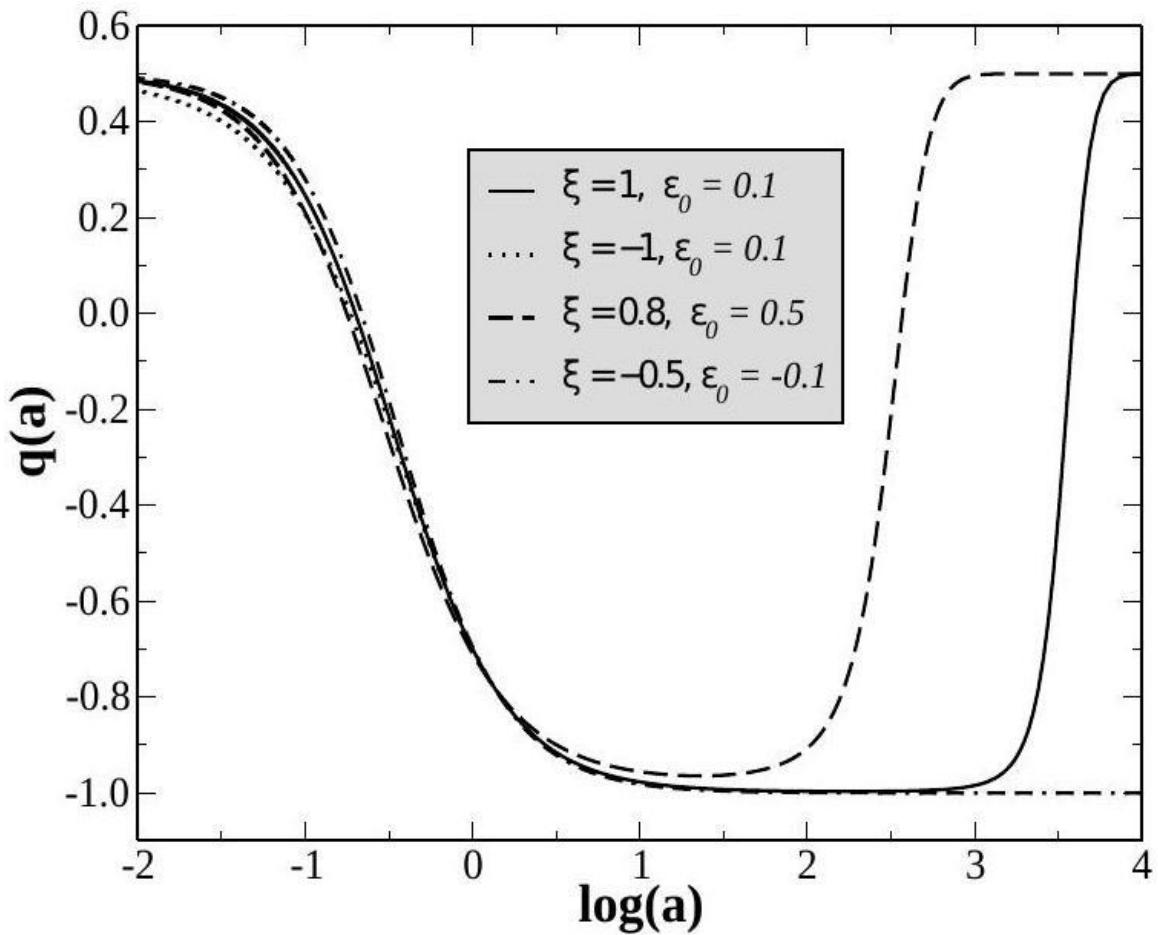


Figura 10. Trayectorias orbitales de la partícula oscura.

$$3M_{Pl}^2 H^2 = \rho_{DE} + \rho_m + \rho_b$$

$$\dot{\rho}_b + 3H\rho_b = 0 \Rightarrow \rho_b = \rho_{b0} \left(\frac{a_0}{a}\right)^3$$

$$\rho_m = \tilde{\rho}_{m0} \left(\frac{a_0}{a}\right)^3 f(a)$$

$$Q = \rho_m \frac{\dot{f}}{f} = \tilde{\rho}_{m0} \left(\frac{a_0}{a}\right)^3 \dot{f}$$

$$f(a) = 1 + g(a)$$

$$\dot{f} = \dot{g} = \frac{dg}{da} \dot{a}$$

$$Q = \tilde{\rho}_{m0} \frac{dg}{da} \dot{a} \left(\frac{a_0}{a}\right)^3$$

$$\rho_m = \tilde{\rho}_{m0} (1 + g) \left(\frac{a_0}{a}\right)^3$$

$$\rho_{m0} = \tilde{\rho}_{m0} (1 + g_0)$$



$$\dot{\rho}_{DE} + 3H(1+w)\rho_{DE} = -\tilde{\rho}_{m0}\frac{dg}{da}\dot{a}\left(\frac{a_0}{a}\right)^3$$

$$\rho_{DE} = (\rho_{m0} + \tilde{\rho}_{m0}g_0)\left(\frac{a_0}{a}\right)^{3(1+w)} - \tilde{\rho}_{m0}\left(\frac{a_0}{a}\right)^3 g + 3w\tilde{\rho}_{m0}a_0^3a^{-3(1+w)}\int_{a_0}^a da g a^{3w-1}$$

$$\begin{aligned} \frac{\ddot{a}}{a} = & -\frac{1}{6}\left\{\tilde{\rho}_{m0}(1+g)\left(\frac{a_0}{a}\right)^3 + \rho_{b0}\left(\frac{a_0}{a}\right)^3 + (1+3w) \times \right. \\ & \left. \times \left[(\rho_{m0} + \tilde{\rho}_{m0}g_0)\left(\frac{a_0}{a}\right)^{3(1+w)} - \tilde{\rho}_{m0}\left(\frac{a_0}{a}\right)^3 g + 3w\tilde{\rho}_{m0}a_0^3a^{-3(1+w)}\int_{a_0}^a da g a^{3w-1}\right]\right\} \end{aligned}$$

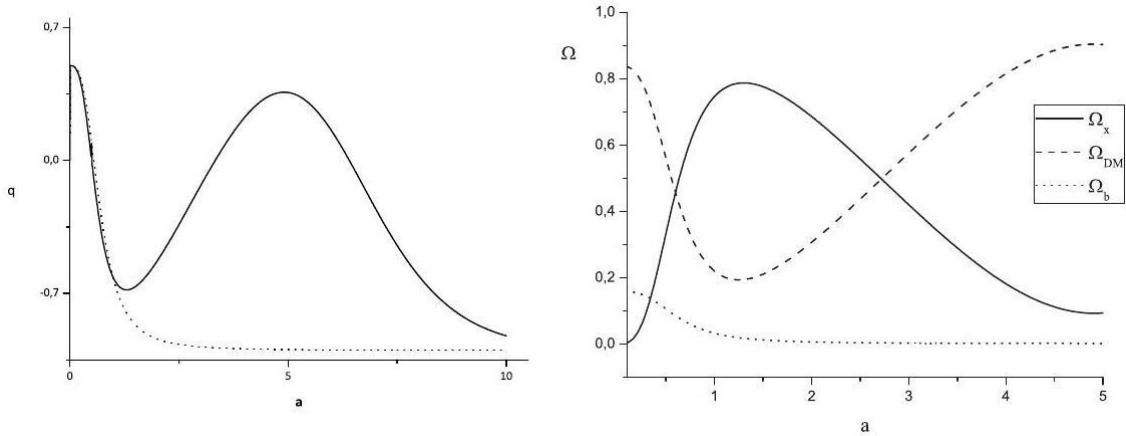


Figura 11. Trayectorias orbitales de la partícula oscura.

$$V(\phi) = \rho_{\phi 0} \left[1 - \frac{\lambda}{6} (1 + \alpha \sqrt{\sigma} \phi)^2 \right] \exp \left[-\lambda \sqrt{\sigma} \left(\phi + \frac{\alpha \sqrt{\sigma}}{2} \phi^2 \right) \right],$$

$$\begin{aligned} h^2 &= \frac{U(y) + x}{1 - \frac{1}{2} \left(\frac{dy}{dN} \right)^2} \\ \frac{d^2y}{dN^2} - \frac{3}{2} \left(\frac{dy}{dN} \right)^3 + 3 \frac{dy}{dN} &= \left(\frac{\Gamma}{dy/dN} + 1.5 \frac{dy}{dN} x - U'(y) \right) h^{-2} \\ \frac{dx}{dN} &= -\Gamma - 3x \end{aligned}$$

$$Q = 3\beta(a)H\rho_{de}$$

$$\beta(a) = \beta_0 a^\xi$$

$$\begin{aligned} \dot{\rho}_{dm} + 3H\rho_{dm} &= Q \\ \dot{\rho}_{de} + 3H(\rho_{de} + p_{de}) &= -Q \end{aligned}$$

$$\rho_{de} = \rho_{de0} a^{-3(1+w_0)} \cdot \exp \left[\frac{3\beta_0(1-a^\xi)}{\xi} \right],$$



$$\rho_{dm} = f(a)\rho_{dm0}$$

$$f(a) \equiv \frac{1}{a^3} \left\{ 1 - \frac{\Omega_{de0}}{\Omega_{dm0}} \frac{3\beta_0 a^{-3w_0} e^{\frac{3\beta_0}{\xi}}}{\xi} \cdot \left[a^\xi E_{\frac{3w_0}{\xi}} \left(\frac{3\beta_0 a^\xi}{\xi} \right) - a^{3w_0} E_{\frac{3w_0}{\xi}} \left(\frac{3\beta_0}{\xi} \right) \right] \right\},$$

$$\begin{aligned}\rho_{de} &= \rho_{de0} a^{-3(1+w_0+\beta_0)} \\ \rho_{dm} &= \rho_{dm0} a^{-3} \left[1 + \frac{\Omega_{de0}}{\Omega_{dm0}} \frac{\beta_0}{w_0 + \beta_0} (1 - a^{-3(w_0+\beta_0)}) \right].\end{aligned}$$

$$E^2(z) \equiv \frac{H^2}{H_0^2} = \Omega_{b0} a^{-3} + \Omega_{dm0} f(a) + \Omega_{de0} a^{-3(1+w_0)} e^{\frac{3\beta_0(1-a^\xi)}{\xi}}$$

$$\begin{aligned}\Omega_b(a) &= \frac{a^{-3}}{a^{-3} + Af(a) + Ba^{-3(1+w_0)} e^{\frac{3\beta_0(1-a^\xi)}{\xi}}} \\ \Omega_{dm}(a) &= \frac{f(a)}{A^{-1}a^{-3} + f(a) + A^{-1}Ba^{-3(1+w_0)} e^{\frac{3\beta_0(1-a^\xi)}{\xi}}} \\ \Omega_{de}(a) &= \frac{a^{-3(1+w_0)} e^{\frac{3\beta_0(1-a^\xi)}{\xi}}}{B^{-1}a^{-3} + AB^{-1}f(a) + a^{-3(1+w_0)} e^{\frac{3\beta_0(1-a^\xi)}{\xi}}},\end{aligned}$$

$$q \equiv -\frac{\ddot{a}}{aH^2} = -1 + \frac{3}{2} \left[\frac{\Omega_b + \Omega_m + (1+w_0)\Omega_{de}}{\Omega_b + \Omega_m + \Omega_{de}} \right]$$

$$q = -1 + \frac{3}{2} \left[\frac{a^{-3} + Af(a) + B(1+w_0)a^{-3(1+w_0)} e^{\frac{3\beta_0(1-a^\xi)}{\xi}}}{a^{-3} + Af(a) + Ba^{-3(1+w_0)} e^{\frac{3\beta_0(1-a^\xi)}{\xi}}} \right].$$

$$q = \frac{1}{2} + \frac{w_0 \Omega_{de0}}{w_0 \Omega_{de0}/(w_0 + \beta_0) + (1 - w_0 \Omega_{de0}/(w_0 + \beta_0)) a^{3(w_0+\beta_0)}}.$$

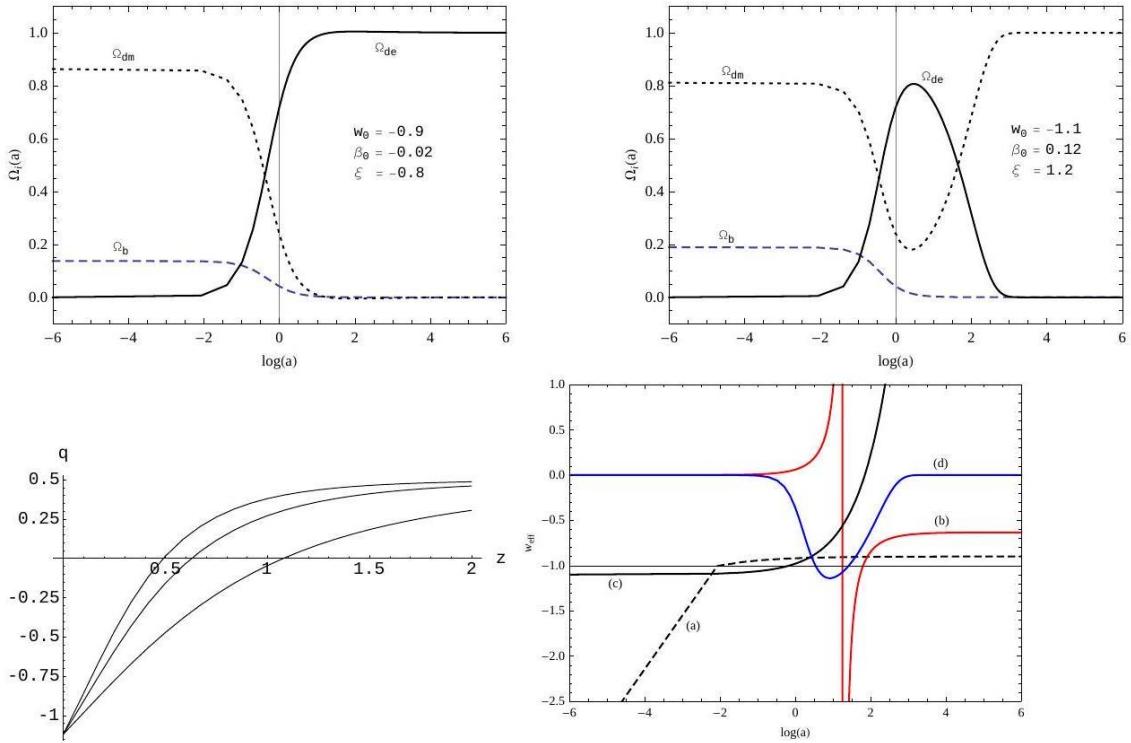


Figura 12. Trayectorias orbitales de la partícula oscura.

$$H^2 = \frac{8\pi G}{3} (\rho_\varphi + \rho_{dm} + \rho_b)$$

$$\dot{H} + H^2 = -\frac{4\pi G}{3} (\rho_\varphi + \rho_{dm} + \rho_b + 3\rho_\varphi)$$

$$\delta(a) = \frac{d \ln m_\psi(a)}{d \ln a}$$

$$\dot{\rho}_{dm} + 3H\rho_{dm} - \delta(a)H\rho_{dm} = 0$$

$$\dot{\rho}_\varphi + 3H(\rho_\varphi + p_\varphi) + \delta(a)H\rho_{dm} = 0$$

$$W(\varphi(a)) = \exp \left(- \int_a^1 \delta(a') d \ln a' \right)$$

$$\ddot{\varphi} + 3H\dot{\varphi} + \left(\frac{dV}{d\varphi} + \frac{\rho_{dm}^{(0)}}{a^3} \frac{dW}{d\varphi} \right) = 0$$

$$\rho_{dm}(a) = \rho_{dm}^{(0)} a^{-3} \exp \left(- \int_a^1 \delta(a') d \ln a' \right)$$

$$\rho_{dm}(u) = \rho_{dm}^{(0)} e^{-3u} \exp \left(- \int_a^1 \delta(u') du' \right)$$

$$\rho'_\varphi(u) + 3(1 + w_\varphi(u))\rho_\varphi(u) + \delta(u)\rho_{dm}(u) = 0$$



$$\frac{H^2(u)}{H_0^2} = \Omega_b e^{-3u} + \Omega_{dm} e^{-3u} \exp\left(-\int_u^0 \delta(u') du'\right) + \Omega_\varphi f(u)$$

$$\left(\frac{d\tilde{\varphi}}{du}\right)^2 = -\frac{1}{4\pi}\left(\frac{d\ln H(u)}{du} + \frac{3}{2}(\Omega_{dm}(u) + \Omega_b(u))\right),$$

$$\Omega_{dm,b}(u) = \frac{\rho_{dm,b}(u)}{\rho_\varphi(u) + \rho_{dm}(u) + \rho_b(u)}$$

$$\tilde{V}(\tilde{\varphi}) \equiv \frac{V(u(\tilde{\varphi}))}{\rho_c^{(0)}} = \frac{1}{3} \frac{H(u)}{H_0} \frac{d(H/H_0)}{du} + \frac{H^2(u)}{H_0^2} - \frac{1}{2} \Omega_b e^{-3u} - \frac{1}{2} \Omega_{dm} e^{-3u} \exp\left(-\int_u^0 \delta(u') du'\right)$$

$$W(u(\tilde{\varphi})) = \exp\left(-\int_u^0 \delta(u') du'\right)$$

$$\rho_{dm} = \rho_{dm}^{(0)} a^{-3+\delta}$$

$$\rho_\varphi(a) = \rho_\varphi^{(0)} a^{-3(1+w_\varphi)} + \frac{\delta}{\delta + 3w_\varphi} \rho_{dm}^{(0)} \mathcal{O}\left(a^{-3(1+w_\varphi)} - a^{-3+\delta}\right)$$

$$W(\tilde{\varphi}(u)) = e^{\delta u}$$

$$\begin{aligned} H(t) &= \frac{1}{a} \frac{da}{dt} \\ q(t) &= -\frac{1}{a} \frac{d^2 a}{dt^2} \left(\frac{1}{a} \frac{da}{dt}\right)^{-2} \\ j(t) &= \frac{1}{a} \frac{d^3 a}{dt^3} \left(\frac{1}{a} \frac{da}{dt}\right)^{-3} \\ s(t) &= \frac{1}{a} \frac{d^4 a}{dt^4} \left(\frac{1}{a} \frac{da}{dt}\right)^{-4} \\ l(t) &= \frac{1}{a} \frac{d^5 a}{dt^5} \left(\frac{1}{a} \frac{da}{dt}\right)^{-5} \end{aligned}$$

$$p = p_{de}, \rho = \rho_{de} + \rho_{dm}$$

$$\frac{1}{a^3} \frac{d}{dt} (\rho_{dm} a^3) + \frac{1}{a^{3(1+w_{de})}} \frac{d}{dt} (\rho_{de} a^{3(1+w_{de})}) = 0$$

$$\frac{1}{a^3} \frac{d}{dt} (\rho_{dm} a^3) = \gamma(t), \frac{1}{a^{3(1+w_{de})}} \frac{d}{dt} (\rho_{de} a^{3(1+w_{de})}) = -\gamma(t).$$

$$\begin{aligned} \rho_{dm} a^3 &= \rho_{dm0} a_0^3 + \int_{t_0}^t \gamma(t) a^3 dt \\ \rho_{de} a^{3(1+w_{de})} &= \rho_{de0} a_0^{3(1+w_{de})} + \int_{t_0}^t \gamma(t) a^{3(1+w_{de})} dt \end{aligned}$$



$$\ddot{a} = \frac{1}{2} \left[-\frac{A(a)}{a^2} - \frac{(1+3w_{de})B(a)}{a^{2+3w_{de}}} \right]$$

$$A(a) \equiv \frac{1}{3}\rho_{dm}a^3, B(a) \equiv \frac{1}{3}a^{3(1+w_{de})}$$

$$qH^2 = \frac{1}{2} \left[\frac{A(a)}{a^3} + \frac{(1+3w_{de})B(a)}{a^{3(1+w_{de})}} \right]$$

$$\nu(t) \equiv \frac{\gamma(t)}{3H^3}$$

$$j - \frac{3}{2}w_{de}\nu = \Omega_{dm} + \frac{1}{2}(1+3w_{de})(2+3w_{de})\Omega_{de}$$

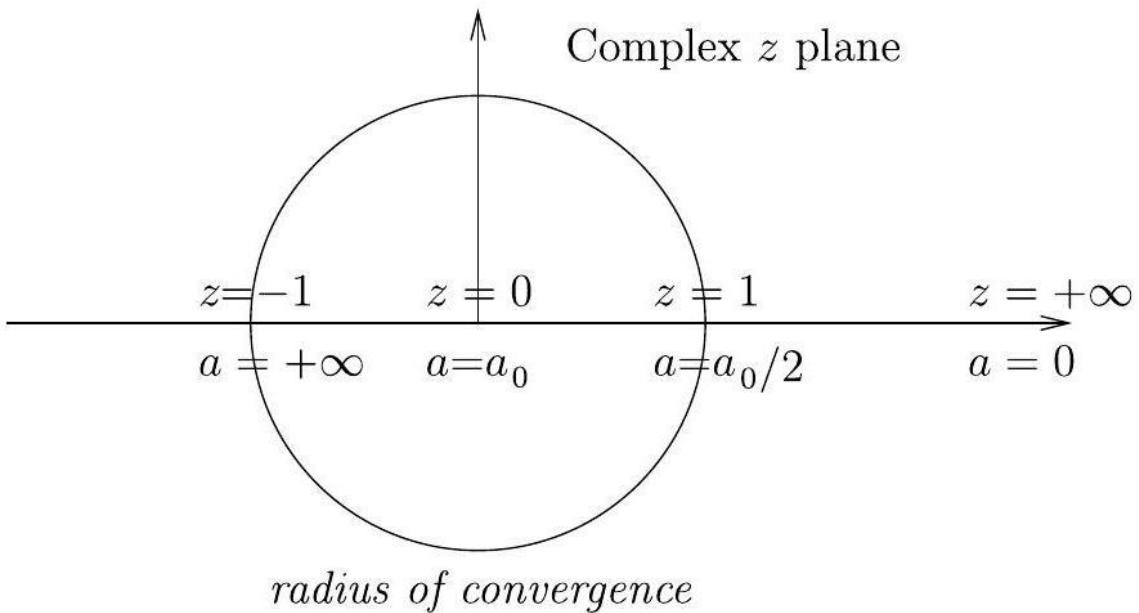
$$j - \frac{3}{2}w_{de}\nu - 1 = \frac{9}{2}w_{de}(1+w_{de})\Omega_{de} - \Omega_c, \Omega_c \equiv -\frac{k}{a^2 H^2}$$

$$q = \frac{1}{2}\Omega_{dm} + \frac{1+3w_{de}}{2}\Omega_{de}$$

$$j - \frac{3}{2}w_{de}\nu + q = \frac{3}{2}\Omega_{dm} + \frac{1}{2}(1+3w_{de})(1+w_{de})\Omega_{de}$$

$$j - \frac{3}{2}w_{de}\nu + q = -\frac{3}{2}\Omega_{dm}(4+3w_{de})w_{de} + \frac{3}{2}(1+3w_{de})(1+w_{de})$$

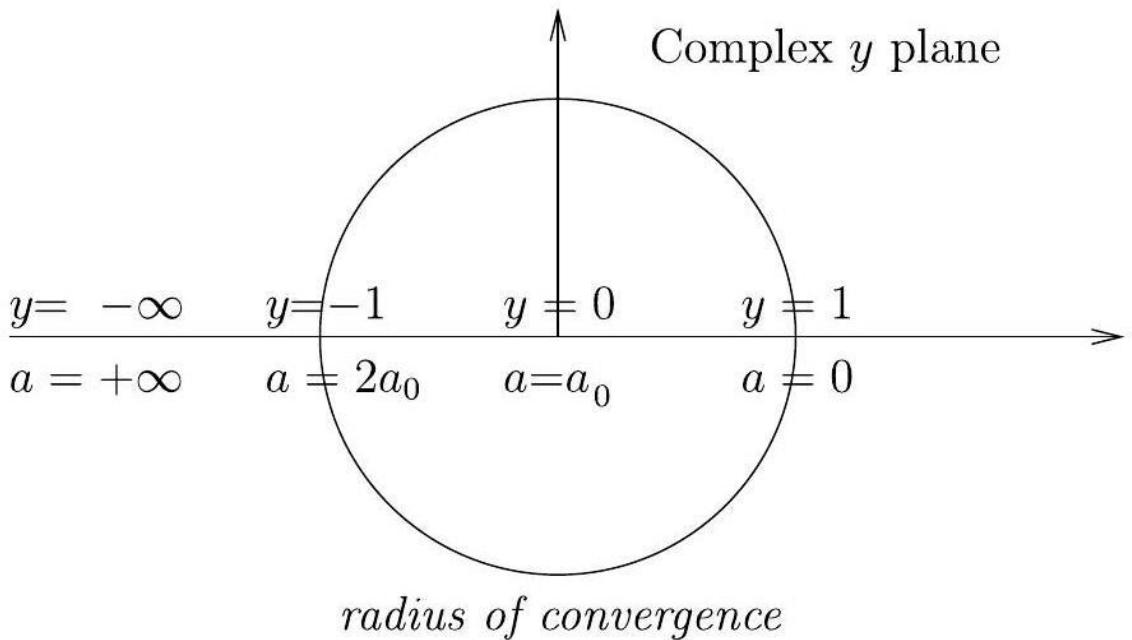
$$\frac{1}{1+z} = \frac{a(t)}{a_0} = 1 + H_0(t-t_0) - \frac{q_0 H_0^2}{2!}(t-t_0)^2 + \frac{j_0 H_0^3}{3!}(t-t_0)^3 + O([t-t_0]^4)$$



$$y = \frac{z}{1+z}; z = \frac{y}{1-y}$$

$$y = \frac{\lambda_0 - \lambda_e}{\lambda_0} = \frac{\Delta\lambda}{\lambda_0}$$





$$0m(x) \equiv \frac{h^2(x) - 1}{x^3 - 1}, x = 1 + z, h(x) = \frac{H(x)}{H_0}$$

$$h^2(x) = \Omega_{m0}x^3 + (1 - \Omega_{m0})x^\alpha, \alpha = 3(1 + w)$$

$$0m(x) = \Omega_{m0} + (1 - \Omega_{m0}) \frac{x^\alpha - 1}{x^3 - 1}$$

$$q = -1 - (\dot{H}/H)^2 = 1/2(1 + 3w_{de}\Omega_{de})$$

$$\frac{\ddot{H}}{H^3} = \frac{9}{2} \left(1 + \frac{p_{de}}{\rho} \right) + \frac{9}{2} \left[w_{de}(1 + w_{de}) \frac{\rho_{de}}{\rho} - w_{de} \frac{\Pi}{\rho} - \frac{\dot{w}_{de}}{3H} \frac{\rho_{de}}{\rho} \right],$$

$$r \equiv \frac{\ddot{a}}{aH^3}, s \equiv \frac{r - 1}{3(q - 1/2)},$$

$$r = 1 + \frac{9}{2}\Omega_{de}w_{de}(1 + w_{de}) - \frac{3}{2}\Omega_{de}\frac{\dot{w}_{de}}{H};$$

$$s = 1 + w_{de} - \frac{1}{3}\frac{\dot{w}_{de}}{w_{de}H}; w_{de} \equiv \frac{p_{de}}{\rho_{de}}$$

$$\{r, s\} = \{1, 0\}$$

$$\{r, s\} = \left\{ 1 + \frac{9}{2}\Omega_{DE}(1 + w_{de}), 1 + w_{de} \right\}$$

$$\{r, s\} = \left\{ 1 + \frac{12\pi G\dot{\phi}^2}{H^2} + \frac{8\pi G\dot{V}}{H^3}, \frac{2\left(\dot{\phi}^2 + \frac{2\dot{V}}{H}\right)}{\dot{\phi}^2 - 2V} \right\}'$$



$$r=1+\frac{9}{2}\frac{w_{de}}{1+R}\Big[1+w_{de}-\frac{\Pi}{\rho_{de}}-\frac{\dot{w}_{de}}{3w_{de}H}\Big], R\equiv \frac{\rho_{dm}}{\rho_{de}}\\ s=1+w-\frac{\Pi}{\rho_{de}}-\frac{\dot{w}_{de}}{3Hw_{de}},$$

$$\Pi=\rho_{de}\left(w_{de}+\frac{\xi}{3}\right)\frac{R_0(1+z)^{\xi}}{1+R_0(1+z)^{\xi}}$$

$$r=1+\frac{9}{2}\frac{w_{de}}{1+R_0(1+z)^{\xi}}\Bigg[1+w_{de}-\bigg(w_{de}+\frac{\xi}{3}\bigg)\frac{R_0(1+z)^{\xi}}{1+R_0(1+z)^{\xi}}\Bigg]\\ s=1+w_{de}-\bigg(w_{de}+\frac{\xi}{3}\bigg)\frac{R_0(1+z)^{\xi}}{1+R_0(1+z)^{\xi}}$$

$$\rho_{de}=3M_p^2(\alpha H^2+\beta \dot{H})$$

$$\dot{r_\rho}=3H\left[w r_\rho+b\bigl(1+r_\rho\bigr)^2\right].$$

$$3M_p^2H^2=\rho_{de}+\rho_m$$

$$\dot{H}=-\frac{3}{2}H^2\left(1+\frac{w}{1+r_\rho}\right)$$

$$\Omega_{de}=\frac{1}{1+r_\rho}$$

$$w=\left(\frac{2\alpha}{3\beta}-1\right)(1+r_\rho)-\frac{2}{3\beta}$$

$$\frac{\ddot{a}}{a}=-\frac{4\pi G}{3}(\rho +3p),$$

$$r=1+\frac{9(\rho+p)}{2\rho}\frac{\dot{p}}{\dot{\rho}}, s=\frac{(\rho+p)}{p}\frac{\dot{p}}{\dot{\rho}},$$

$$q=-\frac{\ddot{a}}{aH^2}=\frac{1}{2}+\frac{3p}{2\rho}$$

$$w^\mathrm{eff}=w+b\bigl(1+r_\rho\bigr)$$

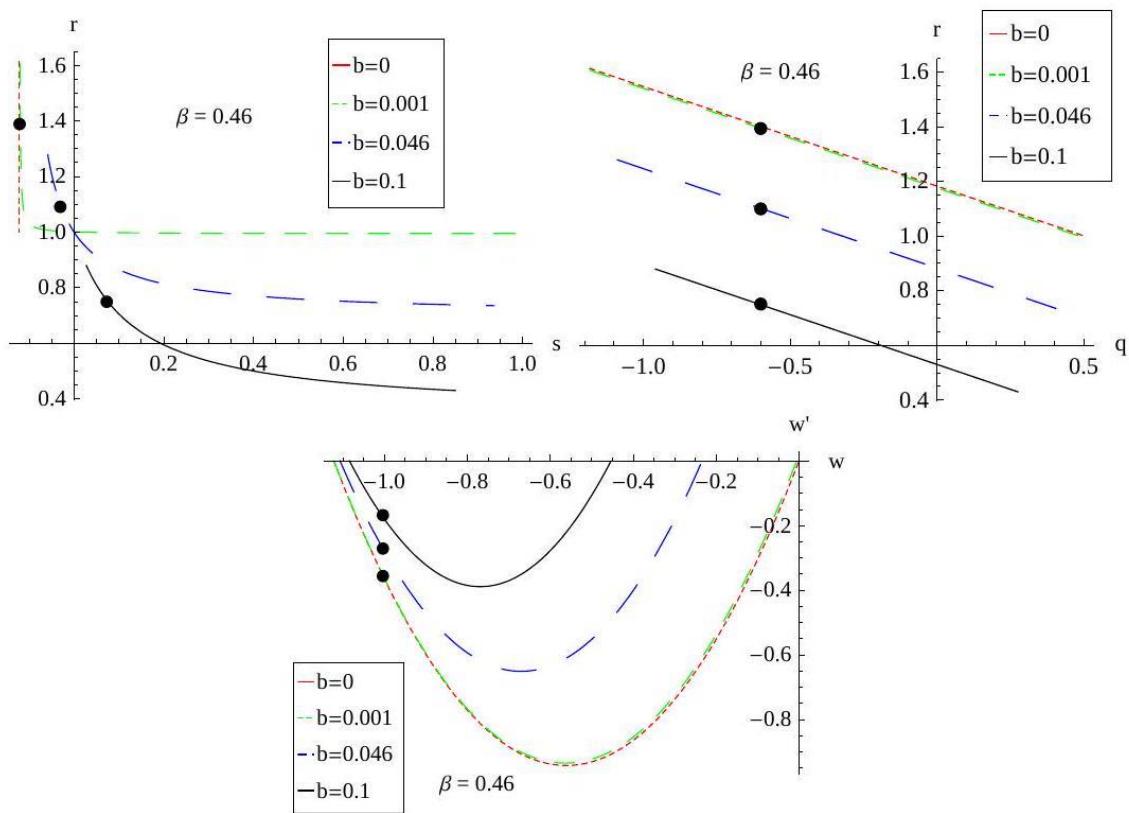
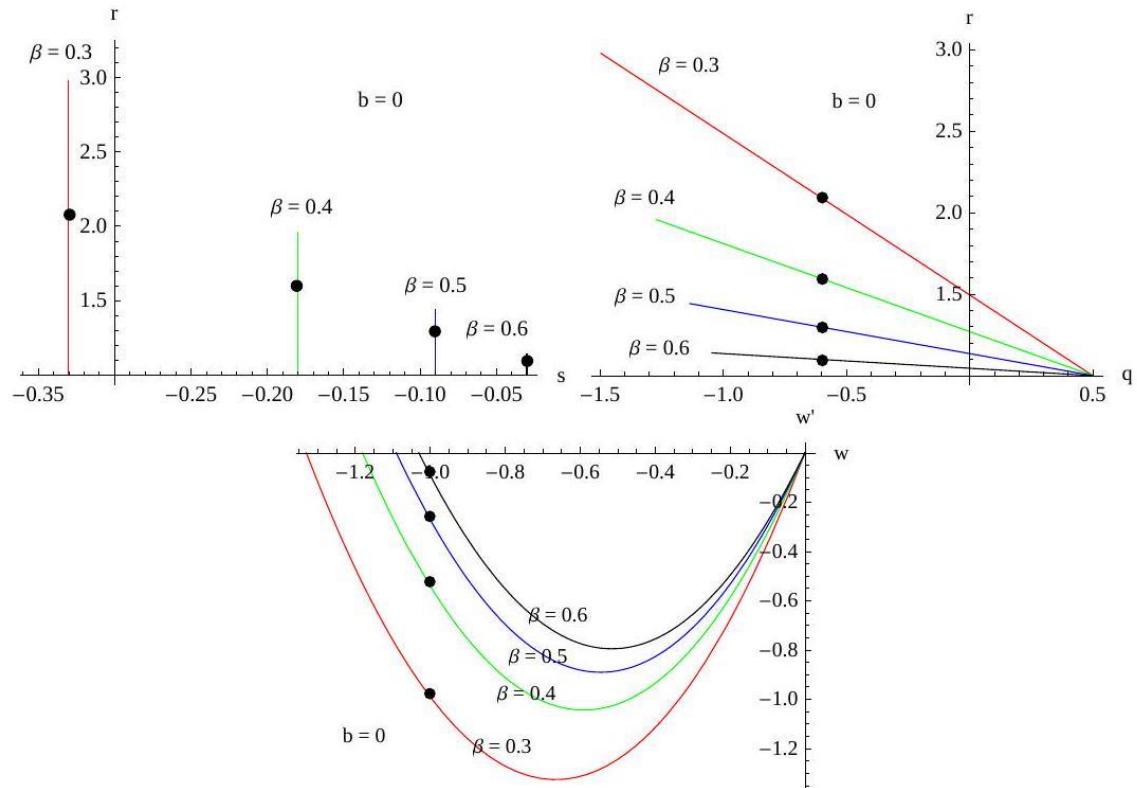
$$\dot{\rho}_{de}+3H\bigl(1+w^\mathrm{eff}\bigr)\rho_{de}=0$$

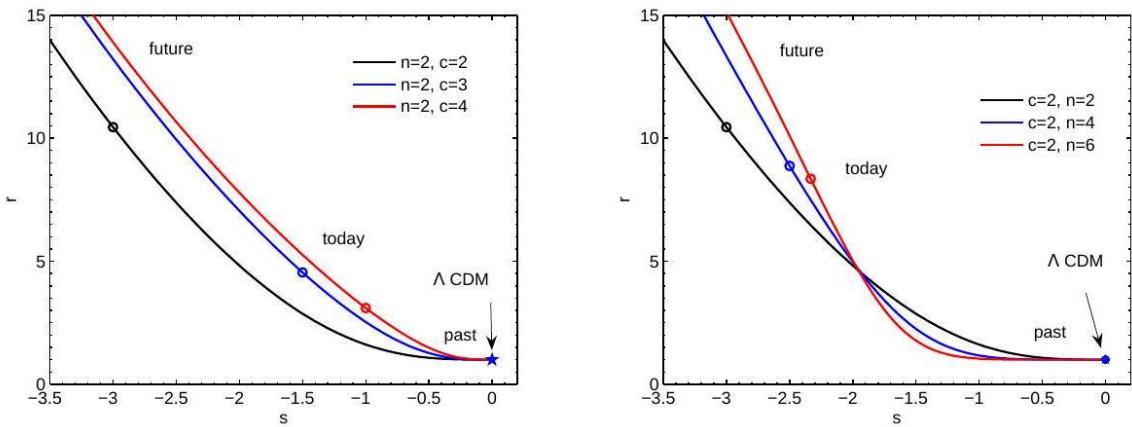
$$r=1-\frac{3}{2}\Omega_{de}\big[w'-3w\big(1+w^\mathrm{eff}\big)\big]\\ s=1+w^\mathrm{eff}-\frac{w'}{3w}\\ q=\frac{1}{2}+\frac{3}{2}w\Omega_{de}$$

$$\{r,s\}|_{\rm LCDM} = \{1,0\}$$



$$H^2 = \frac{1}{3M_p^2} (\rho_m + \rho_\Lambda)$$





Figuras 13, 14 y 15. Trayectorias orbitales de la partícula oscura.

$$\rho_\Lambda = \alpha H$$

$$\Omega_m = \frac{\rho_m}{\rho_c} = \frac{\rho_m}{3M_p^2 H^2}, \Omega_\Lambda = \frac{\rho_\Lambda}{\rho_c} = \frac{\rho_\Lambda}{3M_p^2 H^2}$$

$$r = \frac{\ddot{H}}{H^3} - 3q - 2.$$

$$r = 1 + \frac{9}{4} w_\Lambda \Omega_\Lambda (w_\Lambda \Omega_\Lambda + 1) - \frac{3}{2} \Omega_\Lambda w'_\Lambda$$

$$s = \frac{1}{2} (1 + w_\Lambda \Omega_\Lambda) - \frac{w'_\Lambda}{3w_\Lambda}$$

$$p_\Lambda = K \rho_\Lambda^{1+\frac{1}{n}},$$

$$\rho_\Lambda = \left(\frac{1}{B a^{\frac{3(1+\alpha)}{n}} - \tilde{K}} \right)^n,$$

$$\dot{\rho}_\Lambda + 3H(1+\alpha+w_\Lambda)\rho_\Lambda = 0$$

$$\dot{\rho}_\Lambda = -3BH(1+\alpha)a^{\frac{3(1+\alpha)}{n}} \rho_\Lambda^{1+\frac{1}{n}}$$

$$\begin{aligned} \frac{\ddot{H}}{H^3} &= -\frac{9}{2} \Omega_\Lambda (1+\alpha)(\alpha+w_\Lambda) [(1+\alpha)(-w_\Lambda + \Omega_\Lambda \alpha + \Omega_\Lambda w_\Lambda) - \alpha(\alpha+2)] \\ &\quad - \frac{3}{2} \Omega_\Lambda (1+\alpha) w'_\Lambda + \frac{9}{2} [\Omega_\Lambda (1+\alpha)(\alpha+w_\Lambda) + 1]^2 \end{aligned}$$

$$r = \frac{\ddot{a}}{aH^3} = \frac{\dot{H}}{H^3} - 3q - 2$$

$$r = 1 + \frac{3}{2} \Omega_\Lambda (1+\alpha) [3(1+\alpha)(\alpha+w_\Lambda)(1+\alpha+w_\Lambda) - w'_\Lambda]$$



$$s=\frac{2}{3}\frac{3\alpha(\alpha+1)^2+3\alpha w_{\Lambda}(2\alpha+w_{\Lambda}+3)+3w_{\Lambda}(1+w_{\Lambda})-w_{\Lambda}'}{\alpha+w_{\Lambda}}$$

$$H(z)=-\frac{1}{1+z}\frac{dz}{dt}$$

$$\chi^2_H = \sum_{i=1}^{13} \frac{[H(z_i)-H_{obs}(z_i)]^2}{\sigma^2_{hi}},$$

$$A=\frac{\sqrt{\Omega_m}}{E(z_{\rm BAO})^{1/3}}\bigg[\frac{1}{z_{\rm BAO}}\int_0^{z_{\rm BAO}}\frac{dz'}{E(z')}\bigg]^{2/3}$$

$$\chi^2_{BAO} = \frac{(A - A_{\rm obs})^2}{\sigma_A^2}$$

$$R=\sqrt{\Omega_{m0}}\int_0^{z_{rec}}\frac{dz'}{E(z')}$$

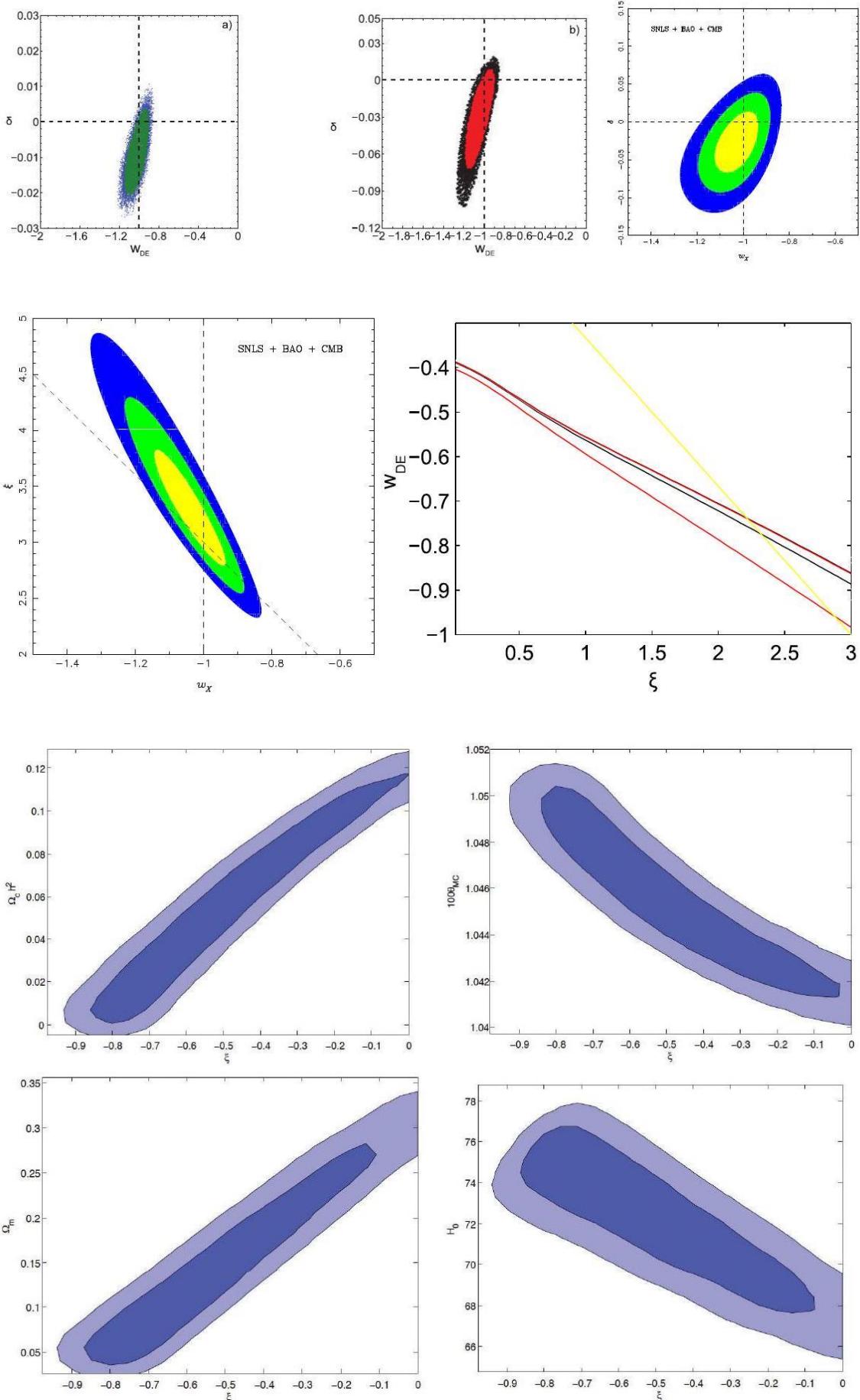
$$\chi^2_{\rm CMB} = \frac{(R-R_{\rm obs})^2}{\sigma_R^2}$$

$$\mu=5\log\frac{d_L}{Mpc}+25=5\log_{10}~H_0d_L-\mu_0$$

$$\chi^2_{SN}=A-\frac{B^2}{C}+\ln\left(\frac{C}{2\pi}\right)$$

$$A=\sum_i^{557}\left(\mu^{\text{data}}-\mu^{\text{th}}\right)^2/\sigma_i^2,B=\sum_i^{557}\mu^{\text{data}}-\mu^{\text{th}}/\sigma_i^2,C=\sum_i^{557}1/\sigma_i^2,\mu^{\text{data}}$$

Model	$\Omega_{m,0}$	w_{DE}	δ
$Q=3\delta H\rho_m$	$0.274^{+0.029}_{-0.029}$	$-1.02^{+0.12}_{-0.13}$	$-0.009^{+0.013}_{-0.012}$
$Q=3\delta H\rho_{DE}$	$0.272^{+0.030}_{-0.030}$	$-1.02^{+0.09}_{-0.09}$	$-0.023^{+0.039}_{-0.040}$
$\rho_m=\rho_{m0}a^{-3+\delta}$	$0.270^{+0.040}_{-0.050}$	$-1.03^{+0.12}_{-0.15}$	$-0.03^{+0.06}_{-0.05}$
ΛCDM	$0.270^{+0.019}_{-0.019}$	$-1.0710^{+0.0775}_{-0.0775}$	0



Figuras 16, 17 y 18. Agujeros negros cuánticos provocados por una partícula oscura.

$$\frac{\rho_{DE}}{\rho_m} = \frac{\rho_{DE0}}{\rho_{m0}} a^\xi$$

$$\begin{aligned}\nabla_\mu T_{(dm)\nu}^\mu &= Qu_\nu^{(dm)}/a \\ T_{(de)\nu}^\mu &= -Qu_\nu^{(dm)}/a \\ Q &= \xi H \rho_{de}\end{aligned}$$

$$\begin{aligned}\dot{\rho}_{dm} + 3\mathcal{H}\rho_{dm} &= \xi\mathcal{H}\rho_{de} \\ \dot{\rho}_{de} + 3\mathcal{H}(1+w)\rho_{de} &= -\xi\mathcal{H}\rho_{de}\end{aligned}$$

$$\begin{aligned}\delta_{dm} &= -\left(kv_{dm} + \frac{1}{2}\dot{h}\right) + \xi\mathcal{H}\frac{\rho_{de}}{\rho_{dm}}(\delta_{de} - \delta_{dm}) \\ &\quad + \xi\frac{\rho_{de}}{\rho_{dm}}\left(\frac{kv_T}{3} + \frac{\dot{h}}{6}\right) \\ \dot{\delta}_{de} &= -(1+w)\left(kv_{de} + \frac{1}{2}\dot{h}\right) - 3\mathcal{H}(1-w) \\ &\quad [\delta_{de} + \mathcal{H}(3(1+w) + \xi)\frac{v_{de}}{k}] - \xi\left(\frac{kv_T}{3} + \frac{\dot{h}}{6}\right) \\ \dot{v}_{dm} &= -\mathcal{H}v_{dm} \\ \dot{v}_{de} &= 2\mathcal{H}\left(1 + \frac{\xi}{1+w}\right)v_{de} + \frac{k}{1+w}\delta_{de} - \xi\mathcal{H}\frac{v_{dm}}{1+w}\end{aligned}$$

$$\begin{aligned}\dot{\rho}_{dm} + 3H\rho_{dm} &= -\beta(\phi)\dot{\phi}\rho_{dm} \\ \dot{\rho}_\phi + 3H\rho_\phi &= +\beta(\phi)\dot{\phi}\rho_{dm}\end{aligned}$$

$$m_c(a) = m_c(a_0)e^{-\int \beta(\phi)d\phi}$$

$$\ddot{\delta}_{dm} + (2H - \beta\dot{\phi})\dot{\delta}_{dm} - \frac{3}{2}H^2[(1 + 2\beta^2)\Omega_{dm}\delta_{dm} + \Omega_b\delta_b] = 0$$

$$-\beta\dot{\phi}\dot{\delta}_{dm}$$

$$\dot{\vec{v}}_i = \beta\dot{\phi}\vec{v}_i + \sum_{j \neq i} \frac{G(1 + 2\beta^2)m_j\vec{r}_{ij}}{|\vec{r}_{ij}|^3},$$

$$\vec{a}_v = \beta\dot{\phi}\vec{v}$$



$$\ddot{\phi} + 3H\dot{\phi} + \frac{dV}{d\phi} = \sqrt{\frac{2}{3}}\beta_c(\phi)\frac{\rho_c}{M_{\text{Pl}}}$$

$$\dot{\rho}_c + 3H\rho_c = -\sqrt{\frac{2}{3}}\beta_c(\phi)\frac{\rho_c\dot{\phi}}{M_{\text{Pl}}}$$

$$\dot{\rho}_b + 3H\rho_b = 0$$

$$\dot{\rho}_r + 4H\rho_r = 0$$

$$3H^2 = \frac{1}{M_{\text{Pl}}^2}(\rho_r + \rho_c + \rho_b + \rho_\phi)$$

$$V(\phi) = A e^{-\alpha\phi}$$

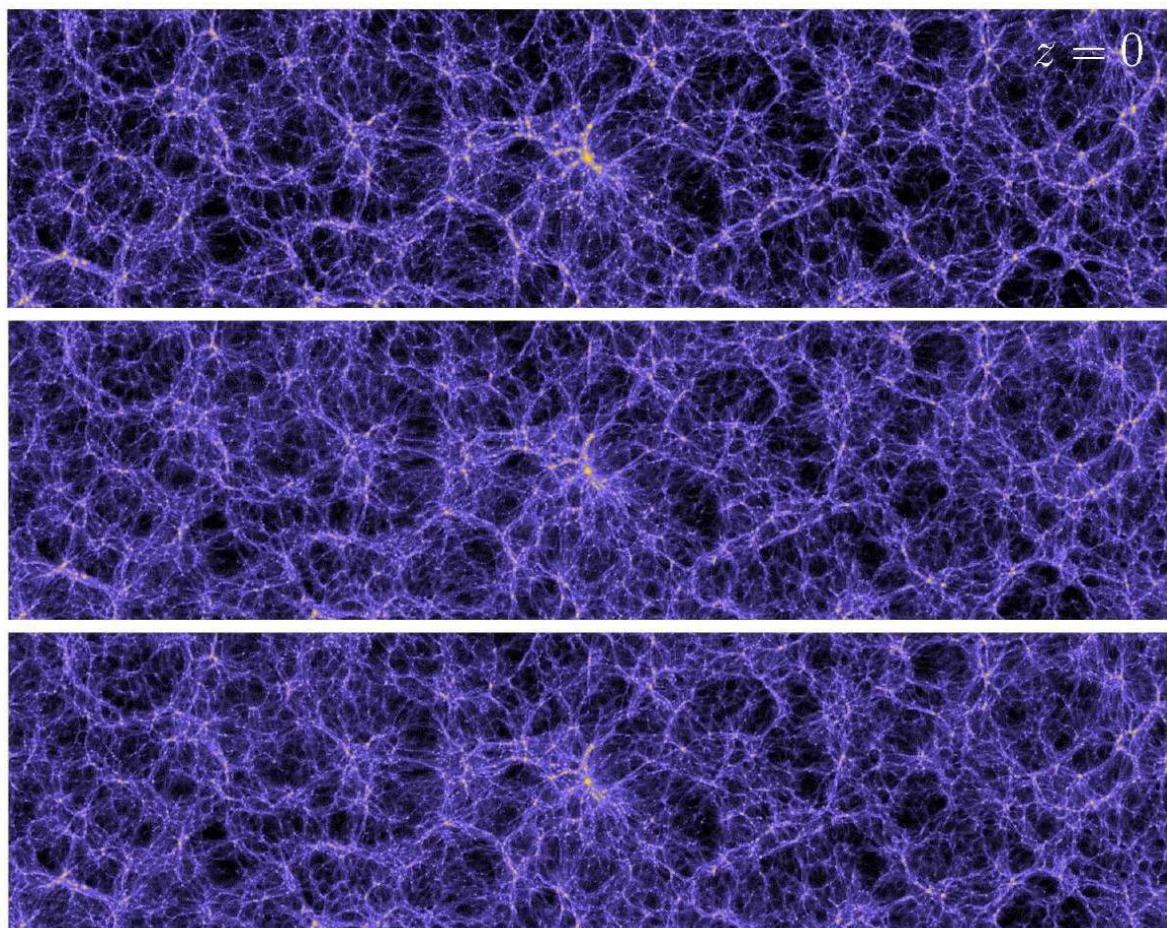
$$V(\phi) = A\phi^{-\alpha}e^{\phi^2/2}$$

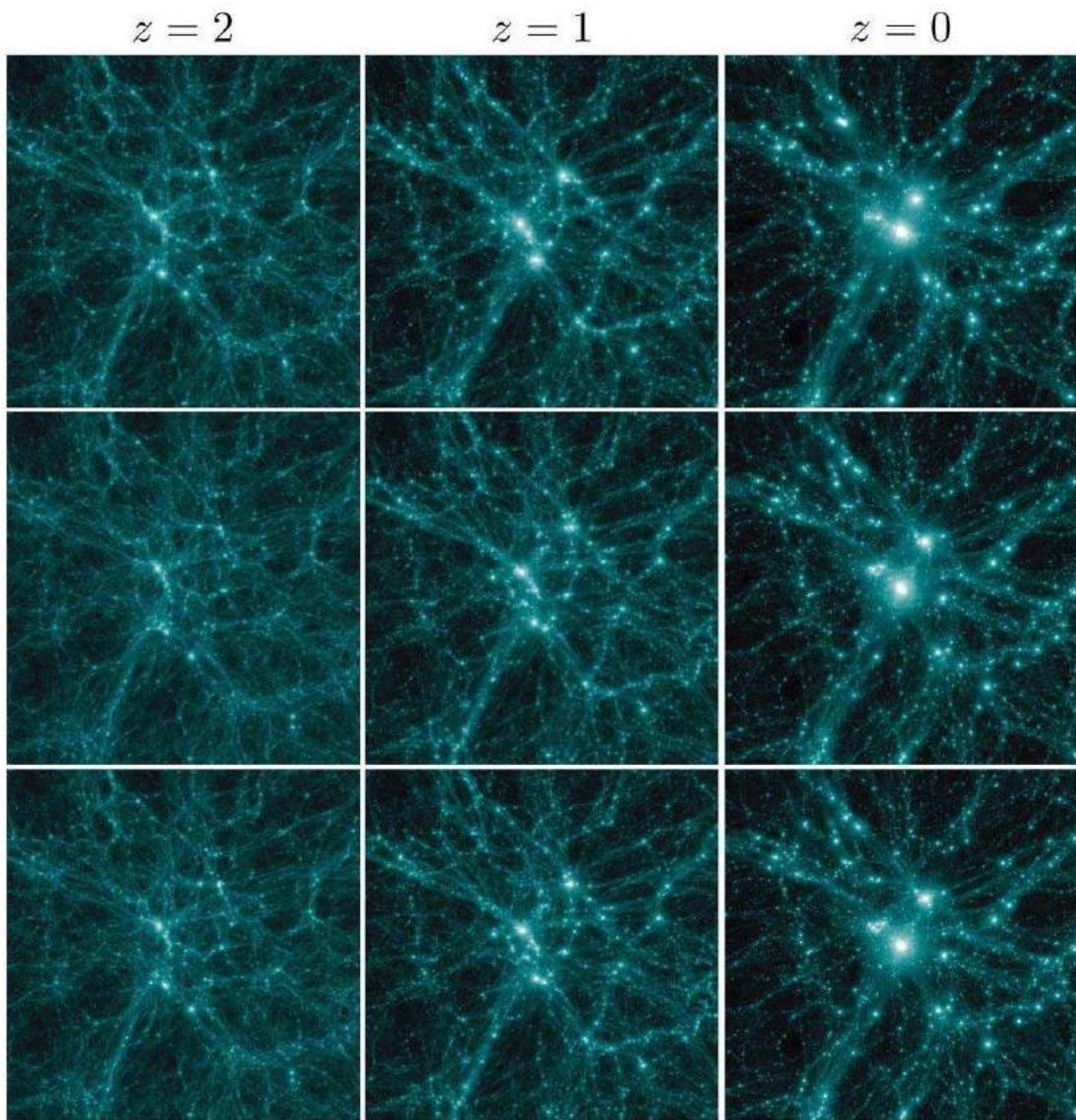
$$\beta_c(\phi) \equiv \beta_0 e^{\beta_1 \phi}$$

Parameter	Value
H_0	70.3 km s ⁻¹ Mpc ⁻¹
Ω_{CDM}	0.226
Ω_{DE}	0.729
\mathcal{A}_s	2.42×10^{-9}
Ω_b	0.0451
n_s	0.966

Model	Potential	α	β_0	β_1	Scalar field normalization	Potential normalization	$w_\phi(z=0)$	$\mathcal{A}_s(z_{\text{CMB}})$	$\sigma_8(z=0)$
ΛCDM	$V(\phi) = A$	—	—	—	—	$A = 0.0219$	-1.0	2.42×10^{-9}	0.809
EXP001	$V(\phi) = A e^{-\alpha\phi}$	0.08	0.05	0	$\phi(z=0) = 0$	$A = 0.0218$	-0.997	2.42×10^{-9}	0.825
EXP002	$V(\phi) = A e^{-\alpha\phi}$	0.08	0.1	0	$\phi(z=0) = 0$	$A = 0.0218$	-0.995	2.42×10^{-9}	0.875
EXP003	$V(\phi) = A e^{-\alpha\phi}$	0.08	0.15	0	$\phi(z=0) = 0$	$A = 0.0218$	-0.992	2.42×10^{-9}	0.967
EXP008e3	$V(\phi) = A e^{-\alpha\phi}$	0.08	0.4	3	$\phi(z=0) = 0$	$A = 0.0217$	-0.982	2.42×10^{-9}	0.895
SUGRA003	$V(\phi) = A\phi^{-\alpha}e^{\phi^2/2}$	2.15	-0.15	0	$\phi(z \rightarrow \infty) = \sqrt{\alpha}$	$A = 0.0202$	-0.901	2.42×10^{-9}	0.806







Figuras 19 y 20. Materia y energía oscuras en un espacio – tiempo cuántico relativista.

SECCIÓN III.

$$ds^2 = -dt^2 + a^2(t) \left[\frac{dr^2}{1 - Kr^2} + r^2 d\theta^2 + r^2 \sin^2 \theta d^2\phi \right]$$

$H_0/[\text{kms}^{-1}\text{Mpc}^{-1}]$	67.74 ± 0.46
$\Omega_{b,0} h^2$	0.02230 ± 0.00014
$\Omega_{c,0} h^2$	0.1188 ± 0.0010
$\Omega_{\Lambda,0}$	0.6911 ± 0.0062
$\Omega_{K,0}$	$0.0008^{+0.0040}_{-0.0039}$
ω_d	$-1.019^{+0.075}_{-0.080}$



$$ds^2=a^2(\tau)\left[-d\tau^2+\frac{dr^2}{1-Kr^2}+r^2d\theta^2+r^2\sin^2\theta d^2\phi\right]$$

$$\begin{gathered}H^2+\frac{K}{a^2}=\frac{8\pi G}{3}\sum \rho_i+\frac{\Lambda}{3}\\\frac{\ddot{a}}{a}=-\frac{4\pi G}{3}\sum (\rho_i+3p_i)+\frac{\Lambda}{3}\end{gathered}$$

$$1+z=\frac{a(t_0)}{a(t_e)}$$

$$E(z)=\frac{H(z)}{H_0}=\left(\sum_i~\Omega_{i,0}f_i(z)\right)^{1/2}, f_i(z)=\exp\left[3\int_0^z\frac{1+\omega_i(z')}{1+z'}dz'\right]$$

$$D_L=(1+z)cH_0^{-1}\int_0^z\frac{dz'}{E(z')}$$

$$D_A=\frac{cH_0^{-1}}{(1+z)}\int_0^z\frac{dz'}{E(z')}$$

$$t_L(z)=H_0^{-1}\int_0^z\frac{dz'}{(1+z')E(z')}\equiv t_0-t_{age}(z)-df$$

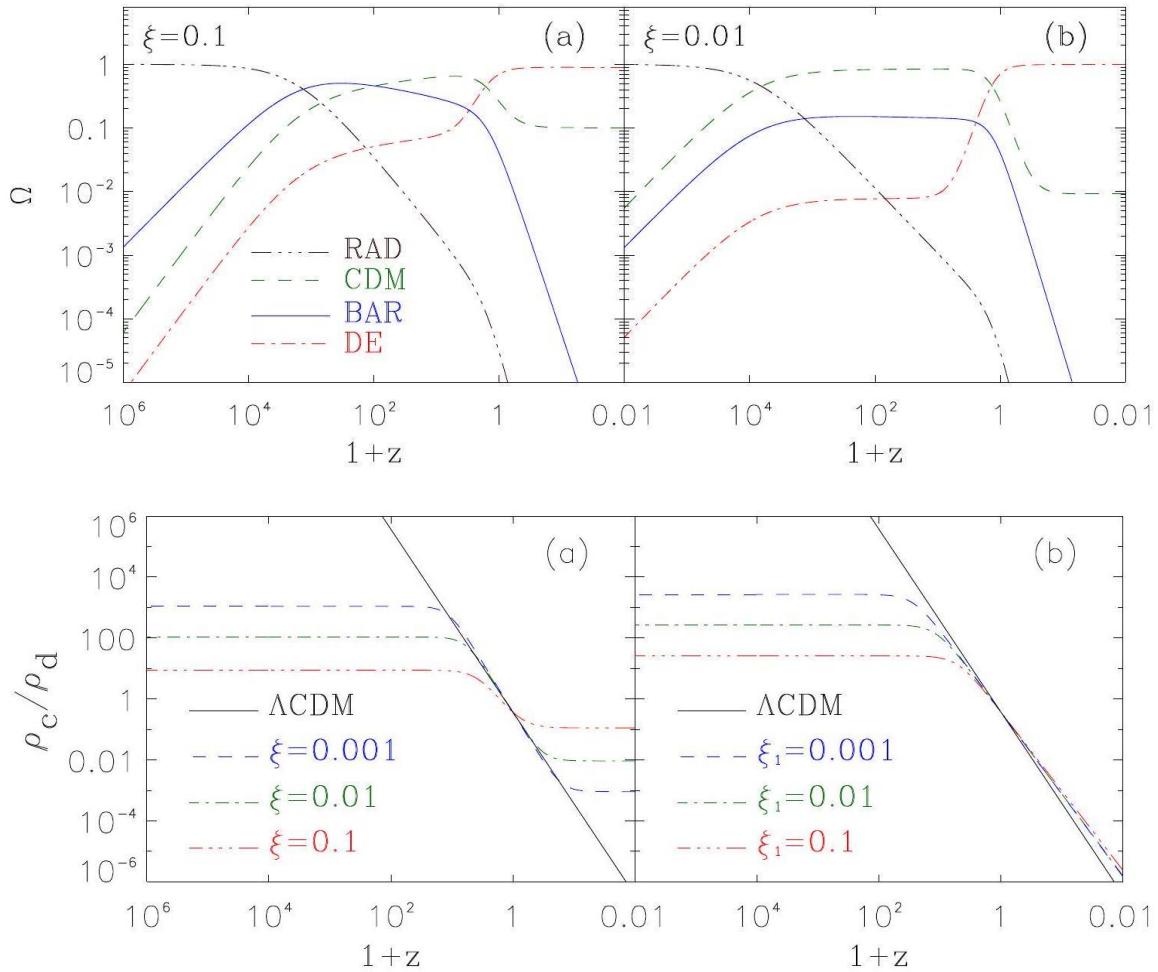
$$\frac{\rho_\Lambda}{\rho_m(t_{\rm Planck})} = \frac{\rho_\Lambda}{\rho_{m,0}} \left(\frac{T_{\rm Planck}}{T_0} \right)^{-3} \simeq 10^{-95}$$

$$\frac{d\rho_i}{dt} + 3H(1+\omega_i)\rho_i = 0$$

Model	Q	DE EoS
I	$\xi_2 H \rho_d$	$-1 < \omega_d < 0$
II	$\xi_2 H \rho_d$	$\omega_d < -1$
III	$\xi_1 H \rho_c$	$\omega_d < -1$
IV	$\xi H(\rho_c + \rho_d)$	$\omega_d < -1$

$$\begin{gathered}\dot{\rho}_c+3H\rho_c=Q\\\dot{\rho}_d+3H(1+\omega_d)\rho_d=-Q\end{gathered}$$





Figuras 21 y 22. Densidades de masa y energía de una partícula oscura.

$$Q = H\xi_1\rho_c; Q = H\xi_2\rho_d; Q = H\xi(\rho_d + \rho_c).$$

$$\frac{dr}{dt} = -3\Gamma Hr, \Gamma = -\omega_d - \xi^2 \frac{(\rho_c + \rho_d)^2}{3\rho_c\rho_d},$$

$$r_s^\pm = -1 + 2b \pm 2\sqrt{b(b-1)}, b = -\frac{3\omega_d}{4\xi} > 1$$

$$\dot{r} = H[\xi_1(1+r) + 3\omega_d]; \\ r = -(3\omega_d + \xi_1)r_0/\{\xi_1r_0 - (1+z)^{-(3\omega_d+\xi_1)}[\xi_1(1+r_0) + 3\omega_d]\},$$

$$\chi = \frac{1}{aH^3} \frac{d^3a}{dt^3}, s = \frac{\chi - 1}{3\left(q - \frac{1}{2}\right)},$$

$$\chi = 1 + \frac{9}{2} \frac{\omega_d}{1 + r_0(1+z)^\zeta} \left[1 + \omega_d - \left(\omega_d + \frac{\zeta}{3}\right) \frac{r_0(1+z)^\zeta}{1 + r_0(1+z)^\zeta} \right] \\ s = 1 + \omega_d - \left(\omega_d + \frac{\zeta}{3}\right) \frac{r_0(1+z)^\zeta}{1 + r_0(1+z)^\zeta}$$

$$\mathcal{L} = \sqrt{-g}\bar{\Upsilon}(i\Big/\mathcal{D}-m)\Upsilon + \text{ } \mathfrak{N}_{\text{non derivative interactions}}$$

$$T_{\mu\nu}=\frac{i}{4}\big(\bar{\Upsilon}\gamma_\mu\nabla_\nu\Upsilon+\bar{\Upsilon}\gamma_\nu\nabla_\mu\Upsilon-\nabla_\mu\bar{\Upsilon}\gamma_\nu\Upsilon-\nabla_\nu\bar{\Upsilon}\gamma_\mu\Upsilon\big)$$

$$\mathcal{L} = \bar{\Upsilon}(i\,\partial)\Upsilon + \mathcal{L}_s(\varphi) + F(\varphi)\bar{\Upsilon}\Upsilon$$

$$\mathcal{L}_s(\varphi)=\ell\frac{1}{2}\partial^\mu\varphi\partial_\mu\varphi-V(\varphi)$$

$$H^2=\frac{8\pi G}{3}\Big(\rho_r+\rho_b+\rho_c+\frac{1}{2}\dot{\varphi}^2+V(\varphi)\Big)$$

$$\begin{aligned}\dot{\rho}_c+3H\rho_c &= -\rho_c\dot{\varphi}\varrho/(1-\varrho\varphi) \\ \ddot{\varphi}+3H\dot{\varphi}+V'(\varphi) &= \rho_c\varrho/(1-\varrho\varphi)\end{aligned}$$

$$\mathcal{L}=p(\varphi,X),X=\frac{1}{2}\big(D_\mu\varphi D^\mu\varphi\big)$$

$$\mathcal{L}_{\text{tach}}=-V(\varphi)\sqrt{(1-\alpha\partial^\mu\varphi\partial_\mu\varphi)}$$

$$\mathcal{L}=\mathcal{L}_{\text{tach}}+\frac{i}{2}\big[\bar{\Upsilon}\gamma^\mu\nabla_\mu\Upsilon-\bar{\Upsilon}\nabla_\mu\gamma^\mu\Upsilon\big]-F(\varphi)\bar{\Upsilon}\Upsilon$$

$$\begin{aligned}i\gamma^\mu\nabla_\mu\Upsilon-(M-\beta\varphi)\Upsilon &= 0 \\ \alpha\nabla_\mu\partial^\mu\varphi+\alpha^2\frac{\partial^\mu\varphi(\nabla_\mu\partial_\sigma\varphi)\partial^\sigma\varphi}{1-\alpha\partial_\mu\varphi\partial^\mu\varphi}+\frac{d\ln V(\varphi)}{d\varphi} &= \frac{\beta\bar{\Upsilon}\Upsilon}{V(\varphi)}\sqrt{1-\alpha\partial^\mu\varphi\partial_\mu\varphi}\end{aligned}$$

$$\ddot{\varphi}=-(1-\alpha\dot{\varphi}^2)\left[\frac{1}{\alpha}\frac{d\ln V(\varphi)}{d\varphi}+3H\dot{\varphi}-\frac{\beta\bar{\Upsilon}\Upsilon}{\alpha V(\varphi)}\sqrt{1-\alpha\dot{\varphi}^2}\right]$$

$$\frac{d(a^3\bar{\Upsilon}\Upsilon)}{dt}=0,\Rightarrow \bar{\Upsilon}\Upsilon=\bar{\Upsilon}_0\Upsilon_0a^{-3}$$

$$\begin{aligned}\rho_\varphi &= \frac{V(\varphi)}{\sqrt{1-\alpha\dot{\varphi}^2}}, & p_\varphi &= -V(\varphi)\sqrt{1-\alpha\dot{\varphi}^2} \\ \rho_\Upsilon &= (M-\beta\varphi)\bar{\Upsilon}\Upsilon, & p_\Upsilon &= 0\end{aligned}$$

$$\begin{aligned}\dot{\rho}_\varphi+3H\rho_\varphi\big(\omega_\varphi+1\big) &= \beta\dot{\varphi}\bar{\Upsilon}_0\Upsilon_0a^{-3} \\ \dot{\rho}_\Upsilon+3H\rho_\Upsilon &= -\beta\dot{\varphi}\bar{\Upsilon}_0\Upsilon_0a^{-3}\end{aligned}$$

$$H^2=\frac{8\pi G}{3}\bigg[\rho_r+\rho_b+(M-\beta\varphi)\bar{\Upsilon}_0\Upsilon_0a^{-3}+\frac{V(\varphi)}{\sqrt{1-\alpha\dot{\varphi}^2}}\bigg]$$

$$\rho_d=\frac{3\wp}{8\pi GL^2}$$

$$\dot{r}=3Hr\left[\omega_d+\frac{1+r}{r}\frac{\xi_1}{3H}\right]$$



$$\omega_d = -\left(1+\frac{1}{r}\right)\left[\frac{\xi_1}{3H} + \frac{\dot{\varphi}}{3H_{\varphi}}\right]$$

$$S=\int~d^4x\sqrt{-g}\Big(-\frac{R}{4}+\frac{1}{2}\partial_\mu\varphi\partial^\mu\varphi-V(\varphi)\\+\frac{i}{2}\big[\bar{Y}\gamma^\mu\nabla_\mu Y-\bar{Y}\nabla_\mu\gamma^\mu Y\big]-F(\varphi)\bar{Y}Y\Big)$$

$$\ddot{\varphi}+3H\dot{\varphi}+V'=-F'\bar{Y}Y\\ H^2=\frac{1}{3M_p^2}\biggl\{\frac{\dot{\varphi}^2}{2}+V(\varphi)+F(\varphi)\bar{Y}Y\biggr\}$$

$$\dot{H}=-\frac{1}{2M_p^2}\{\dot{\varphi}^2+F(\varphi)\bar{Y}Y\}$$

$$-W_\varphi\dot{\varphi}2M_p^2=\dot{\varphi}^2+F(\varphi)\frac{\bar{Y}_0Y_0}{a^3}$$

$$\dot{\varphi}^{n-1}+\frac{\left[\dot{\varphi}+2M_p^2W_\varphi\right]}{F(\varphi)\bar{Y}_0Y_0\sigma J(\varphi)}=0$$

$$\ddot{\varphi}+3H_0\dot{\varphi}+V'=\frac{\beta\bar{Y}_0Y_0}{a^3}$$

$$\varphi(t)=K_1+K_2e^{-3H_0t}+K_3e^{-\frac{3}{2}H_0t}$$

$$\varphi(t)=Y_1+Y_2\left[\frac{(\ln t)^2}{2}+Y_3\ln t\right]$$

$$F(\varphi)=-C_1\frac{{\rm e}^{\left(3\varphi^2/8M_p^2\right)}}{\varphi^4}, V_3(\varphi)=\frac{3\mu^8}{4M_p^2\varphi^2}\\ \varphi(t)=(6\mu^4t)^{1/3}, a(t)=\left(\frac{\bar{Y}_0Y_0C_1}{2\mu^8}\right)^{1/3}e^{\left(6\mu^4t\right)^{2/3}/8M_p^2}$$

$$\rho_d(a)=\frac{\mu^8}{\left(32M_p^4\right)}\frac{1}{\ln^2\left(\gamma a\right)}\left(1+3\ln\left(\gamma a\right)\right)\\ p_a(a)=\frac{\mu^8}{\left(32M_p^4\right)}\frac{1}{\left(\ln\left(\gamma a\right)\right)^2}\left(1-3\ln\left(\gamma a\right)\right)\\ w_d(a)=\frac{1-3\ln\left(\gamma a\right)}{1+3\ln\left(\gamma a\right)}$$

$$H^2=\frac{8\pi G}{3}[X(f+2f_1)+\rho_c+\rho_r]\\ \dot{H}=-4\pi G\left[2X(f+f_1)+\rho_c+\frac{4}{3}\rho_r\right]\\ \ddot{\varphi}=-3AH(f+f_1)\dot{\varphi}-\lambda X\big(1-A(f+2f_1)\big)-AQ\rho_c$$



$$Q=-\frac{1}{\rho_c \sqrt{-g}}\frac{\partial \mathcal{L}_m}{\partial \varphi}$$

$$\begin{array}{l}x=\frac{\sqrt{4\pi G}\dot{\varphi}}{3H},y=\frac{\sqrt{8\pi G}e^{-\lambda\varphi/2}}{3H},z=\frac{\sqrt{8\pi G\rho_r}}{3H},\\\Omega_c=\frac{8\pi G\rho_c}{9H^2}=1-\Omega_{\varphi}-z^2,\Omega_{\varphi}=x^2(f+2f_1).\end{array}$$

$$ds^2=a^2(\tau)\bigl[-(1+2\psi)d\tau^2+2\partial_iBd\tau dx^i+(1+2\phi)\delta_{ij}dx^idx^j+D_{ij}Edx^idx^j\bigr]$$

$$\begin{aligned}\delta\nabla_\mu T_{(\lambda)}^{\mu 0}=&\frac{1}{a^2}\{-2[\rho'_\lambda+3\mathcal{H}(p_\lambda+\rho_\lambda)]\psi+\delta\rho'_\lambda+(p_\lambda+\rho_\lambda)\theta_\lambda\\&+3\mathcal{H}(\delta p_\lambda+\delta\rho_\lambda)+3(p_\lambda+\rho_\lambda)\phi'\}=\delta Q_\lambda^0\\\partial_i\delta\nabla_\mu T_{(\lambda)}^{\mu i}=&\frac{1}{a^2}\{[p'_\lambda+\mathcal{H}(p_\lambda+\rho_\lambda)]\nabla^2B+[(p'_\lambda+\rho'_\lambda)+4\mathcal{H}(p_\lambda+\rho_\lambda)]\theta_\lambda\\&+(p_\lambda+\rho_\lambda)\nabla^2B'+\nabla^2\delta p_\lambda+(p_\lambda+\rho_\lambda)\theta'_\lambda+(p_\lambda+\rho_\lambda)\nabla^2\psi\}\\&=\partial_i\delta Q_{(\lambda)}^i\\-4\pi Ga^2\delta\rho=&\nabla^2\phi+3\mathcal{H}(\mathcal{H}\psi-\phi')+ \mathcal{H}\nabla^2B-\frac{1}{6}[\nabla^2]^2E\\-4\pi Ga^2(\rho+p)\theta=&\mathcal{H}\nabla^2\psi-\nabla^2\phi'+2\mathcal{H}^2\nabla^2B-\frac{a''}{a}\nabla^2B+\frac{1}{6}[\nabla^2]^2E'\\8\pi Ga^2\Pi_j^i=&-\partial^i\partial_j\psi-\partial^i\partial_j\phi+\frac{1}{2}\partial^i\partial_jE''+\mathcal{H}\partial^i\partial_jE'+\frac{1}{6}\partial^i\partial_j\nabla^2E\\-2\mathcal{H}\partial^i\partial_jB-\partial^i\partial_jB\end{aligned}$$

$$\tilde{x}^\mu=x^\mu+\delta x^\mu, \delta x^0=\xi^0(x^\mu), \delta x^i=\partial^i\beta(x^\mu)+v_*^i(x^\mu)$$

$$\begin{array}{ll}\tilde{\psi}=\psi-\xi^{0'}-\dfrac{a'}{a}\xi^0,&\tilde{B}=B+\xi^0-\beta',\\\tilde{\phi}=\phi-\dfrac{1}{3}\nabla^2\beta-\dfrac{a'}{a}\xi^0,&\tilde{E}=E-2\beta\\\tilde{\nu}=\nu+\beta',&\tilde{\theta}=\theta+\nabla^2\beta'\end{array}$$

$$\tilde{\delta Q^0} = \delta Q^0 - Q^{0'} \xi^0 + Q^0 \xi^{0'}, \tilde{\delta Q_p} = \delta Q_p + Q^0 \beta'$$

$$\delta Q^i = \partial^i \delta Q_p + \delta Q^i_*$$

$$\mathcal{L}_{\delta x}Q^\nu=\delta x^\sigma Q^\nu_{,\sigma}-Q^\sigma\delta x^\nu_{,\sigma},\delta\tilde{Q}^\nu=\delta Q^\nu-\mathcal{L}_{\delta x}Q^\nu$$

$$\begin{array}{ll}\tilde{\psi}Y^{(s)}=\left(\psi-\xi^{0'}-\dfrac{a'}{a}\xi^0\right)Y^{(s)}&\tilde{B}Y_i^{(s)}=(B-k\xi^0-\beta')Y_i^{(s)}\\\tilde{\phi}Y^{(s)}=\left(\phi-\dfrac{1}{3}k\beta-\dfrac{a'}{a}\xi^0\right)Y^{(s)}&\tilde{E}Y_{ij}^{(s)}=(E+2k\beta)Y_{ij}^{(s)}\\\tilde{\theta}Y^{(s)}=(\theta+k\beta')Y^{(s)}&\end{array}$$

$$\begin{aligned}
& \delta_\lambda' + 3\mathcal{H} \left(\frac{\delta p_\lambda}{\delta \rho_\lambda} - \omega_\lambda \right) \delta_\lambda = -(1 + \omega_\lambda) k v_\lambda - 3(1 + \omega_\lambda) \phi' \\
& + (2\psi - \delta_\lambda) \frac{a^2 Q_\lambda^0}{\rho_\lambda} + \frac{a^2 \delta Q_\lambda^0}{\rho_\lambda} \\
& (v_\lambda + B)' + \mathcal{H}(1 - 3\omega_\lambda)(v_\lambda + B) = \frac{k}{1 + \omega_\lambda} \frac{\delta p_\lambda}{\delta \rho_\lambda} \delta_\lambda - \frac{\omega_\lambda'}{1 + \omega_\lambda} (v_\lambda + B) \\
& + k\psi - \frac{a^2 Q_\lambda^0}{\rho_\lambda} v_\lambda - \frac{\omega_\lambda a^2 Q_\lambda^0}{(1 + \omega_\lambda) \rho_\lambda} B + \frac{a^2 \delta Q_{p\lambda}}{(1 + \omega_\lambda) \rho_\lambda} \\
& \Psi = \psi - \frac{1}{k} \mathcal{H} \left(B + \frac{E'}{2k} \right) - \frac{1}{k} \left(B' + \frac{E''}{2k} \right), \Phi = \phi + \frac{1}{6} E - \frac{1}{k} \mathcal{H} \left(B + \frac{E'}{2k} \right), \\
& \delta Q_\lambda^{0I} = \delta Q_\lambda^0 - \frac{Q_\lambda^{0'}}{\mathcal{H}} \left(\phi + \frac{E}{6} \right) + Q_\lambda^0 \left[\frac{1}{\mathcal{H}} \left(\phi + \frac{E}{6} \right) \right]', V_\lambda = v_\lambda - \frac{E'}{2k}, \\
& \delta Q_{p\lambda}^I = \delta Q_{p\lambda} - Q_\lambda^0 \frac{E'}{2k}, D_\lambda = \delta_\lambda - \frac{\rho_\lambda'}{\rho_\lambda \mathcal{H}} \left(\phi + \frac{E}{6} \right), \\
& D_c' + \left\{ \left(\frac{a^2 Q_c^0}{\rho_c \mathcal{H}} \right)' + \frac{\rho_c'}{\rho_c \mathcal{H}} \frac{a^2 Q_c^0}{\rho_c} \right\} \Phi + \frac{a^2 Q_c^0}{\rho_c} D_c + \frac{a^2 Q_c^0}{\rho_c \mathcal{H}} \Phi' = -kV_c \\
& + 2\Psi \frac{a^2 Q_c^0}{\rho_c} + \frac{a^2 \delta Q_c^{0I}}{\rho_c} + \frac{a^2 Q_c^{0'}}{\rho_c \mathcal{H}} \Phi - \frac{a^2 Q_c^0}{\rho_c} \left(\frac{\Phi}{\mathcal{H}} \right)', \\
& V_c' + \mathcal{H} V_c = k\Psi - \frac{a^2 Q_c^0}{\rho_c} V_c + \frac{a^2 \delta Q_{pc}^I}{\rho_c}, \\
& D_d' + \left\{ \left(\frac{a^2 Q_d^0}{\rho_d \mathcal{H}} \right)' - 3\omega_d' + 3(C_e^2 - \omega_d) \frac{\rho_d'}{\rho_d} + \frac{\rho_d'}{\rho_d \mathcal{H}} \frac{a^2 Q_d^0}{\rho_d} \right\} \Phi \\
& + \left\{ 3\mathcal{H}(C_e^2 - \omega_d) + \frac{a^2 Q_d^0}{\rho_d} \right\} D_d + \frac{a^2 Q_d^0}{\rho_d \mathcal{H}} \Phi', \\
& = -(1 + \omega_d) k V_d + 3\mathcal{H}(C_e^2 - C_a^2) \frac{\rho_d'}{\rho_d} \frac{V_d}{k} + 2\Psi \frac{a^2 Q_d^0}{\rho_d} + \frac{a^2 \delta Q_d^{0I}}{\rho_d} \\
& + \frac{a^2 Q_d^{0'}}{\rho_d \mathcal{H}} \Phi - \frac{a^2 Q_d^0}{\rho_d} \left(\frac{\Phi}{\mathcal{H}} \right)', \\
& V_d' + \mathcal{H}(1 - 3\omega_d) V_d = \frac{k C_e^2}{1 + \omega_d} D_d + \frac{k C_e^2}{1 + \omega_d} \frac{\rho_d'}{\rho_d \mathcal{H}} \Phi - \frac{a^2 Q_d^0}{\rho_d} V_d \\
& - (C_e^2 - C_a^2) \frac{V_d}{1 + \omega_d} \frac{\rho_d'}{\rho_d} - \frac{\omega_d'}{1 + \omega_d} V_d + k\Psi + \frac{a^2 \delta Q_{pd}^I}{(1 + \omega_d) \rho_d} \\
& \frac{\delta p_d}{\rho_d} = C_e^2 \delta_d - (C_e^2 - C_a^2) \frac{\rho_d'}{\rho_d} \frac{v_d + B}{k}
\end{aligned}$$

$$U_\lambda = (1 + \omega_d) V_\lambda$$



$$D'_c + \left\{ \left(\frac{a^2 Q_c^0}{\rho_c \mathcal{H}} \right)' + \frac{\rho'_c}{\rho_c \mathcal{H}} \frac{a^2 Q_c^0}{\rho_c} \right\} \Phi + \frac{a^2 Q_c^0}{\rho_c} D_c + \frac{a^2 Q_c^0}{\rho_c \mathcal{H}} \Phi' = \\ -k U_c + 2\Psi \frac{a^2 Q_c^0}{\rho_c} + \frac{a^2 \delta Q_c^{0I}}{\rho_c} + \frac{a^2 Q_c^{0'}}{\rho_c \mathcal{H}} \Phi - \frac{a^2 Q_c^0}{\rho_c} \left(\frac{\Phi}{\mathcal{H}} \right)' \\ U'_c + \mathcal{H} U_c = k\Psi - \frac{a^2 Q_c^0}{\rho_c} U_c + \frac{a^2 \delta Q_{pc}^I}{\rho_c}$$

$$D'_d + \left\{ \left(\frac{a^2 Q_d^0}{\rho_d \mathcal{H}} \right)' - 3\omega'_d + 3(C_e^2 - \omega_d) \frac{\rho'_d}{\rho_d} + \frac{\rho'_d}{\rho_d \mathcal{H}} \frac{a^2 Q_d^0}{\rho_d} \right\} \Phi \\ + \left\{ 3\mathcal{H}(C_e^2 - \omega_d) + \frac{a^2 Q_d^0}{\rho_d} \right\} D_d + \frac{a^2 Q_d^0}{\rho_d \mathcal{H}} \Phi' = 3\mathcal{H}(C_e^2 - C_a^2) \frac{\rho'_d}{\rho_d} \frac{U_d}{(1 + \omega_d)k} \\ -k U_d + 2\Psi \frac{a^2 Q_d^0}{\rho_d} + \frac{a^2 \delta Q_d^{0I}}{\rho_d} + \frac{a^2 Q_d'}{\rho_d \mathcal{H}} \Phi - \frac{a^2 Q_d^0}{\rho_d} \left(\frac{\Phi}{\mathcal{H}} \right)' \\ U'_d + \mathcal{H}(1 - 3\omega_d) U_d = kC_e^2 D_d + kC_e^2 \frac{\rho'_d}{\rho_d \mathcal{H}} \Phi - (C_e^2 - C_a^2) \frac{U_d}{1 + \omega_d} \frac{\rho'_d}{\rho_d} \\ +(1 + \omega_d)k\Psi - \frac{a^2 Q_d^0}{\rho_d} U_d + \frac{a^2 \delta Q_{pd}^I}{\rho_d}$$

$$\Phi = \frac{4\pi G a^2 \sum \rho_i \{ D^i + 3\mathcal{H} U^i / k \}}{k^2 - 4\pi G a^2 \sum \rho'_i / \mathcal{H}}.$$

$$Q_{(\lambda)}^\mu = Q_{(\lambda\eta)} U_{(\eta)}^\mu + F_{(\lambda\eta)}^\mu$$

$$Q_{(\lambda)}^\nu = \left[\frac{Q_{(\lambda)}}{a}, 0, 0, 0 \right]^T$$

$$Q_{(\lambda)}{}^\mu Q_{(\lambda)\mu} = g_{00} \left(Q_{(\lambda)}{}^0 \right)^2 = -Q_{(\lambda)}{}^2$$

$$\delta Q_{(\lambda)}^0 = -\frac{\psi}{a} Q_{(\lambda)} + \frac{1}{a} \delta Q_{(\lambda)}$$

$$\delta Q_{p\lambda} = \delta Q_{p\lambda}^I \Big|_t + Q_{(\lambda)}^0 v_t$$

$$Q_c = -Q_d = 3H(\xi_1 \rho_c + \xi_2 \rho_d)$$

$$\delta Q_c = -\delta Q_d = 3H(\xi_1 \delta \rho_c + \xi_2 \delta \rho_d) \\ \delta Q_c^0 = -\delta Q_d^0 = -3H(\xi_1 \rho_c + \xi_2 \rho_d) \frac{\psi}{a} + 3H(\xi_1 \delta \rho_c + \xi_2 \delta \rho_d) \frac{1}{a}$$

$$\frac{a^2 \delta Q_c^{0I}}{\rho_c} = -3\mathcal{H}(\xi_1 + \xi_2/r) \Psi + 3\mathcal{H}\{\xi_1 D_c + \xi_2 D_d/r\} + 3 \left(\xi_1 \frac{\rho'_c}{\rho_c} + \frac{\xi_2}{r} \frac{\rho'_d}{\rho_d} \right) \Phi \\ - \frac{a^2 Q_c^{0'}}{\rho_c \mathcal{H}} \Phi + \frac{a^2 Q_c^0}{\rho_c} \left[\frac{\Phi}{\mathcal{H}} \right]', \\ \frac{a^2 \delta Q_d^{0I}}{\rho_d} = 3\mathcal{H}(\xi_1 r + \xi_2) \Psi - 3\mathcal{H}\{\xi_1 D_c r + \xi_2 D_d\} - 3 \left(\xi_1 r \frac{\rho'_c}{\rho_c} + \frac{\rho'_d}{\rho_d} \xi_2 \right) \Phi$$

$$-\frac{a^2 Q_d^{0'}}{\rho_d} \Phi + \frac{a^2 Q_d^0}{\rho_d} \left[\frac{\Phi}{\mathcal{H}} \right]'$$

$$\begin{aligned} D'_c &= -kU_c + 3\mathcal{H}\Psi(\xi_1 + \xi_2/r) - 3(\xi_1 + \xi_2/r)\Phi' + 3\mathcal{H}\xi_2(D_d - D_c)/r \\ U'_c &= -\mathcal{H}U_c + k\Psi - 3\mathcal{H}(\xi_1 + \xi_2/r)U_c \\ D'_d &= -3\mathcal{H}(C_e^2 - \omega_d)D_d + \{3\omega_d' - 9\mathcal{H}(\omega_d - C_e^2)(\xi_1 r + \xi_2 + 1 + \omega_d)\}\Phi \\ &\quad - 9\mathcal{H}^2(C_e^2 - C_a^2)\frac{U_d}{k} + 3(\xi_1 r + \xi_2)\Phi' - 3\Psi\mathcal{H}(\xi_1 r + \xi_2) \\ &\quad - 9\mathcal{H}^2(C_e^2 - C_a^2)(\xi_1 r + \xi_2)\frac{U_d}{(1 + \omega_d)k} - kU_d + 3\mathcal{H}\xi_1 r(D_d - D_c) \\ U'_d &= -\mathcal{H}(1 - 3\omega_d)U_d - 3kC_e^2(\xi_1 r + \xi_2 + 1 + \omega_d)\Phi \\ &\quad + 3\mathcal{H}(C_e^2 - C_a^2)(\xi_1 r + \xi_2)\frac{U_d}{(1 + \omega_d)} + 3(C_e^2 - C_a^2)\mathcal{H}U_d \\ &\quad + kC_e^2D_d + (1 + \omega_d)k\Psi + 3\mathcal{H}(\xi_1 r + \xi_2)U_d \end{aligned}$$

$$\frac{D_c}{1 - \xi_1 - \xi_2/r} = \frac{D_d}{1 + \omega_d + \xi_1 r + \xi_2}$$

$$\begin{aligned} D'_d &\approx (-1 + \omega_d + \xi_1 r)3\mathcal{H}D_d - 9\mathcal{H}^2(1 - \omega_d)\left(1 + \frac{\xi_1 r + \xi_2}{1 + \omega_d}\right)\frac{U_d}{k} \\ U'_d &\approx 2\left[1 + \frac{3}{1 + \omega_d}(\xi_1 r + \xi_2)\right]\mathcal{H}U_d + kD_d \end{aligned}$$

$$D'_d \approx -3\mathcal{H}D_d - 9\mathcal{H}^2\frac{1 - \omega_d}{1 + \omega_d}\frac{U_d}{k}, U'_d \approx 2\frac{1 - 2\omega_d}{1 + \omega_d}\mathcal{H}U_d + kD_d$$

$$D''_d \approx \left(2\frac{\mathcal{H}'}{\mathcal{H}} - \frac{1 + 7\omega_d}{1 + \omega_d}\mathcal{H}\right)D'_d + 3(\mathcal{H}' - \mathcal{H}^2)D_d$$

$$\mathcal{H} \sim \tau^{-1}, \mathcal{H}' \sim -\tau^{-2}, (\mathcal{H}'/\mathcal{H}) \sim -\tau^{-1} \quad D''_d \approx -3\frac{1+3\omega_d}{1+\omega_d}\frac{D'_d}{\tau} - \frac{6}{\tau^2}D_d, \quad D_d \approx C_1\tau^{r_1} + C_2\tau^{r_2}, \quad r_1 =$$

$$-\frac{1+4\omega_d-\sqrt{-5-4\omega_d+10\omega_d^2}}{1+\omega_d}, r_2 = -\frac{1+4\omega_d+\sqrt{-5-4\omega_d+10\omega_d^2}}{1+\omega_d}$$

$$\begin{aligned} D'_d &\approx (-1 + \omega_d)3\mathcal{H}D_d - 9\mathcal{H}^2(1 - \omega_d)\left(1 + \frac{\xi_2}{1 + \omega_d}\right)\frac{U_d}{k}, \\ U'_d &\approx 2\left(1 + \frac{3\xi_2}{1 + \omega_d}\right)\mathcal{H}U_d + kD_d. \end{aligned}$$

$$\begin{aligned} D''_d &= \left[\left(-1 + 3\omega_d + \frac{6\xi_2}{1 + \omega_d}\right)\mathcal{H} + 2\frac{\mathcal{H}'}{\mathcal{H}}\right]D'_d \\ &\quad + 3(1 - \omega_d)\left[\mathcal{H}' + \mathcal{H}^2\left(-1 + \frac{3\xi_2}{1 + \omega_d}\right)\right]D_d \end{aligned}$$

$$D''_d = \left(-3 + 3\omega_d + \frac{6\xi_2}{1 + \omega_d}\right)\frac{D'_d}{\tau} + 3(1 - \omega_d)\left(-2 + \frac{3\xi_2}{1 + \omega_d}\right)\frac{D_d}{\tau^2}$$



$$D_d \sim C_1 \tau^{r_1} + C_2 \tau^{r_2}$$

$$r_1 = \frac{1}{2} \frac{\Gamma}{1 + \omega_d} + \frac{1}{2} \frac{\sqrt{\Delta}}{1 + \omega_d}, r_2 = \frac{1}{2} \frac{\Gamma}{1 + \omega_d} - \frac{1}{2} \frac{\sqrt{\Delta}}{1 + \omega_d}$$

$$D_d \sim C_1 \tau^{\frac{1}{2(1+\omega_d)}} \cos \frac{1}{2} \frac{\sqrt{|\Delta|}}{1 + \omega_d} \ln \tau + C_2 \tau^{\frac{1}{2(1+\omega_d)}} \sin \frac{1}{2} \frac{\sqrt{|\Delta|}}{1 + \omega_d} \ln \tau$$

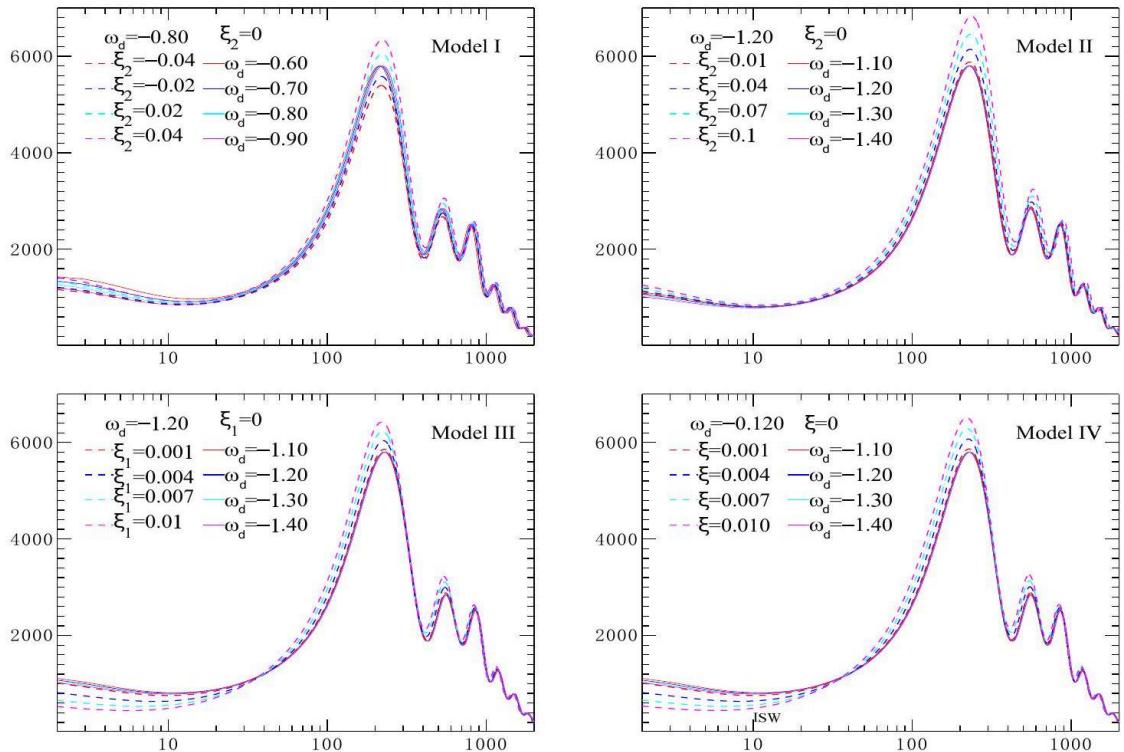
$$\mathcal{C}_\ell = 4\pi \int \frac{dk}{k} \mathcal{P}_\chi(k) |\Delta_\ell(k, \tau_0)|^2$$

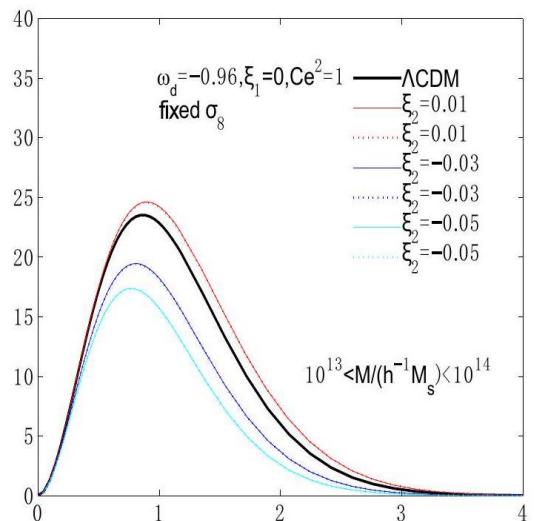
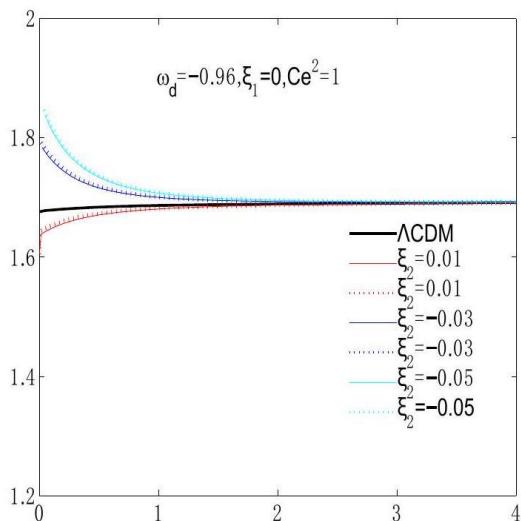
$$\Delta_\ell(k, \tau_0) = \Delta_\ell^{SW}(k) + \Delta_\ell^{ISW}(k),$$

$$\Delta_\ell^{ISW} = \int_{\tau_i}^{\tau_0} d\tau j_\ell(k[\tau_0 - \tau]) e^{\kappa(\tau_0) - \kappa(\tau)} [\Psi' - \Phi']$$

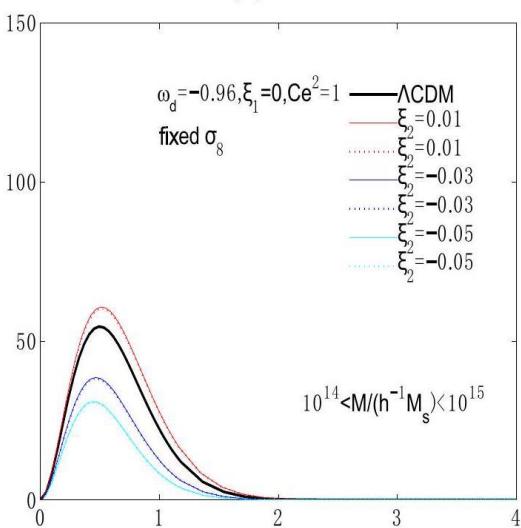
$$\Psi' - \Phi' = 2\mathcal{H} \left[\Phi + 4\pi G a^2 \sum_i U^i \rho^i / (\mathcal{H} k) + \mathcal{T} \right] - \mathcal{T}'$$

$$\Phi' = -\mathcal{H}\Phi - \mathcal{H}\mathcal{T} - 4\pi G a^2 \sum_i U^i \rho^i / k, \mathcal{T} = \frac{8\pi G a^2}{k^2} \{ p^\gamma \Pi^\gamma + p^\nu \Pi^\nu \}$$

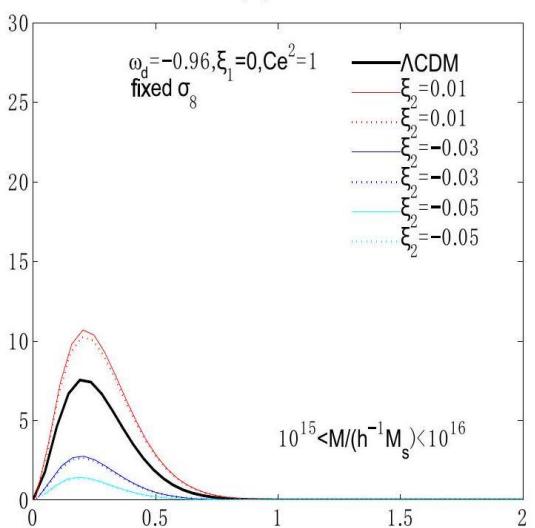




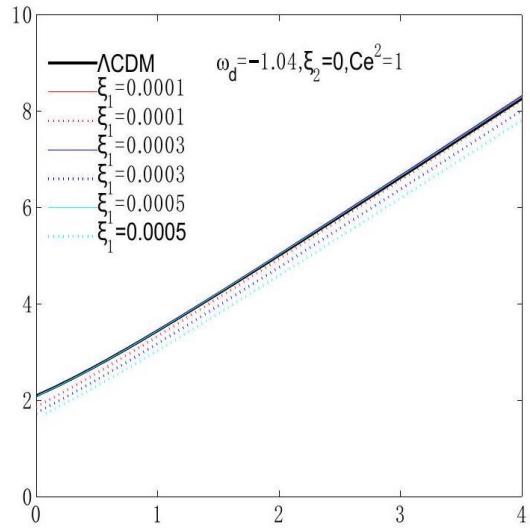
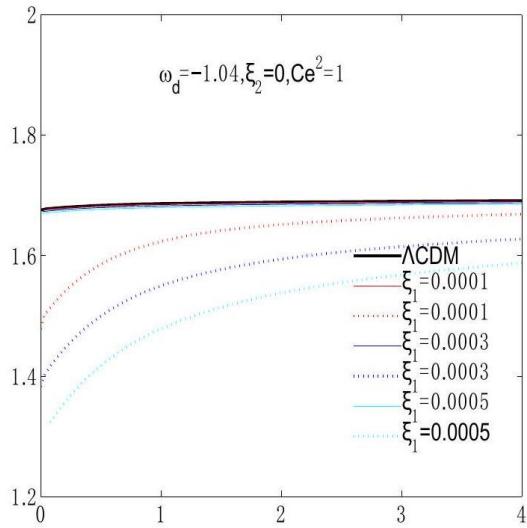
(a)



(c)

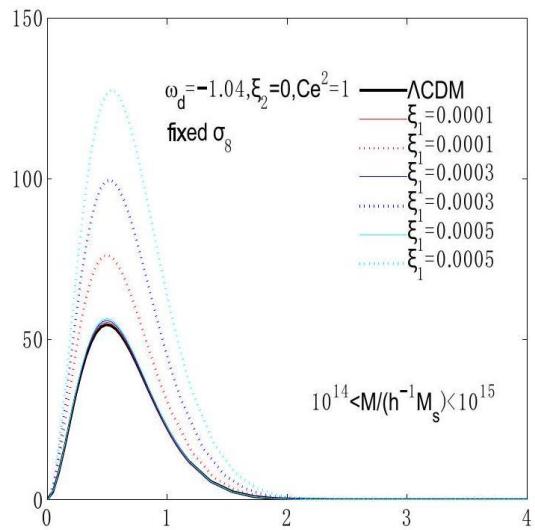
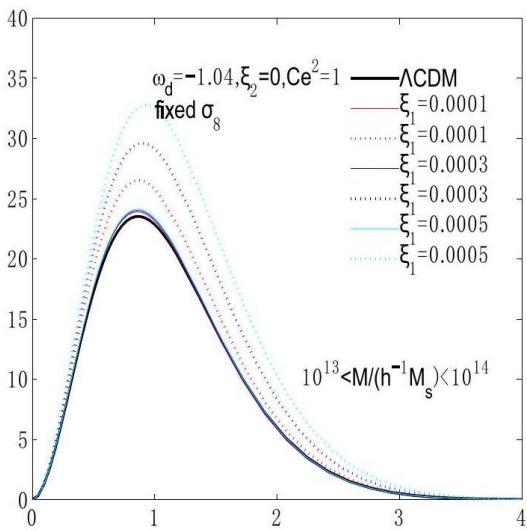


(d)



(a)

(b)



(c)

(d)

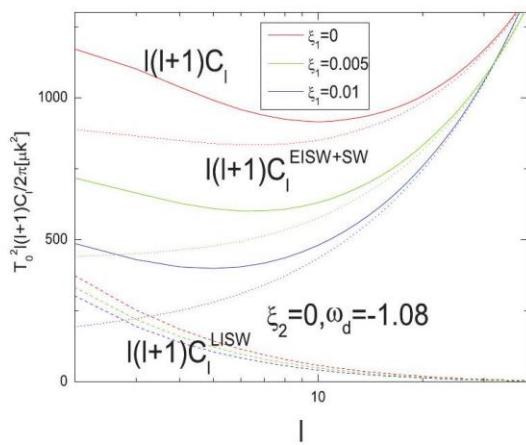
Figuras 23, 23A y 23B. Fluctuaciones espectrales de una partícula oscura en relación a sectores oscuros.

$$\begin{aligned} C_l^{gg} &= 4\pi \int \frac{dk}{k} \mathcal{P}_\chi(k) I_l^g(k) I_l^g(k) \\ C_l^{gI} &= 4\pi \int \frac{dk}{k} \mathcal{P}_\chi(k) I_l^g(k) \Delta_l^{ISW}(k), \end{aligned}$$

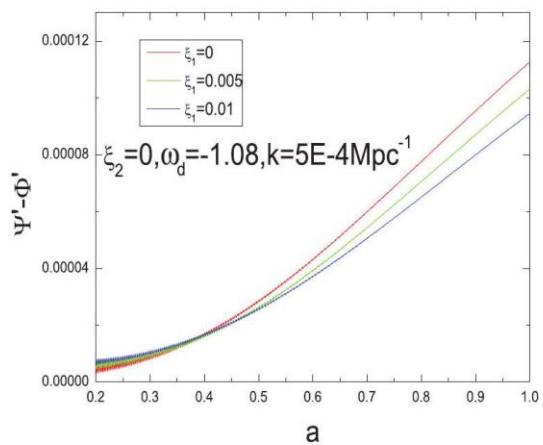
$$I_l^g(k) = \int dz b_g(z) n(z) (D_c + D_b) j_l[k\chi(z)], \quad z = 0, \quad \chi(z) = \int_0^z \frac{dz'}{H(z')} = \int_{\tau_i(z)}^{\tau_0} d\tau = \tau_0 - \tau_i(z) \quad n(z) =$$

$$\frac{3}{2} \frac{z^2}{z_0^3} \exp \left[- \left(\frac{z}{z_0} \right)^{3/2} \right] \int n(z) dz = 1$$

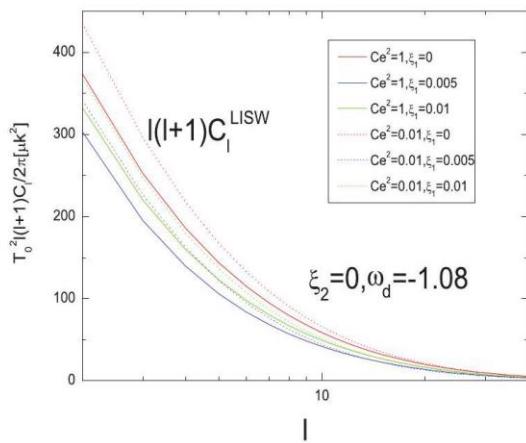




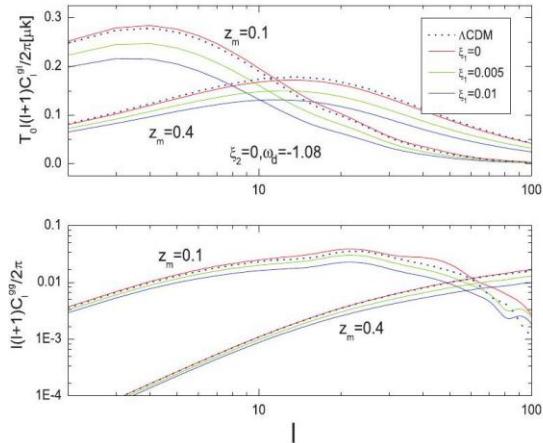
(a)



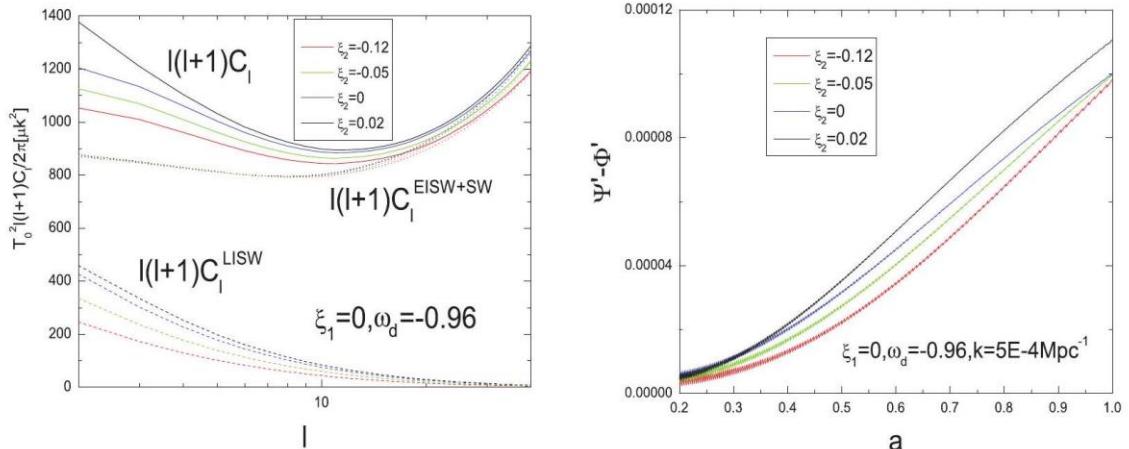
(b)



(c)

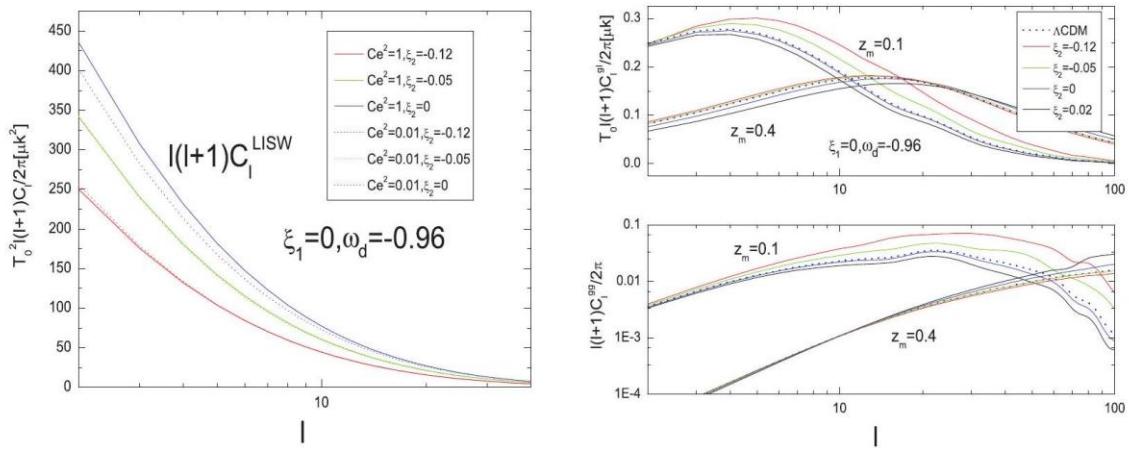


(d)



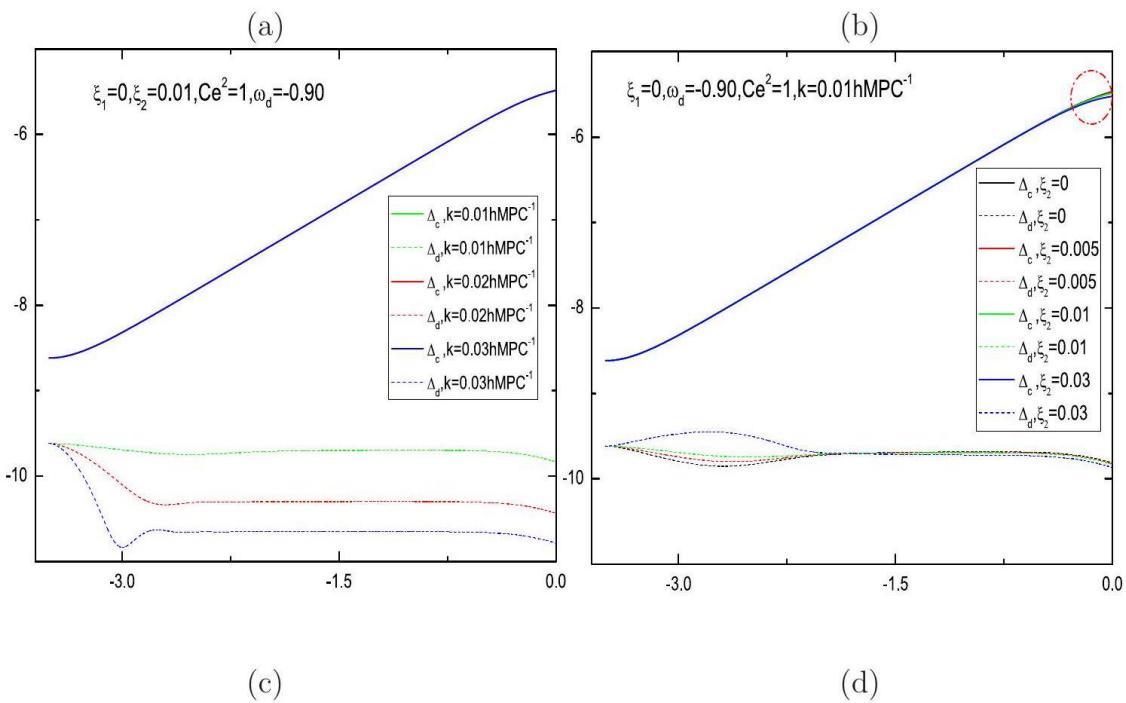
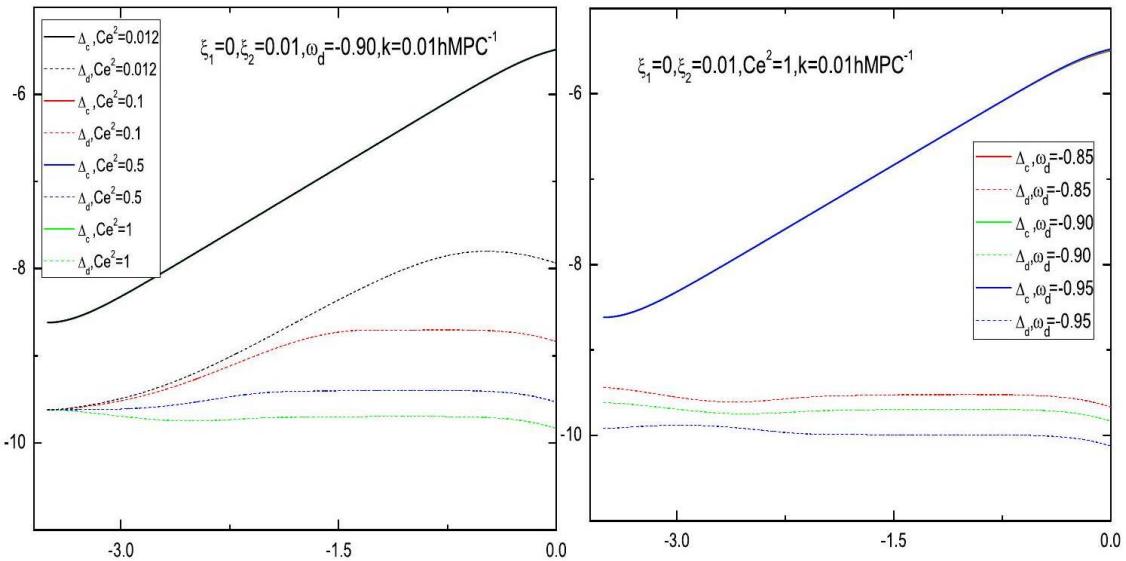
(a)

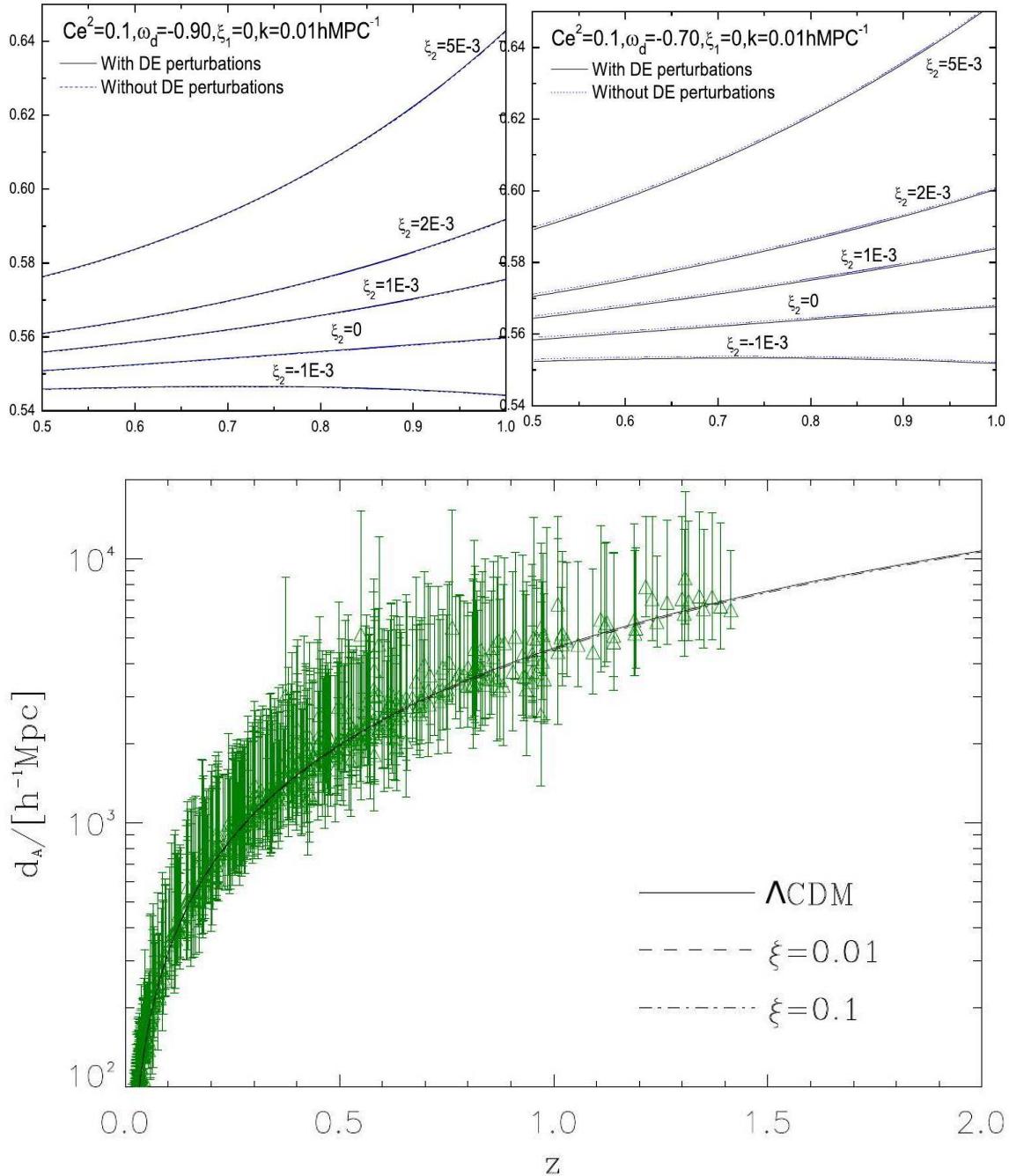
(b)



(c)

(d)





Figuras 24, 25, 26, 27 y 28. Campos oscuros perturbativos por interacción de una partícula oscura.

$$\delta\rho_\lambda^I = \delta\rho_\lambda - \rho'_\lambda \frac{v_\lambda + B}{k}, \delta p_\lambda^I = \delta p_\lambda - p'_\lambda \frac{v_\lambda + B}{k}$$

$$\Delta_\lambda = \delta_\lambda - \frac{\rho'_\lambda}{\rho_\lambda} \frac{v_\lambda + B}{k}, V_\lambda = v_\lambda - \frac{E'}{2k}$$

$$\Phi = 4\pi G \frac{a^2}{k^2} \sum_\lambda \left(\Delta_\lambda + \frac{a^2 Q_\lambda^0 V_\lambda}{\rho_\lambda} \right) \rho_\lambda,$$

$$k(\mathcal{H}\Psi - \Phi') = 4\pi G a^2 \sum_\lambda (\rho_\lambda + p_\lambda) V_\lambda, \Psi = -\Phi,$$

$$\begin{aligned}\Delta'_c + \left[\frac{\rho'_c V_c}{\rho_c k} \right]' &= -kV_c - 3\Phi' + 2\Psi \frac{a^2 Q_c^0}{\rho_c} - \Delta_c \frac{a^2 Q_c^0}{\rho_c} - \frac{\rho'_c V_c}{\rho_c k} \frac{a^2 Q_c^0}{\rho_c} \\ &+ \frac{a^2 \delta Q_c^{0I}}{\rho_c} + \frac{a^2 Q_c^{0I}}{\rho_c \mathcal{H}} \Phi - \frac{a^2 Q_c^0}{\rho_c} \left[\frac{\Phi}{\mathcal{H}} \right]' \\ V'_c &= -\mathcal{H}V_c + k\Psi - \frac{a^2 Q_c^0}{\rho_c} V_c + \frac{a^2 \delta Q_{pc}^I}{\rho_c}.\end{aligned}$$

$$\begin{aligned}\Delta'_d + \left[\frac{\rho'_d V_d}{\rho_d k} \right]' + 3\mathcal{H}C_e^2 \left(\Delta_d + \frac{\rho'_d V_d}{\rho_d k} \right) - 3\mathcal{H}(C_e^2 - C_a^2) \frac{\rho'_d V_d}{\rho_d k} \\ - 3\omega_d \mathcal{H} \left(\Delta_d + \frac{\rho'_d V_d}{\rho_d k} \right) = -k(1 + \omega_d)V_d - 3(1 + \omega_d)\Phi' + 2\Psi \frac{a^2 Q_d^0}{\rho_d} \\ - \left(\Delta_d + \frac{\rho'_d V_d}{\rho_d k} \right) \frac{a^2 Q_d^0}{\rho_d} + \frac{a^2 \delta Q_d^{0I}}{\rho_d} + \frac{a^2 Q_d^{0I}}{\rho_d \mathcal{H}} \Phi - \frac{a^2 Q_d^0}{\rho_d} \left[\frac{\Phi}{\mathcal{H}} \right]'\end{aligned}$$

$$\begin{aligned}V'_d + \mathcal{H}(1 - 3\omega_d)V_d &= \frac{k}{1 + \omega_d} \left[C_e^2 \left(\Delta_d + \frac{\rho'_d V_d}{\rho_d k} \right) - (C_e^2 - C_a^2) \frac{\rho'_d V_d}{\rho_d k} \right] \\ &- \frac{\omega'_d}{1 + \omega_d} V_d + k\Psi - \frac{a^2 Q_d^0}{\rho_d} V_d + \frac{a^2 \delta Q_{pd}^I}{(1 + \omega_d)\rho_d}\end{aligned}$$

$$\begin{aligned}\Delta'_c &= -kV_c - \Delta_c \frac{a^2 Q_c^0}{\rho_c} + \frac{a^2 \delta Q_c^{0I}}{\rho_c} \\ V'_c &= -\mathcal{H}V_c + k\Psi - \frac{a^2 Q_c^0}{\rho_c} V_c + \frac{a^2 \delta Q_{pc}^I}{\rho_c};\end{aligned}$$

$$\begin{aligned}\Delta'_d &= 3\mathcal{H}(\omega_d - C_e^2)\Delta_d - k(1 + \omega_d)V_d - \Delta_d \frac{a^2 Q_d^0}{\rho_d} + \frac{a^2 \delta Q_d^{0I}}{\rho_d} \\ V'_d + \mathcal{H}(1 - 3\omega_d)V_d &= \frac{kC_e^2}{1 + \omega_d} \Delta_d + \frac{C_a^2}{1 + \omega_d} \frac{\rho'_d}{\rho_d} V_d + k\Psi - \frac{\omega'_d}{1 + \omega_d} V_d \\ &- \frac{a^2 Q_d^0}{\rho_d} V_d + \frac{a^2 \delta Q_{pd}^I}{(1 + \omega_d)\rho_d}\end{aligned}$$

$$\begin{aligned}\Delta''_c &= - \left(\mathcal{H} + \frac{2a^2 Q_c^0}{\rho_c} \right) \Delta'_c + \left(-\Delta_c \frac{a^2 Q_c^0}{\rho_c} + \frac{a^2 \delta Q_c^{0I}}{\rho_c} \right) \left(\mathcal{H} + \frac{a^2 Q_c^0}{\rho_c} \right) \\ &- \Delta_c \left(\frac{a^2 Q_c^0}{\rho_c} \right)' + \left(\frac{a^2 \delta Q_c^{0I}}{\rho_c} \right)' - \frac{a^2 k \delta Q_{pc}^I}{\rho_c} - k^2 \Psi.\end{aligned}$$

$$\begin{aligned}\Delta''_d &= -3\mathcal{H}'C_e^2 \Delta_d - \left(\frac{a^2 Q_d^0}{\rho_d} \right)' \Delta_d + \left\{ \mathcal{H}(1 - 3\omega_d) - \frac{\omega_d}{1 + \omega_d} \frac{\rho'_d}{\rho_d} + \frac{a^2 Q_d^0}{\rho_d} \right\} \\ &\times \left\{ -3\mathcal{H}C_e^2 + 3\omega_d \mathcal{H} - \frac{a^2 Q_d^0}{\rho_d} \right\} \Delta_d \\ &- \left[\mathcal{H} + 3\mathcal{H}C_e^2 - 6\omega_d \mathcal{H} + \frac{2a^2 Q_d^0}{\rho_d} - \frac{\omega_d}{1 + \omega_d} \frac{\rho'_d}{\rho_d} \right] \Delta'_d \\ &- k \left(\frac{a^2 \delta Q_{pd}^I}{\rho_d} \right) - k^2 C_e^2 \Delta_d - k^2 (1 + \omega_d) \Psi + 3(\omega'_d \mathcal{H} + \omega_d \mathcal{H}') \Delta_d \\ &+ \frac{a^2 \delta Q_d^{0I}}{\rho_d} \left[\mathcal{H}(1 - 3\omega_d) - \frac{\omega_d}{1 + \omega_d} \frac{\rho'_d}{\rho_d} + \frac{a^2 Q_d^0}{\rho_d} \right] + \left(\frac{a^2 \delta Q_d^{0I}}{\rho_d} \right)'\end{aligned}$$



$$-\frac{k^2}{a^2}\Psi=\frac{3}{2}H^2\{\Omega_c\Delta_c+(1-\Omega_c)\Delta_d\}$$

$$\frac{a^2\delta Q_c^{0I}}{\rho_c}\approx 3\mathcal{H}(\xi_1\Delta_c+\xi_2\Delta_d/r), \frac{a^2\delta Q_d^{0I}}{\rho_d}\approx -3\mathcal{H}(\xi_1\Delta_cr+\xi_2\Delta_d),$$

$$\begin{aligned}\Delta'_c + \nabla_{\bar{x}} \cdot V_c &= 3\mathcal{H}\xi_2(\Delta_d - \Delta_c)/r \\ V'_c + \mathcal{H}V_c &= -\nabla_{\bar{x}}\Psi - 3\mathcal{H}(\xi_1 + \xi_2/r)V_c \\ \Delta'_d + (1 + \omega_d)\nabla_{\bar{x}} \cdot V_d &= 3\mathcal{H}(\omega_d - C_e^2)\Delta_d + 3\mathcal{H}\xi_1r(\Delta_d - \Delta_c) \\ V'_d + \mathcal{H}V_d &= -\nabla_{\bar{x}}\Psi - \frac{C_e^2}{1 + \omega_d}\nabla_{\bar{x}}\Delta_d - \frac{\omega'_d}{1 + \omega_d}V_d \\ &\quad + 3\mathcal{H}\left\{(\omega_d - C_a^2) + \frac{1 + \omega_d - C_a^2}{1 + \omega_d}(\xi_1r + \xi_2)\right\}V_d\end{aligned}$$

$$\begin{aligned}\dot{\sigma}_c + 3H\sigma_c + \nabla_x(\rho_c V_c) &= 3H(\xi_1\sigma_c + \xi_2\sigma_d) \\ \frac{\partial}{\partial t}(aV_c) &= -\nabla_x(a\Psi) - 3H(\xi_1 + \xi_2/r)(aV_c) \\ \dot{\sigma}_d + 3H(1 + C_e^2)\sigma_d + (1 + \omega_d)\nabla_x(\rho_d V_d) &= -3H(\xi_1\sigma_c + \xi_2\sigma_d) \\ \frac{\partial}{\partial t}(aV_d) &= -\nabla_x(a\Psi) - \frac{C_e^2}{1 + \omega_d}\nabla_x \cdot (a\Delta_d) \\ &\quad + 3H\left[(\omega_d - C_a^2) + \frac{1 + \omega_d - C_a^2}{1 + \omega_d}(\xi_1r + \xi_2)\right](aV_d)\end{aligned}$$

$$\nabla^2\psi_\lambda = 4\pi G(1+3\omega_\lambda)\sigma_\lambda$$

$$\psi_c = -4\pi G \int dV' \frac{\sigma_c}{|x-x'|}, \psi_d = -4\pi G \int dV' \frac{(1+3\omega_d)\sigma_d}{|x-x'|}$$

$$\frac{\partial}{\partial t}(aV_c) = -\nabla_x(a\psi_c + a\psi_d) - 3H(\xi_1 + \xi_2/r)(aV_c)$$

$$\frac{\partial}{\partial t}(a^2T_c) - a^23H(\xi_1 + \xi_2/r)T_c$$

$$\begin{aligned}-\int aV_c\nabla_x(a\psi_c + a\psi_d)\rho_c\hat{\epsilon} &= a^2\int \nabla_x(\rho_c V_c)\psi_c\hat{\epsilon} + a^2\int \nabla_x(\rho_c V_c)\psi_d\hat{\epsilon} \\ -\int aV_c\nabla_x(a\psi_c + a\psi_d)\rho_c\hat{\epsilon} &= -a^2(\dot{U}_{cc} + HU_{cc}) - a^2\int \psi_d\frac{\partial}{\partial t}(\sigma_c\hat{\epsilon}) \\ &\quad + 3a^2H\{\xi_1U_{cd} + \xi_2U_{dc} + 2\xi_1U_{cc} + 2\xi_2U_{dd}\}\end{aligned}$$

$$\begin{aligned}U_{cc} &= \frac{1}{2}\int \sigma_c\psi_c\hat{\epsilon}, U_{dc} = \int \sigma_d\psi_c\hat{\epsilon}, U_{cd} = \int \sigma_c\psi_d\hat{\epsilon}, U_{dd} = \frac{1}{2}\int \sigma_d\psi_d\hat{\epsilon} \\ -\int (aV_c)^23H(\xi_1 + \xi_2/r)\rho_c\hat{\epsilon} &= -a^26H(\xi_1 + \xi_2/r)T_c \\ -\int (aV_c)^23H(\xi_1 + \xi_2/r)\rho_c\hat{\epsilon} &= -a^26H(\xi_1 + \xi_2/r)T_c\end{aligned}$$

$$\begin{aligned}\dot{T}_c + \dot{U}_{cc} + H(2T_c + U_{cc}) &= -\int \psi_d\frac{\partial}{\partial t}(\sigma_c\hat{\epsilon}) - 3H(\xi_1 + \xi_2/r)T_c \\ &\quad + 3H\{\xi_1U_{cd} + \xi_2U_{dc} + 2\xi_1U_{cc} + 2\xi_2U_{dd}\}\end{aligned}$$



$$\dot{T}_c + \dot{U}_{cc} + H(2T_c + U_{cc}) = -3H(\xi_1 + \xi_2/r)T_c + 6H\xi_1U_{cc}$$

$$\frac{\partial}{\partial t}(a^2T_d)+3a^2H(\omega_d+\xi_1r+\xi_2)T_d$$

$$\begin{aligned} & -\int aV_d\nabla_x(a\psi_c+a\psi_d)\rho_d\hat{\varepsilon}=-\frac{a^2}{1+\omega_d}(\dot{U}_{dd}+HU_{dd}) \\ & -\frac{a^2}{1+\omega_d}3H\{2(C_e^2+\xi_2)U_{dd}+2\xi_1U_{cc}+\xi_1U_{cd}+(C_e^2+\xi_2)U_{dc}\} \\ & -\frac{a^2}{1+\omega_d}\int\psi_c\frac{\partial}{\partial t}(\sigma_d\hat{\varepsilon}) \\ & 3H\left[(w-c_a^2)+\frac{1+w-C_a^2}{1+\omega_d}(\xi_1r+\xi_2)\right]\int(aV_d)^2\rho_d\hat{\varepsilon} \\ & -\frac{c_e^2}{1+\omega_d}\int aV_d\nabla_x(a\Delta_d)\rho_d\hat{\varepsilon} \\ & =6a^2H\left[(\omega_d-C_a^2)+\frac{1+\omega_d-c_a^2}{1+\omega_d}(\xi_1r+\xi_2)\right]T_d \\ & -\frac{c_e^2}{1+\omega_d}a^2\int V_d\nabla_x(\sigma_d)\hat{\varepsilon} \end{aligned}$$

$$\begin{aligned} (1+\omega_d)\dot{T}_d + \dot{U}_{dd} + H[2(1+\omega_d)T_d + U_{dd}] &= -3H\{2(C_e^2+\xi_2)U_{dd}+2\xi_1U_{cc} \\ & +\xi_1U_{cd}+(C_e^2+\xi_2)U_{dc}\}-\int\psi_c\frac{\partial}{\partial t}(\sigma_d\hat{\varepsilon})-c_e^2\int V_d\nabla_x(\sigma_d)\hat{\varepsilon} \\ & +3H[(1+\omega_d)(\omega_d-2C_a^2)+(1+\omega_d-2C_a^2)(\xi_1r+\xi_2)]T_d \end{aligned}$$

$$\dot{\theta}=-\frac{1}{3}\theta^2-4\pi G\sum_{\lambda}\left(\rho_{\lambda}+3p_{\lambda}\right)$$

$$\frac{\ddot{R}}{R}=-\frac{4\pi G}{3}\sum_{\lambda}\left(\rho_{\lambda}+3p_{\lambda}\right)$$

$$\dot{\rho}_c^{cl}+3\frac{\dot{R}}{R}\rho_c^{cl}=3H\big(\xi_1\rho_c^{cl}+\xi_2\rho_d\big),$$

$$\ddot{R}=-\frac{4\pi G}{3}[\rho_c^{cl}+(1+3\omega_d)\rho_d]R,$$

$$2a^2\Big(1+\frac{1}{r}\Big)\frac{d^2R}{da^2}-\frac{dR}{da}[1+(3\omega_d+1)/r]a=-R[(3\omega_d+1)/r+\zeta]$$

$$\frac{d\zeta}{da}=\frac{3}{a}(1-\xi_2/r)\zeta-3\frac{1}{R}\frac{dR}{da}\zeta+\frac{3}{a}\xi_2/r$$

$$u_{(d)}^a=\gamma(u_{(c)}^a+v_d^a)$$

$$\begin{aligned} T_{(c)}^{ab} &= \rho_c u_{(c)}^a u_{(c)}^b \\ T_{(d)}^{ab} &= \rho_d u_{(d)}^a u_{(d)}^b + p_d h_{(d)}^{ab} \end{aligned}$$



$$T^{ab}_{(d)}=\rho_d u^a u^b + p_d h^{ab} + 2 u^a q^b_{(d)}$$

$$\nabla_a T^{ab}_{(\lambda)} = Q^b_{(\lambda)}$$

$$\begin{aligned}\dot{\rho}_c^{cl}+3h\rho_c^{cl}&=3H\big(\xi_1\rho_c^{cl}+\xi_2\rho_d^{cl}\big)\\\dot{\rho}_d^{cl}+3h(1+\omega_a)\rho_d^{cl}&=-\vartheta(1+\omega_d)\rho_d^{cl}-3H\big(\xi_1\rho_c^{cl}+\xi_2\rho_d^{cl}\big)\end{aligned}$$

$$\dot{q}^a_{(d)}+4hq^a_{(d)}=0$$

$$\dot{\vartheta}+h(1-3\omega_d)\vartheta=3H(\xi_1\Gamma+\xi_2)\vartheta.$$

$$\begin{aligned}\vartheta'+\frac{R'}{R}(1-3\omega_d)\vartheta&=\frac{3}{a}(\xi_1\Gamma+\xi_2)\vartheta\\\zeta'_c=\frac{3}{a}\{1-\xi_2/r\}\zeta_c-3\frac{R'}{R}\zeta_c+\frac{3}{a}\xi_2\zeta_d/r\end{aligned}$$

$$\zeta'_d=\frac{3}{a}(1+\omega_d+\xi_1r)\zeta_d-3(1+\omega_d)\frac{R'}{R}\zeta_d-\frac{3}{a}\xi_1\zeta_cr-\vartheta(1+\omega_d)\zeta_d$$

$$2a^2\left(1+\frac{1}{r}\right)R''-R'[1+(3\omega_d+1)/r]a=-R[(3\omega_d+1)\zeta_d/r+\zeta_c]$$

$$\frac{dn(M,z)}{dM}=\sqrt{\frac{2}{\pi}}\frac{\rho_m}{3M^2}\frac{\delta_c}{\sigma}e^{-\delta_c^2/2\sigma^2}\left[-\frac{R}{\sigma}\frac{d\sigma}{dR}\right]$$

$$\sigma(R,z)=\sigma_8 \Bigl({R\over 8h^{-1}{\rm Mpc}}\Bigr)^{-\gamma(R)} D(z)$$

$$\gamma(R)=(0.3\Gamma+0.2)\left[2.92+\log_{10}\left(\frac{R}{8h^{-1}{\rm Mpc}}\right)\right]$$

$$R=0.951h^{-1}{\rm Mpc}\biggl(\frac{M}{\Omega_m10^{12}h^{-1}M_\odot}\biggr)^{1/3}$$

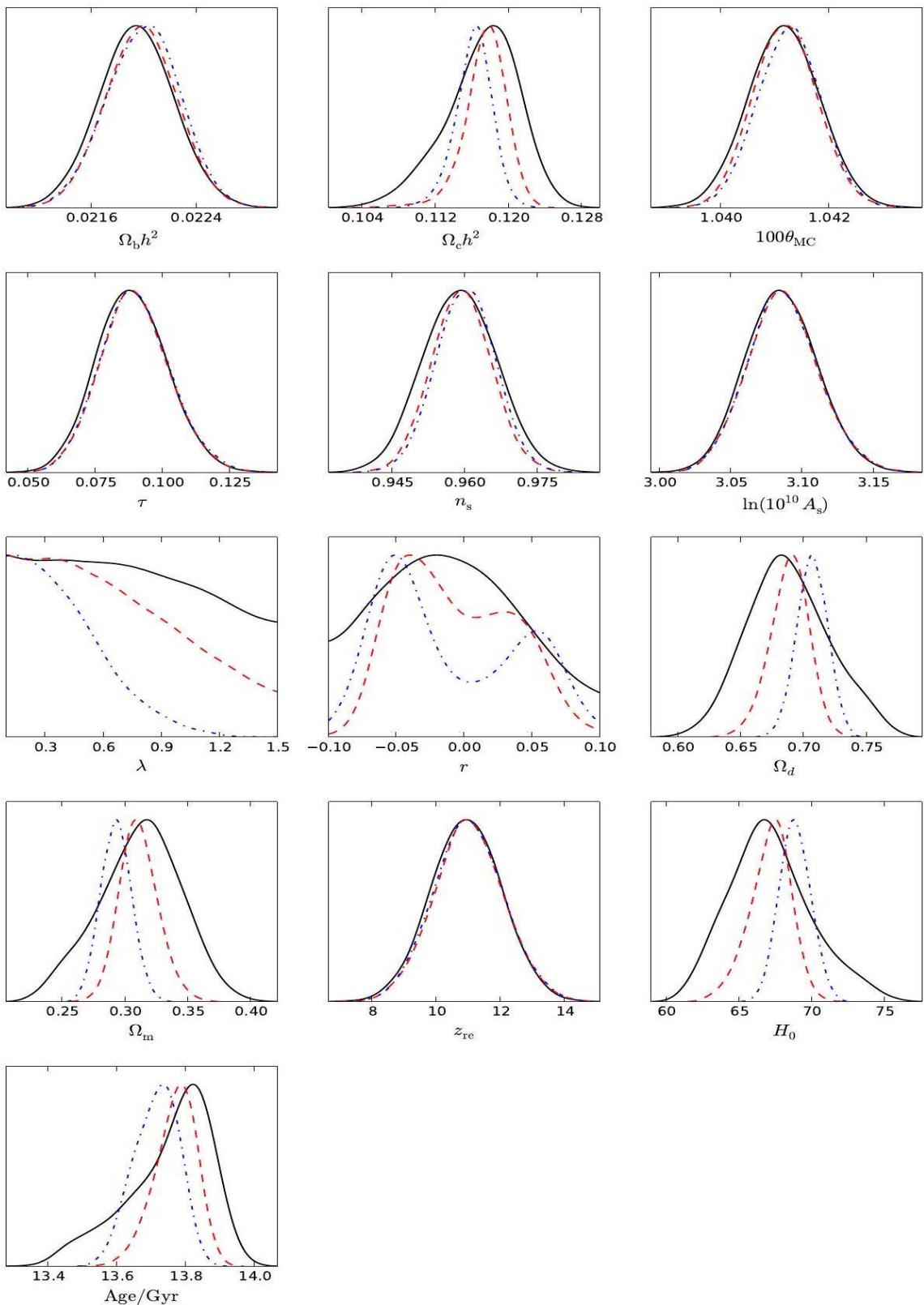
$$\frac{dN}{dz}=\int\,\,d\Omega\frac{dV}{dzd\Omega}\int\,\,n(M)dM$$

$$dV/dzd\Omega=r^2(z)/H(z)\; r(z)=\int_0^z\frac{dz'}{H(z)}$$

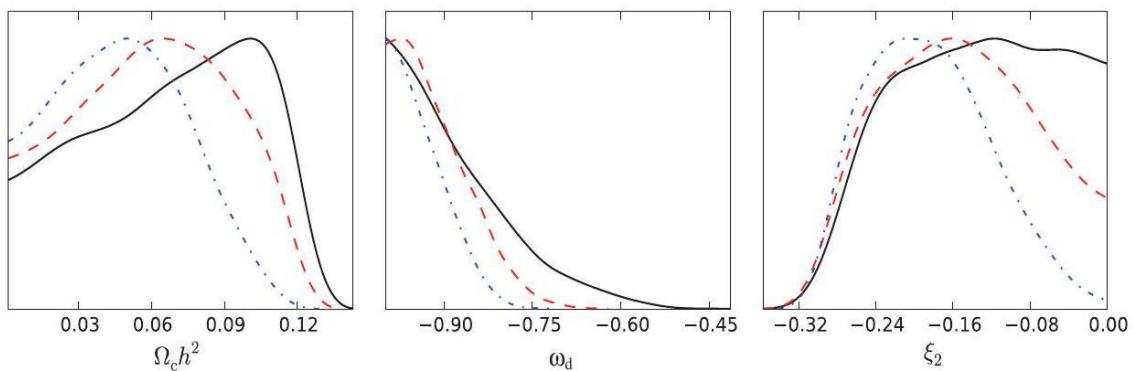
$$\varphi(\hat{n})=-2\int_0^{\varsigma_{rec}}d\varsigma\frac{\varsigma_{rec}-\varsigma}{\varsigma_{rec}\varsigma}\Psi(\varsigma\hat{n};\tau_0-\tau)$$

$$AIC=2k-2\text{ln }\hat{\mathcal{L}}(\theta);\;BIC=-2\text{ln }\hat{\mathcal{L}}(\theta)+k\text{ln }N.$$

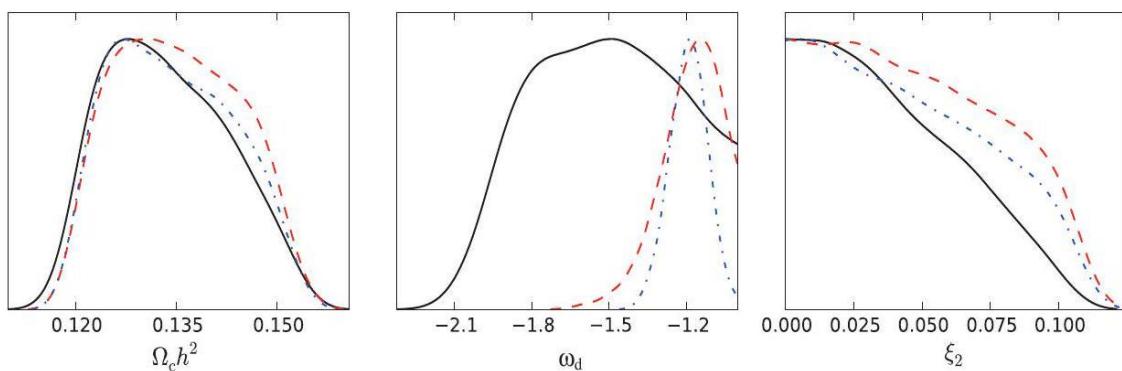




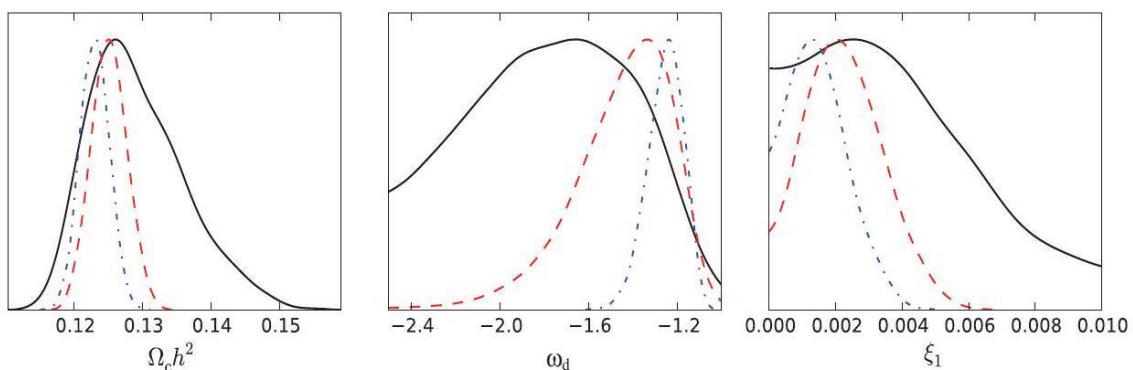
Model I



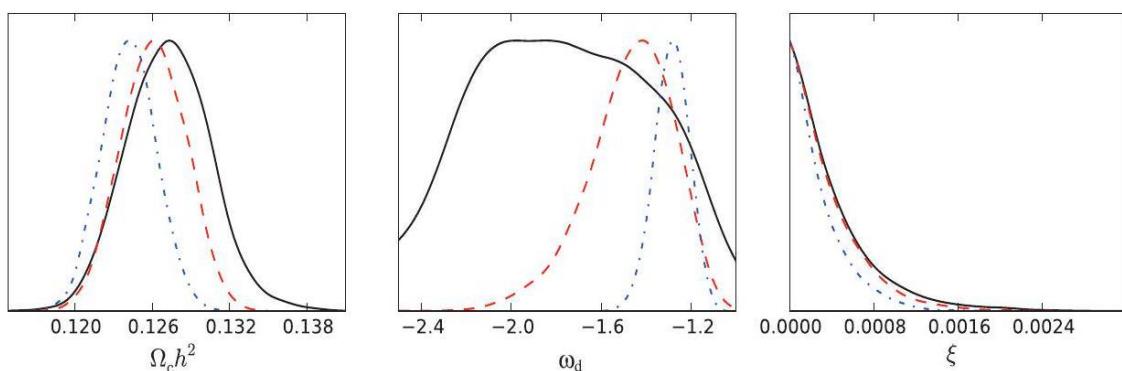
Model II

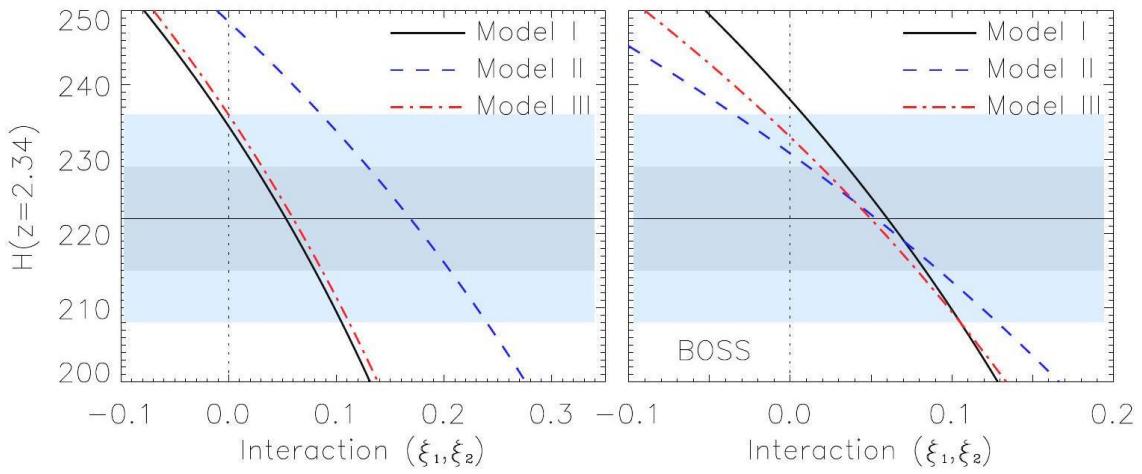


Model III



Model IV





Figuras 29, 30 y 31. Fluctuaciones de gauge en un espacio – tiempo cuántico relativista o curvo, en el que interactúa una partícula oscura.

$$\rho_d = (1+z)^{3(-1+\omega_d+\xi_2)} \rho_d^0$$

$$\rho_c = (1+z)^3 \left\{ \frac{\xi_2 [1 - (1+z)^{3(\xi_2+\omega_d)}] \rho_d^0}{\xi_2 + \omega_d} + \rho_c^0 \right\}$$

$$\rho_d = (1+z)^{3(1+\omega_d)} \left(\rho_d^0 + \frac{\xi_1 \rho_c^0}{\xi_1 + \omega_d} \right) - \frac{\xi_1}{\xi_1 + \omega_d} (1+z)^{3(1-\xi_1)} \rho_c^0$$

$$\rho_c = \rho_c^0 (1+z)^{3(1-\xi_1)}$$

$$r_s^+ \sim \frac{1}{\xi} \text{ and } r_s^- \sim \xi$$



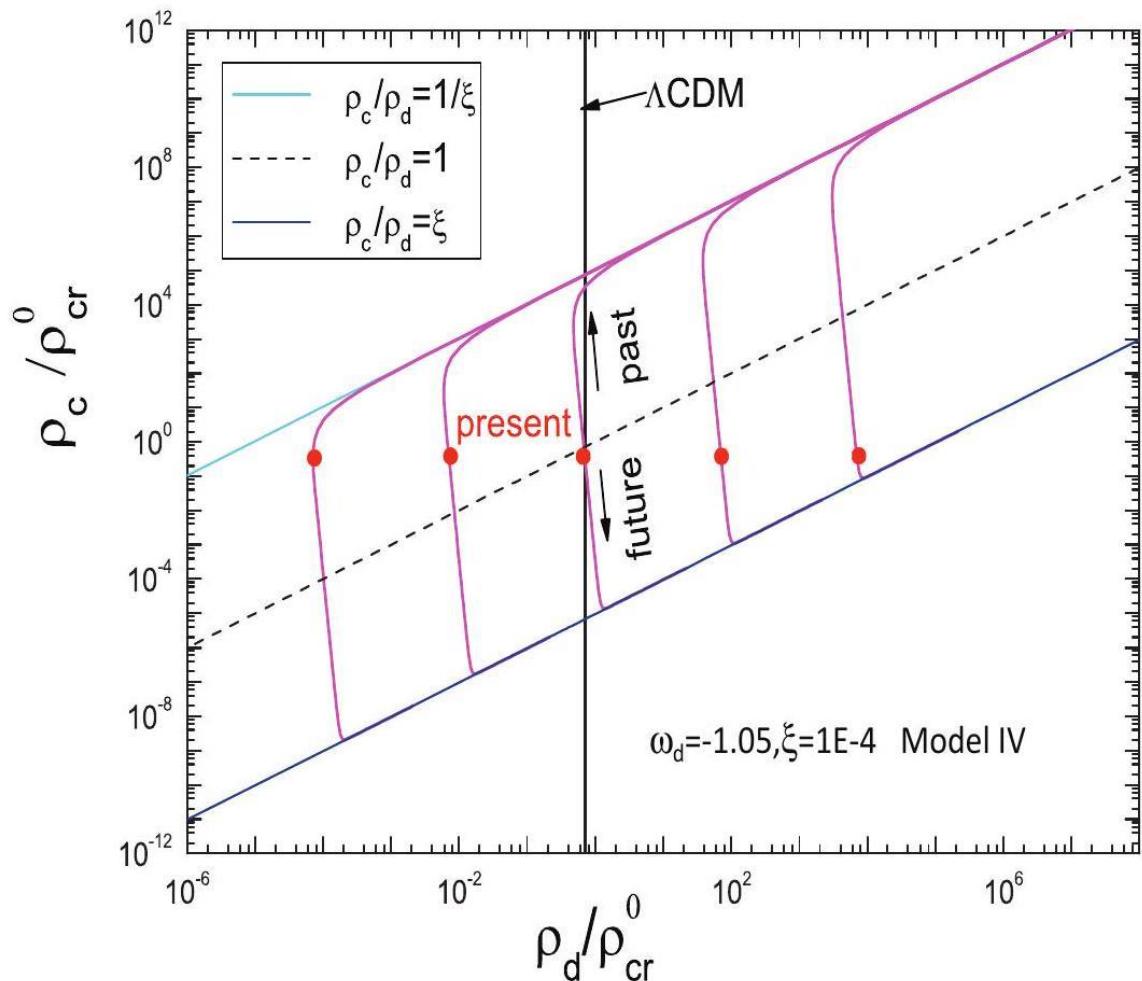


Figura 32. Comportamiento de un agujero negro provocado por una partícula oscura. Crisis inflacionaria y perturbativa de campo.

SECCIÓN IV.

$$\dot{\rho}_c + 3\mathcal{H}\rho_c = Q, \dot{\rho}_d + 3\mathcal{H}(1 + \omega_d)\rho_d = -Q,$$

$$\begin{aligned} Q_1 &= \mathcal{H}\xi\rho_c, & Q_2 &= \mathcal{H}\xi\rho_d, & Q_3 &= \mathcal{H}\xi(\rho_c + \rho_d) \\ Q_4 &= \mathcal{H}\xi(\rho_c + (1 + \omega_d)\rho_d), & Q_5 &= \mathcal{H}\xi\sqrt{\rho_d\rho_c}, & Q_6 &= \mathcal{H}\xi \frac{\rho_c\rho_d}{\rho_c + \rho_d} \end{aligned}$$



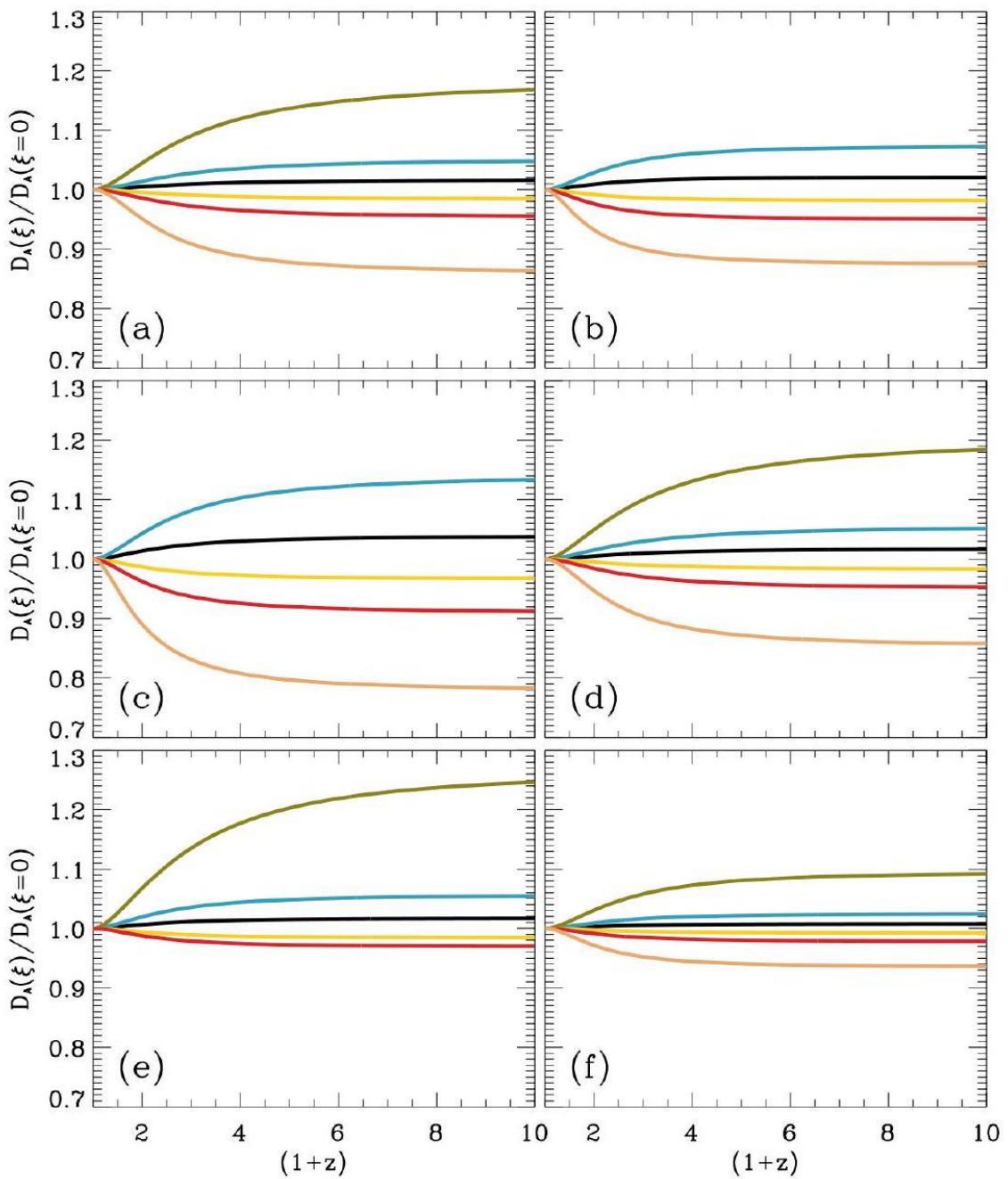


Figura 33. Kernels de campo a propósito de la interacción de una partícula oscura.

$$\begin{aligned}
 S &= S_g(g_{\mu\nu}) + S_\phi(\phi_1, \phi_2) + S_m \\
 S_\phi(\phi_1, \phi_2) &= - \int d^4x \sqrt{-g} \left(\frac{1}{2} (\partial\phi_1)^2 + \frac{1}{2} (\partial\phi_2)^2 + V(\phi_1, \phi_2) \right) \\
 V(\phi_1, \phi_2) &= V_1(\phi_1) + V_2(\phi_2) + V_{\text{int}}(\phi_1, \phi_2) \\
 V_1(\phi_1) &= \mu_1^4 \left[1 - \cos \left(\frac{\phi_1}{f_1} \right) \right], \quad V_2(\phi_2) = \mu_2^4 \left[1 - \cos \left(\frac{\phi_2}{f_2} \right) \right],
 \end{aligned}$$

$$V_{\rm int}(\phi_1,\phi_2)=\mu_3^4\left[1-\cos\left(\frac{\phi_1}{f_1}-n\frac{\phi_2}{f_2}\right)\right].$$

$$T_{\mu\nu}=-\frac{2}{\sqrt{-g}}\frac{\delta(S_m+S_\phi)}{\delta g^{\mu\nu}}$$

$$\begin{aligned}3H^2=\rho+\frac{1}{2}\dot{\phi_1}^2+\frac{1}{2}\dot{\phi_2}^2\\+2\mu_1^4\sin^2\left(\frac{\phi_1}{2f_1}\right)+2\mu_2^4\sin^2\left(\frac{\phi_2}{2f_2}\right)\\+2\mu_3^4\sin^2\left(\frac{\phi_1}{2f_1}-n\frac{\phi_2}{2f_2}\right)\end{aligned}$$

$$2\dot{H}=-\rho-\dot{\phi_1}^2-\dot{\phi_2}^2.$$

$$\begin{aligned}\ddot{\phi}_1+3H\dot{\phi}_1+\frac{\mu_1^4}{f_1}\sin\left(\frac{\phi_1}{f_1}\right)+\frac{\mu_3^4}{f_1}\sin\left(\frac{\phi_1}{f_1}-n\frac{\phi_2}{f_2}\right)=0,\\\ddot{\phi}_2+3H\dot{\phi}_2+\frac{\mu_2^4}{f_2}\sin\left(\frac{\phi_2}{f_2}\right)-n\frac{\mu_3^4}{f_2}\sin\left(\frac{\phi_1}{f_1}-n\frac{\phi_2}{f_2}\right)=0.\end{aligned}$$

$$\dot\rho+3H\rho=0.$$

$$\nabla_\mu T^{\mu\nu}_{(\lambda)} = Q^\nu_{(\lambda)}$$

$$\sum_{\lambda}\,\nabla_\mu T^{\mu\nu}_{(\lambda)}=0$$

$$T^{\mu\nu}=\rho U^\mu U^\nu+p(g^{\mu\nu}+U^\mu U^\nu)$$

$$\begin{aligned}ds^2=&a^2(\tau)[-(1+2\psi)d\tau^2+2\partial_iBd\tau dx^i\\&+(1+2\phi)\delta_{ij}dx^idx^j+D_{ij}Edx^idx^j]\end{aligned}$$

$$D_{ij}=(\,\partial_i\partial_j-\tfrac{1}{3}\delta_{ij}\nabla^2\,)$$

$$\begin{aligned}\Psi&=\psi-\frac{1}{k}\mathcal{H}\left(B+\frac{E'}{2k}\right)-\frac{1}{k}\left(B'+\frac{E''}{2k}\right)\\\Phi&=\phi+\frac{1}{6}E-\frac{1}{k}\mathcal{H}\left(B+\frac{E'}{2k}\right)\\D_\lambda&=\delta_\lambda-\frac{\rho'_\lambda}{\rho_{\lambda\mathcal{H}}}\Big(\phi+\frac{E}{6}\Big)\\V_\lambda&=v_\lambda-\frac{E'}{2k}\end{aligned}$$

$$\begin{aligned}\Psi&=\psi\\\Phi&=\phi\\D_\lambda&=\delta_\lambda-\frac{\rho'_\lambda}{\rho_\lambda\mathcal{H}}\Phi\\V_\lambda&=v_\lambda\end{aligned}$$



$$\begin{aligned} D'_c &= -kU_c + 6\mathcal{H}\Psi(\xi_1 + \xi_2/r) - 3(\xi_1 + \xi_2/r)\Phi' \\ &\quad + 3\mathcal{H}\xi_2(D_d - D_c)/r \\ U'_c &= -\mathcal{H}U_c + k\Psi - 3\mathcal{H}(\xi_1 + \xi_2/r)U_c \end{aligned}$$

$$\begin{aligned} D'_d &= -3(C_e^2 - w_d)D_d - 9\mathcal{H}^2(C_e^2 - C_a^2)\frac{U_d}{k} \\ &\quad + [3w'_d - 9\mathcal{H}(w_d - C_e^2)(\xi_1 r + \xi_2 + 1 + w_d)]\Phi \\ &\quad + 3(\xi_1 r + \xi_2)\Phi' - 3\Psi\mathcal{H}(\xi_1 r + \xi_2) \\ &\quad - 9\mathcal{H}^2(C_e^2 - C_a^2)(\xi_i r + \xi_2)\frac{U_d}{(1 + w_d)k} \\ &\quad - kU_d + 3\mathcal{H}\xi_1 r(D_d - D_c), \\ U'_d &= -\mathcal{H}(1 - 3w_d)U_d + 3(C_e^2 - C_a^2)\mathcal{H}U_d \\ &\quad - 3kC_e^2(\xi_1 r + \xi_2 + 1 + w_d)\Phi + kC_e^2D_d \\ &\quad + 3\mathcal{H}(C_e^2 - C_a^2)(\xi_1 r + \xi_2)\frac{U_d}{1 + w_d} \\ &\quad + (1 + w_d)k\Psi + 3\mathcal{H}(\xi_1 r + \xi_2)U_d, \\ -k^2\Psi &= \frac{3}{2}\mathcal{H}^2[\Omega_c \Delta_c + (1 - \Omega_c) \Delta_d] \end{aligned}$$

$$\Delta_\lambda = \delta_\lambda - \frac{\rho'_\lambda}{\rho_\lambda} \frac{V_\lambda}{k}, \Omega_\lambda = \frac{\rho_\lambda}{\rho_{\text{crit}}}$$

$$\nabla^2\Psi = -\frac{3}{2}\mathcal{H}^2[\Omega_c \Delta_c + (1 - \Omega_c) \Delta_d]$$

$$V'_c + [\mathcal{H} + 3\mathcal{H}(\xi_1 + \xi_2/r)]V_c - k\Psi = 0.$$

$$\begin{aligned} \nabla V'_c &+ [\mathcal{H} + 3\mathcal{H}(\xi_1 + \xi_2/r)]\nabla V_c \\ &+ \frac{3}{2}\mathcal{H}^2[\Omega_c \Delta_c + (1 - \Omega_c) \Delta_d] = 0 \end{aligned}$$

$$\alpha(a) = -3\mathcal{H}(\xi_1 + \xi_2/r)a.$$

$$\mathbf{x}(\mathbf{q}) = \mathbf{q} + \nabla_q \Psi^{(1)} + \nabla_q \Psi^{(2)}$$

$$\mathbf{v}(\mathbf{q}) = f_1 H \nabla_q \Psi^{(1)} + f_2 H \nabla_q \Psi^{(2)}$$

$$f_{1,2}=\frac{d\ln\left(D_{1,2}\right)}{d\ln\,a}$$



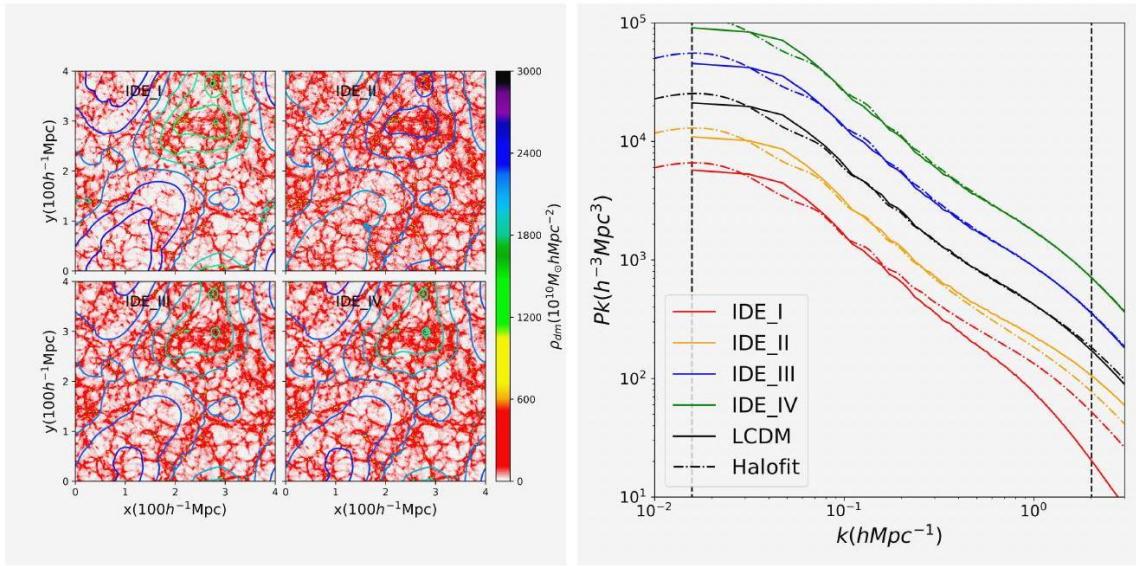


Figura 34. Energía y materia oscuras inherentes a un sector de campo cuántico – relativista oscuro.

$$\begin{aligned}\bar{T}_b(z) &= \frac{3}{32\pi} \frac{(h_p c)^3 \bar{n}_{HI} A_{10}}{k_B E_{21}^2 (1+z) H(z)} \\ &= 0.188 h \Omega_{HI}(z) \frac{(1+z)^2}{E(z)} K\end{aligned}$$

$$ds^2 = a^2(\eta)[(1+2\Psi)d\eta^2 - (1-2\Phi)\delta_{ij} dx^i dx^j]$$

$$\dot{\mathbf{v}} + \mathcal{H}\mathbf{v} + \nabla\Psi = -\mathbf{v} \frac{aQ}{\rho_m},$$

$$\begin{aligned}\Delta_{T_b}(z, \hat{\mathbf{n}}) &= \delta_n - \frac{1}{\mathcal{H}} \hat{\mathbf{n}} \cdot (\hat{\mathbf{n}} \cdot \nabla \mathbf{v}) + \left(\frac{d\ln(a^3 \bar{n}_{HI})}{d\eta} - \frac{\dot{\mathcal{H}}}{\mathcal{H}} - 2\mathcal{H} \right) \delta\eta \\ &\quad + \frac{1}{\mathcal{H}} \dot{\Phi} + \Psi - \frac{1}{\mathcal{H}} \hat{\mathbf{n}} \cdot \mathbf{v} \frac{aQ}{\rho_m}\end{aligned}$$

$$\delta_n = b_{HI} \delta_m^{\text{syn}} + \left(\frac{d\ln(a^3 \bar{n}_{HI})}{d\eta} - 3\mathcal{H} \right) \frac{v_m}{k},$$

$$\Delta_{T_b}(z, \hat{\mathbf{n}}) = \sum_{\ell m} \Delta_{T_b, \ell m}(z) Y_{\ell m}(\hat{\mathbf{n}})$$

$$\Delta_{T_b, \ell m}(z) = 4\pi i^\ell \int \frac{d^3 \mathbf{k}}{(2\pi)^{3/2}} \Delta_{T_b, \ell}(\mathbf{k}, z) Y_{\ell m}^*(\hat{\mathbf{k}})$$

$$\begin{aligned}\Delta_{T_b,\ell}(\mathbf{k}, z) = & \delta_n j_\ell(k\chi) + \frac{kv}{\mathcal{H}} j''_\ell(k\chi) + \left(\frac{1}{\mathcal{H}} \dot{\Phi} + \Psi \right) j_\ell(k\chi) \\ & - \left(\frac{1}{\mathcal{H}} \frac{d\ln(a^3 \bar{n}_{HI})}{d\eta} - \frac{\dot{\mathcal{H}}}{\mathcal{H}^2} - 2 \right) [\Psi j_\ell(k\chi) \\ & + v j'_\ell(k\chi) + \int_0^\chi (\dot{\Psi} + \dot{\Phi}) j_\ell(k\chi') d\chi'] \\ & + \frac{1}{\mathcal{H}} v j'_\ell(k\chi) \frac{aQ}{\rho_m}\end{aligned}$$

$$\Delta_{T_b,\ell}^W(\mathbf{k}) = \int_0^\infty dz W(z) \Delta_{T_b,\ell}(\mathbf{k}, z)$$

$$W(z) = \begin{cases} \frac{1}{\Delta z}, & z - \frac{\Delta z}{2} \leq z \leq z + \frac{\Delta z}{2} \\ 0, & \text{otherwise} \end{cases}$$

$$C_\ell^{WW'} = 4\pi \int dz \ln k \mathcal{P}_R(k) \Delta_{T_b,\ell}^W(k) \Delta_{T_b,\ell}^{W'}(k)$$

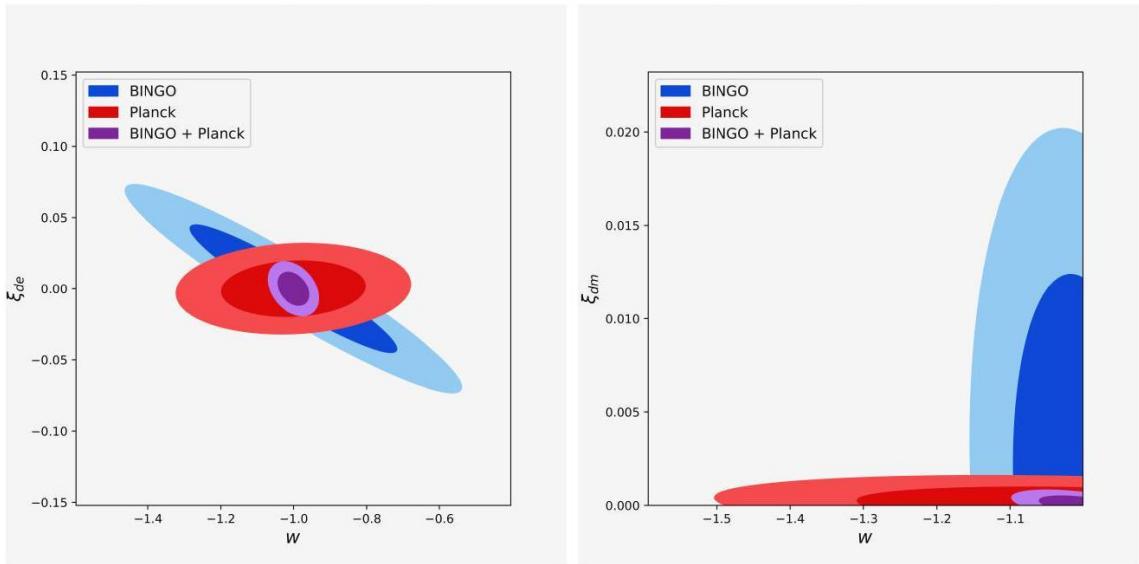


Figura 35. Deformación del espacio – tiempo causada por una partícula supermasiva.

$$\widehat{\mathcal{DM}} = \widehat{\mathcal{DM}}_{MW} + \widehat{\mathcal{DM}}_{\text{halo}} + \widehat{\mathcal{DM}}_{IGM} + \frac{\widehat{\mathcal{DM}}_{\text{host}} + \widehat{\mathcal{DM}}_{\text{src}}}{1+z}$$

$$\langle \widehat{\mathcal{D}}_{IGM}(z) \rangle = \frac{3cH_0\Omega_b f_{IGM}}{8\pi G m_p} \int_0^z \frac{\chi(z')(1+z')dz'}{E(z')}$$

$$\chi(z) \simeq \frac{3}{4} \chi_{e,H}(z) + \frac{1}{8} \chi_{e,He}(z)$$

$$\begin{aligned}h_+(t) &= \left(\frac{G\mathcal{M}_z}{c^2}\right)^{5/3} \left(\frac{\pi f(t)}{c}\right)^{2/3} \frac{2(1+\cos^2 \iota)}{D_L} \cos(\Psi(t, \mathcal{M}_z, \eta)) \\ h_\times(t) &= \left(\frac{G\mathcal{M}_z}{c^2}\right)^{5/3} \left(\frac{\pi f(t)}{c}\right)^{2/3} \frac{4\cos \iota}{D_L} \sin(\Psi(t, \mathcal{M}_z, \eta))\end{aligned}$$



SECCIÓN V.

$$S_{\text{ESGB}} = \int dx^4 \sqrt{-g} \left[\frac{1}{2\kappa} (R - 2\Lambda) + f(\phi) \mathcal{G} + \mathcal{L}_\phi(\phi) + \mathcal{L}_m(g_{\mu\nu}, \phi; \psi) \right]$$

$$\begin{aligned}\mathcal{L}_\phi &= -\frac{1}{2} (D^\mu \phi) (D_\mu \phi) - V(\phi) \\ \mathcal{G} &= R^2 - 4R_{\mu\nu}R^{\mu\nu} + R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma}\end{aligned}$$

$$\begin{aligned}R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R + \Lambda g_{\mu\nu} &= T_{\mu\nu}^{GB} + \kappa T_{\mu\nu}^\phi + \kappa T_{\mu\nu}^m \\ D^2 \phi &= \frac{dV}{d\phi} - \alpha \frac{df}{d\phi} G - Q\end{aligned}$$

$$\begin{aligned}T_{\mu\nu}^{GB} &\equiv 2(D_\mu D_\nu f)R - 2g_{\mu\nu}(D_\rho D^\rho f)R - 4(D^\rho D_\nu f)R_{\mu\rho} - 4(D^\rho D_\mu f)R_{\nu\rho} \\ &\quad + 4(D^\rho D_\rho f)R_{\mu\nu} + 4g_{\mu\nu}(D^\rho D^\sigma f)R_{\rho\sigma} - 4(D^\rho D^\sigma f)R_{\mu\rho\nu\sigma} \\ T_{\mu\nu}^\phi &\equiv \frac{1}{2}(D_\mu \phi)D_\nu \phi - \frac{1}{4}g_{\mu\nu}[(D^\alpha \phi)D_\alpha \phi + 2V(\phi)] \\ T_{\mu\nu}^m &\equiv -\frac{2}{\sqrt{-g}} \frac{\delta(\sqrt{-g}\mathcal{L}_m)}{\delta g^{\mu\nu}} \\ Q &\equiv 2\kappa \frac{\delta\mathcal{L}_m}{\delta\phi}\end{aligned}$$

$$3H^2 = \rho_m + \rho_\phi + 24f'(\phi)\dot{\phi}H^3$$

$$\begin{aligned}\dot{\rho}_m + 3H\rho_m &= Q\rho_m\dot{\phi} \\ \ddot{\phi} + V_{,\phi} + 3H\dot{\phi} &= -Q\rho_m - 24f'H^2(H^2 + \dot{H})\end{aligned}$$

$$V(\phi) = V_0 e^{-\lambda\phi}, f'(\phi) = \alpha e^{\mu\phi}$$

$$x = \frac{\dot{\phi}}{\sqrt{6}H}, y = \frac{\sqrt{V}}{\sqrt{3}H}, z = f'H^2, \Omega_M = \frac{\rho_m}{3H^2}, \Omega_\phi = \frac{\rho_\phi}{3H^2}.$$

$$\Omega_M = 1 - \Omega_\phi - \Omega_{\text{GB}}.$$

$$0 \leq \Omega_M \leq 1, \Omega_\phi > 0, 0 < \Omega_\phi + \Omega_{\text{GB}} < 1$$

$$\frac{\dot{H}}{H^2} = \frac{12Q\Omega_M z + 24\mu x^2 z - 3x^2 - 12\sqrt{6}xz - 4\sqrt{6}xz + 12\lambda y^2 z - \frac{3\Omega_M}{2} - 96z^2}{-8\sqrt{6}xz + 96z^2 + 1}.$$

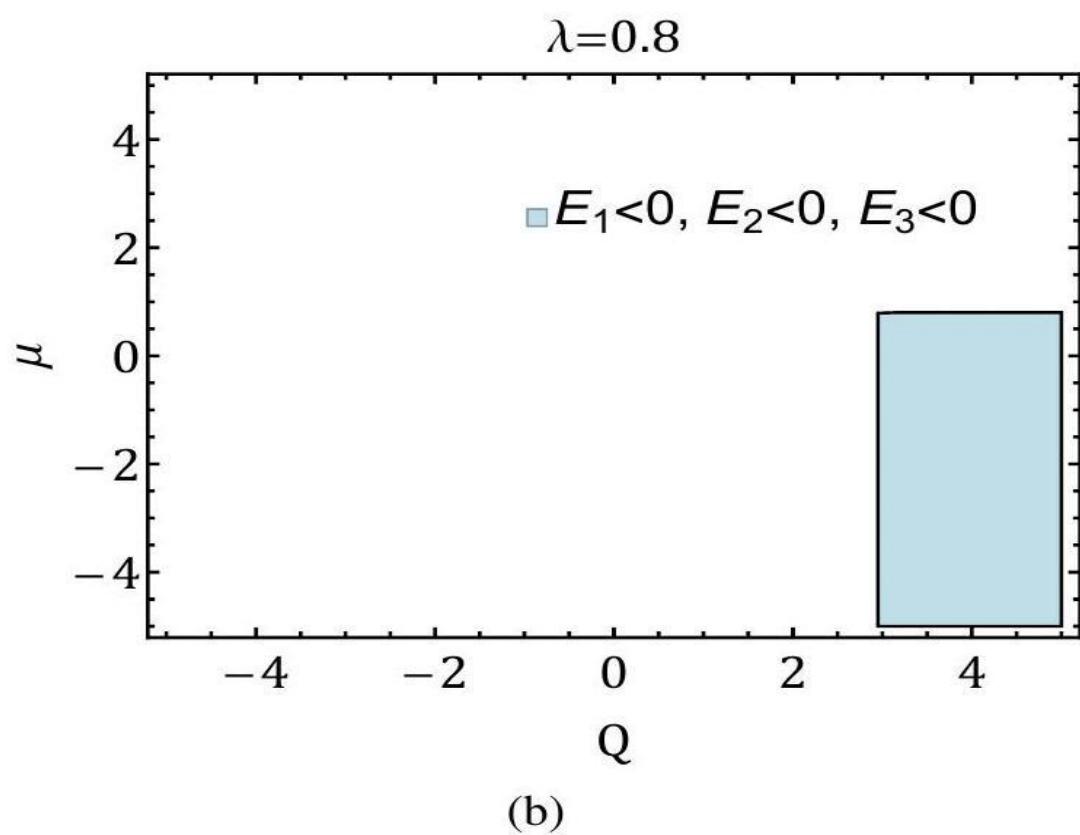
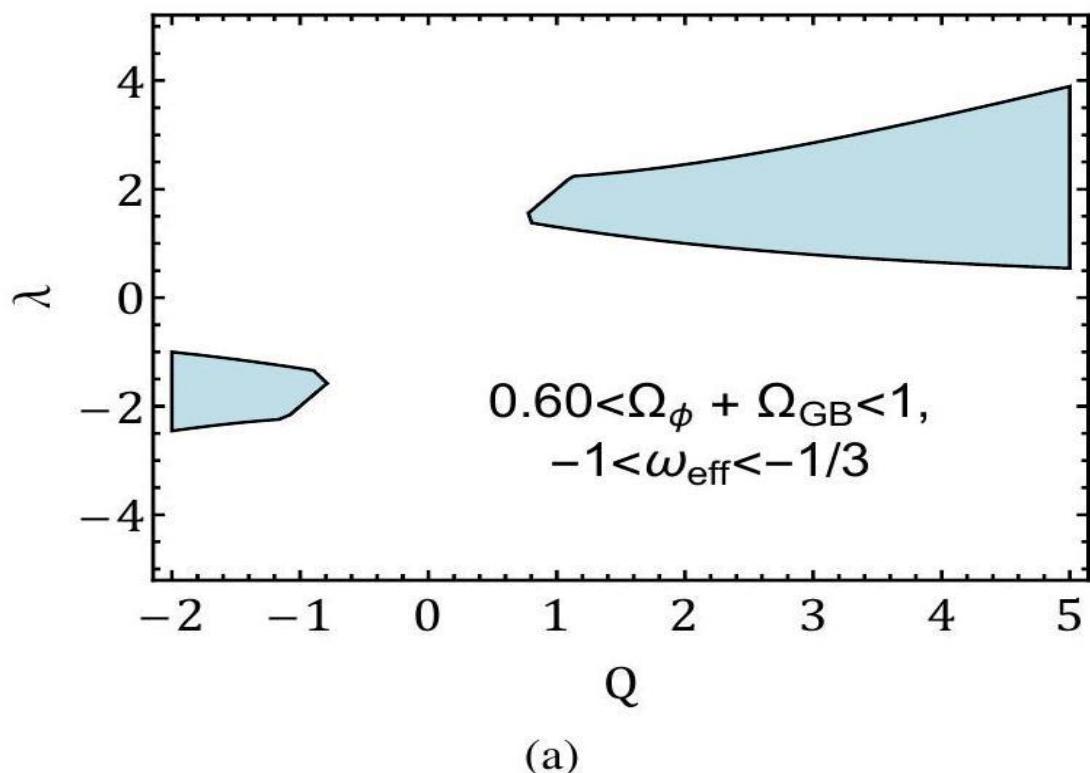
$$\begin{aligned}x' &= \frac{\dot{x}}{H} = -\frac{\dot{H}}{H^2} \left(x + \frac{24z}{\sqrt{6}} \right) - \frac{3Q\Omega_M}{\sqrt{6}} - 3x + \frac{3\lambda y^2}{\sqrt{6}} - \frac{24z}{\sqrt{6}} \\ y' &= \frac{\dot{y}}{H} = \frac{-1}{2}\sqrt{6}\lambda xy - y\frac{\dot{H}}{H^2} \\ z' &= \frac{\dot{z}}{H} = \sqrt{6}\mu xz + 2z\frac{\dot{H}}{H^2}\end{aligned}$$

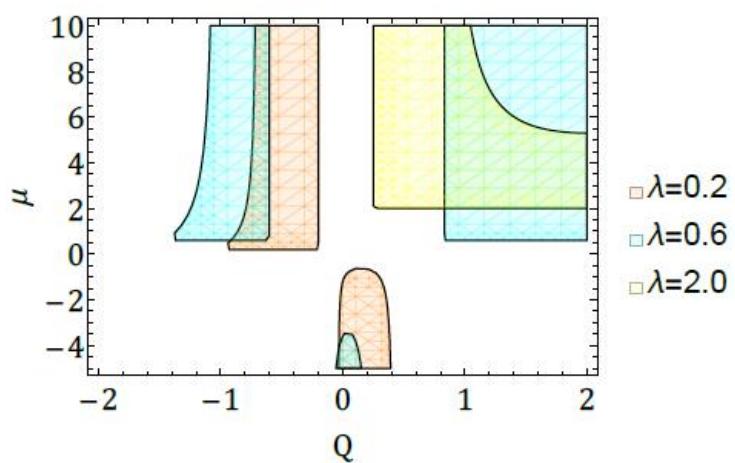
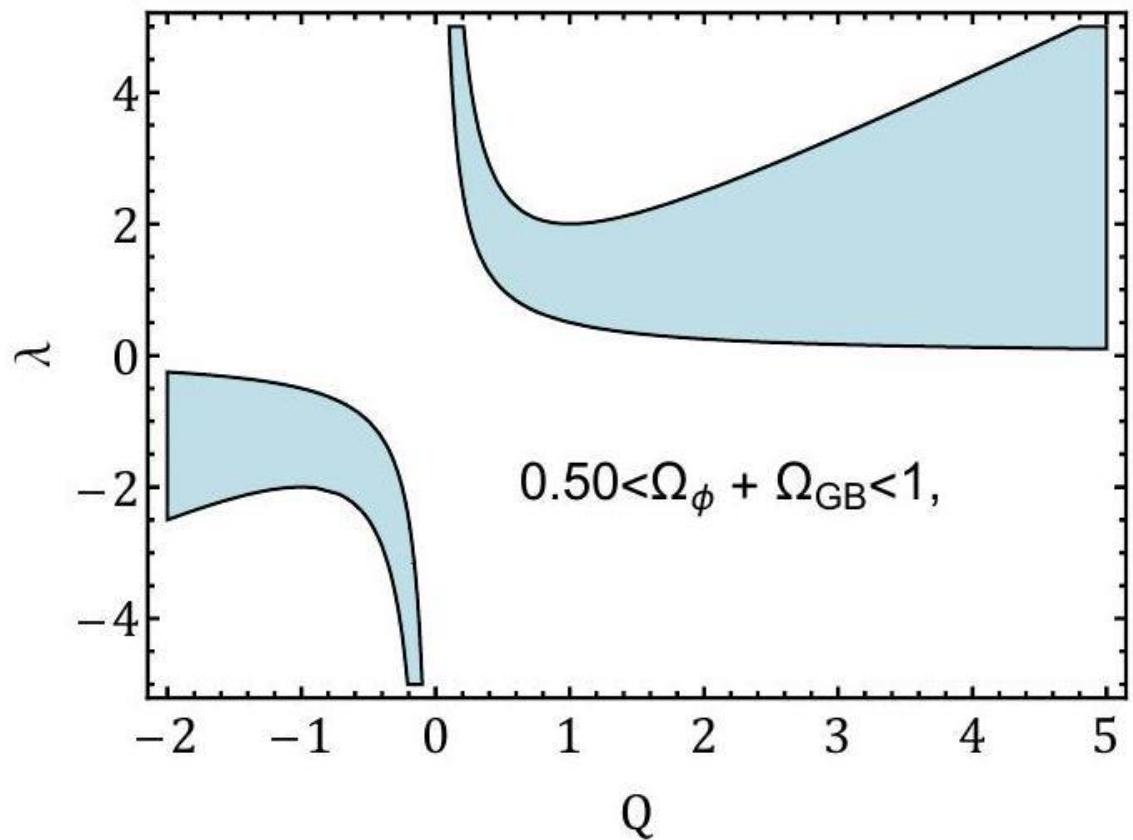


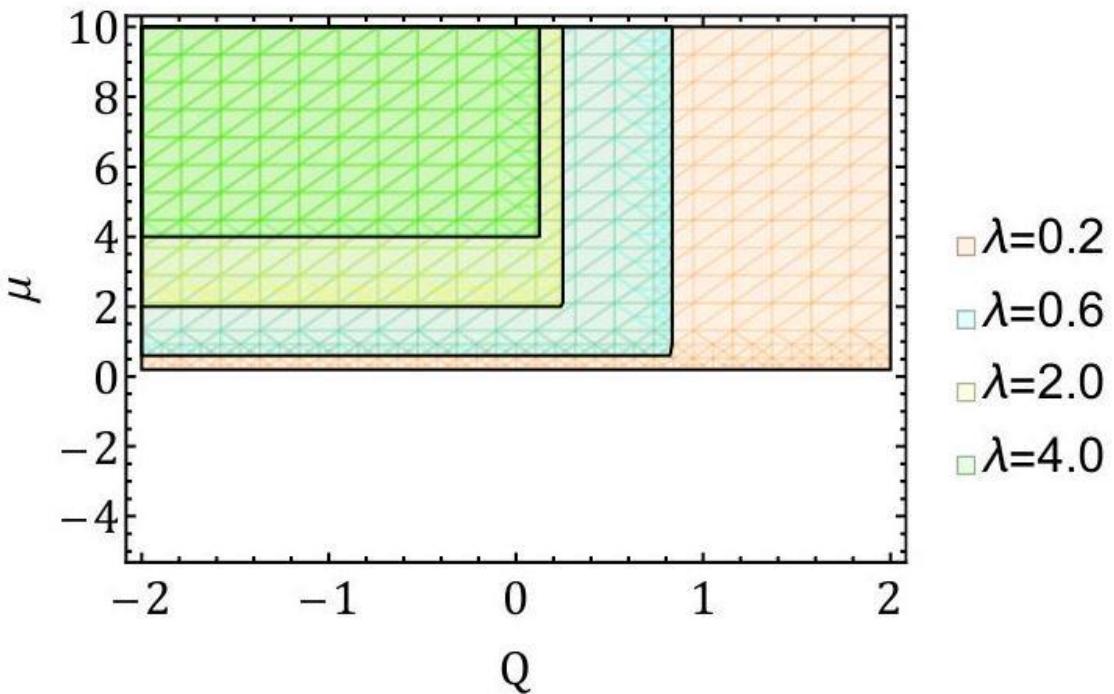
Points	(x_*, y_*, z_*)	Ω_ϕ	Ω_{GB}	ω_{eff}	Eigenvalues
$P_{1,2}$	$(\mp 1, 0, 0)$	1	0	1	$\left(3 \mp \sqrt{6}Q, 3 \pm \sqrt{\frac{3}{2}}\lambda, \sqrt{6}(\mp\mu) - 6\right)$
P_3	$\left(-\sqrt{\frac{2}{3}}Q, 0, 0\right)$	$\frac{2Q^2}{3}$	0	$\frac{2Q^2}{3}$	$\left(Q^2 - \frac{3}{2}, Q(\lambda + Q) + \frac{3}{2}, -2Q(\mu + Q) - 3\right)$
P_4	$\left(\frac{\lambda}{\sqrt{6}}, \left \frac{\sqrt{-((\lambda^2 - 6)(\lambda + Q))}}{\sqrt{6}\sqrt{\lambda + Q}}\right , 0\right)$	1	0	$\frac{1}{3}(\lambda^2 - 3)$	$\left(\frac{1}{2}(\lambda^2 - 6), \lambda(\lambda + Q) - 3, \lambda(\mu - \lambda)\right)$
P_5	$\left(\frac{\sqrt{\frac{3}{2}}}{\lambda + Q}, \left \frac{\sqrt{Q^2 + \lambda Q + \frac{3}{2}}}{\lambda + Q}\right , 0\right)$	$\frac{Q^2 + \lambda Q + 3}{(\lambda + Q)^2}$	0	$-\frac{Q}{\lambda + Q}$	$(E_1, E_2, E_3) - \boxed{}$
P_6	$\left(0, \frac{\sqrt{2}Q^2 + 1}{\sqrt{2}\sqrt{Q}\sqrt{\lambda + Q}}, \frac{1}{16Q}\right)$	$\frac{2Q^2 + 1}{2Q^2 + 2\lambda Q}$	0	-1	$(E_1, E_2, E_3) - \boxed{}$
P_7	$(0, 1, \lambda/8)$	1	0	-1	$6\lambda Q - 3, -\frac{3}{2} \mp \frac{\sqrt{3}\sqrt{(3\lambda^2 + 2)(17\lambda^2 - 8\lambda\mu + 6)}}{6\lambda^2 + 4}$
$P_{8,9}$	-	-	-	-	-

$$\omega_{\text{eff}} = -1 - \frac{2}{3} \frac{\dot{H}}{H^2}.$$





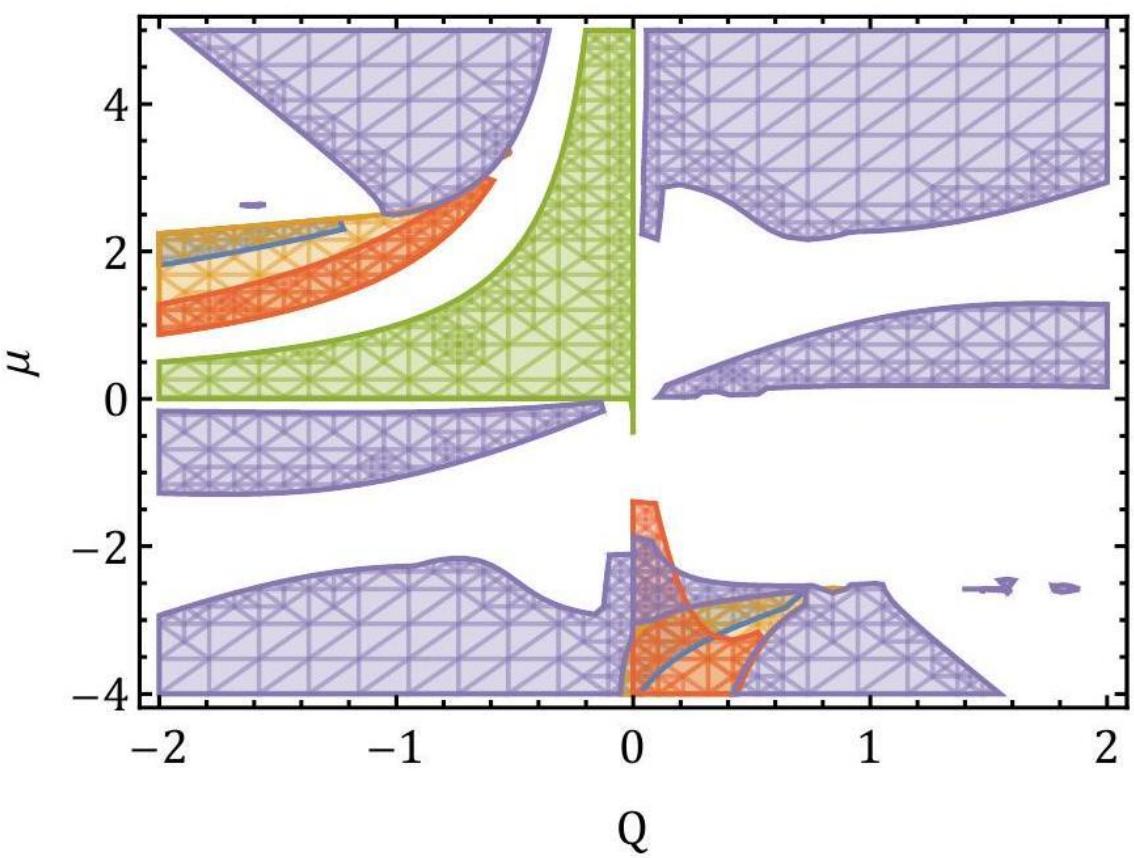
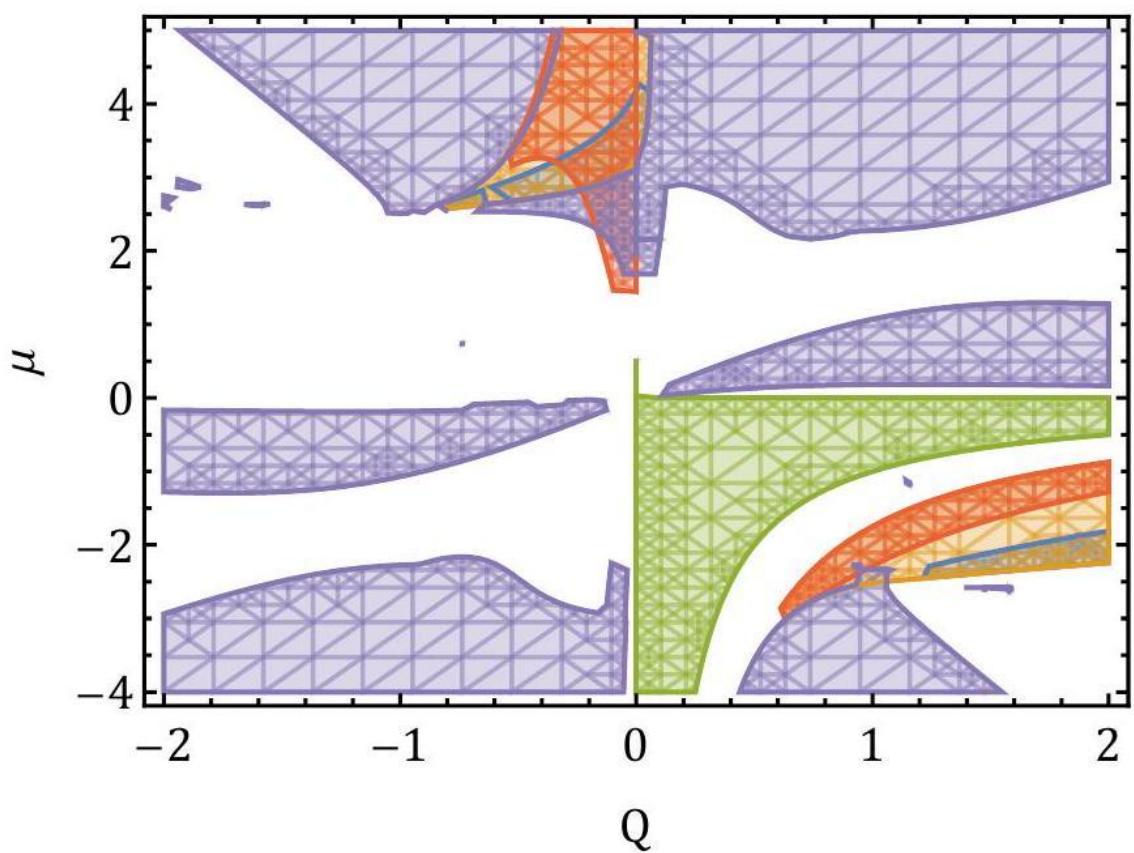




Figuras 36, 37, 38 y 39. Densidades de energía oscura.

$$(x_*, y_*, z_*) = \left(\frac{5\mu + \sqrt{\mu^2 - 48\mu Q^3 + (16 - 48\mu^2)Q^2 - 64\mu Q} - 4Q}{2\sqrt{6}\mu(\mu + Q)}, \frac{\mu - \sqrt{\mu^2 - 48\mu Q^3 + (16 - 48\mu^2)Q^2 - 64\mu Q} + 4Q}{-9 - 96\mu Q} \right).$$



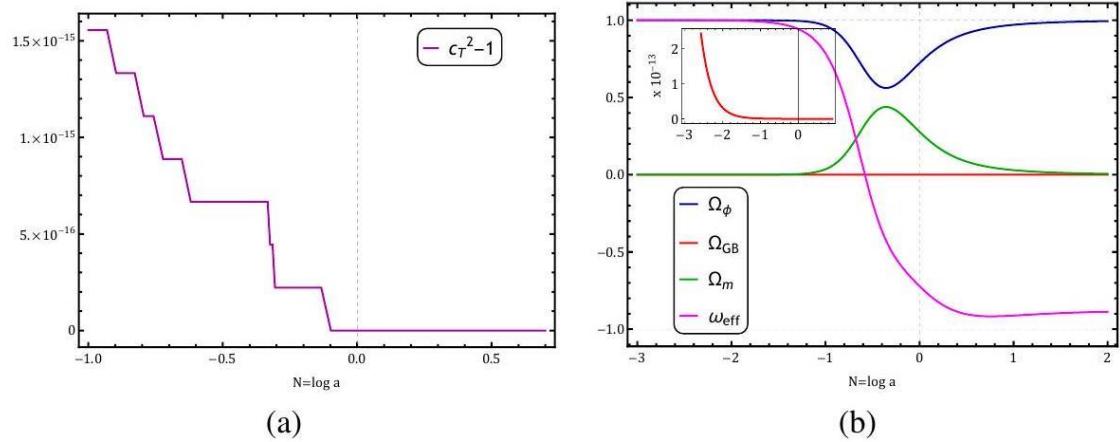
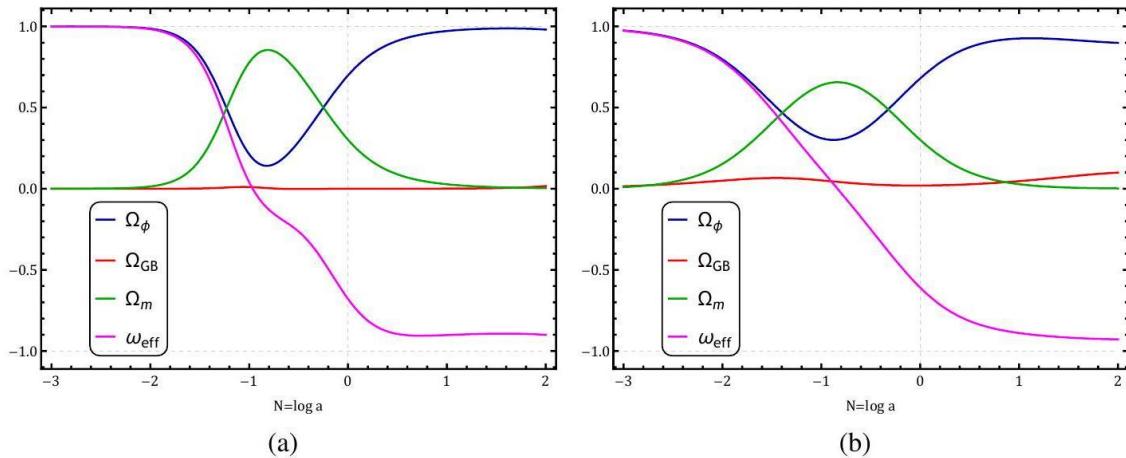


Figuras 39 y 40. Puntos y efectos gravitacionales críticos.

$$\Omega_\phi = \frac{\left(5\mu + \sqrt{\mu^2 - 48\mu Q^3 + (16 - 48\mu^2)Q^2 - 64\mu Q} - 4Q\right)^2}{24\mu^2(\mu + Q)^2}$$

$$(x_*, y_*, z_*) = \begin{pmatrix} \frac{5\mu - \sqrt{\mu^2 - 48\mu Q^3 + (16 - 48\mu^2)Q^2 - 64\mu Q} - 4Q}{2\sqrt{6}\mu(\mu + Q)} \\ \frac{\mu + \sqrt{\mu^2 - 48\mu Q^3 + (16 - 48\mu^2)Q^2 - 64\mu Q} + 4Q}{96\mu Q} \end{pmatrix}$$

$$\Omega_\phi = \frac{(-5\mu + \sqrt{\mu^2 - 48\mu Q^3 + (16 - 48\mu^2)Q^2 - 64\mu Q} + 4Q)^2}{24\mu^2(\mu + Q)^2}$$



Figuras 41 y 42. Parámetro cosmológico – cuántico en un campo escalar oscuro de materia.

$$c_T^2 = \frac{1 - 8\ddot{f}}{1 - 8H\dot{f}}.$$

$$\left| c_T^2 - 1 \right| \leq 5 \times 10^{-16},$$

$$c_T^2=\frac{1-8\left(6\mu zx^2+\sqrt{6}z(x'+x\dot{H}/H^2)\right)}{1-8\sqrt{6}zx}.$$

$$\mu = 5\log_{10}\left(D_L\right) + 25$$

$$D_L=c(1+z)\int_0^z\frac{1}{H}dz$$

$$c_T^2=\frac{1-8\ddot{f}}{1-8H\dot{f}}=1\Longrightarrow \dot{f}=a/f_0$$

$$f_{,\phi}=\frac{a}{f_0}\frac{1}{\dot{\phi}}, f_{,\phi\phi}=\left(\frac{\dot{a}}{f_0}-\frac{a}{f_0}\frac{\ddot{\phi}}{\dot{\phi}}\right)\frac{1}{\dot{\phi}^2}.$$



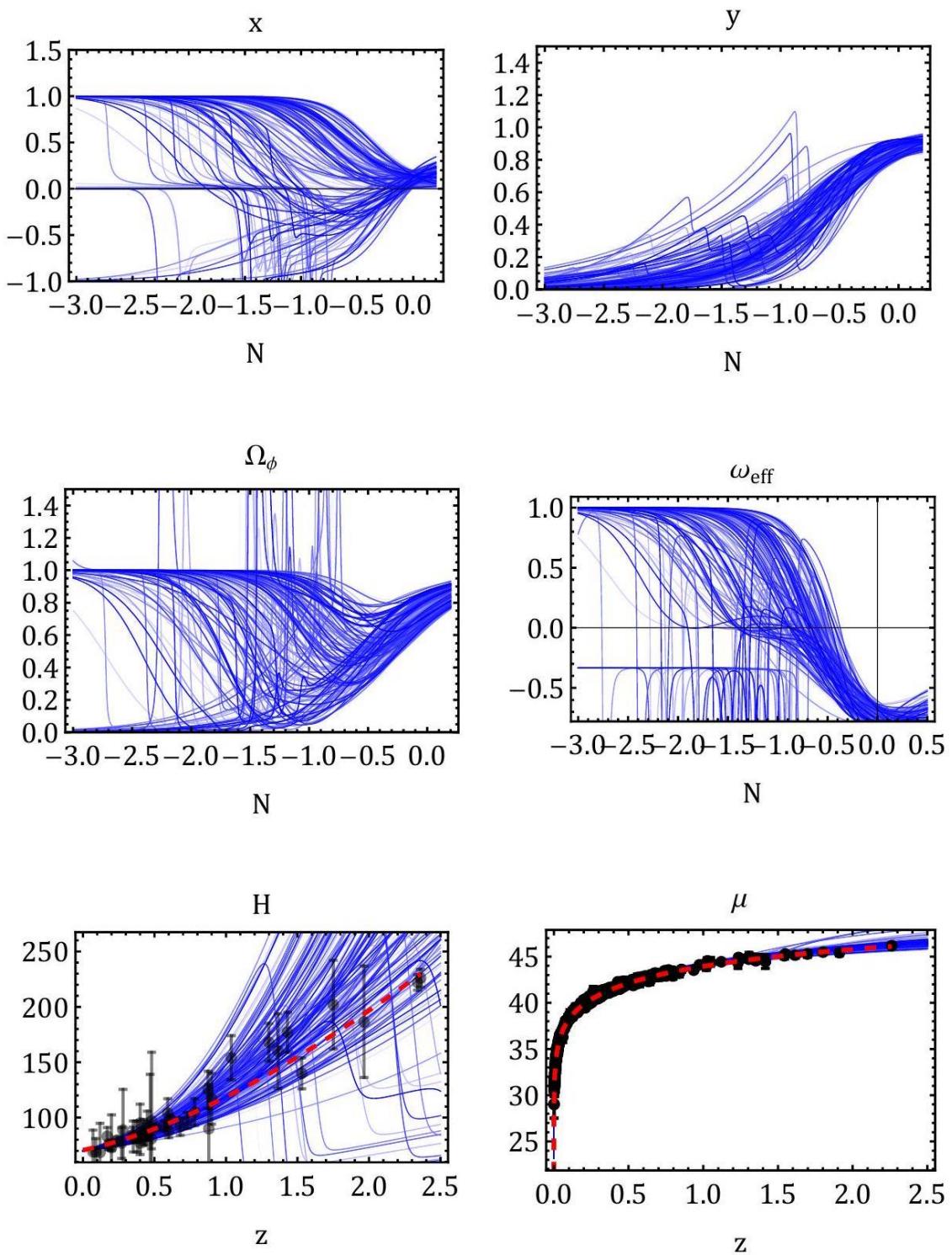


Figura 43. Fluctuaciones gravitacionales de campo causadas por una partícula supermasiva – sector oscuro.

$$\xi = \frac{aH}{a_0 H_0 + aH}$$



$$\xi = \begin{cases} 0, & aH \rightarrow 0 \\ \frac{1}{2}, & aH = a_0 H_0 \\ 1, & aH \gg a_0 H_0 \end{cases}$$

$$\dot{a} = \frac{-\xi \mathcal{M}_0 f_0}{(1-\xi)} \left(\frac{\dot{H}}{H^2} \right) + \frac{\mathcal{M}_0 f_0 \xi'}{(1-\xi)^2}$$

$$\xi' = \xi(1-\xi) \left(1 + \frac{\dot{H}}{H^2} \right)$$

$$\frac{\dot{H}}{H^2} \left(1 - \frac{4\mathcal{M}_0 \xi}{1-\xi} \right) = \frac{-3}{2} \Omega_M - \frac{4\mathcal{M}_0 \xi}{1-\xi} - 3x^2 + \frac{4\mathcal{M}_0 \xi'}{(1-\xi)^2}$$

$$\frac{\dot{H}}{H^2} = \frac{1}{1 - \frac{8\mathcal{M}_0 \xi}{1-\xi}} \left(\frac{-3}{2} \Omega_M - 3x^2 \right)$$

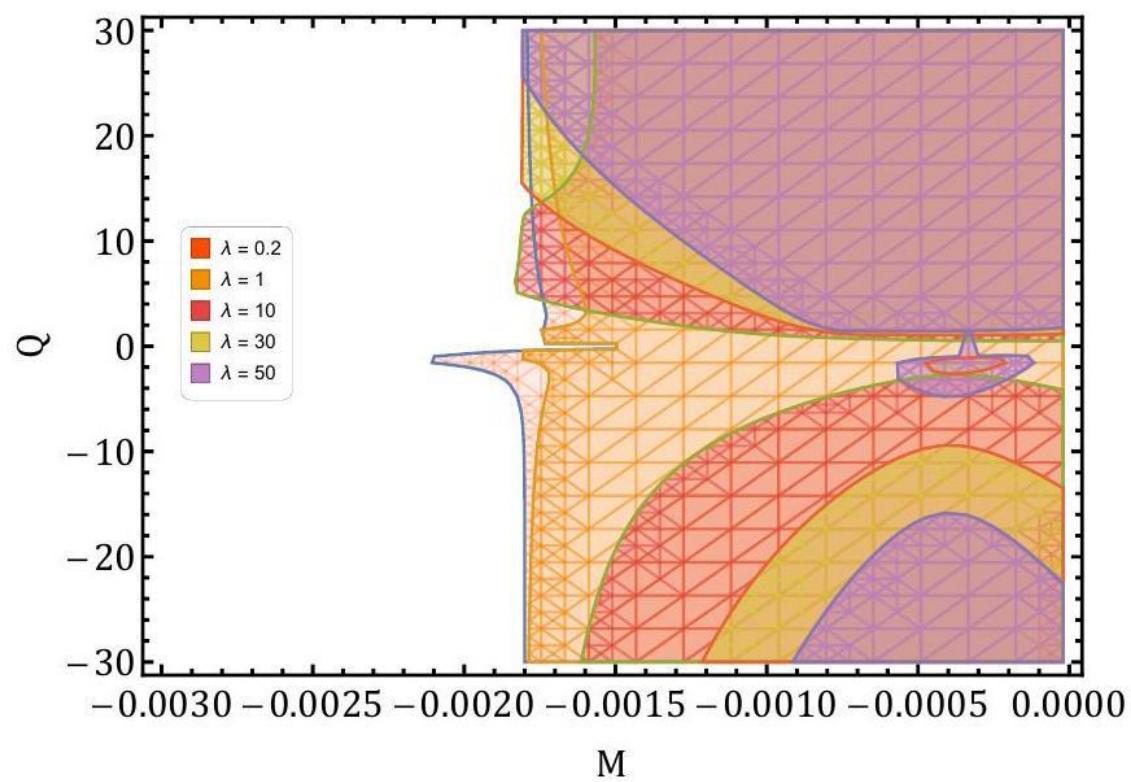
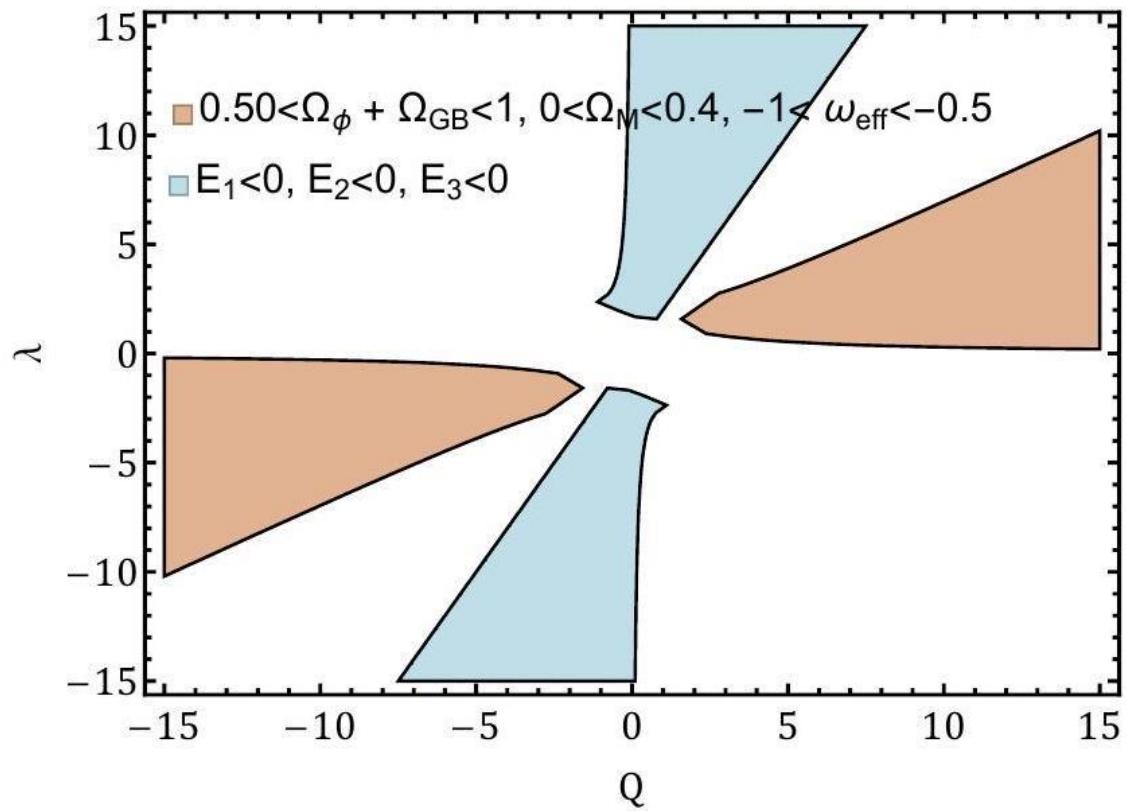
$$\begin{aligned} x' &= \frac{-3Q\Omega_M}{\sqrt{6}} + \frac{3\lambda y^2}{\sqrt{6}} - 3x - \frac{4\mathcal{M}_0 \xi}{(1-\xi)x} - \frac{\dot{H}}{H^2} \left(x + \frac{4\mathcal{M}_0 \xi}{(1-\xi)x} \right) \\ y' &= \frac{-1}{2} \sqrt{6} \lambda xy - y \frac{\dot{H}}{H^2} \\ \xi' &= \xi(1-\xi) \left(1 + \frac{\dot{H}}{H^2} \right) \end{aligned}$$

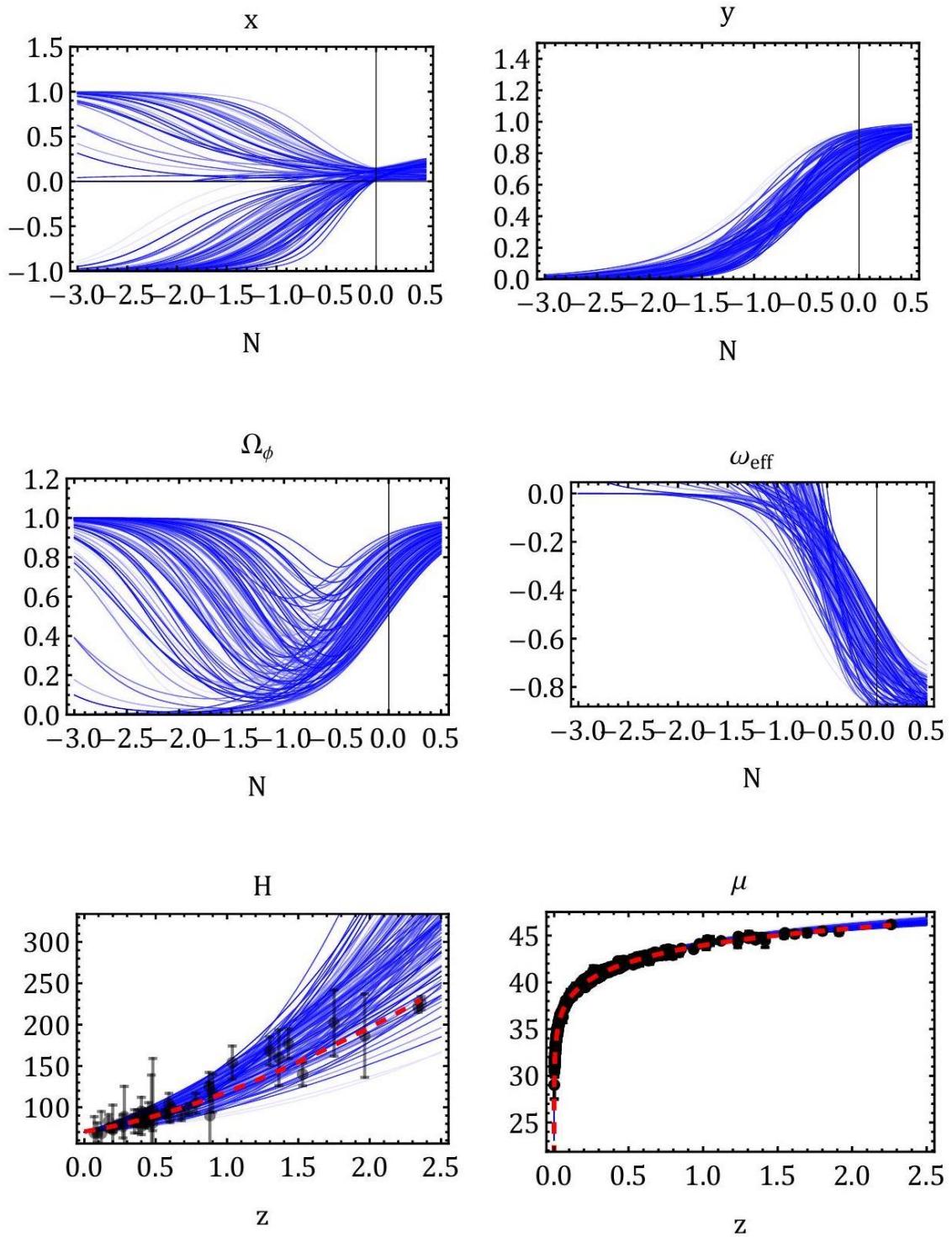
Points	(x_*, y_*, ξ_*)	Ω_ϕ	Ω_{GB}	ω_{eff}
$P_{1,2}$	$(\mp 1, 0, 0)$	1	0	1
P_3	$\left(-\sqrt{\frac{2}{3}}Q, 0, 0 \right)$	$\frac{2Q^2}{3}$	0	$\frac{2Q^2}{3}$
P_4	$\left(\frac{\lambda}{\sqrt{6}}, \frac{\sqrt{-((\lambda^2-6)(\lambda+Q))}}{\sqrt{6}\sqrt{\lambda+Q}}, 0 \right)$	1	0	$\frac{1}{3}(\lambda^2 - 3)$
P_5	$\left(\frac{\sqrt{\frac{3}{2}}}{\lambda+Q}, \frac{\sqrt{Q^2+\lambda Q+\frac{3}{2}}}{\lambda+Q}, 0 \right)$	$\frac{Q^2+\lambda Q+3}{(\lambda+Q)^2}$	0	$-\frac{Q}{\lambda+Q}$
P_6	$\left(\frac{\sqrt{\frac{2}{3}}}{\lambda}, \frac{2}{\sqrt{3}\lambda}, \frac{\lambda^2-2}{\lambda^2(8\mathcal{M}_0+1)-2} \right)$	$\frac{2}{\lambda^2}$	$1 - \frac{2}{\lambda^2}$	$-\frac{1}{3}$
P_7	$\left(-\frac{1}{2\sqrt{6}Q}, \text{Any}, 1 \right)$	$\frac{1}{24Q^2} + y_*^2$	$-\frac{8M\xi_*}{\xi_*-1}$	$-\frac{(\xi_*-1)(24Q^2y_*^2-1)}{24Q^2(8\mathcal{M}_0\xi_*+\xi_*-1)}$

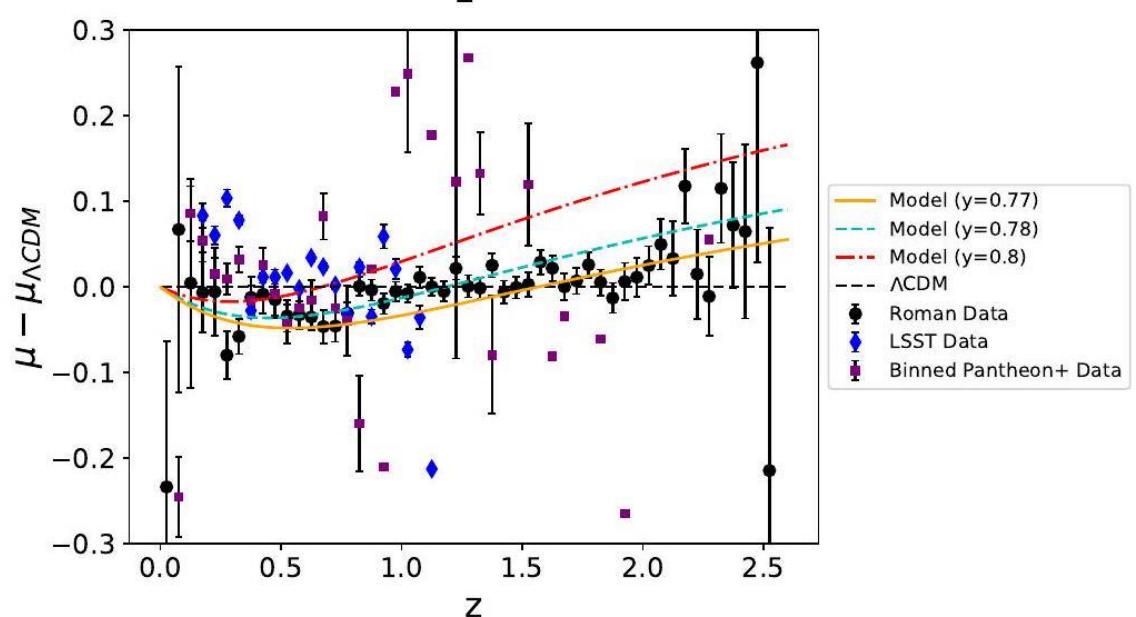
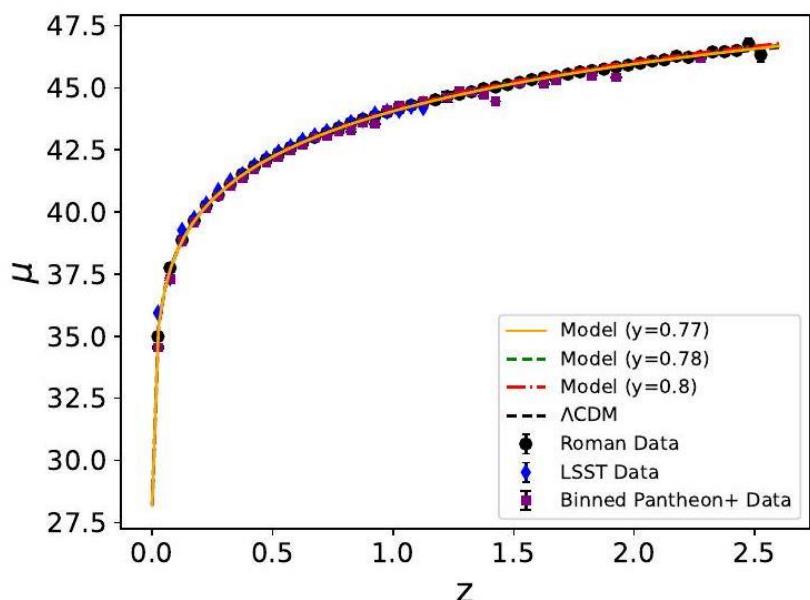
$$\begin{aligned} E_1 &< 0 \rightarrow \\ \left(\lambda > 0, Q < \frac{\lambda}{2} \right) \end{aligned}$$



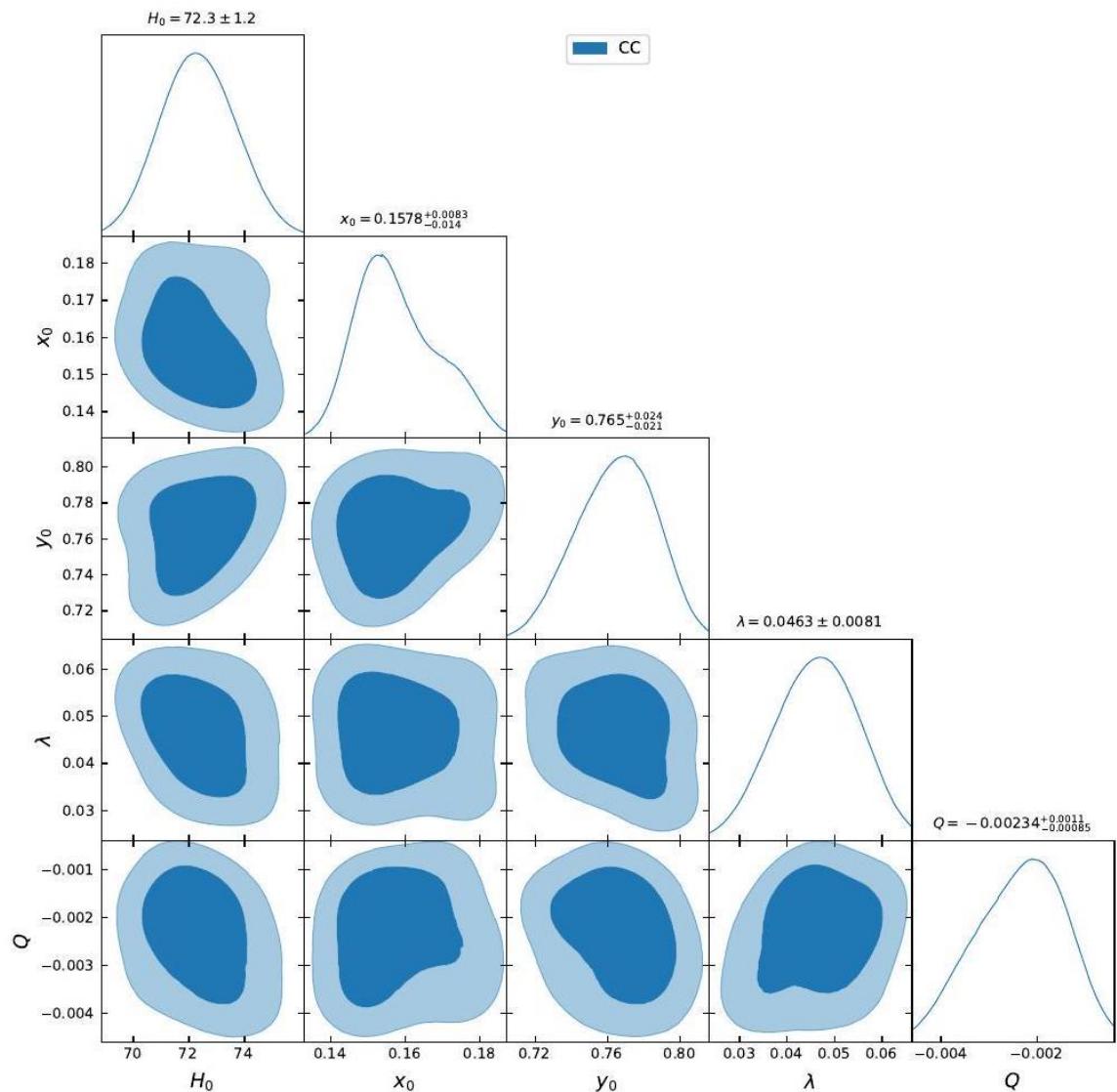
$$\Omega_M = 1 - x^2 - y^2 - \frac{8\mathcal{M}_0\xi}{(1-\xi)}$$

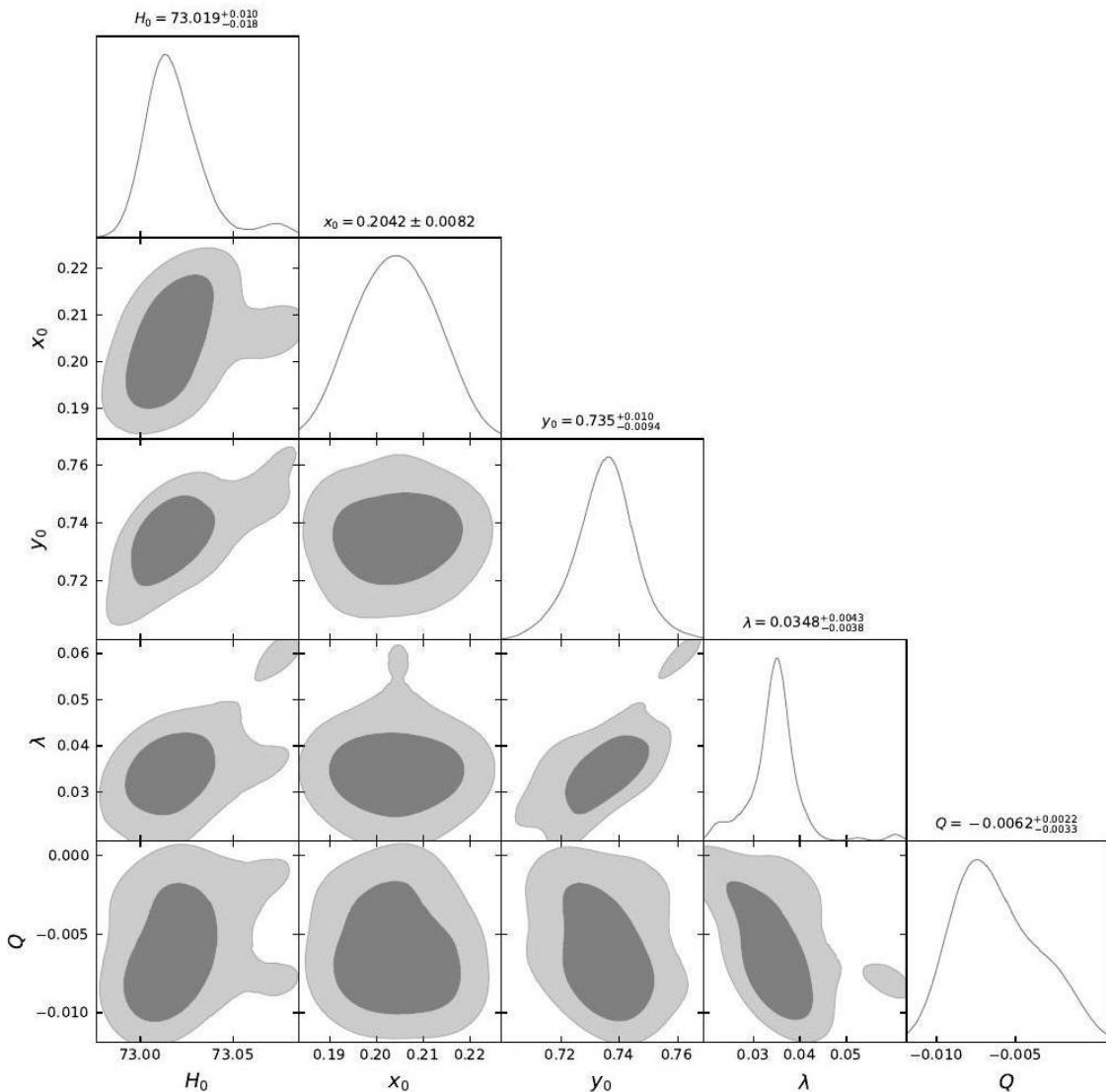






Figuras 44, 45, 46 y 47. Curvatura perturbativa y fluctuaciones gravitacionales extraordinarias.





Figuras 48 y 49. Deformaciones gravitacionales extraordinarias. Parametrización.

SECCIÓN VI.

$$\begin{aligned} \dot{\rho}_{\text{dm}} + 3H\rho_{\text{dm}} &= Q & ; \quad \dot{\rho}_{\text{de}} + 3H\rho_{\text{de}}(1 + \omega) &= -Q \\ \dot{\rho}_{\text{bm}} + 3H\rho_{\text{bm}} &= 0 & ; \quad \dot{\rho}_{\text{r}} + 3H\rho_{\text{r}}(1 + 1/3) &= 0 \end{aligned}$$

$$Q = \begin{cases} > 0 & \text{Dark Energy} \rightarrow \text{Dark Matter (iDEDM regime)} \\ < 0 & \text{Dark Matter} \rightarrow \text{Dark Energy (iDMDE regime)} \\ = 0 & \text{No interaction.} \end{cases}$$

$$\begin{aligned} H^2(a) &= \left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G}{3}(\rho_{\text{r}} + \rho_{\text{bm}} + \rho_{\text{dm}} + \rho_{\text{de}}) - \frac{kc^2}{a^2} \\ q &= \Omega_{\text{r}} + \frac{1}{2}(\Omega_{\text{bm}} + \Omega_{\text{dm}}) + \frac{1}{2}\Omega_{\text{de}}(1 + 3\omega) \\ \omega^{\text{eff}} &= \frac{P_{\text{tot}}}{\rho_{\text{tot}}} = \frac{\frac{1}{3}\Omega_{\text{r}} + \omega_{\text{de}}\Omega_{\text{de}}}{\Omega_{\text{r}} + \Omega_{\text{bm}} + \Omega_{\text{dm}} + \Omega_{\text{de}}} \end{aligned}$$



$$\omega_{\text{dm}}^{\text{eff}} = -\frac{Q}{3H\rho_{\text{dm}}} ; \omega_{\text{de}}^{\text{eff}} = \omega_{\text{de}} + \frac{Q}{3H\rho_{\text{de}}}.$$

$$Q > 0 (\text{iDEDMD}) \begin{cases} \omega_{\text{dm}}^{\text{eff}} < 0 & \text{Dark matter redshifts slower than } a^{-3} (\text{ less DM in past }), \\ \omega_{\text{de}}^{\text{eff}} > \omega_{\text{de}} & \text{Dark energy has less accelerating pressure,} \end{cases}$$

$$Q < 0 (\text{iDMDE}) \begin{cases} \omega_{\text{dm}}^{\text{eff}} > 0 & \text{Dark matter redshifts faster than } a^{-3} (\text{ more DM in past }), \\ \omega_{\text{de}}^{\text{eff}} < \omega_{\text{de}} & \text{Dark energy has more accelerating pressure.} \end{cases}$$

$$\omega_{\text{dm}}^{\text{eff}} = \omega_{\text{dm}} = 0 \text{ and } \omega_{\text{de}}^{\text{eff}} = \omega_{\text{de}}$$

$$r_{\text{IDE}} = \frac{\rho_{\text{dm}}}{\rho_{\text{de}}} = \frac{\rho_{(\text{dm},0)} a^{-3(1+\omega_{\text{dm}}^{\text{eff}})}}{\rho_{(\text{de},0)} a^{-3(1+\omega_{\text{de}}^{\text{eff}})}} = r_0 a^{-\zeta_{\text{IDE}}} ; \text{ with } \zeta_{\text{IDE}} = 3(\omega_{\text{dm}}^{\text{eff}} - \omega_{\text{de}}^{\text{eff}})$$

$$\zeta_{\text{IDE}} = 3(\omega_{\text{dm}}^{\text{eff}} - \omega_{\text{de}}^{\text{eff}}) \begin{cases} Q > 0 (\text{iDEDMD}): & \zeta_{\text{IDE}} < \zeta_{\Lambda\text{CDM}} \text{ alleviates coincidence problem,} \\ Q < 0 (\text{iDMDE}): & \zeta_{\text{IDE}} > \zeta_{\Lambda\text{CDM}} \text{ worsens coincidence problem.} \end{cases}$$

$$\mathbf{d} = \frac{Q}{3H\rho_{\text{de}}(1+\omega)}$$

$$\dot{\Omega}_{\text{x}} = \left[\frac{8\pi\dot{G}}{3H^2} \rho_{\text{x}} \right] = \frac{8\pi G}{3} \left[\frac{\dot{\rho}_{\text{x}}}{H^2} - \rho \frac{2\dot{H}}{H^3} \right] = \frac{8\pi G}{3H^2} \left[\dot{\rho}_{\text{x}} - \rho \frac{2\dot{H}}{H} \right]$$

$$\dot{\rho}_{\text{x}} = -3H\rho_{\text{x}}(1 + \omega_{\text{x}}) \pm Q$$

$$\dot{\Omega}_{\text{x}} = \frac{8\pi G}{3H^2} \left[-3H\rho_{\text{x}}(1 + \omega_{\text{x}}) \pm Q - \rho_{\text{x}} \frac{2\dot{H}}{H} \right] = \frac{8\pi G}{3H^2} \rho_{\text{x}} H \left[-3 \left(1 + \omega_{\text{x}} - \frac{2\dot{H}}{H^2} \right) \pm \frac{8\pi G}{3H^2} Q \right]$$

$$\dot{H} = \frac{d}{dt} \dot{a} a^{-1} = \frac{\ddot{a}}{a} - \left(\frac{\dot{a}}{a} \right)^2 \rightarrow \frac{\dot{H}}{H^2} = \frac{\ddot{a}a}{a^2} - 1 = -q - 1.$$

$$\dot{\Omega}_{\text{x}} = \Omega_{\text{x}} H \left[-3(1 + \omega_{\text{x}}) + 2q + 1 \right] \pm \frac{8\pi G}{3H^2} Q = \Omega_{\text{x}} H [2q - 1 - 3\omega_{\text{x}}] \pm \frac{8\pi G}{3H^2} Q.$$

$$\begin{aligned} \dot{\Omega}_{\text{x}} &= \Omega_{\text{x}} H \left[2 \left(\Omega_{\text{r}} + \frac{1}{2} \Omega_{\text{bm}} + \frac{1}{2} \Omega_{\text{dm}} + \frac{1}{2} \Omega_{\text{de}} (1 + 3\omega_{\text{de}}) \right) - 1 - 3\omega_{\text{x}} \right] \pm \frac{8\pi G}{3H^2} Q \\ &= \Omega_{\text{x}} H [2\Omega_{\text{r}} + \Omega_{\text{bm}} + \Omega_{\text{dm}} + \Omega_{\text{de}} (1 + 3\omega_{\text{de}}) - 1 - 3\omega_{\text{x}}] \pm \frac{8\pi G}{3H^2} Q \end{aligned}$$

$$\dot{\Omega}_{\text{de}} = \Omega_{\text{de}} H [2\Omega_{\text{r}} + \Omega_{\text{bm}} + \Omega_{\text{dm}} + \Omega_{\text{de}} (1 + 3\omega_{\text{de}}) - 1 - 3\omega_{\text{de}}] - \frac{8\pi G}{3H^2} Q$$

$$\dot{\Omega}_{\text{dm}} = \Omega_{\text{dm}} H [2\Omega_{\text{r}} + \Omega_{\text{bm}} + \Omega_{\text{dm}} + \Omega_{\text{de}} (1 + 3\omega_{\text{de}}) - 1] + \frac{8\pi G}{3H^2} Q$$

$$\dot{\Omega}_{\text{bm}} = \Omega_{\text{bm}} H [2\Omega_{\text{r}} + \Omega_{\text{bm}} + \Omega_{\text{dm}} + \Omega_{\text{de}} (1 + 3\omega_{\text{de}}) - 1]$$

$$\dot{\Omega}_{\text{r}} = \Omega_{\text{r}} H [2\Omega_{\text{r}} + \Omega_{\text{bm}} + \Omega_{\text{dm}} + \Omega_{\text{de}} (1 + 3\omega_{\text{de}}) - 2]$$

$$\frac{d\Omega_{\text{de}}}{d\Omega_{\text{dm}}} = \frac{\Omega_{\text{de}} H [2\Omega_{\text{r}} + \Omega_{\text{bm}} + \Omega_{\text{dm}} + \Omega_{\text{de}} (1 + 3\omega_{\text{de}}) - 1 - 3\omega_{\text{de}}] - \frac{8\pi G}{3H^2} Q}{\Omega_{\text{dm}} H [2\Omega_{\text{r}} + \Omega_{\text{bm}} + \Omega_{\text{dm}} + \Omega_{\text{de}} (1 + 3\omega_{\text{de}}) - 1] + \frac{8\pi G}{3H^2} Q}$$



$$Q = \delta H \rho_{\text{de}} ; Q = \delta H \rho_{\text{dm}}$$

$$\frac{d\Omega_{\text{de}}}{d\Omega_{\text{dm}}} = \frac{\Omega_{\text{de}}[2\Omega_r + \Omega_{\text{bm}} + \Omega_{\text{dm}} + \Omega_{\text{de}}(1 + 3\omega_{\text{de}}) - 1 - 3\omega_{\text{de}} - \delta]}{\Omega_{\text{dm}}[2\Omega_r + \Omega_{\text{bm}} + \Omega_{\text{dm}} + \Omega_{\text{de}}(1 + 3\omega_{\text{de}}) - 1] + \delta\Omega_{\text{de}}}$$

$$\frac{d\Omega_{\text{de}}}{d\Omega_m} = \frac{\Omega_{\text{de}}[\Omega_m + \Omega_{\text{de}}(1 + 3\omega_{\text{de}}) - 1 - 3\omega_{\text{de}} - \delta]}{\Omega_m[\Omega_m + \Omega_{\text{de}}(1 + 3\omega_{\text{de}}) - 1] + \delta\Omega_{\text{de}}}$$

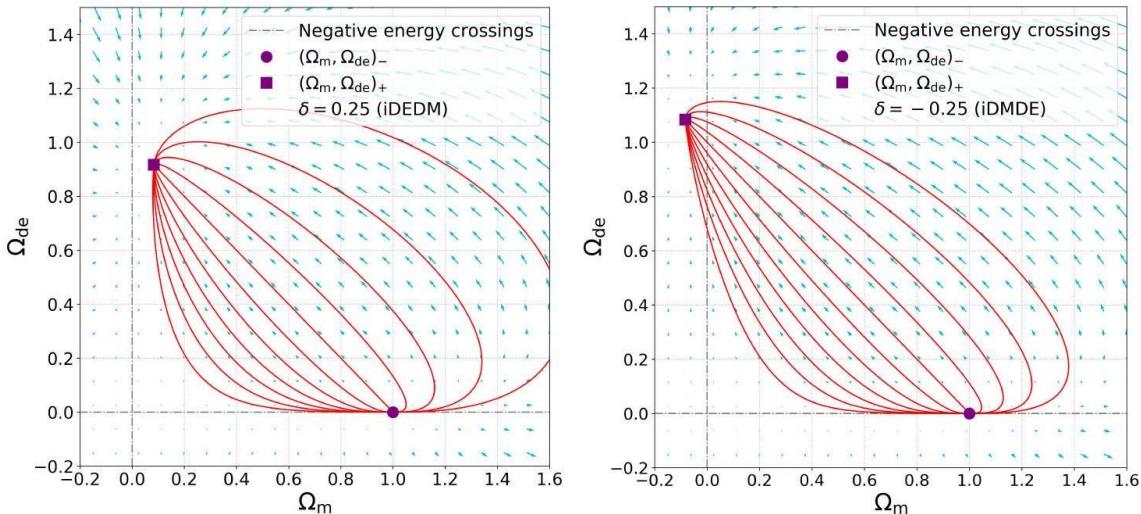


Figura 50. Deformaciones gravitacionales en pluridimensionalidad.

$$(\Omega_m, \Omega_{\text{de}})_- = (1, 0); (\Omega_m, \Omega_{\text{de}})_+ = \left(-\frac{\delta}{3\omega}, 1 + \frac{\delta}{3\omega}\right).$$

$$\left(-\frac{\delta}{3\omega}, 1 + \frac{\delta}{3\omega}\right)$$

$$r_- = \frac{\Omega_{(m,-)}}{\Omega_{(de,-)}} = \frac{1}{0} \rightarrow \infty; r_+ = \frac{\Omega_{(dm,+)}}{\Omega_{(de,-)}} \approx \frac{\Omega_{(m,+)}}{\Omega_{(de,+)}} = \frac{-\frac{\delta}{3\omega}}{1 + \frac{\delta}{3\omega}} \rightarrow -\frac{\delta}{\delta + 3\omega}$$

$$\begin{aligned} \rho_{\text{dm}} &= \left(\rho_{(\text{dm},0)} + \rho_{(\text{de},0)} \frac{\delta}{\delta + 3\omega} [1 - a^{-(\delta + 3\omega)}]\right) a^{-3} \\ \rho_{\text{de}} &= \rho_{(\text{de},0)} a^{-(\delta + 3\omega + 3)} \end{aligned}$$

$$\omega_{\text{dm}}^{\text{eff}} = -\frac{Q}{3H\rho_{\text{de}}} = -\frac{\delta H \rho_{\text{de}}}{3H\rho_{\text{dm}}} = -\frac{\delta}{3} \frac{\rho_{\text{de}}}{\rho_{\text{dm}}} = -\frac{\delta}{3} \frac{1}{r}$$

$$\omega_{\text{de}}^{\text{eff}} = \omega + \frac{Q}{3H\rho_{\text{de}}} = \omega + \frac{\delta H \rho_{\text{de}}}{3H\rho_{\text{de}}} = \omega + \frac{\delta}{3}$$



$$\Omega_{\text{dm}} = \frac{H_0^2}{H^2} \left(\Omega_{(\text{dm},0)} + \Omega_{(\text{de},0)} \frac{\delta}{\delta + 3\omega} [1 - (1+z)^{(\delta+3\omega)}] \right) (1+z)^3$$

$$\Omega_{\text{de}} = \frac{H_0^2}{H^2} \Omega_{(\text{de},0)} (1+z)^{(\delta+3\omega+3)}$$

$$\Omega_{\text{bm}} = \frac{H_0^2}{H^2} \Omega_{(\text{bm},0)} (1+z)^3$$

$$\Omega_r = \frac{H_0^2}{H^2} \Omega_{(r,0)} (1+z)^4$$

$$a^{-(\delta+3\omega)} = 1 + r_0 \left(\frac{\delta + 3\omega}{\delta} \right)$$

$$z_{(\text{dm}=0)} = \left[1 + r_0 \left(\frac{\delta + 3\omega}{\delta} \right) \right]^{\frac{1}{\delta+3\omega}} - 1$$

$$a^{-(\delta+3\omega)} = 1 + r_0 \left(\frac{\delta + 3\omega}{\delta} \right) \text{ where } (\delta + 3\omega < 0)$$

$$\delta < 0 \Rightarrow \begin{cases} \text{Past} & (a < 1) \rightarrow (0 < \text{L.H.S.} < 1; \text{R.H.S.} > 1) \\ \text{Future} & (a > 1) \rightarrow (\text{L.H.S.} > 1; \text{R.H.S.} > 1) \end{cases}$$

$$\delta > 0 \Rightarrow \begin{cases} \text{Past} & (a < 1) \rightarrow (0 < \text{L.H.S.} < 1; \text{R.H.S.} < 1) \\ \text{Future} & (a > 1) \rightarrow (\text{L.H.S.} > 1; \text{R.H.S.} < 1) \end{cases}$$

$$1 + r_0 \left(\frac{\delta + 3\omega}{\delta} \right) < 0 \rightarrow \delta < -\frac{3\omega}{\left(1 + \frac{1}{r_0} \right)}$$

$$0 < \delta < -\frac{3\omega}{\left(1 + \frac{1}{r_0} \right)}$$

$$r(z) = \frac{\rho_{\text{dm}}(z)}{\rho_{\text{de}}(z)} = \left(r_0 + \frac{\delta}{\delta + 3\omega} \right) (1+z)^{-(\delta+3\omega)} - \frac{\delta}{\delta + 3\omega}$$

$$r \propto a^{(\delta+3\omega)} \rightarrow \zeta_{Q_1} = \zeta_Q = -3\omega - \delta$$

$$\zeta_Q = -3\omega - \delta \rightarrow \begin{cases} \text{if } \delta > 0 \text{(iDEDM)} & \rightarrow \zeta_Q < \zeta \\ \text{if } \delta < 0 \text{(iDMDE)} & \rightarrow \zeta_Q > \zeta \end{cases} \text{ worsens coincidence problem.}$$

$$\lim_{(1+z) \rightarrow \infty} r_- \rightarrow \infty; \quad \lim_{(1+z) \rightarrow 0} r_+ = \rightarrow -\frac{\delta}{\delta + 3\omega}$$

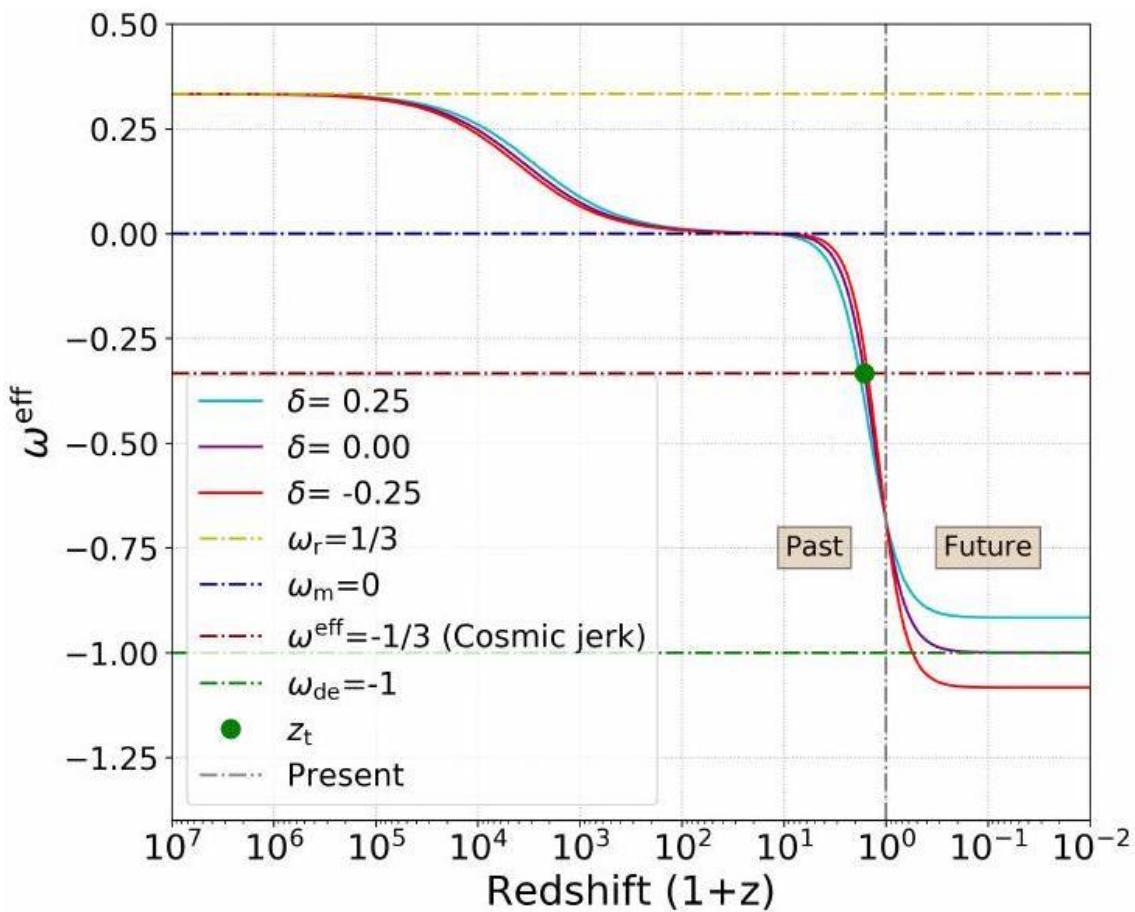
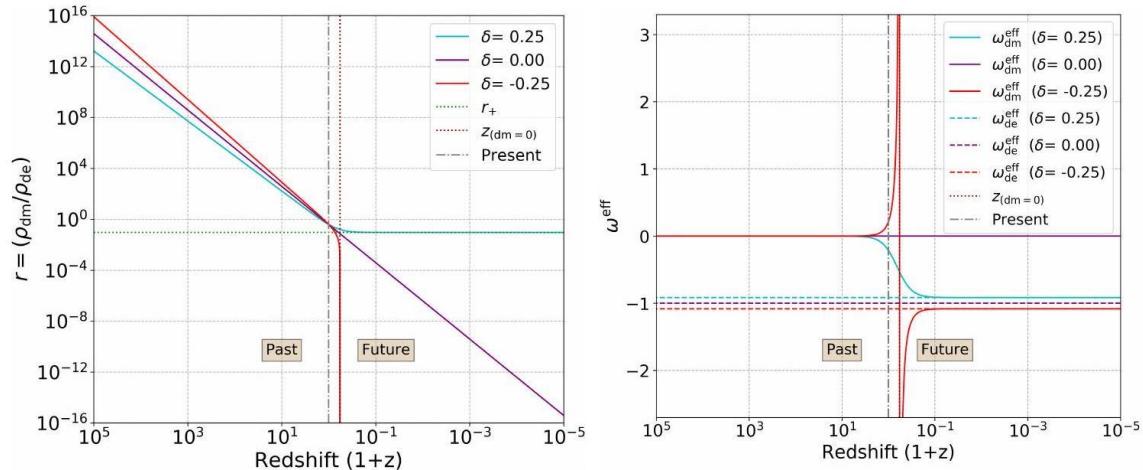
$$\lim_{(1+z) \rightarrow 0} r_+ \propto a^0 \rightarrow \zeta_{(Q,-)} = 0$$

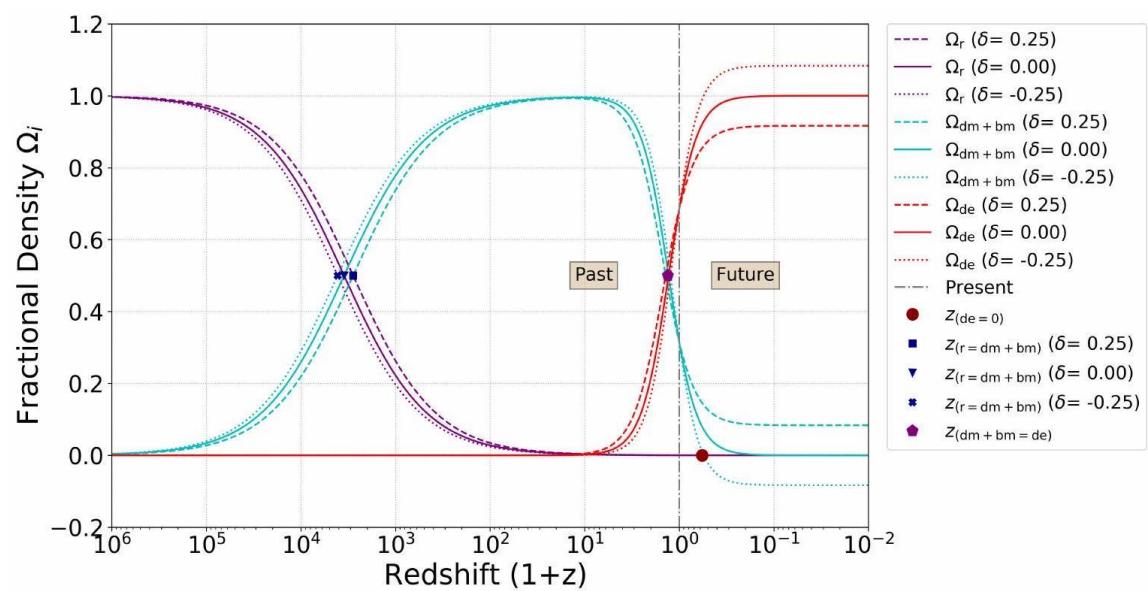
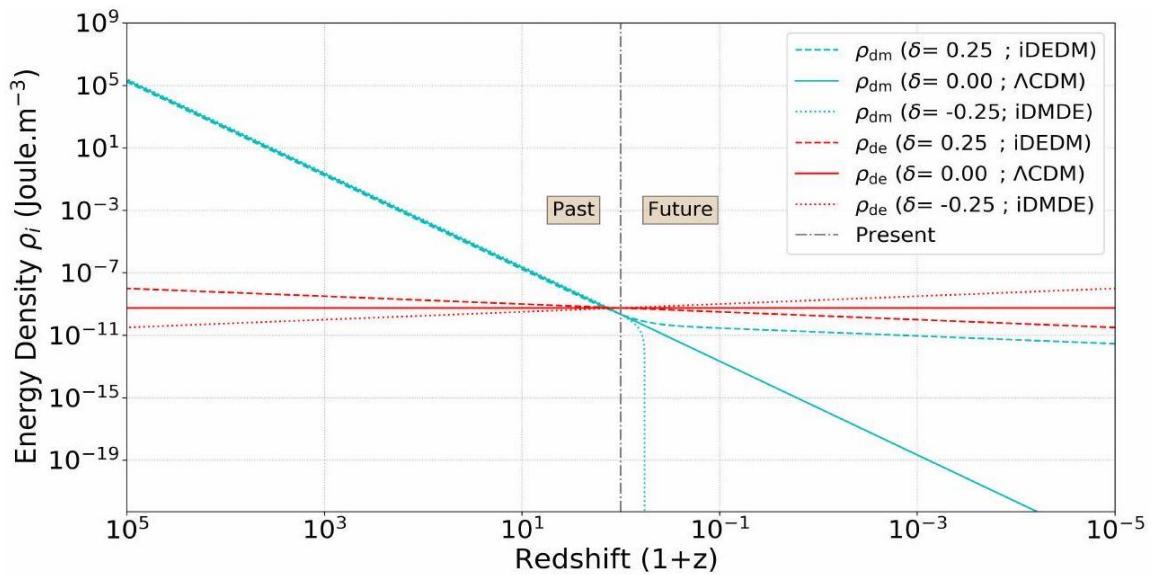
$$\lim_{(1+z) \rightarrow 0} \zeta_Q = 0 \begin{cases} \text{if } \delta > 0 \rightarrow r_- = +\text{constant} & \text{solves coincidence problem} \\ \text{if } \delta < 0 \rightarrow r_- = -\text{constant} & \text{negative energy densities (unphysical).} \end{cases}$$

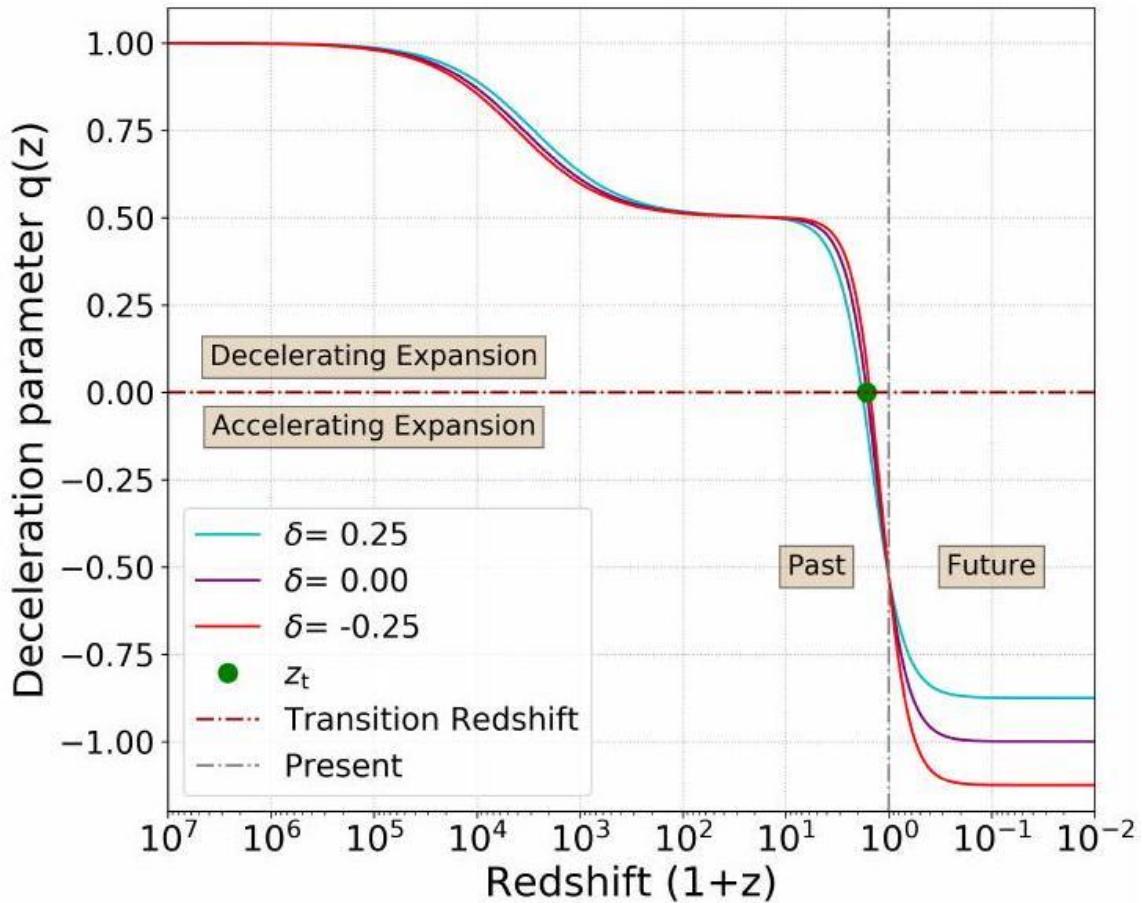


$$\omega_{\text{dm}}^{\text{eff}} = -\frac{\delta}{3} \frac{1}{r} = -\frac{\delta}{3} \frac{1}{(r_0 + \frac{\delta}{\delta + 3\omega}) (1+z)^{-(\delta+3\omega)}} - \frac{\delta}{\delta + 3\omega}$$

$$\omega_{\text{dm}}^{\text{eff}} = -\frac{\delta}{3} \frac{1}{r} \begin{cases} \text{Distant past } (r = r_-): & \omega_{\text{dm}}^{\text{eff}} = -\frac{\delta}{3} \frac{1}{\infty} = 0 = \omega_{\text{dm}} \\ \text{Distant future } (r = r_+): & \omega_{\text{dm}}^{\text{eff}} = -\frac{\delta}{3} \left(\frac{\delta}{\delta + 3\omega} \right) = \omega + \frac{\delta}{3} = \omega_{\text{de}}^{\text{eff}} \end{cases}$$







Figuras 51, 52, 53, 54 y 55. Multidimensiones por aceleración y desaceleración inflacionarias de un campo cuántico – relativista oscuro.

$$\delta > 0 \text{ (iDEDM)} \begin{cases} \text{Past expansion: } \omega_{de}^{eff} > \omega_{de}(\zeta_Q < \zeta) & \text{alleviates coincidence problem} \\ \text{Future expansion: } \omega_{dm}^{eff} = \omega_{de}^{eff}(\zeta_Q = 0) & \text{solves coincidence problem,} \end{cases}$$

$$\delta < 0 \text{ (iDMDE)} \begin{cases} \text{Past expansion: } \omega_{de}^{eff} < \omega_{de}(\zeta_Q > \zeta) & \text{worsens coincidence problem} \\ \text{Future expansion: } \omega_{dm}^{eff} = \omega_{de}^{eff}(\rho_{de} < 0) & \text{negative energy densities.} \end{cases}$$

$$z_{(r=dm+bm)} \approx \left(\frac{\Omega_{(bm,0)} + \Omega_{(dm,0)} + \Omega_{(de,0)} \frac{\delta}{\delta + 3\omega}}{\Omega_{(r,0)}} \right)^{\frac{1}{\delta + 3\omega}} - 1.$$

$$z_{(dm+bm=de)} = \left(\frac{\frac{\Omega_{(bm,0)} + \Omega_{(dm,0)}}{\Omega_{(de,0)}} + \frac{\delta}{\delta + 3\omega}}{\left(1 + \frac{\delta}{\delta + 3\omega}\right)} \right)^{\frac{1}{\delta + 3\omega}} - 1.$$

$$\delta > 0 \text{ (iDEDM)} \begin{cases} \text{Radiation-matter equality: } z_{IDE} < z_{\Lambda CDM} & \text{happens later than } \Lambda CDM \\ \text{Matter-dark energy equality: } z_{IDE} > z_{\Lambda CDM} & \text{happens earlier than } \Lambda CDM, \end{cases}$$

$$\delta < 0 \text{ (iDMDE)} \begin{cases} \text{Radiation-matter equality: } z_{IDE} > z_{\Lambda CDM} & \text{happens earlier than } \Lambda CDM \\ \text{Matter-dark energy equality: } z_{IDE} < z_{\Lambda CDM} & \text{happens later than } \Lambda CDM, \end{cases}$$



$$(\Omega_{\text{dm}}, \Omega_{\text{de}})_- = (1,0), (\Omega_{\text{dm+bm}}, \Omega_{\text{de}})_+ = \left(-\frac{\delta}{3\omega}, 1 + \frac{\delta}{3\omega}\right)$$

$$\rightarrow z_t = \left(-\frac{\frac{\Omega_{(\text{bm},0)} + \Omega_{(\text{dm},0)}}{\Omega_{(\text{de},0)}} + \frac{\delta}{\delta + 3\omega}}{1 + 3\omega + \frac{\delta}{\delta + 3\omega}} \right)^{\left(\frac{1}{\delta + 3\omega}\right)} - 1$$

Cosmic jerk (z_t) $\begin{cases} \delta > 0 \text{(iDEDM): } z_{\text{IDE}} > z_{\Lambda\text{CDM}} & \text{happens earlier than } \Lambda\text{CDM,} \\ \delta < 0 \text{(iDMDE): } z_{\text{IDE}} < z_{\Lambda\text{CDM}} & \text{happens later than } \Lambda\text{CDM.} \end{cases}$

$$\omega_+^{\text{eff}} = \frac{\frac{1}{3}\Omega_r + \omega\Omega_{\text{de}}}{\Omega_r + \Omega_{\text{bm}} + \Omega_{\text{dm}} + \Omega_{\text{de}}} = \frac{\frac{1}{3}(0) + \omega\left(1 + \frac{\delta}{3\omega}\right)}{(0) + \left(-\frac{\delta}{3\omega}\right) + \left(1 + \frac{\delta}{3\omega}\right)} = \frac{\omega + \frac{\delta}{3}}{1} = \omega + \frac{\delta}{3} = \omega_{\text{de}}^{\text{eff}}$$

$$\mathbf{d} = \frac{Q}{3H\rho_{\text{de}}(1+\omega)} = \frac{\delta H\rho_{\text{de}}}{3H\rho_{\text{de}}(1+\omega)} = \frac{\delta}{3(1+\omega)},$$

$\mathbf{d} < 0 \begin{cases} \delta < 0 & ; \quad \omega > -1 \quad (\text{Quintessence regime}) \\ \delta > 0 & ; \quad \omega < -1 \quad (\text{Phantom regime}) \end{cases} \rightarrow \begin{array}{ccccc} & & \text{No} & \text{instabilities} & \text{expected} \end{array}$

$\mathbf{d} > 0 \begin{cases} \delta > 0 & ; \quad \omega > -1 \quad (\text{Quintessence regime}) \\ \delta < 0 & ; \quad \omega < -1 \end{cases} \rightarrow \text{Instabilities can develop if } \mathbf{d} > 1$

$$\rho_{\text{de}} = \rho_{(\text{de},0)} a^{-3\left(1+\omega+\frac{\delta}{3}\right)}, -3\left(1+\omega+\frac{\delta}{3}\right) > 0 \text{ if } \omega_{\text{de}}^{\text{eff}} = \omega + \frac{\delta}{3} < -1$$

$$t_{rip} \approx -\frac{2}{3H_0\left(1+\omega+\frac{\delta}{3}\right)\sqrt{\left(1-\frac{\delta}{\delta+3\omega}\right)(1-\Omega_{(\text{dm+bm},0)})}}$$

$$3(\omega + 1) < \delta < -\frac{3\omega}{\left(1 + \frac{1}{r_0}\right)} \text{ with } \omega_{\text{de}} < -1$$

$$\begin{aligned} \rho_{\text{dm}} &= \rho_{(\text{dm},0)} a^{(\delta-3)} \\ \rho_{\text{de}} &= \left[\rho_{(\text{de},0)} + \rho_{(\text{dm},0)} \frac{\delta}{\delta + 3\omega} (1 - a^{\delta+3\omega}) \right] a^{-3(1+\omega)} \end{aligned}$$

$$\omega_{\text{dm}}^{\text{eff}} = -\frac{\delta}{3}, \omega_{\text{de}}^{\text{eff}} = \omega_{\text{de}} + \frac{\delta}{3}r$$

$$r(z) = \frac{1}{\left(\frac{1}{r_0} + \frac{\delta}{\delta + 3\omega}\right)(1+z)^{(\delta+3\omega)} - \frac{\delta}{\delta + 3\omega}},$$

$$0 < \delta < -\frac{3\omega}{(1+r_0)}$$



$$z_{(\text{de}=0)} = \left[1 + \frac{1}{r_0} \left(\frac{\delta + 3\omega}{\delta} \right) \right]^{-\frac{1}{\delta+3\omega}} - 1$$

$$\begin{aligned} \delta > 0 (\text{iDEDM}) & \left\{ \begin{array}{ll} \text{Past expansion:} & \omega_{\text{dm}}^{\text{eff}} = \omega_{\text{de}}^{\text{eff}}(\zeta_Q = 0) \\ \text{Future expansion:} & \omega_{\text{dm}}^{\text{eff}} < \omega_{\text{dm}}^{\text{eff}}(\zeta_Q < \zeta) \end{array} \right. \begin{array}{l} \text{solves coincidence problem} \\ \text{alleviates coincidence problem,} \end{array} \\ \delta < 0 (\text{iDMDE}) & \left\{ \begin{array}{ll} \text{Past expansion:} & \omega_{\text{dm}}^{\text{eff}} = \omega_{\text{de}}^{\text{eff}}(\rho_{\text{de}} < 0) \\ \text{Future expansion:} & \omega_{\text{dm}}^{\text{eff}} > \omega_{\text{dm}}^{\text{eff}}(\zeta_Q > \zeta) \end{array} \right. \begin{array}{l} \text{negative energy densities} \\ \text{worsens coincidence problem.} \end{array} \end{aligned}$$

$$\mathbf{d} = \frac{Q}{3H\rho_{\text{de}}(1+\omega)} = \frac{\delta H\rho_{\text{dm}}}{3H\rho_{\text{de}}(1+\omega)} = \frac{\delta}{3(1+\omega)} \frac{\rho_{\text{dm}}}{\rho_{\text{de}}},$$

$$t_{\text{rip}} \approx -\frac{2}{3H_0(1+\omega)\sqrt{1-\Omega_{(\text{bm},0)}-\left(1-\frac{\delta}{\delta+3\omega}\right)\Omega_{(\text{dm},0)}}},$$

$$\begin{aligned} \left(\frac{\dot{a}}{a}\right) &\approx H_0 \sqrt{\left(\Omega_{(\text{dm},0)} + \Omega_{(\text{de},0)} \frac{\delta}{\delta+3\omega} [1-a^{-(\delta+3\omega)}]\right) a^{-3} + \Omega_{(\text{de},0)} a^{-3\left(1+\omega+\frac{\delta}{3}\right)}} \\ &= H_0 \sqrt{\left(\Omega_{(\text{dm},0)} + \Omega_{(\text{de},0)} \frac{\delta}{\delta+3\omega}\right) a^{-3} + \left(1 - \frac{\delta}{\delta+3\omega}\right) \Omega_{(\text{de},0)} a^{-3\left(1+\omega+\frac{\delta}{3}\right)}} \end{aligned}$$

$$\left(\frac{\dot{a}}{a}\right) \approx H_0 \sqrt{\left(1 - \frac{\delta}{\delta+3\omega}\right) \Omega_{(\text{de},0)} a^{-3\left(1+\omega+\frac{\delta}{3}\right)}}$$

$$\left(\frac{\dot{a}}{a}\right) \approx H_0 \sqrt{\left(1 - \frac{\delta}{\delta+3\omega}\right) (1 - \Omega_{(\text{dm+bm},0)})} a^{-\frac{3}{2}(1+\omega+\frac{\delta}{3})}$$

$$\begin{aligned} \frac{da}{dt} &= H_0 \sqrt{\left(1 - \frac{\delta}{\delta+3\omega}\right) (1 - \Omega_{(\text{dm+bm},0)})} a^{-\frac{1}{2}(1+3\omega+\delta)} \\ \int_{t_0}^{t_{\text{rip}}} dt &= \frac{1}{H_0} \frac{1}{\sqrt{\left(1 - \frac{\delta}{\delta+3\omega}\right) (1 - \Omega_{(\text{dm+bm},0)})}} \int_1^\infty a^{-\frac{1}{2}(1+3\omega+\delta)} da \\ t_{\text{rip}} - t_0 &= \frac{1}{H_0} \frac{1}{\sqrt{\left(1 - \frac{\delta}{\delta+3\omega}\right) (1 - \Omega_{(\text{dm+bm},0)})}} \frac{2}{3\left(1+\omega+\frac{\delta}{3}\right)} a^{\frac{3(1+\omega+\delta/3)}{2}} \Big|_1^\infty \end{aligned}$$

$$\begin{aligned} t_{\text{rip}} - t_0 &= \frac{2}{3H_0\left(1+\omega+\frac{\delta}{3}\right)\sqrt{\left(1 - \frac{\delta}{\delta+3\omega}\right) (1 - \Omega_{(\text{dm+bm},0)})}} \left(0 - 1^{\frac{3(1+\omega+\delta/3)}{2}}\right) \\ \rightarrow t_{\text{rip}} - t_0 &= -\frac{2}{3H_0\left(1+\omega+\frac{\delta}{3}\right)\sqrt{\left(1 - \frac{\delta}{\delta+3\omega}\right) (1 - \Omega_{(\text{dm+bm},0)})}} \end{aligned}$$

SECCIÓN VII.



$$\rho_{\Lambda}(a)=\frac{3H_0^2}{8\pi G_N}\Omega_{\Lambda}(a)=\frac{\Lambda(a)}{8\pi G_N}=\int_{\tilde{E}_{\text{IR}}(a)}^{\tilde{E}_{\text{UV}}(a)}d\tilde{E}I(\tilde{E})$$

$$S = - \int_x \int_{\tilde{x}} \sqrt{-g(x)} \sqrt{-\tilde{g}(\tilde{x})} [R(x) + \tilde{R}(\tilde{x}) + \cdots] \\ = - \int_x \sqrt{-g(x)} \left[R(x) \int_{\tilde{x}} \sqrt{-\tilde{g}(\tilde{x})} + \int_{\tilde{x}} \sqrt{-\tilde{g}(\tilde{x})} \tilde{R}(\tilde{x}) + \cdots \right]$$

$$1/G_N \sim \int_{\tilde{x}} \sqrt{-\tilde{g}(\tilde{x})}$$

$$\Lambda/G_N \sim \int_{\tilde{x}} \sqrt{-\tilde{g}(\tilde{x})} \tilde{R}(\tilde{x})$$

$$\rho_{\Lambda}(a)=\frac{\Lambda(a)}{8\pi G_N}$$

$$\rho_{\Lambda}(a)=\int_{\tilde{E}_{\text{IR}}(a)}^{\tilde{E}_{\text{UV}}(a)}d\tilde{E}I(\tilde{E})$$

$$w(a)=w_0+w_a(1-a)$$

$$\rho_{\Lambda}(a)=\frac{3H_0^2}{8\pi G_N}\Omega_{\Lambda}(a)=\frac{\Lambda(a)}{8\pi G_N}=\int_{\tilde{E}_{\text{IR}}(a)}^{\tilde{E}_{\text{UV}}(a)}d\tilde{E}I(\tilde{E})$$

$$w(a)=-1-\frac{a}{3}\frac{d}{da}\log\,\Omega_{\Lambda}(a)\\=-1-\frac{a}{3\Lambda(a)}\frac{d\Lambda(a)}{da}\\=-1-\frac{8\pi G_Na}{3\Lambda(a)}\frac{d}{da}\left[\int_{\tilde{E}_{\text{IR}}(a)}^{\tilde{E}_{\text{UV}}(a)}d\tilde{E}I(\tilde{E})\right]\\=-1-\frac{8\pi G_Na}{3\Lambda(a)}\left[I\big(\tilde{E}_{\text{UV}}(a)\big)\frac{d\tilde{E}_{\text{UV}}(a)}{da}-I\big(\tilde{E}_{\text{IR}}(a)\big)\frac{d\tilde{E}_{\text{IR}}(a)}{da}\right]$$

$$I\big(\tilde{E}_{\text{UV}}(a)\big)\frac{d\tilde{E}_{\text{UV}}(a)}{da}>I\big(\tilde{E}_{\text{IR}}(a)\big)\frac{d\tilde{E}_{\text{IR}}(a)}{da}$$

$$I\big(\tilde{E}_{\text{UV}}(a)\big)\frac{d\tilde{E}_{\text{UV}}(a)}{da}<I\big(\tilde{E}_{\text{IR}}(a)\big)\frac{d\tilde{E}_{\text{IR}}(a)}{da}$$

$$E_{\text{IR}}\tilde{E}_{\text{UV}}=E_{\text{UV}}\tilde{E}_{\text{IR}}=\mu$$

$$E_{\text{IR}}(a)=\frac{E_0}{a}\rightarrow \tilde{E}_{\text{UV}}(a)=\frac{\mu}{E_{\text{IR}}(a)}=\frac{\mu a}{E_0},$$

$$\frac{d\tilde{E}_{\text{UV}}(a)}{da}=\frac{\mu}{E_0}>0$$

$$I_{\text{DE}}\big(\tilde{E},E_0\big)=A\tilde{E}^3e^{-B\tilde{E}/E_0},$$

$$\rho_\Lambda(a) = \frac{\Lambda(a)}{8\pi G_N} = \int_{\tilde{E}_{\text{IR}}(a)}^{\tilde{E}_{\text{UV}}(a)} d\tilde{E} I_{\text{DE}}(\tilde{E}, E_0)$$

$$\rho_\Lambda(a) = \frac{\Lambda(a)}{8\pi G_N} = \rho_*[b(\xi_{\text{IR}}(a)) - b(\xi_{\text{UV}}(a))],$$

$$\xi_{\text{UV}}(a) = \frac{B\tilde{E}_{\text{UV}}(a)}{E_0}, \xi_{\text{IR}}(a) = \frac{B\tilde{E}_{\text{IR}}(a)}{E_0}$$

$$\rho_* = \frac{6A}{B^4} E_0^4, b(\xi) = \left(1 + \xi + \frac{1}{2}\xi^2 + \frac{1}{6}\xi^3\right)e^{-\xi}$$

$$\begin{aligned} w(a) &= -1 - \frac{a}{3\rho_\Lambda(a)} \left[I(\tilde{E}_{\text{UV}}(a)) \frac{d\tilde{E}_{\text{UV}}(a)}{da} - I(\tilde{E}_{\text{IR}}(a)) \frac{d\tilde{E}_{\text{IR}}(a)}{da} \right] \\ &= -1 - \frac{1}{18[b_{\text{IR}}(a) - b_{\text{UV}}(a)]} \left[\xi_{\text{UV}}^3 e^{-\xi_{\text{UV}}} \left(a \frac{d\xi_{\text{UV}}}{da} \right) - \xi_{\text{IR}}^3 e^{-\xi_{\text{IR}}} \left(a \frac{d\xi_{\text{IR}}}{da} \right) \right] \end{aligned}$$

$$\tilde{E}_{\text{UV}}(a) = \frac{\mu}{E_0} a, \tilde{E}_{\text{IR}}(a) = 0$$

$$\xi_{\text{UV}}(a) = \frac{B\mu}{E_0^2} a \equiv \xi_0 a, \xi_{\text{IR}}(a) = 0$$

$$a \frac{d\xi_{\text{UV}}(a)}{da} = \xi_0 a = \xi_{\text{UV}}(a), a \frac{d\xi_{\text{IR}}(a)}{da} = 0$$

$$\rho_\Lambda(a) = \frac{\Lambda(a)}{8\pi G_N} = \rho_*[1 - b(\xi_0 a)]$$

$$w(a) = -1 - \frac{\xi^4 e^{-\xi}}{18\{1 - b(\xi)\}}, \xi = \xi_0 a$$

$$\frac{dw(a)}{da} = \frac{d\xi}{da} \frac{dw(\xi)}{d\xi} = -\frac{\xi_0}{18} \left[\frac{(4-\xi)\xi^3 e^{-\xi}}{\{1 - b(\xi)\}} - \frac{\xi^7 e^{-2\xi}}{6\{1 - b(\xi)\}^2} \right].$$

$$\begin{aligned} w_0 &= w(1) = -1 - \frac{\xi_0^4 e^{-\xi_0}}{18\{1 - b(\xi_0)\}} \\ w_a &= -w'(1) = \frac{\xi_0^4 e^{-\xi_0}}{18\{1 - b(\xi_0)\}} \left[(4 - \xi_0) - \frac{\xi_0^4 e^{-\xi_0}}{6\{1 - b(\xi_0)\}} \right] \\ &= -(4 - \xi_0)(w_0 + 1) - 3(w_0 + 1)^2 \end{aligned}$$

$$\begin{aligned} w_0 &= -\frac{7}{3} + \frac{4}{15}\xi_0 - \frac{2}{225}\xi_0^2 - \frac{2}{2625}\xi_0^3 + O(\xi_0^4) \\ w_a &= -\frac{4}{15}\xi_0 + \frac{4}{225}\xi_0^2 + \frac{2}{875}\xi_0^3 + O(\xi_0^4) \end{aligned}$$

$$\tilde{E}_{\text{UV}} = \frac{\mu}{E_0} [\alpha + (1 - \alpha)a], \tilde{E}_{\text{IR}} = \frac{\mu}{E_{\text{Planck}}} \frac{1}{[\beta + (1 - \beta)a]}$$



$$\alpha\beta=\frac{E_0}{E_{\rm Planck}}$$

$$\tilde E_{\rm UV}(0)=\tilde E_{\rm IR}(0)$$

$$\alpha = \beta = \sqrt{\frac{E_0}{E_{\rm Planck}}} \equiv k, \tilde E_{\rm UV}(0) = \tilde E_{\rm IR}(0) = \frac{\mu}{\sqrt{E_0 E_{\rm Planck}}}$$

$$\xi_{\rm UV}(a) = \frac{B\mu}{E_0^2} [k + (1-k)a] = \xi_0 [k + (1-k)a]$$

$$\begin{aligned}\xi_{\rm IR}(a) &= \frac{B\mu}{E_0 E_{\rm Planck}} [k + (1-k)a]^{-1} \\ &= \xi_0 \frac{E_0}{E_{\rm Planck}} [k + (1-k)a]^{-1} = \xi_0 k^2 [k + (1-k)a]^{-1}\end{aligned}$$

$$E_{\rm IR}\tilde E_{\rm IR}=\mu_{\rm IR}, E_{\rm UV}\tilde E_{\rm UV}=\mu_{\rm UV},$$

$$E_{\rm IR}(a)=\frac{E_0}{a}, E_{\rm UV}(a)=E_{\rm Planck}a$$

$$\tilde E_{\rm IR}(a)=\frac{\mu_{\rm IR}a}{E_0}, \tilde E_{\rm UV}(a)=\frac{\mu_{\rm UV}}{E_{\rm Planck}a}$$

$$\tilde E_{\rm IR}(0)<\tilde E_{\rm UV}(0)$$

$$\tilde E_{\rm IR}(a)=\tilde E_{\rm UV}(a)$$

$$\begin{gathered}\frac{\mu_{\rm IR}a}{E_0}=\frac{\mu_{\rm UV}}{E_{\rm Planck}a}\\\downarrow\\1< a^2=\frac{\mu_{\rm UV}}{\mu_{\rm IR}}\frac{E_0}{E_{\rm Planck}},\end{gathered}$$

$$\frac{\mu_{\rm UV}}{\mu_{\rm IR}}>\frac{E_{\rm Planck}}{E_0}\gg 1$$

$$\begin{aligned}\xi_{\rm IR}(a) &= \frac{B\tilde E_{\rm IR}(a)}{E_0} = \frac{B\mu_{\rm IR}}{E_0^2}a \equiv \xi_{\rm IR}^{(1)}a \\ \xi_{\rm UV}(a) &= \frac{B\tilde E_{\rm UV}(a)}{E_0} = \frac{B\mu_{\rm UV}}{E_0^2a} \equiv \frac{\xi_{\rm UV}^{(1)}}{a}\end{aligned}$$

$$\xi_{\rm IR}(a)\stackrel{a\rightarrow 0}{\rightarrow}0,\xi_{\rm UV}(a)\stackrel{a\rightarrow 0}{\rightarrow}\infty$$

$$b(\xi_{\rm IR}(a))\stackrel{a\rightarrow 0}{\rightarrow}1,b(\xi_{\rm UV}(a))\stackrel{a\rightarrow 0}{\rightarrow}0$$

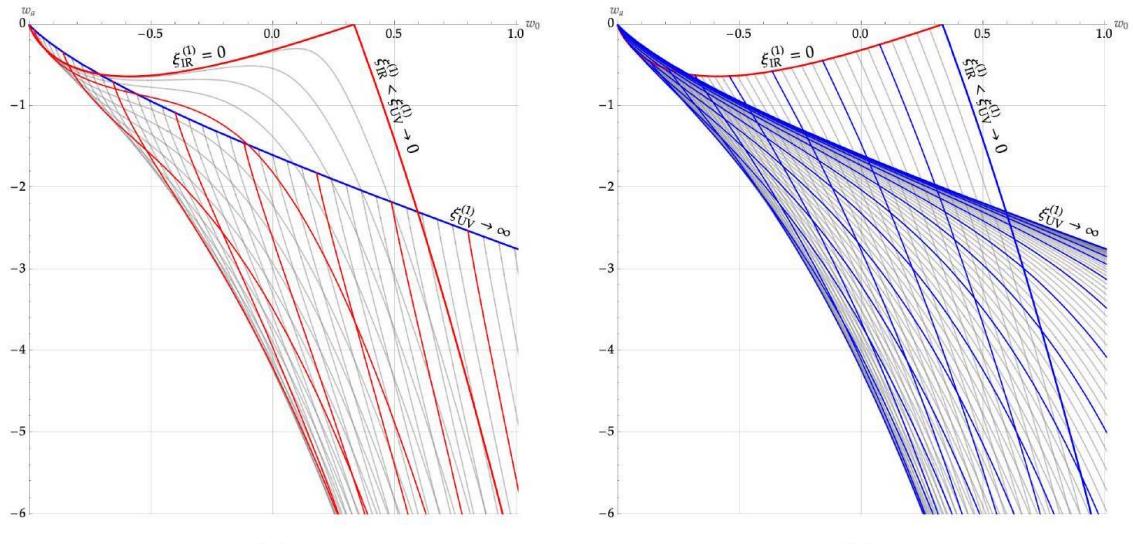
$$\rho_\Lambda(0)/\rho_* = 1, w(0) = -1$$



$$\begin{aligned}
& \rho_\Lambda(a)/\rho_* \begin{cases} \xi_{\text{IR}}^{(1)} \rightarrow 0 \\ \xi_{\text{UV}}^{(1)} \rightarrow \infty \end{cases} \begin{cases} 1 - b(\xi_{\text{UV}}) \\ b(\xi_{\text{IR}}) \end{cases} \\
& w(a) \begin{cases} \xi_{\text{IR}}^{(1)} \rightarrow 0 \\ \xi_{\text{UV}}^{(1)} \rightarrow \infty \end{cases} \begin{cases} -1 - \frac{1}{18[1-b(\xi_{\text{UV}})]} \left[\xi_{\text{UV}}^3 e^{-\xi_{\text{UV}}} \left(a \frac{d\xi_{\text{UV}}}{da} \right) \right] \\ -1 + \frac{1}{18b(\xi_{\text{IR}})} \left[\xi_{\text{IR}}^3 e^{-\xi_{\text{IR}}} \left(a \frac{d\xi_{\text{IR}}}{da} \right) \right] \end{cases} \\
& w_0 = w(1) = -1 + \frac{\left(\xi_{\text{UV}}^{(1)} \right)^4 e^{-\xi_{\text{UV}}^{(1)}} + \left(\xi_{\text{IR}}^{(1)} \right)^4 e^{-\xi_{\text{IR}}^{(1)}}}{18 \left[b \left(\xi_{\text{IR}}^{(1)} \right) - b \left(\xi_{\text{UV}}^{(1)} \right) \right]} \\
& \quad \xrightarrow{\xi_{\text{IR}}^{(1)} \rightarrow 0} -1 + \frac{\left(\xi_{\text{UV}}^{(1)} \right)^4 e^{-\xi_{\text{UV}}^{(1)}}}{18 \left[1 - b \left(\xi_{\text{UV}}^{(1)} \right) \right]} \\
& \xrightarrow{\xi_{\text{UV}}^{(1)} \rightarrow \infty} -1 + \frac{\left(\xi_{\text{IR}}^{(1)} \right)^4 e^{-\xi_{\text{IR}}^{(1)}}}{18b \left(\xi_{\text{IR}}^{(1)} \right)} = -1 + \frac{\left(\xi_{\text{IR}}^{(1)} \right)^4}{\left[1 + \left(\xi_{\text{IR}}^{(1)} \right) + \frac{1}{2} \left(\xi_{\text{IR}}^{(1)} \right)^2 + \frac{1}{6} \left(\xi_{\text{IR}}^{(1)} \right)^3 \right]} \\
& w_a = -w'(1) = \frac{\left(4 - \xi_{\text{UV}}^{(1)} \right) \left(\xi_{\text{UV}}^{(1)} \right)^4 e^{-\xi_{\text{UV}}^{(1)}} - \left(4 - \xi_{\text{IR}}^{(1)} \right) \left(\xi_{\text{IR}}^{(1)} \right)^4 e^{-\xi_{\text{IR}}^{(1)}}}{18 \left[b \left(\xi_{\text{IR}}^{(1)} \right) - b \left(\xi_{\text{UV}}^{(1)} \right) \right]}
\end{aligned}$$

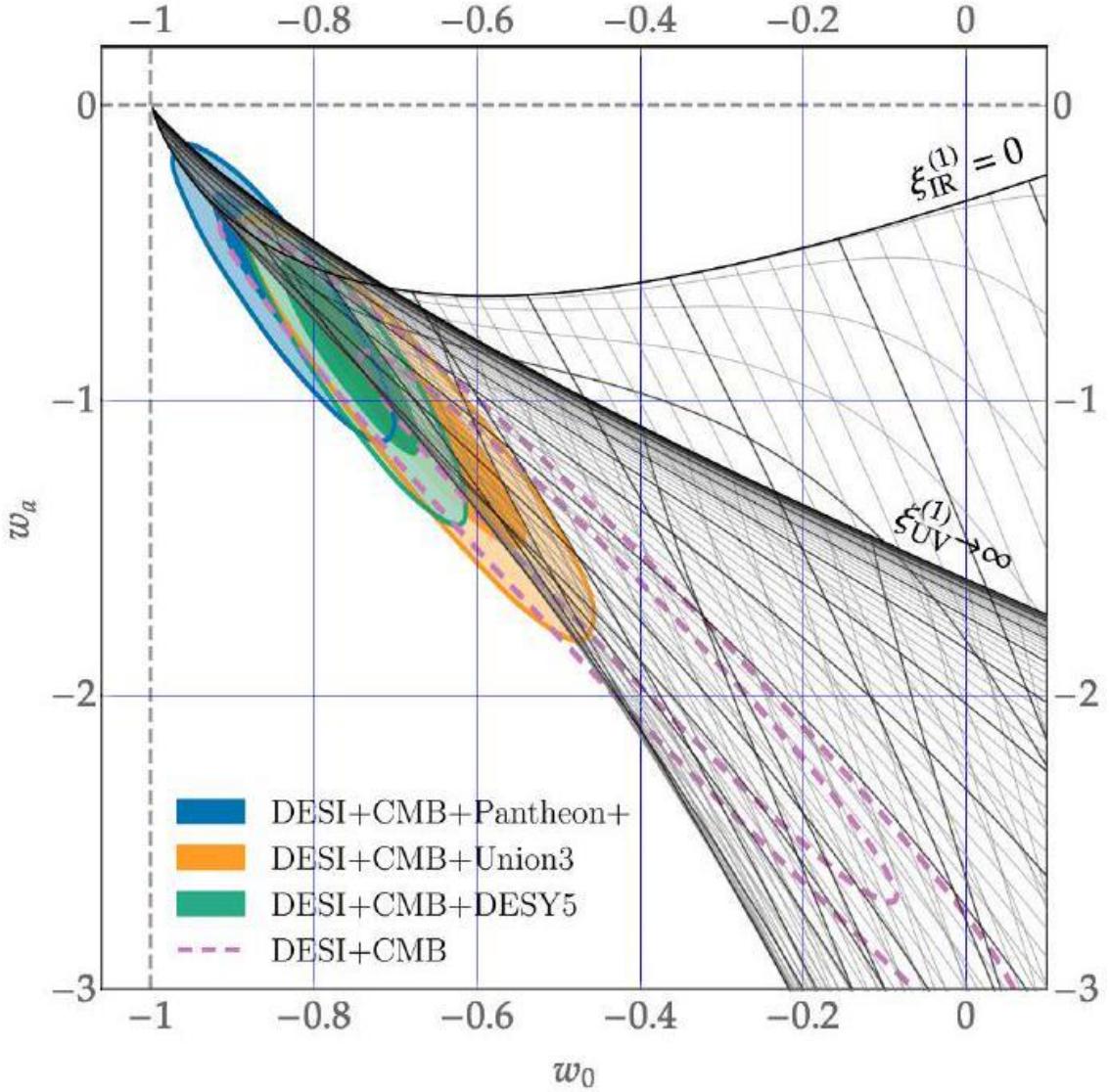
$$\begin{aligned}
& -3 \left[\frac{\left(\xi_{\text{UV}}^{(1)} \right)^4 e^{-\xi_{\text{UV}}^{(1)}} + \left(\xi_{\text{IR}}^{(1)} \right)^4 e^{-\xi_{\text{IR}}^{(1)}}}{18 \left[b \left(\xi_{\text{IR}}^{(1)} \right) - b \left(\xi_{\text{UV}}^{(1)} \right) \right]} \right]^2 \\
& = \left[\frac{\left(4 - \xi_{\text{UV}}^{(1)} \right) \left(\xi_{\text{UV}}^{(1)} \right)^4 e^{-\xi_{\text{UV}}^{(1)}} - \left(4 - \xi_{\text{IR}}^{(1)} \right) \left(\xi_{\text{IR}}^{(1)} \right)^4 e^{-\xi_{\text{IR}}^{(1)}}}{\left(\xi_{\text{UV}}^{(1)} \right)^4 e^{-\xi_{\text{UV}}^{(1)}} + \left(\xi_{\text{IR}}^{(1)} \right)^4 e^{-\xi_{\text{IR}}^{(1)}}} \right] (w_0 + 1) - 3(w_0 + 1)^2 \\
& \quad \xrightarrow{\xi_{\text{IR}}^{(1)} \rightarrow 0} \left(4 - \xi_{\text{UV}}^{(1)} \right) (w_0 + 1) - 3(w_0 + 1)^2 \\
& \quad \xrightarrow{\xi_{\text{UV}}^{(1)} \rightarrow \infty} - \left(4 - \xi_{\text{IR}}^{(1)} \right) (w_0 + 1) - 3(w_0 + 1)^2 \\
& w_0 = -1 + \frac{\left(\xi_{\text{IR}}^{(1)} \right)^4 \left[\kappa^4 e^{-\kappa \xi_{\text{IR}}^{(1)}} + e^{-\xi_{\text{IR}}^{(1)}} \right]}{18 \left[b \left(\xi_{\text{IR}}^{(1)} \right) - b \left(\kappa \xi_{\text{IR}}^{(1)} \right) \right]} \\
& \quad \xrightarrow{\xi_{\text{IR}}^{(1)} \rightarrow 0} -1 + \frac{4(\kappa^4 + 1)}{3(\kappa^4 - 1)} = \frac{\kappa^4 + 7}{3(\kappa^4 - 1)} \xrightarrow{\kappa \rightarrow \infty} \frac{1}{3} \\
& w_a = \left[\frac{\left(4 - \kappa \xi_{\text{IR}}^{(1)} \right) \left(\kappa \xi_{\text{IR}}^{(1)} \right)^4 e^{-\kappa \xi_{\text{IR}}^{(1)}} - \left(4 - \xi_{\text{IR}}^{(1)} \right) \left(\xi_{\text{IR}}^{(1)} \right)^4 e^{-\xi_{\text{IR}}^{(1)}}}{\left(\kappa \xi_{\text{IR}}^{(1)} \right)^4 e^{-\kappa \xi_{\text{IR}}^{(1)}} + \left(\xi_{\text{IR}}^{(1)} \right)^4 e^{-\xi_{\text{IR}}^{(1)}}} \right] (w_0 + 1) - 3(w_0 + 1)^2 \\
& \quad \xrightarrow{\xi_{\text{IR}}^{(1)} \rightarrow 0} \frac{4(\kappa^4 - 1)}{\kappa^4 + 1} (w_0 + 1) - 3(w_0 + 1)^2 = \frac{16}{3} - 3(w_0 + 1)^2 = -\frac{64\kappa^4}{3(\kappa^4 - 1)^2} \xrightarrow{\kappa \rightarrow \infty} 0
\end{aligned}$$

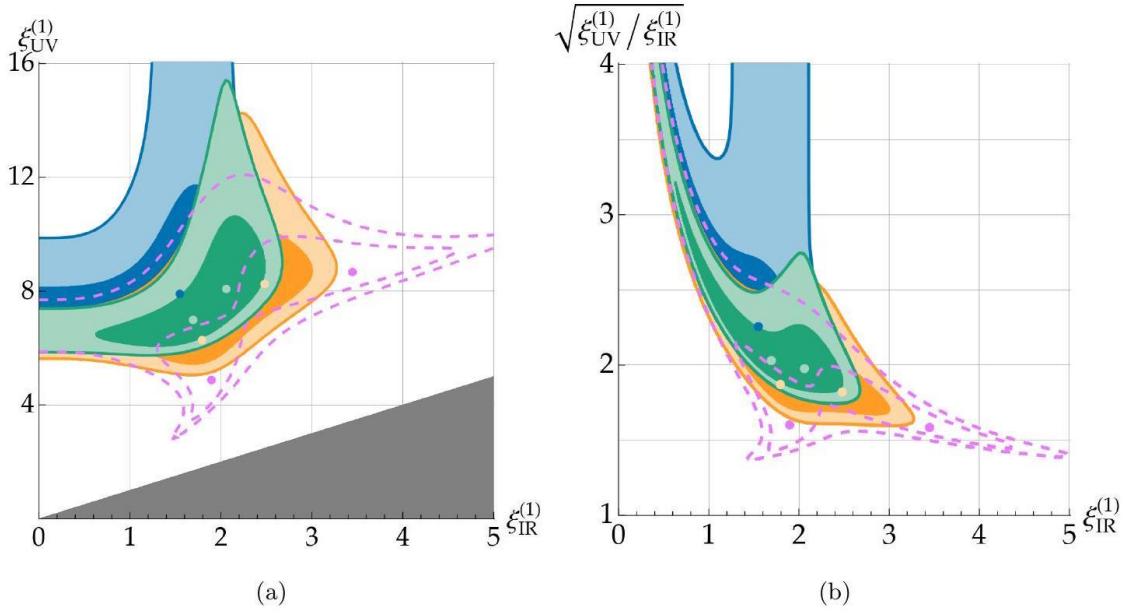




(a)

(b)





Figuras 56, 57 y 58. Curvatura crítica.

$$(w_0, w_a) = (-0.42, -1.75) \rightarrow \left(\xi_{\text{IR}}^{(1)}, \xi_{\text{UV}}^{(1)} \right) = (1.90, 4.90), (3.45, 8.71).$$

$$(w_0, w_a) = (-0.838, -0.62) \rightarrow (\xi_{\text{IR}}^{(1)}, \xi_{\text{UV}}^{(1)}) = (1.55, 7.93).$$

$$(w_0, w_a) = (-0.667, -1.09) \rightarrow \left(\xi_{\text{IR}}^{(1)}, \xi_{\text{UV}}^{(1)} \right) = (1.79, 6.31), (2.48, 8.28).$$

$$(w_0, w_a) = (-0.752, -0.86) \rightarrow \left(\xi_{\text{IR}}^{(1)}, \xi_{\text{UV}}^{(1)} \right) = (1.69, 7.02), (2.06, 8.10).$$

$$w_a = -\frac{dw}{da}\Big|_{a=1}$$

$$H(a)^2 = \left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G_N}{3} \sum_i \rho_i(a) \approx \frac{8\pi G_N}{3} [\rho_m(a) + \rho_\Lambda(a)]$$

$$\rho_\Lambda(a) = \frac{3H(a)^2}{8\pi G_N} - \rho_m(a) = \frac{3H(a)^2}{8\pi G_N} [1 - \Omega_m(a)],$$

$$\Omega_i(a) = \frac{8\pi G_N}{3H(a)^2} \rho_i(a)$$

$$q(a) = -\frac{\ddot{a}}{aH(a)^2} = \frac{4\pi G_N}{3H(a)^2} \sum_i [\rho_i(a) + 3p_i(a)] \approx \frac{1}{2} + \frac{4\pi G_N}{H(a)^2} p_\Lambda(a)$$

$$p_\Lambda(a) = \frac{H(a)^2}{4\pi G_N} \left[q(a) - \frac{1}{2} \right]$$

$$w(a) = \frac{p_\Lambda(a)}{\rho_\Lambda(a)} = \frac{2q(a) - 1}{3[1 - \Omega_m(a)]}$$



$$\Omega_m(a)=\left[\frac{H_0}{H(a)}\right]^2\frac{\Omega_{m,0}}{a^3}$$

$$H(a) = H_0 \sqrt{\frac{\Omega_{m,0}}{a^3} + \Omega_\Lambda(a)}$$

$$3[1-\Omega_m(a)] = \frac{3H_0^2}{H(a)^2}\Omega_\Lambda(a)$$

$$\frac{dH}{dt}=\frac{\ddot{a}}{a}-\left(\frac{\dot{a}}{a}\right)^2=-H(a)^2[q(a)+1]$$

$$q(a)=-\frac{1}{H(a)^2}\frac{dH(a)}{dt}-1=-\frac{a}{H(a)}\frac{dH(a)}{da}-1$$

$$2q(a)-1=\frac{3H_0^2}{H(a)^2}\Bigl[-\frac{a}{3}\frac{d\Omega_\Lambda(a)}{da}-\Omega_\Lambda(a)\Bigr]$$

$$w(a)=\frac{-\frac{a}{3}\frac{d\Omega_\Lambda(a)}{da}-\Omega_\Lambda(a)}{\Omega_\Lambda(a)}=-1-\frac{a}{3}\frac{d}{da}\log\,\Omega_\Lambda(a)$$

$$S_{2d}=\frac{1}{4\pi}\int_{\Sigma}[\partial_{\tau}\mathbb{X}^A(\eta_{AB}(\mathbb{X})+\omega_{AB}(\mathbb{X}))-\partial_{\sigma}\mathbb{X}^AH_{AB}(\mathbb{X})]\partial_{\sigma}\mathbb{X}^B$$

$$\left[\hat{\mathbb{X}}^A,\hat{\mathbb{X}}^B\right]=i\omega^{AB}$$

$$S\sim \int~~d\tau d\sigma {\rm Tr}\big[\partial_{\tau}\hat{\mathbb{X}}^A\partial_{\sigma}\hat{\mathbb{X}}^B\big(\omega_{AB}(\hat{\mathbb{X}})+\eta_{AB}(\hat{\mathbb{X}})\big)-\partial_{\sigma}\hat{\mathbb{X}}^AH_{AB}(\hat{\mathbb{X}})\partial_{\sigma}\hat{\mathbb{X}}^B\big]$$

$$\partial_{\sigma}\hat{\mathbb{X}}^A~\rightarrow~\big[\hat{\mathbb{X}}^{26},\hat{\mathbb{X}}^A\big], A=0,1,\cdots,25$$

$$S\sim \int~~d\tau {\rm Tr}\big(\partial_{\tau}\hat{\mathbb{X}}^a\big[\hat{\mathbb{X}}^b,\hat{\mathbb{X}}^c\big]\eta_{abc}(\hat{\mathbb{X}})-H_{ac}\big[\hat{\mathbb{X}}^a,\hat{\mathbb{X}}^b\big]\big[\hat{\mathbb{X}}^c,\hat{\mathbb{X}}^d\big]H_{bd}(\hat{\mathbb{X}})\big)$$

$$\mathbb{S}_{nc\textbf{M}}=\frac{1}{4\pi}\int_{\tau}\big(\partial_{\tau}\hat{\mathbb{X}}^i\big[\hat{\mathbb{X}}^j,\hat{\mathbb{X}}^k\big]g_{ijk}(\hat{\mathbb{X}})-\big[\hat{\mathbb{X}}^i,\hat{\mathbb{X}}^j\big]\big[\hat{\mathbb{X}}^k,\hat{\mathbb{X}}^{\ell}\big]h_{ijk\ell}(\mathbb{X})\big)$$

$$S_{\mathrm{eff}}^{nc}=\int_x\int_{\tilde{x}}\mathrm{Tr}\sqrt{g(x,\tilde{x})}[R(x,\tilde{x})+L_m(x,\tilde{x})+\cdots]$$

$$\left[\hat{x}^a,\hat{\tilde{x}}_b\right]=2\pi i\lambda^2\delta^a_b,[\hat{x}^a,\hat{x}^b]=\left[\hat{\tilde{x}}_a,\hat{\tilde{x}}_b\right]=0$$

$$\begin{aligned} S_{d=4}=&-\int_x\int_{\tilde{x}}\sqrt{-g(x)}\sqrt{-\tilde{g}(\tilde{x})}[R(x)+\tilde{R}(\tilde{x})]\\ &=-\int_x\sqrt{-g(x)}\left[R(x)\int_{\tilde{x}}\sqrt{-\tilde{g}(\tilde{x})}+\int_{\tilde{x}}\sqrt{-\tilde{g}(\tilde{x})}\tilde{R}(\tilde{x})\right] \end{aligned}$$

$$1/G_N \sim \int_{\tilde{x}}\sqrt{-\tilde{g}(\tilde{x})}$$



$$\Lambda/G_N \sim \int_{\tilde{x}} \sqrt{-\tilde{g}(\tilde{x})} \tilde{R}(\tilde{x})$$

$$[\hat{a}_i, \hat{a}_j^\dagger]_q = \hat{a}_i \hat{a}_j^\dagger - q \hat{a}_j^\dagger \hat{a}_i = \delta_{ij}$$

$$\hat{a}_i |0\rangle = 0, \hat{a}_i \hat{a}_j^\dagger = \delta_{ij}, \sum_i \hat{a}_i^\dagger \hat{a}_i = \mathbf{1} - |0\rangle\langle 0|$$

$$\hat{N} = \sum_{k=1}^{\infty} \sum_{i_1} \hat{a}_{i_1}^\dagger \left(\sum_{i_2} \hat{a}_{i_2}^\dagger \dots \left(\sum_{i_k} \hat{a}_{i_k}^\dagger \hat{a}_{i_k} \right) \dots \hat{a}_{i_2} \right) \hat{a}_{i_1}$$

$$\hat{N}_1 = \sum_{k=1}^{\infty} (\hat{a}^\dagger)^k (\hat{a})^k = \frac{\hat{a}^\dagger \hat{a}}{1 - \hat{a}^\dagger \hat{a}}$$

Sección VIII.

$$ds^2 = -dt^2 - a^2(t)[dr^2 + r^2(d\theta^2 + \sin^2 \theta d\phi^2)],$$

$$H^2 = \frac{8\pi G}{3}(\rho_b + \rho_{dm} + \rho_r + \rho_{de})$$

$$\begin{aligned}\rho_{dim} + 3H(1 + \omega_{dm})\rho_{dm} &= Q \\ \dot{\rho_{de}} + 3H(1 + \omega_{de})\rho_{de} &= -Q \\ \dot{\rho_r} + 3H(1 + \omega_r)\rho_r &= 0 \\ \dot{\rho_b} + 3H(1 + \omega_b)\rho_b &= 0\end{aligned}$$

$$\rho_{dm} = \rho_{dm0} a^{-I}$$

$$\begin{aligned}\frac{d\rho_{dm}}{da} + \frac{3}{a}(1 + \omega_{dm})\rho_{dm} &= \kappa \left(\frac{d\rho_{dm}}{da} + \frac{d\rho_{de}}{da} \right) \\ \frac{d\rho_{de}}{da} + \frac{3}{a}(1 + \omega_{de})\rho_{de} &= -\kappa \left(\frac{d\rho_{dm}}{da} + \frac{d\rho_{de}}{da} \right)\end{aligned}$$

$$\rho_{de} = \frac{\rho_{dm0}}{\kappa} \left[1 - \frac{3}{I} - \kappa \right] a^{-I} + \tilde{\rho}_{de}$$

$$\omega_{de} = -1 + \left[\frac{\frac{\rho_{dm0}}{\kappa} \left(1 - \frac{3}{I} - \frac{3\kappa}{I} \right) a^{-I}}{\frac{\rho_{dm0}}{\kappa} \left(1 - \frac{3}{I} - \kappa \right) a^{-I} + \tilde{\rho}_{de}} \right] \frac{I}{3}$$

$$\Omega_i = \frac{8\pi G}{3H^2} \rho_i$$

$$\begin{aligned}\Omega_{dm} &= \frac{8\pi G}{3H^2} \rho_{dm0} a^{-I} \\ \Omega_{de} &= \frac{8\pi G}{3H^2} \left[\frac{\rho_{dm0}}{\kappa} \left(1 - \frac{3}{I} - \kappa \right) a^{-I} + \tilde{\rho}_{de} \right]\end{aligned}$$

$$\Omega_{de} = \Omega_{dm} \left(\frac{1}{\kappa} - \frac{3}{Ik} - 1 \right) + \tilde{\Omega}_{de}$$



$$H^2(a) = H_0^2 \left[\Omega_{b0} a^{-3} + \Omega_{r0} a^{-4} + \Omega_{dm0} a^{-I} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} + \tilde{\Omega}_{deo} \right]$$

$$\Omega_b + \Omega_r + \Omega_{dm} + \Omega_{de} = 1$$

$$\tilde{\Omega}_{deo} = 1 - \Omega_{b0} - \Omega_{r0} - \Omega_{dm0} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa}$$

$$H^2(z) = H_0^2 [\Omega_{b0}(1+z)^3 + \Omega_{r0}(1+z)^4 + \Omega_{dm0}(1+z)^I \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} + 1 - \Omega_{b0} - \Omega_{r0} - \Omega_{dm0} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa}]$$

$$q = -1 - \frac{\dot{H}}{H^2} = -1 - \frac{\dot{a}}{H^2} \frac{dH}{da}$$

$$\dot{a}^2 = H_0^2 \left[\Omega_{bo} a^{-1} + \Omega_{ro} a^{-2} + \Omega_{dm0} a^{2-I} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} + \tilde{\Omega}_{deo} a^2 \right]$$

$$\begin{aligned} \frac{dH}{da} &= \frac{H_0}{2} \left[\Omega_{b0} a^{-3} + \Omega_{r0} a^{-4} + \Omega_{dm0} a^{-I} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} + \tilde{\Omega}_{deo} \right]^{-1/2} \\ &\quad \times \left[(-3)\Omega_{b0} a^{-4} + (-4)\Omega_{r0} a^{-5} + (-I)\Omega_{dm0} a^{-I-1} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} \right] \end{aligned}$$

$$\begin{aligned} q(a) &= -1 - \left(\frac{H_0 \left[\Omega_{bo} a^{-1} + \Omega_{ro} a^{-2} + \Omega_{dm0} a^{2-I} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} + \tilde{\Omega}_{deo} a^2 \right]^{1/2}}{H_0^2 \left[\Omega_{b0} a^{-3} + \Omega_{r0} a^{-4} + \Omega_{dm0} a^{-I} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} + \tilde{\Omega}_{deo} \right]} \right. \\ &\quad \times \left. \frac{H_0}{2} \left[\Omega_{b0} a^{-3} + \Omega_{r0} a^{-4} + \Omega_{dm0} a^{-I} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} + \tilde{\Omega}_{deo} \right]^{-1/2} \right. \\ &\quad \times \left. \left[(-3)\Omega_{b0} a^{-4} + (-4)\Omega_{r0} a^{-5} + (-I)\Omega_{dm0} a^{-I-1} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} \right] \right) \end{aligned}$$

$$\begin{aligned} q(z) &= -1 - \left(\frac{H_0 \left[\Omega_{bo}(1+z)^1 + \Omega_{ro}(1+z)^2 + \Omega_{dm0}(1+z)^{I-2} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} + \tilde{\Omega}_{deo}(1+z)^{-2} \right]^{1/2}}{H_0^2 \left[\Omega_{b0}(1+z)^3 + \Omega_{r0}(1+z)^4 + \Omega_{dm0}(1+z)^I \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} + \tilde{\Omega}_{deo} \right]} \right. \\ &\quad \times \left. \frac{H_0}{2} \left[\Omega_{b0}(1+z)^3 + \Omega_{r0}(1+z)^4 + \Omega_{dm0}(1+z)^I \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} + \tilde{\Omega}_{deo} \right]^{-1/2} \right. \\ &\quad \times \left. \left[(-3)\Omega_{b0}(1+z)^4 + (-4)\Omega_{r0}(1+z)^5 + (-I)\Omega_{dm0}(1+z)^{I+1} \left(1 - \frac{3}{I} \right) \frac{1}{\kappa} \right] \right) \end{aligned}$$

$$\begin{aligned} d_L &= \chi(1+z) \\ d_A &= \frac{\chi}{(1+z)} \end{aligned}$$



$$d_L = \frac{1+z}{H_0} \int_0^z \frac{dz}{\left[\Omega_{b0}(1+z)^3 + \Omega_{r0}(1+z)^4 + \Omega_{dm0}(1+z)^I \left(1-\frac{3}{I}\right) \frac{1}{\kappa} \right.} \\ \left. + 1 - \Omega_{b0} - \Omega_{r0} - \Omega_{dm0} \left(1-\frac{3}{I}\right) \frac{1}{\kappa} \right]^{1/2}$$

$$d_A = \frac{(1+z)^{-1}}{H_0} \int_0^z \frac{dz}{\left[\Omega_{b0}(1+z)^3 + \Omega_{r0}(1+z)^4 + \Omega_{dm0}(1+z)^I \left(1-\frac{3}{I}\right) \frac{1}{\kappa} \right.} \\ \left. + 1 - \Omega_{b0} - \Omega_{r0} - \Omega_{dm0} \left(1-\frac{3}{I}\right) \frac{1}{\kappa} \right]^{1/2}$$

$$t_0-t=\int_0^z\frac{dz}{(1+z)H(z)}$$

$$t_0-t=\int_0^z \frac{dz}{(1+z)H_0\left[\Omega_{b0}(1+z)^3+\Omega_{r0}(1+z)^4+\Omega_{dm0}(1+z)^I\left(1-\frac{3}{I}\right)\frac{1}{\kappa}\right.} \\ \left.+1-\Omega_{b0}-\Omega_{r0}-\Omega_{dm0}\left(1-\frac{3}{I}\right)\frac{1}{\kappa}\right]^{1/2}$$

$$t_0=\int_0^\infty \frac{dz}{(1+z)H_0\left[\Omega_{b0}(1+z)^3+\Omega_{r0}(1+z)^4+\Omega_{dm0}(1+z)^I\left(1-\frac{3}{I}\right)\frac{1}{\kappa}\right.} \\ \left.+1-\Omega_{b0}-\Omega_{r0}-\Omega_{dm0}\left(1-\frac{3}{I}\right)\frac{1}{\kappa}\right]^{1/2}$$

$$m_{\rm model}=5\log_{10}\left(\frac{d_L}{1Mpc}\right)+25+M$$

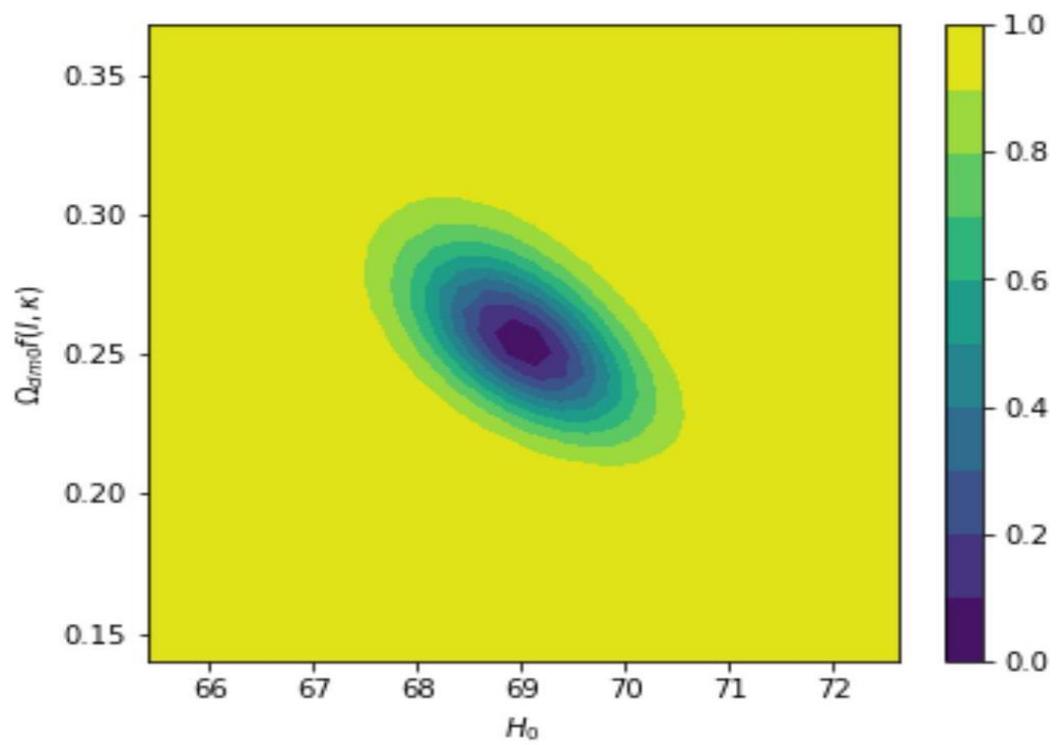
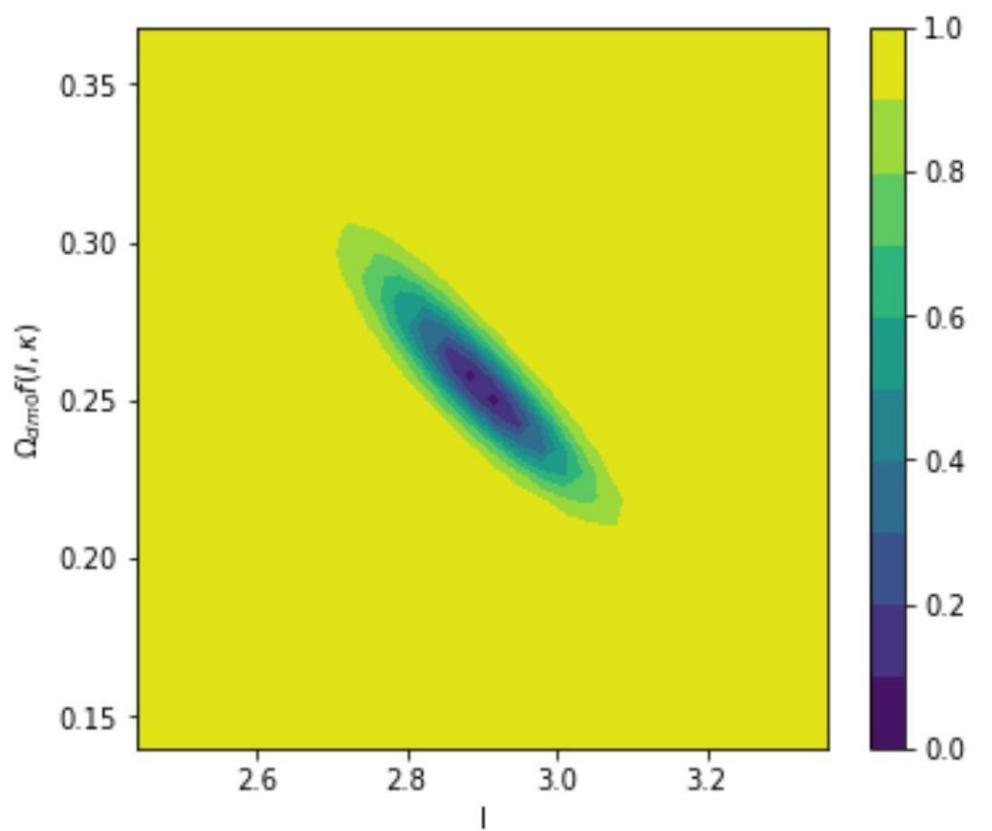
$$\chi^2_{SN} = \frac{\left(m_{obs_i} - m_{\rm model}\right)^2}{\sigma_i^2}$$

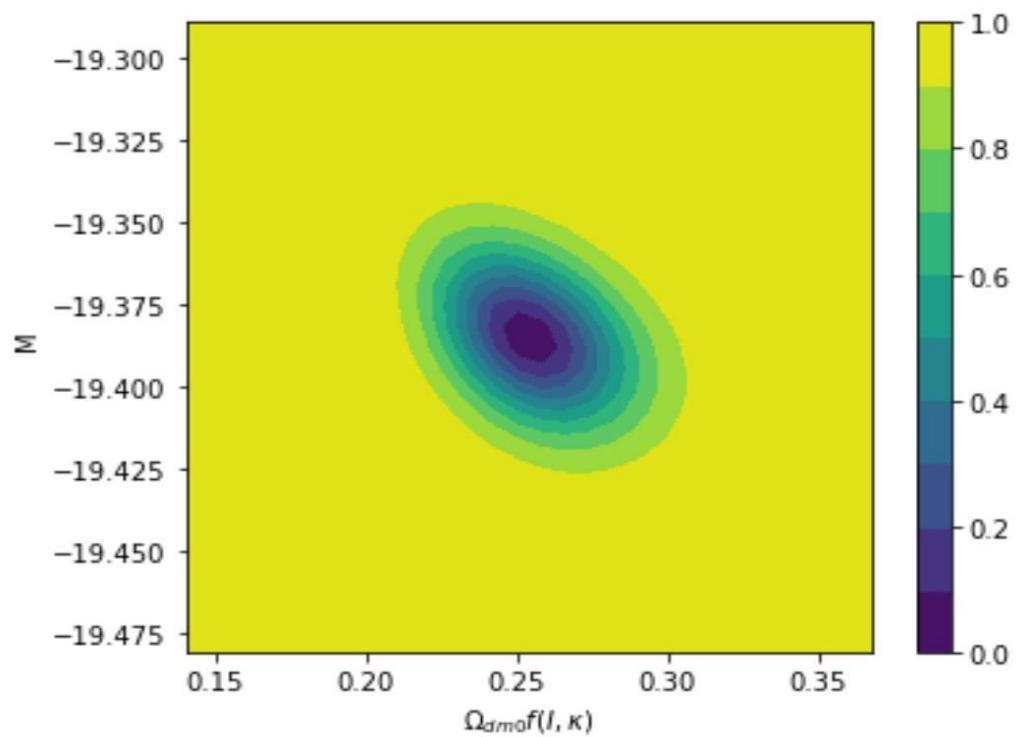
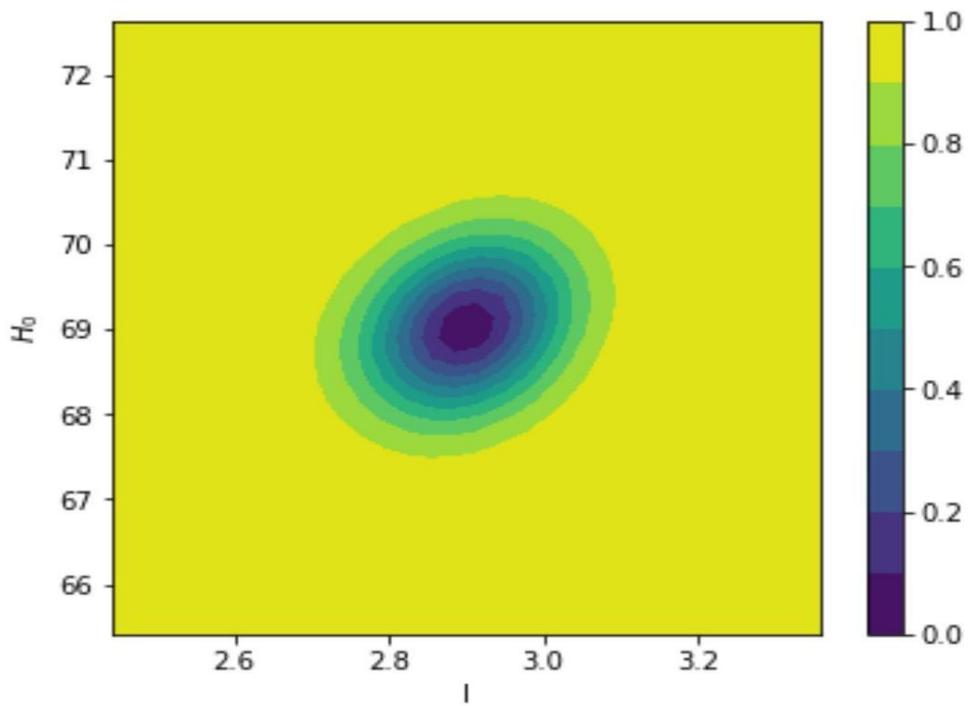
$$\chi^2_{H(z)} = \frac{\left(H(z)_{obs_i} - H(z)_{model}\right)^2}{\sigma_i^2}$$

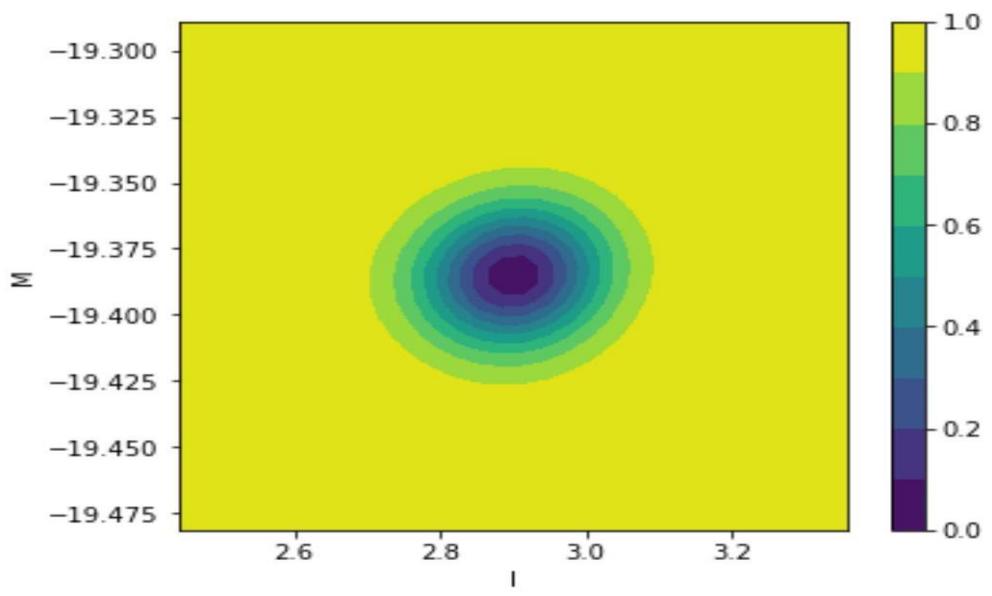
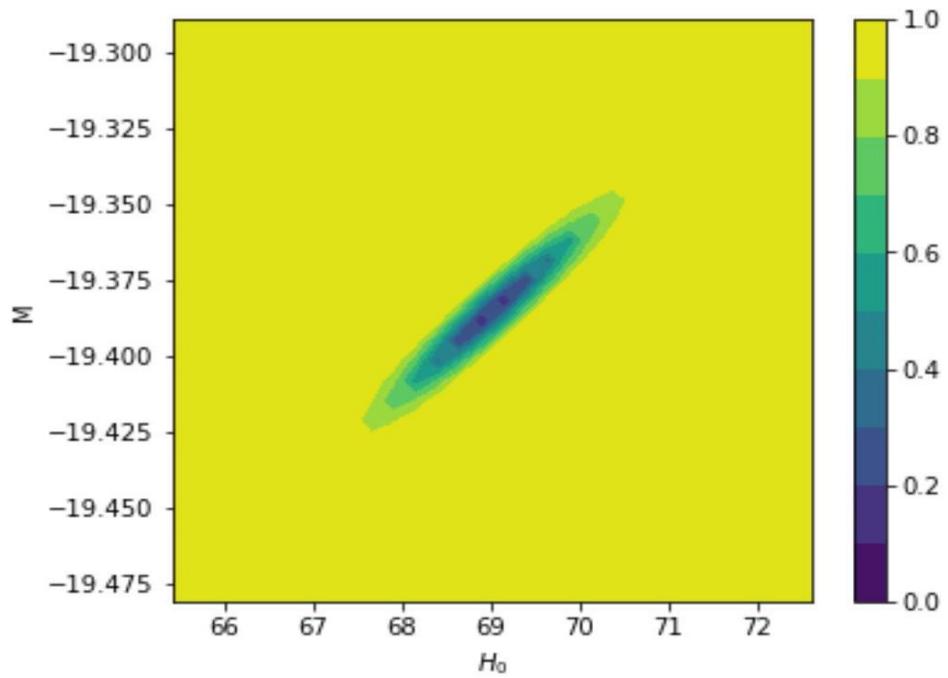
$$\chi^2 = \chi^2_{SN} + \chi^2_{H(z)}$$

$$\left(\frac{da}{d\tilde{t}}\right)^2 = \left[\Omega_{bo}a^{-1} + \Omega_{dm0}a^{2-I}\left(1-\frac{3}{I}\right)\frac{1}{\kappa} + \tilde{\Omega}_{deo}a^2\right]$$









Figuras 59, 60, 61, 62, 63 y 64. Agujero negro cuántico. Mareas gravitacionales.

$$q_0^{IDE} = -0.561; z_t^{IDE} = 0.744$$

$$q_0^{\Lambda CDM} = -0.580; z_t^{\Lambda CDM} = 0.727$$

$$t_0^{IDE} = 14.032 \text{Gyr}$$



$$t_0^{\Lambda CDM} = 13.885 \text{Gyr}$$

$$\frac{d(m_{dm}n_{dm})}{da} + \frac{3}{a}(1 + \omega_{dm})m_{dm}n_{dm} = \kappa \left[\frac{d(m_{dm}n_{dm})}{da} + \frac{d\rho_{de}}{da} \right]$$

$$m_{dm} \frac{dn_{dm}}{da} + \frac{3}{a} m_{dm} n_{dm} = \kappa \left(m_{dm} \frac{dn_{dm}}{da} + \frac{d\rho_{de}}{da} \right)$$

$$\frac{dn_{dm}}{da} + \frac{3}{a(1-\kappa)} n_{dm} = \frac{\kappa}{1-\kappa} \frac{n_{dm}}{\rho_{dm}} \frac{d\rho_{de}}{da}$$

$$n_{dm} = n_{dm0} a^{-l}$$

$$m_{dm} \frac{dn_{dm}}{da} + n_{dm} \frac{dm_{dm}}{da} + \frac{3}{a} m_{dm} n_{dm} = \kappa \left[m_{dm} \frac{dn_{dm}}{da} + n_{dm} \frac{dm_{dm}}{da} + \frac{d\rho_{de}}{da} \right]$$

$$\frac{dn_{dm}}{da} + \frac{3}{a} n_{dm} = 0$$

$$m_{dm} = m_{dm0} a^{3-l}$$

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